

# What the biggest things in the universe can tell us about the oldest and the smallest

After the talk, explore our different interactive activities and the telescopes at the McLennan Physical Laboratories, until 10:00pm EST.

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Image Credit: Nathan J. Carlson





# Land Acknowledgment

Inupiat

Holikachuk

Alutiiq (Sugpiaq)

Ungazigmit

(Chaplino) Yupik

**Unangam Tanangin** 

(Unangax /Aleut)

Gwich'in Nanh Kasho Gotine

Hän

Sahtú Got'ine

Koyukon

Upper Tanana

Shita Got'ine

Carcross/Tagish First Nation (Yukon)

Lingit Aani (Tlingit)

**Gitanyow Laxyip** 

xà'isla wawis (Haisla)

https://native-land.ca/

Yup'ik/Cup'ik

https://www.whose.land/en/

Quatsino

Twana/Skokomish

Chepenefa Tututni Wailaki

Inuvialuit

Tłicho Ndè

Dehcho Dene **NWT Métis Nation** 

Dene Tha'

Beaver

Beaver Lake Cree

Tsuu Tina

Salish

Goshute

Hunkpapa

Ponca

Kiikaapoi (Kickapoo)

Anishinabewaki

Inuit

Kalaallit Nur

St'aschinuw (Naskapi)

Nitassinan (Innu)

Beothuk

Penobscot

Aucocisco

Quinnipiac

Moose Cree

Niúachi Petun-



★ Lecture by Nathan J. Carlson.

interactive activities and visit the telescopes, until 10:00pm EST.

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# Tonight's Plan

# Go to the McLennan Physical Laboratories and explore our different



January AstroTour:

 Speaker James Lane Thurs. January 11th, 8:00PM EST Location: TBD



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# Jpcoming Events







# Upcoming Events

# ◆ Variety of online and in person events being held this month by the Royal Astronomical Society of Canada, visit their website for more information: <u>https://rascto.ca/events</u>



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# Tonight's Speaker Nathan J Carlson

Nathan is 5th year PhD candidate in the Department of Physics and at the Canadian Institute for Theoretical Astrophysics at the University of Toronto. His research focuses on simulating the distribution of energy in the early Universe to model the evolution of large scale structure (the clumping together of galaxies over cosmic time) in order to better understand the epoch of inflation (a period of time when the very early Universe expanded very, very rapidly). Nathan was born and raised in Ottawa, completing his BSc in Physics at UOttawa before coming to UofT to start his PhD. In his free time, Nathan is very outdoorsy, and likes to do things outside, like hiking and canoeing. In fact, Nathan has seen at least four different species of wild deer.

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# What the biggest things in the universe can tell us about the oldest and Smalest

Nathan J. Carlson - CITA/UofT - AstroTours December 6, 2023





Physics IVERSITY OF TORONTO





### What is cosmology? We care about things that are a huge range of sizes!

Quantum fluctuations

Credit: Wagyx (2022)

**Stars** 

Credit: NASA/SDO

(2013)

500,000,000 m

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0.0000000005 m

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# Galaxies

Leros after

A HNO WIT



(that's a one with 20 zeros after it!)

Credit: A. Block/ U. Arizona (2022)



# What is cosmology? We also care about things that happened over a huge span of time (13.8 billion years to be precise!)



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Years after the Big Bang



# What interests me as a cosmologist? Why does the universe look like this and not like something else?



# Light has a speed, so it takes time to get to us, and so we see far, far away things as they were a long time ago

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Credit: presumably Disney, they own everything now



# Let's do some cosmic detective work. Can we figure out what the universe looked like in the distant past?







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We are here







Credit: A. Block/U. Arizona (2022)









# The Doppler effect tells us that things that look red are actually moving away from us!

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Credit: own work







# Let's do some cosmic detective work. Can we figure out what the universe looked like in the distant past?

Cosmology - Nathan J. Carlson @ UofT AstroTours, 6 Dec 2023

We are here

Credit: A. Block/U. Arizona (2022)





# The universe is expanding like a loaf of raisin bread baking in the oven.

### (and whether we like it or not, we're probably not in the middle of it )

Credit: NASA/ WMAP (2013)

MAP990404

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## The farthest back we can see ...





### The farthest back we can see is the CMB! (Cosmic Microwave Background)



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Credit: ESA/Planck Collaboration (2013)

Both are flattened

out spheres



# How could you get a CMB? All matter emits light with a colour that depends only on its temperature.



Iron or steel glow orange when heated up by a blacksmith to around 1000°C.

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Animals at around 37°C (like you or these dinosaurs) emit *infrared* light! We can't see it, but it's made up of the same stuff as the light we can see!





### Our atmosphere is made of neutral gas, so we can see through it.

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Credit: own work

# Stars are made of more ionized gas, so we can't see through them.



# All this evidence points to the *Big Bang*. But there are still some physics problems that the Big Bang alone can't explain.

