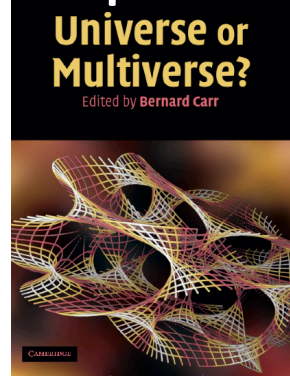
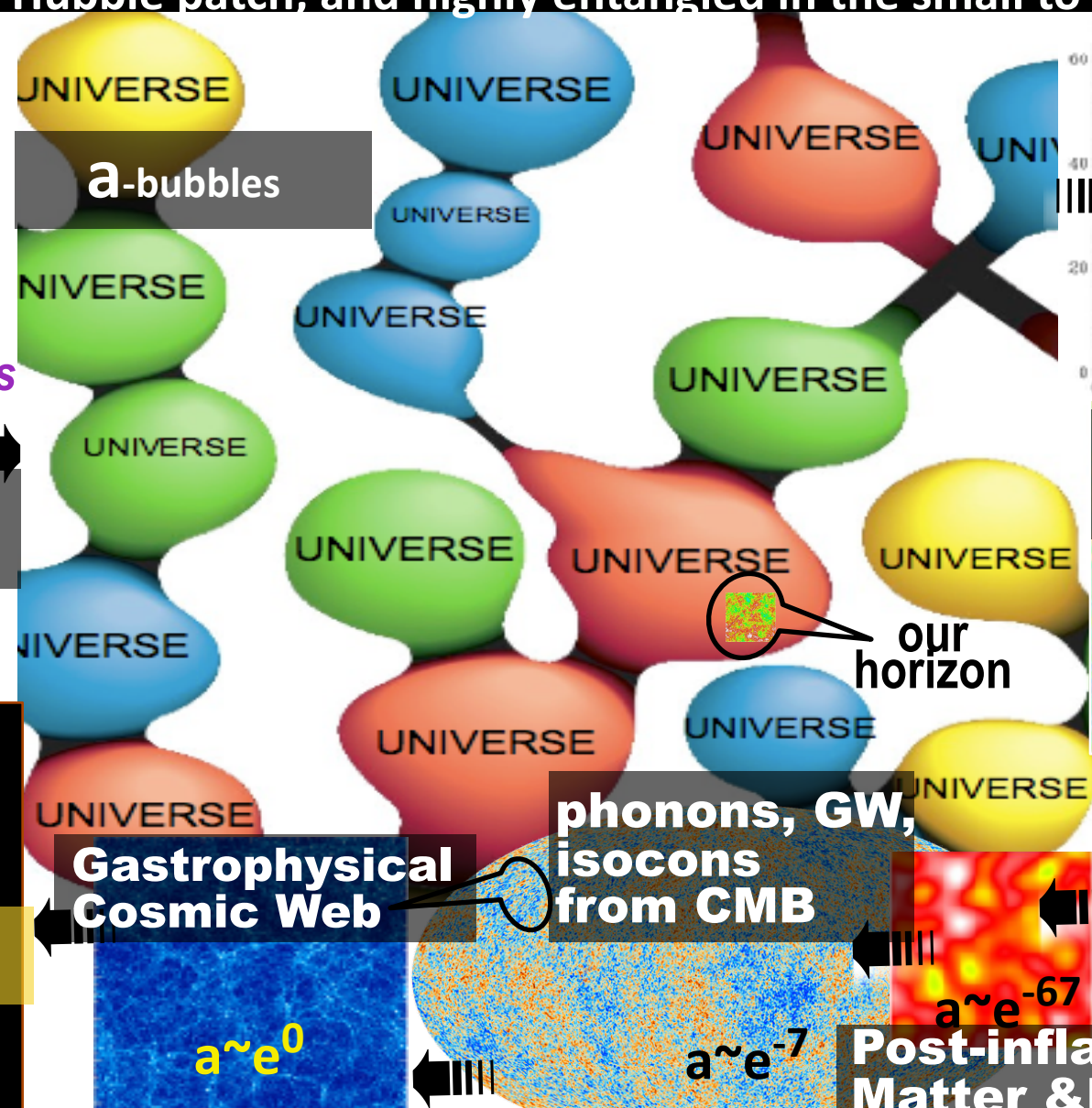


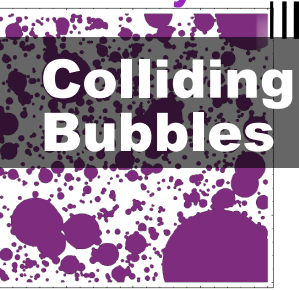
Dick Bond CITA & CIFAR: Aspects of Inflation ζ -Phenomenology

SuperWeb of ultra-Ultra Large Scale Structure of the Universe

a highly strained & stressed state in the universe at large (*very, very*), randomly simple in our Hubble patch, and highly entangled in the small to medium scale



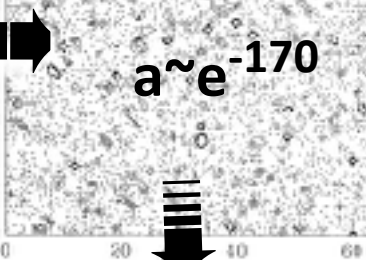
quantum tunnels = bubbly-U



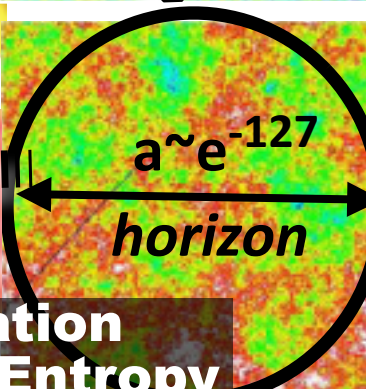
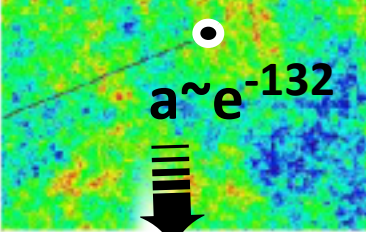
END
a future DE-Void

Dark Energy Trajectories
 $a \sim e^{+++}$

a-bubbles

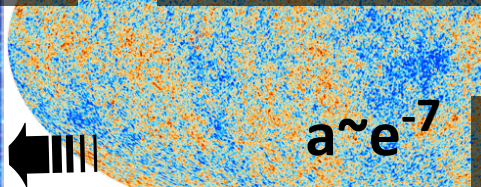
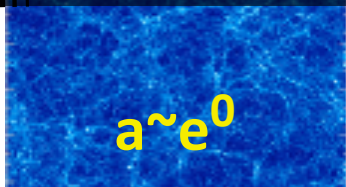


Phonons Isocons Gravitons



Gastrophysical Cosmic Web

phonons, GW, isocons from CMB

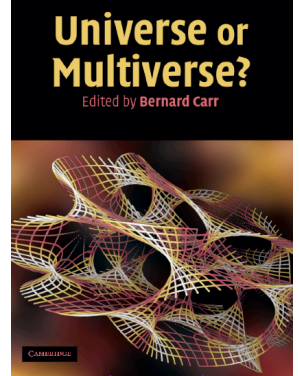


Post-inflation Matter & Entropy

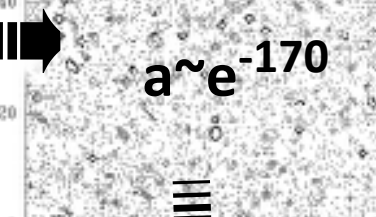
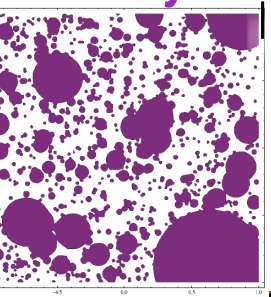
Dick Bond CITA & CIFAR: Aspects of Inflation ζ -Phenomenology

SuperWeb: a highly strained & stressed state at large (*very*), randomly simple in our Hubble patch, and (*now*) highly entangled in the small to medium scale

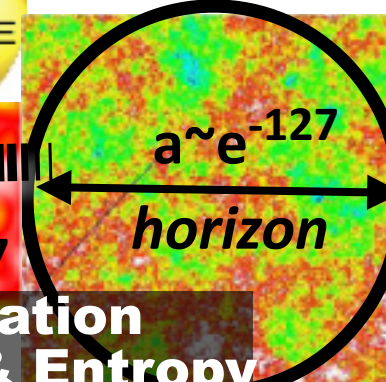
topography of the ζ -scape = entropy -scape



quantum tunnels = bubbly-U



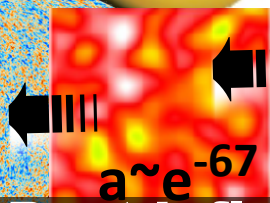
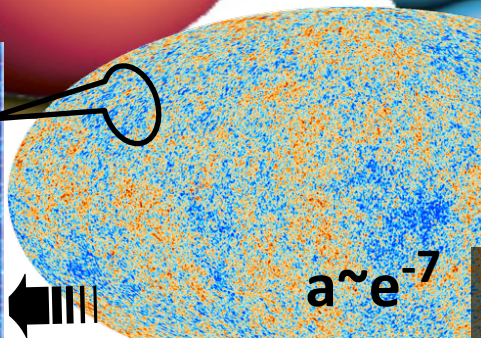
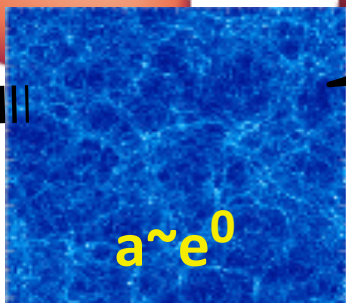
Phonons
Gravitons
 ζ -maps
trajectories



our horizon

END
a future DE-Void

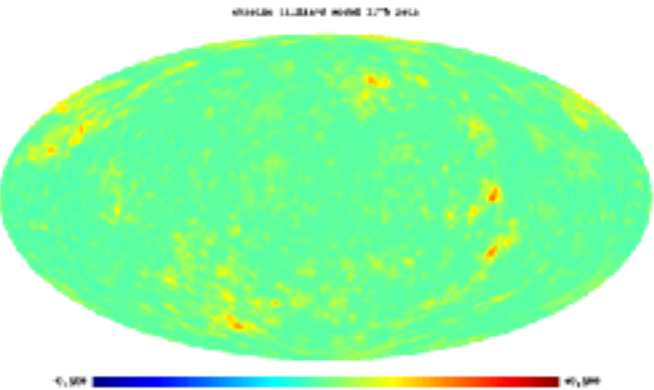
Dark Energy / modified gravity Trajectories
 $a \sim e^{+++}$



Post-inflation Matter & Entropy

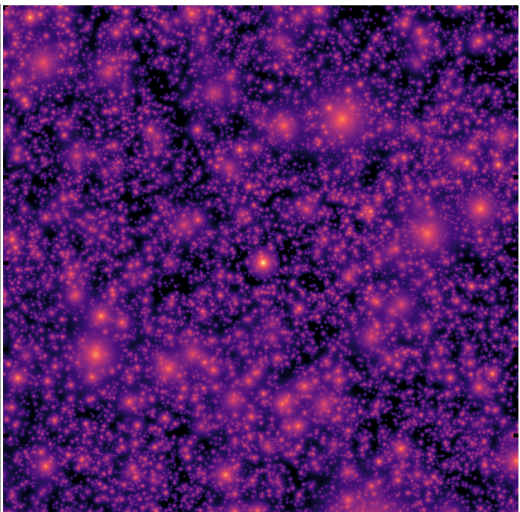
ζ & entropy & adiabatic trajectories stochastic “coarse grain” S ballistics caustics corrugated shock-in-time => S intermittent nonG