# WHYREFIS

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Credit: ESA/Planck Collaboration (2013)



# The horizon problem

### The CMB 300,000 years after the Big Bang

Photons from the CMB have been travelling 13.8B years to reach Earth

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Credits: ESA/*Planck Collaboration* (2013), NASA/Apollo 17 (1972), own work



# Before Inflation



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# How could inflation happen? Theoretical physics might be able to tell us. There are four fundamental forces in nature.









Holds atoms together.

Involved in nuclear decay.

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Atom

**Nucleus** 

Proton

Neutron

Electron

**Strong Nuclear** Force

Gravity





Holds atomic nuclei together.

Makes apples fall on Newton's head.

Credits: NASA/Apollo 17 (1972), own work





# If things get hot enough, the forces can merge. If a hot thing cools, the unmerging releases lots of energy.





### These forces become one at around 1 quadrillion °C

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Atom

Nucleus

Proton

Neutron

Electron









Holds atomic nuclei together.

Makes apples fall on Newton's head.

Credits: NASA/Apollo 17 (1972), own work



# If things get hot enough, the forces can merge. If a hot thing cools, the unmerging releases lots of energy.

## **Grand Unified Theory**



### These forces become one at around 10 trillion times hotter than that!

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Gravity



Makes apples fall on Newton's head.

Credits: NASA/Apollo 17 (1972), own work



# If things get hot enough, the forces can merge. If a hot thing cools, the unmerging releases lots of energy.

## Theory of Everything?





We have no possible...

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We have no idea if this is

Credits: NASA/Apollo 17 (1972), own work



# What proof do we have that inflation actually happened?



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The often bizarre seeming theory of Quantum Mechanics offers one piece of evidence.



## What proof do we have that inflation actually happened?

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Credit: D. B. Leinweber (2003)/ ESA/Planck Collaboration (2013)







### Inflation predicts a CMB and the CMB has a lot that it can tell us



# colour that we see with our microwave telescopes.

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Inflation predicts that we should see the CMB and that it should have the exact same





# We see different things when we look at the Earth and focus on objects of different sizes.













![](_page_34_Picture_0.jpeg)

![](_page_35_Picture_1.jpeg)

# Now, instead of zooming in, let's look at a region of the CMB and gradually increase the resolution of the image, resolving finer and finer detail.

![](_page_36_Picture_1.jpeg)

### **Size of objects**

Credit: W. Hu & M. White (2004)

![](_page_36_Picture_5.jpeg)

# We can't see inflation, the CMB is blocking our view, so we don't know exactly what happened during inflation.

![](_page_37_Figure_1.jpeg)

The CMB like many things in nature, very closely follows the same "Bell Curve" that you might be familiar from grading in schools.

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![](_page_37_Figure_4.jpeg)

![](_page_37_Picture_5.jpeg)

7

7

## The "Bell Curve" is also called a Gaussian. A non-Gaussian is any distribution that differs from the Bell Curve.

![](_page_38_Picture_1.jpeg)

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# Gausstan-Gaussian Irregular curves Heavy tailed curves

Credits: own work

![](_page_38_Picture_5.jpeg)

![](_page_38_Picture_6.jpeg)

# I make simulations of the the the Cosmic Web starting from different types of inflation.

![](_page_39_Picture_1.jpeg)

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Make the Cosmic Web

![](_page_39_Picture_4.jpeg)

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# I make simulations of the the the Cosmic Web starting from different types of inflation.

The distribution of energy in the universe just after inflation (light means lots of energy, dark means little energy):

![](_page_40_Picture_2.jpeg)

### Gaussian (following a bell curve)

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![](_page_40_Picture_5.jpeg)

Non-Gaussian (not following a bell curve) Credit: own work.

# My simulations let us map out the matter in a universe that would form after a given type of inflation. Galaxies form earlier

![](_page_41_Figure_2.jpeg)

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![](_page_41_Figure_4.jpeg)

![](_page_41_Picture_8.jpeg)

## My simulations let us map out the matter in a universe that would form after a given type of inflation.

![](_page_42_Figure_2.jpeg)

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# We can compare this to actual observations of the universe to check whether our inflation model is reasonable.

 $\eta_0$ 

### Spiderweb-like structure

20h

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![](_page_43_Figure_4.jpeg)

![](_page_43_Picture_5.jpeg)

# You might have seen news about NASA discovering very distant galaxies, further away and older than we expected. Non-Gaussianity could explain these!

f 🎽 🔊

![](_page_44_Picture_2.jpeg)