CMB: std inflaton ζ +
 subdominant uncorrelated ζ from
 modulated preheating



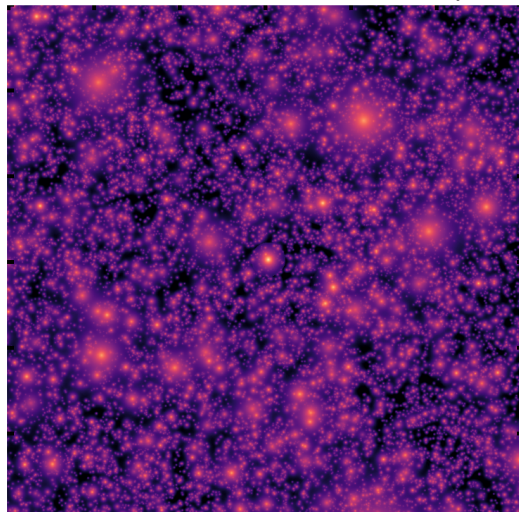
BBFH, b+braden+frolov+huang

LSS tSZ: Gaussian std



ABSB+FH, alvarez+b+stein+frolov+huang

LSS tSZ: Gaussian std +
 subdominant uncorrelated ζ



the Super-WEB aka the primordial 3-curvature web aka the phonon/isotropic strain= volume deformation web

$$\ln \rho(\mathbf{x}, t) / \langle \rho \rangle |_{\mathbf{v}} \quad \ln V / \langle V \rangle |_{\rho} = 3 \ln \mathbf{a}(\mathbf{x}, t) / \langle \mathbf{a} \rangle |_{\rho} \quad d\zeta \sim T dS / 3(E+PV)$$

$$\zeta(\mathbf{x}, t) = \int (dE+pdV) / E \quad / \langle 3(1+p/\rho) \rangle (t) \quad \text{BST83, SBB89, SB90,91, B95, Bond+Braden2017 } \zeta \text{ for preheating}$$

$$\zeta(\mathbf{x}, t) = \ln \rho(\mathbf{x}, t) / \langle 3(1+p/\rho) \rangle (t) + \int (1+p/\rho)(\mathbf{x}, t) d \ln \mathbf{a}(\mathbf{x}, t) / \langle 1+p/\rho \rangle (t)$$

$$\text{or: } \zeta(\mathbf{x}, t) = \ln \rho(\mathbf{x}, t) / \rho_b / 3(1+p_b/\rho_b) + \ln \mathbf{a}(\mathbf{x}, t) / a_b$$

gradient / Morse flow +stochastic jitter, simple Hamilton principle function $S \sim H(\phi_b)$

along coarse-grain trajectories $d\zeta = d \ln \rho / \rho_b / 3(1+p/\rho) + d \ln \mathbf{a} / a_b = [d \bar{\zeta}] (fg \rightarrow cg)$

early preheating: gradient / Morse flow, complicated Hamilton principle function S

ballistic /caustic phase => ΔS nonlinear ζ lattice sims

cf. late-time density web ~ strain web - $\ln \rho / \langle \rho \rangle = \text{Trace } \ln \mathbf{e}_J^J = \ln V / \langle V \rangle |_{\rho}$

cold $\langle p/\rho \rangle \sim 0 \Rightarrow \zeta(\mathbf{x}, t | cdm)$ conserved before shell crossing (preheating)

Reconstructing the Early Universe

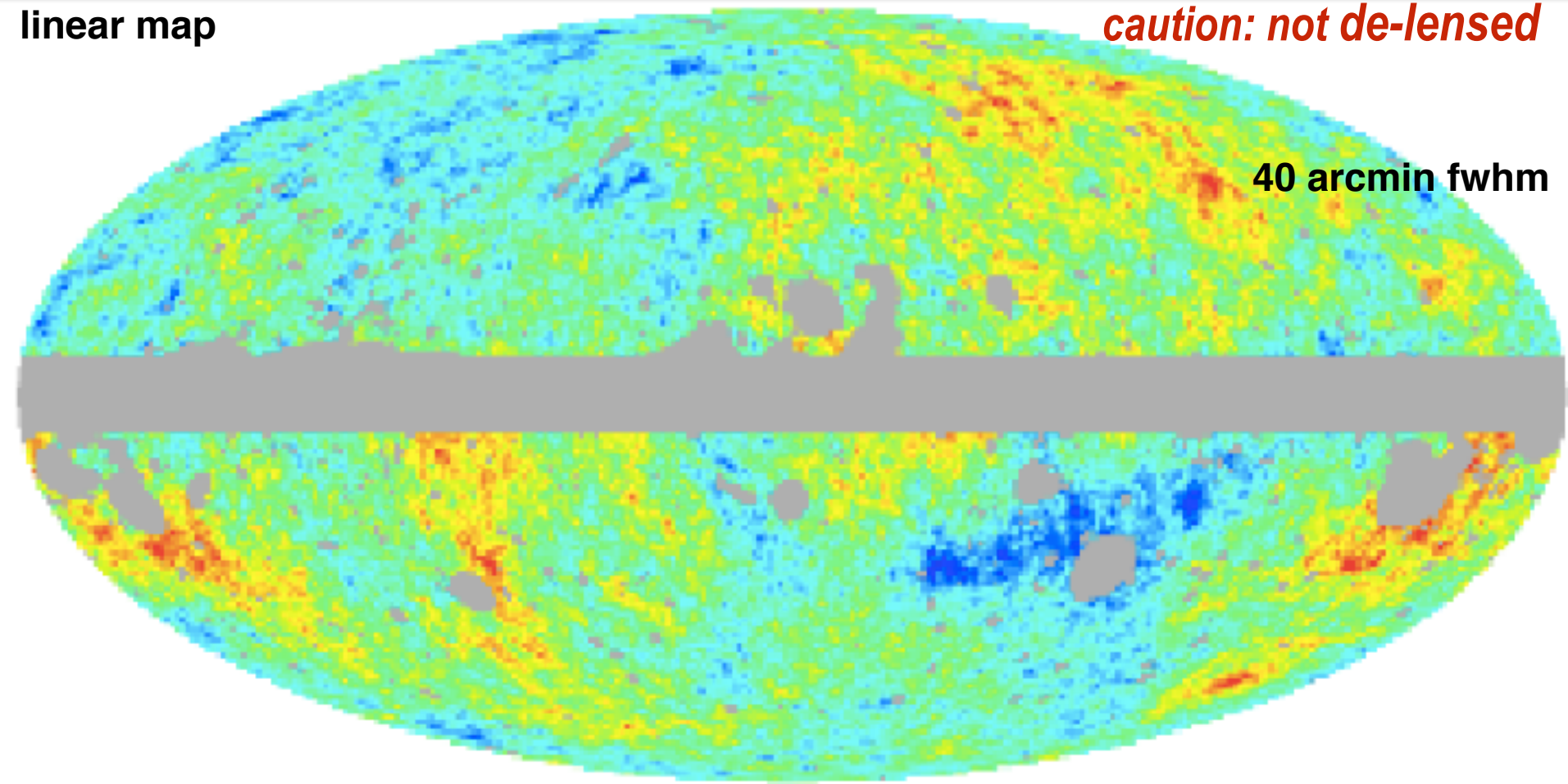
$$\int d\text{visibility}(\text{distance}) \langle \zeta | \text{Temp}, E \text{ pol} \rangle \quad (\text{angles, distance})$$

$$\text{sb89, bb15 } \zeta_{NL} = \ln(\rho a^{3(1+w)}) / 3(1+w) \leftarrow dE + pdV \sim d\text{Entropy} \quad \text{phonons / strain}$$

linear map

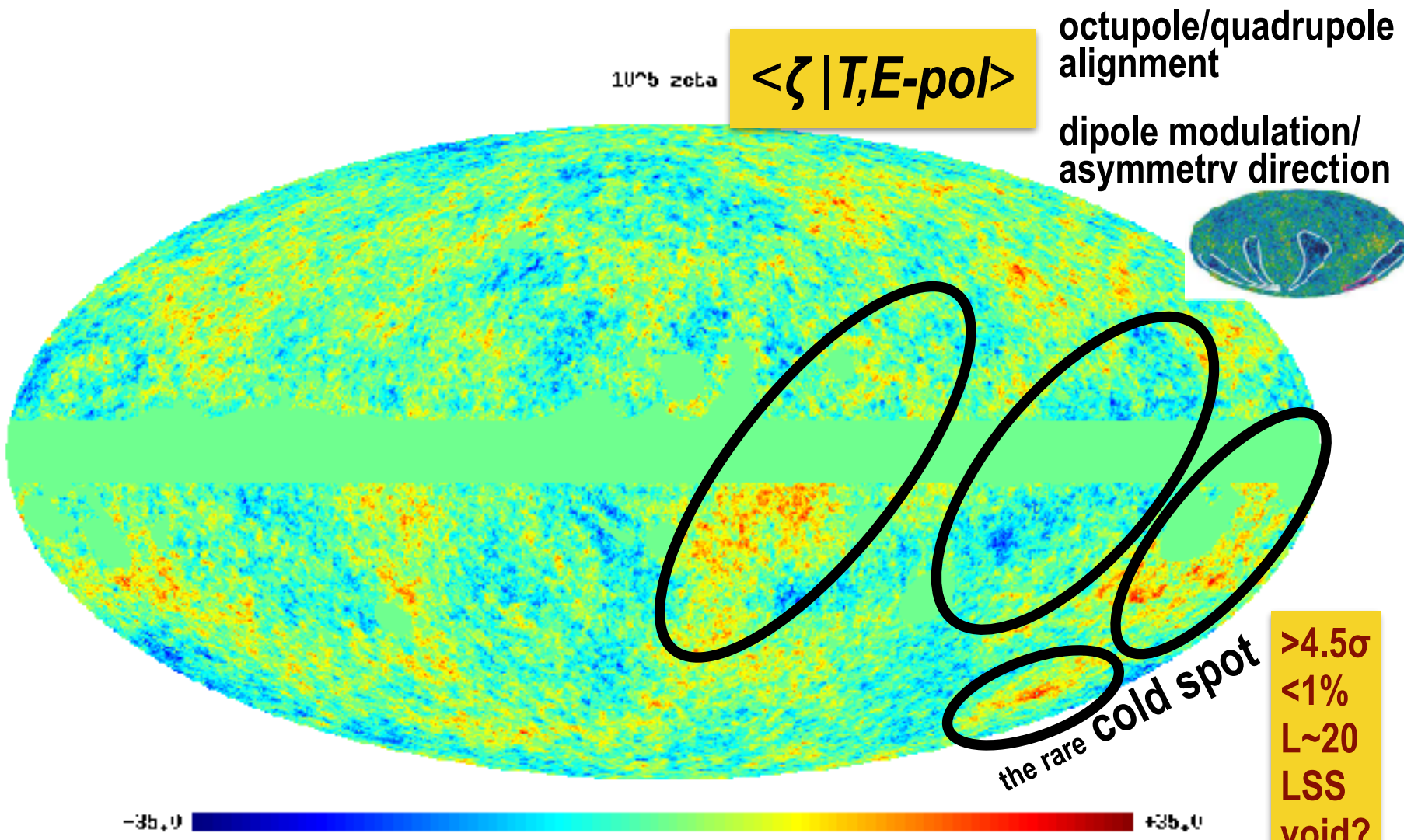
caution: not de-lensed

40 arcmin fwhm



visibility mask

BSMc = SMc + primordial anomalies



ζ -Maps of the Early Universe

a Map is an ensemble = mean-map + fluctuation-maps, encoding correlated errors

Maps = (radical) **compressions** of the **time ordered information Tol** onto a parameterized space q^A : **Linear maps, Quadratic maps (power), cosmic parameter maps**
 $\text{Prob}(q | \text{Data}, \text{Th prior}) \Rightarrow \langle q^A | D, \text{Th} \rangle, \langle \Delta q^A \Delta q^B | D, \text{Th} \rangle, \dots$ or $q_{\text{max}L}$

TOPOGRAPHY & CARTOGRAPHY

of our Hubble-patch aka our bit of the universe

reconstructing $\zeta = \ln a(\mathbf{x}, t)$ @uniform density,
aka primordial **scalar curvature** ${}^{(3)}R = -4 {}^{(3)}\Delta \ln a$