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### JWST smashes the record for the earliest galaxy

by Zili Shen | Dec 16, 2022 | Daily Paper Summaries | 5 comments

### Title: Discovery and properties of the earliest galaxies with confirmed distances

Authors: \*B. E. Robertson, \*S. Tacchella, B. D. Johnson, K. Hainline, L. Whitler, D. J. Eisenstein, R. Endsley, M. Rieke, D. P. Stark, S. Alberts, A. Dressler, E. Egami, R. Hausen, G. Rieke, I. Shivaei, C. C. Williams, C. N. A. Willmer, S. Arribas, N. Bonaventura, A. Bunker, A. J. Cameron, S. Carniani, S. Charlot, J. Chevallard, M. Curti, E. Curtis-Lake, F. D'Eugenio, P. Jakobsen, T. J. Looser, N. Lützgendorf, R. Maiolino, M. V. Maseda, T. Rawle, H.-W. Rix, R. Smit, H. Übler, C. Willott, J. Witstok, S. Baum, R. Bhatawdekar, K. Boyett, Z. Chen, A. de Graaff, M. Florian, J. M. Helton, R. E. Hviding, Z. Ji, N. Kumari, J. Lyu, E. Nelson, L. Sandles, A. Saxena, K. A. Suess, F. Sun, M. Topping, I. E. B. Wallace (\* equal contribution)

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![](_page_44_Picture_14.jpeg)

Figure 1: JWST NIRCam image of the GOODS-South field. b-e) are spectroscopically confirmed high-redshift galaxies. Reproduced from Fig 1 of the paper.

spectroscopy supernovae

Course Assignments Daily Paper Summaries Historical Astronomy Personal Experiences Undergraduate Research

Astrobitos (Astrobites in Astropontos (Astrobites in Portugese)

![](_page_44_Picture_20.jpeg)

# My simulations also let us make maps of the sky as they would be seen by certain telescopes.

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Credit: own work

Here, I'm showing a series of inflation models with increasing amounts of non-Gaussianity.

![](_page_45_Picture_5.jpeg)

# We cam compare simulations to observations from the Planck satellite

45°

200

These colours look different because a different "colour map" is used here.

Observations don't look exactly like simulations because actual observations pick up some accidental signals from things between us and the CMB. We can compare statistically though...

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![](_page_46_Figure_4.jpeg)

### Everything the maps of the sky have to tell us is in their statistics. They tell us how much stuff there is of each size. The observational data is:

![](_page_47_Figure_1.jpeg)

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Credit: Planck Collaboration (2015)

![](_page_47_Picture_4.jpeg)

### Everything the maps of the sky have to tell us is in their statistics. They tell us how much stuff there is of each size. We can compare with my simulations compare:

![](_page_48_Figure_1.jpeg)

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These types of inflation are definitely impossible!

These types of inflation are allowed!

Credit: own work.

![](_page_48_Picture_6.jpeg)

![](_page_48_Picture_7.jpeg)

![](_page_48_Picture_8.jpeg)

![](_page_48_Picture_9.jpeg)

# What we've learned today

The Cosmic Web is the biggest structure in the universe, it's made up of trillions of galaxies

2. Inflation is a period of rapid expansion very early in the universe. It stretches tiny quantum fluctuations into the seeds of galaxies.

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3. Different models of inflation leave measurable differences in the statistics of the universe.

4. By using these statistics in simulations of the Cosmic Web, we can test inflation models and see back into the depths of cosmic time.

![](_page_49_Picture_7.jpeg)

![](_page_50_Picture_0.jpeg)

![](_page_50_Picture_1.jpeg)

Ask me anything about cosmology!

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### Image Credit: Nathan J. Carlson

![](_page_50_Picture_5.jpeg)

![](_page_51_Picture_0.jpeg)

# Thank-you to our volunteers:

Organization:

James<br/>Nolan<br/>JacobEmma<br/>Jenny<br/>Aryanna<br/>AliciaMarkJenny<br/>Aryanna<br/>Alicia

Telescopes: Aryana Geetam Shokoofa Isabella Ethan

![](_page_51_Picture_5.jpeg)

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# **Photo and Video:** Nolan

# Demos: Jacob Isabela Diana Jenny

General: Laura Utkarsh Anupa Louis Raina

![](_page_52_Picture_1.jpeg)

David A. Dunlap Department of Astronomy & Astrophysics UNIVERSITY OF TORONTO

![](_page_52_Picture_3.jpeg)

![](_page_52_Picture_5.jpeg)

instagram.com/uoftastrotours

youtube.com/c/UofTAstroTours

# AstroTours is sponsored by:

![](_page_52_Picture_9.jpeg)

### Dunlap Institute for Astronomy & Astrophysics UNIVERSITY OF TORONTO

Canadian Institute for Theoretical Astrophysics

L'institut Canadien d'astrophysique théorique

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![](_page_52_Picture_14.jpeg)

astro.utoronto.ca/astrotours

![](_page_52_Picture_16.jpeg)

tours@astro.utoronto.ca

![](_page_52_Picture_18.jpeg)

![](_page_53_Picture_0.jpeg)

Go to the McLennan Physical Laboratories lobby and roof to explore our different interactive activities, they will take place until 10:00pm EST.

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![](_page_53_Picture_3.jpeg)

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# Visit our astronomy activities and the telescopes!

# Activities will include:

- Telescopes
- Tactile Astronomy
- World Wide Telescope
  - 3D printed telescopes

![](_page_54_Picture_0.jpeg)

# Visit our astronomy activities and the telescopes!

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![](_page_54_Picture_4.jpeg)