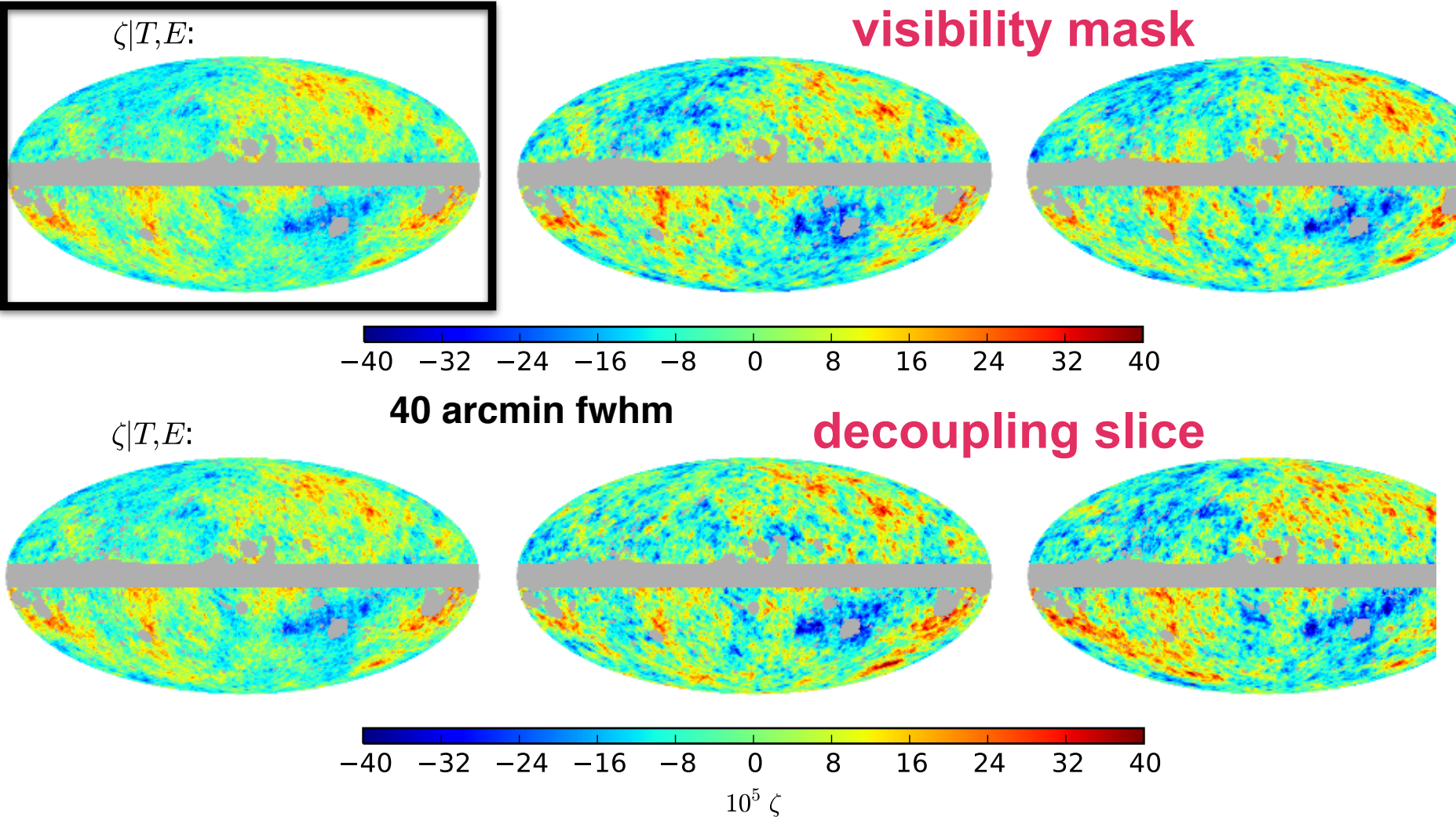
Wiener-filtered ζ maps make $\zeta_{LM}(\chi)$, $\chi = |\mathbf{x}|$ instead of $\zeta(\mathbf{x})$

$\mathbf{T}_{LM c,s} / \mathbf{E}_{LM c,s} = \int \mathbf{e}_{\zeta}^{T/E} L_{\chi c,s} \zeta_{LM c,s}(\chi) d\chi$, susceptibility \mathbf{e} depends on cosmic parameters

\Rightarrow Linear response $\zeta_{LM c,s}(\chi) = \mathbf{e} * \zeta^T L_{\chi c,s} \mathbf{T}_{LM c,s} + \mathbf{e} * \zeta^E L_{\chi c,s} \mathbf{E}_{LM c,s} + \delta \zeta_{LM c,s}$

susceptibility of ζ to T/E : $\mathbf{e} * \zeta^{T/E}$ interpolates T/E to ζ , if no info relax to $\delta \zeta$

project ζ to minimize fluctuations: $\int d\text{visibility}(\text{distance}) (\langle \zeta | \text{Temp}, E \text{ pol} \rangle + \delta \zeta)$



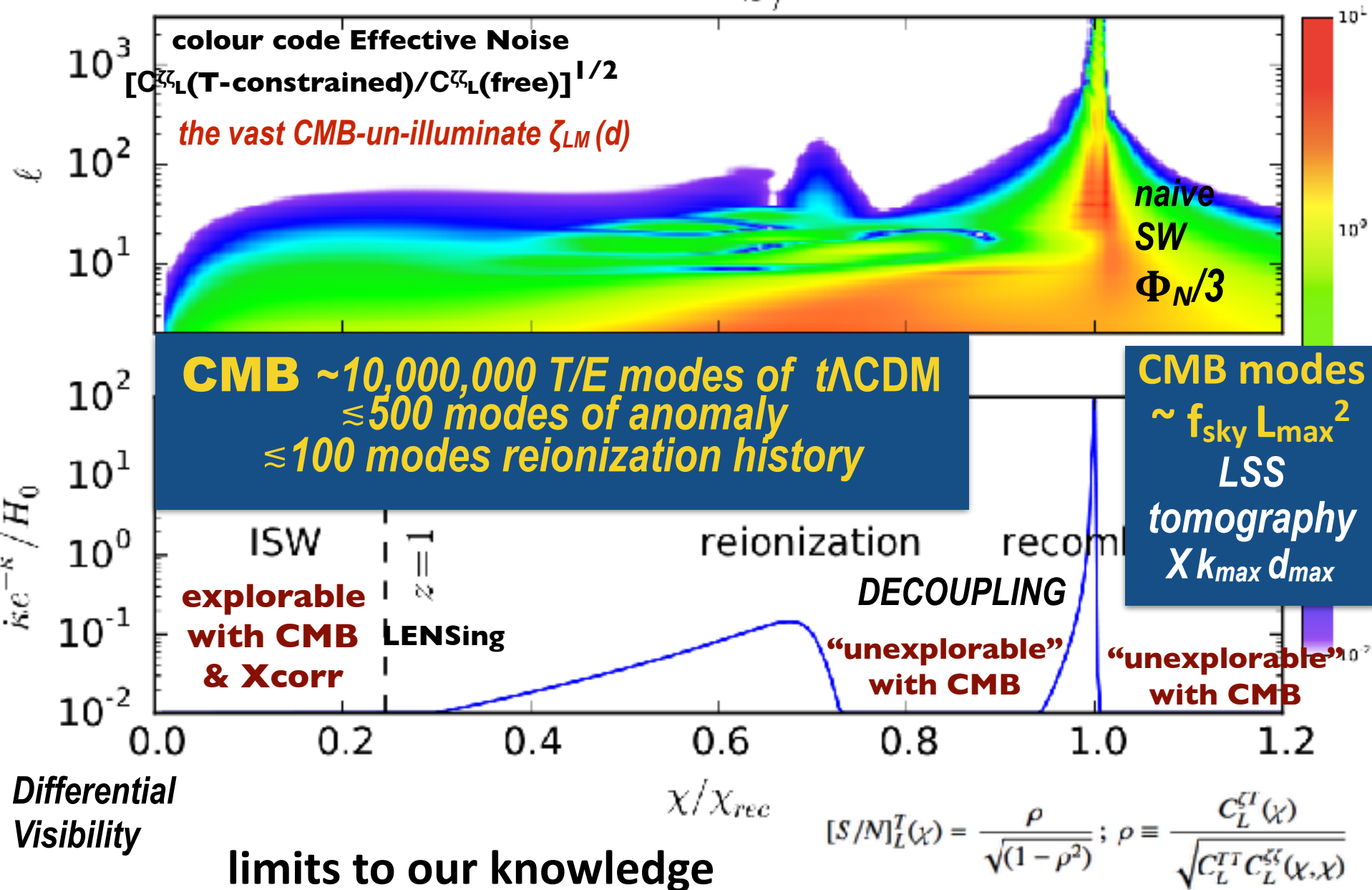
allowed fluctuations are less noisy with E pol (extra mode/LM)

caution: not de-lensed, but the Wiener filter does partially de-lens

$$\langle \zeta_{LM}(\chi) | T_{LM} E_{LM} \rangle$$

the unexplorable ζ -scape,
 explore with landscape++ ideas
 our Hubble Bit will reveal all?

$T + E S/N$



Reconstructing the Early Universe

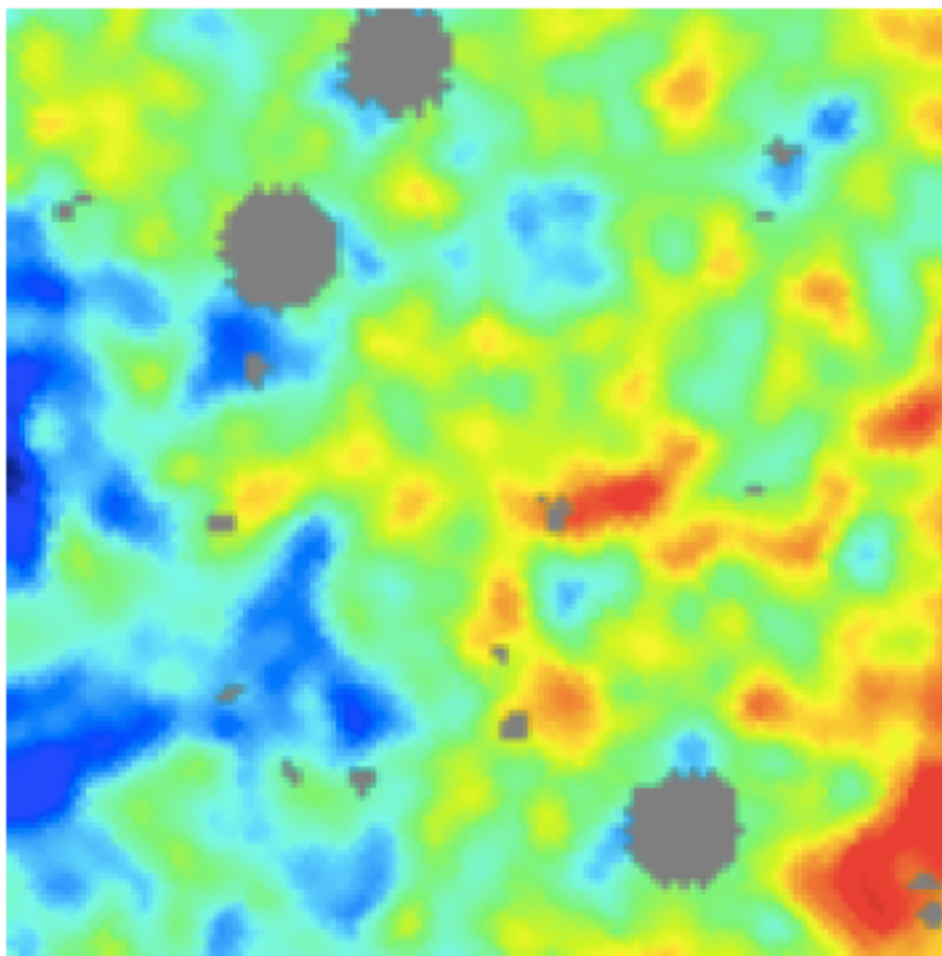
$\int d\text{visibility}(\text{distance}) \langle \zeta | \text{Temp}, E \text{ pol} \rangle$ (angles, distance)

sb89, bb15 $\zeta_{NL} = \ln(\rho a^{3(1+w)}) / 3(1+w) \leftarrow dE + pdV \sim d\text{Entropy}$ phonons / strain

linear map

$\zeta | T, E:$

caution: not de-lensed

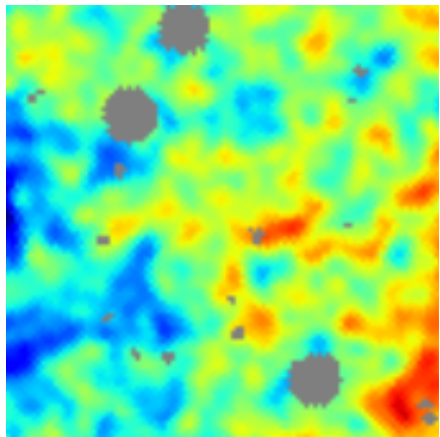


20x20 sq deg

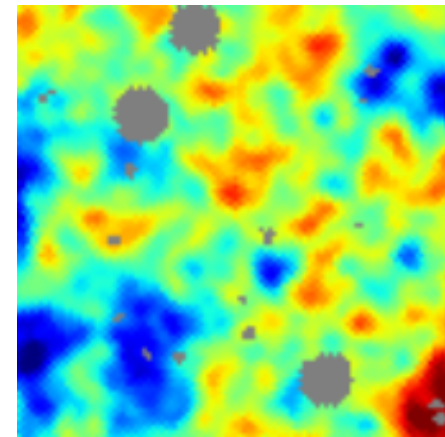
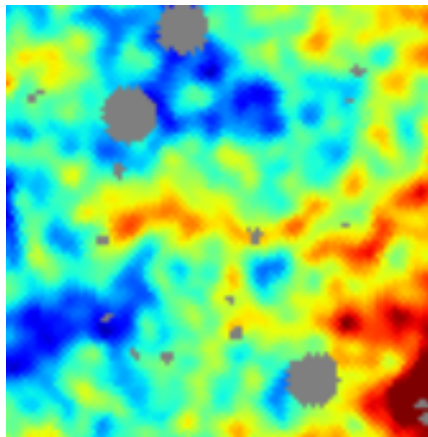
20 arcmin fwhm

visibility mask

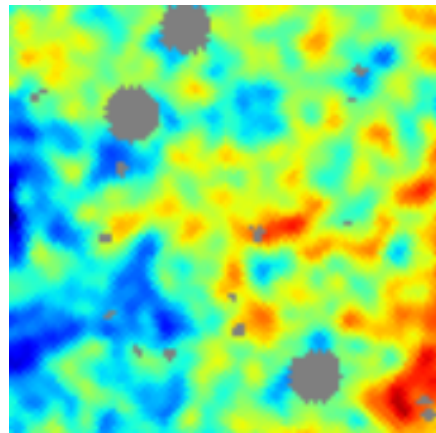
$\zeta|T,E:$



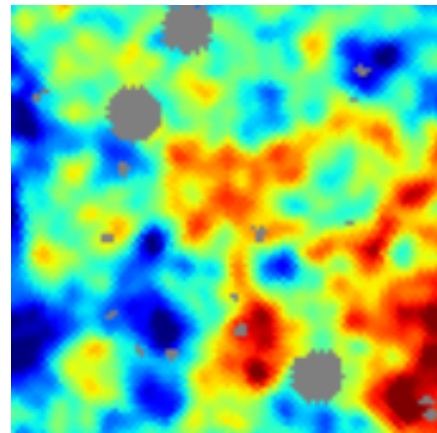
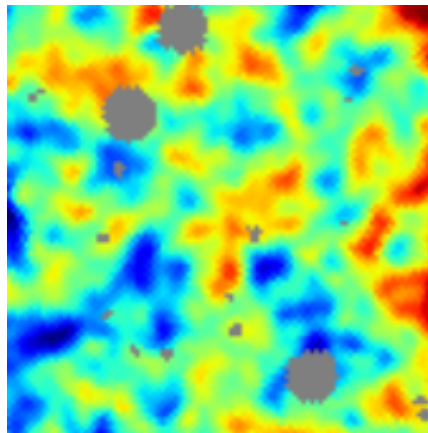
visibility mask



$\zeta|T,E:$



decoupling slice



20 arcmin fwhm

20x20 sq deg

allowed fluctuations are less noisy with E pol (extra mode/LM)

caution: not de-lensed, but the Wiener filter does partially de-lens

SIMPLICITY

at $a \sim e^{-7} \sim 1/1100 \Rightarrow$

at $a \sim e^{-67-60} \sim 1/10^{30+25}$

stacked linear map aka
mean-field map

stacked
 $\langle \zeta_{dv} | \zeta_{dv-pk} \rangle$

Planck2015 early U structure map

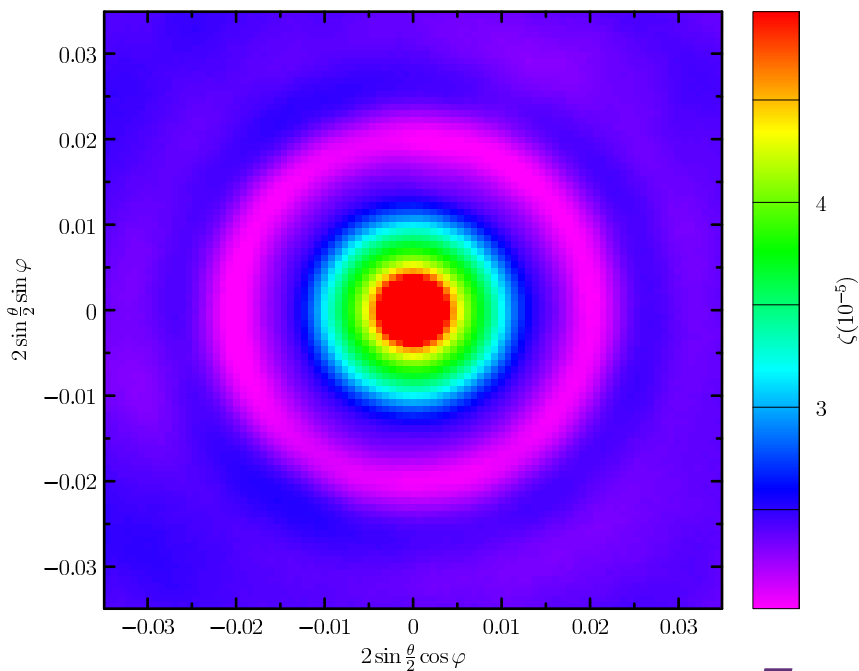
reveals *primordial sound waves in matter*

\Rightarrow learn **contents & structure** at 380000 yr, $a \sim e^{-7}$

\Rightarrow infer the sound structure far far earlier $a \sim e^{-67-60}$

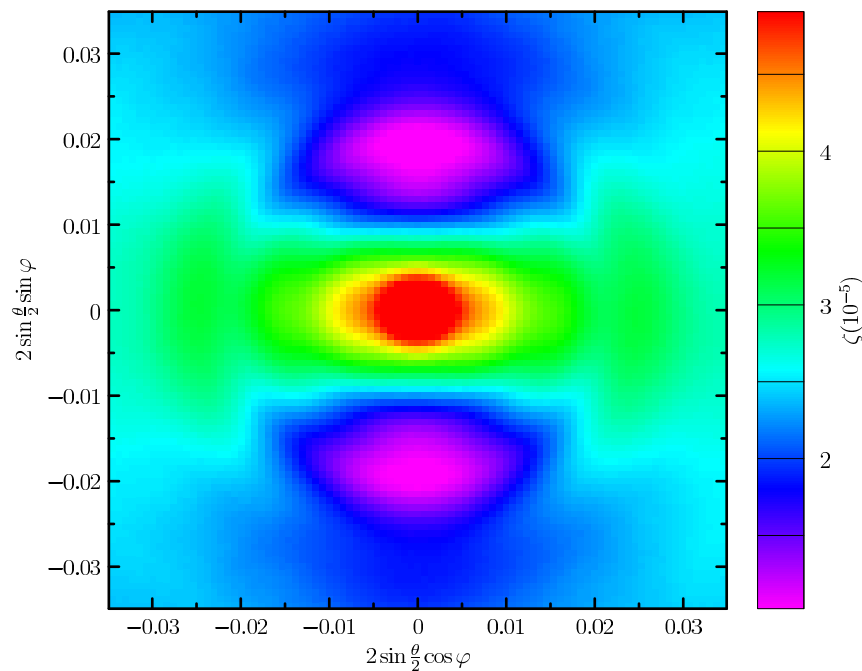
stacked **2⁺ numbers**
 $\langle \zeta_{dv} | \text{oriented } \zeta_{dv-pk} \rangle$

20857 patches on ζ maxima, random orientation, threshold $\nu=0$



BFH, b+frolov+huang

20854 patches on ζ maxima, oriented, threshold $\nu=0$



ζ stacks of P13 & WMAP9 look similar
simulations look very similar

quadratic map of the ζ -scape

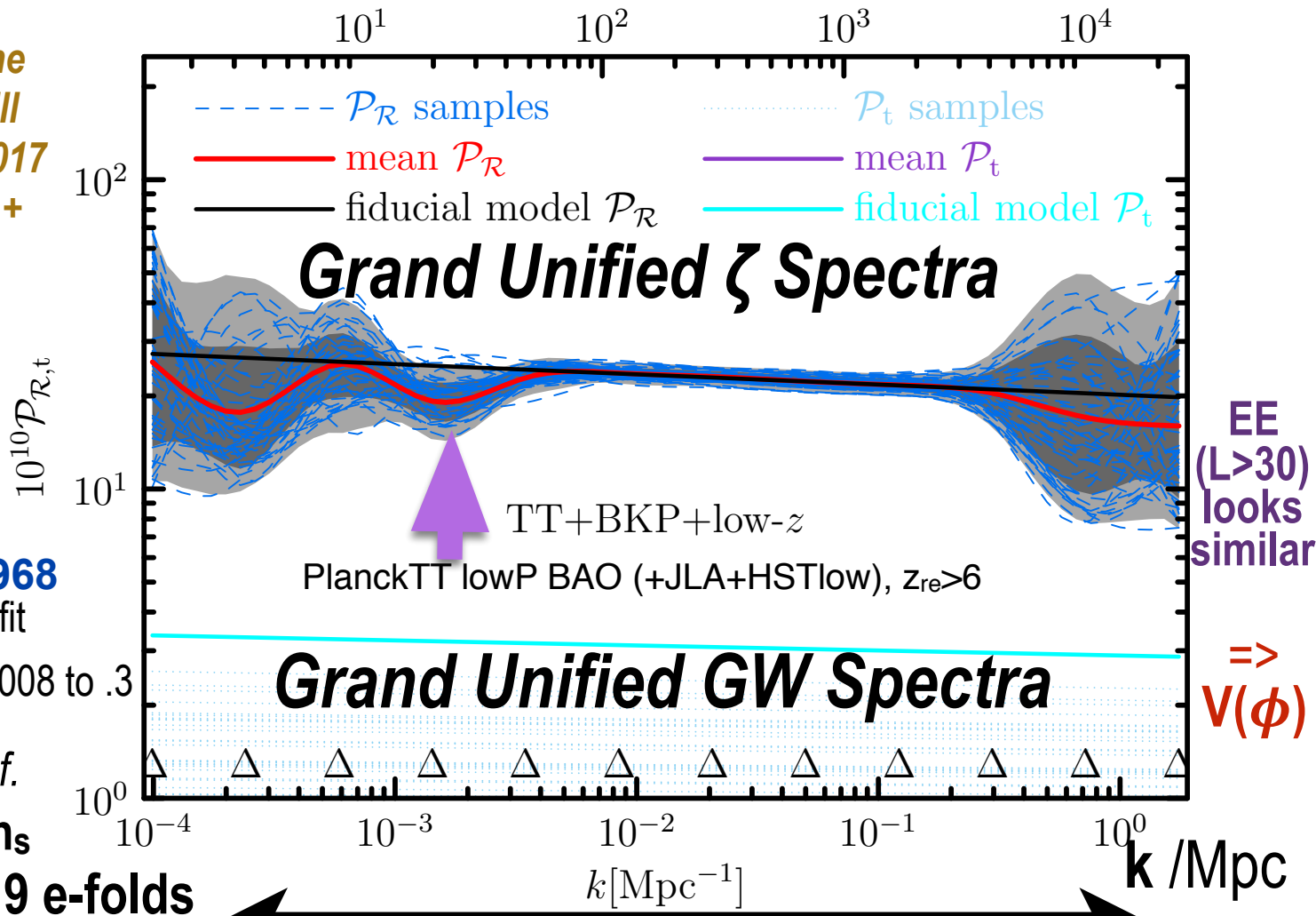
even more (radical) **compression** in quadratic space, using Planck likelihood rather than linear ($\langle \zeta | \text{Temp}, E \text{ pol} \rangle + \delta\zeta$) maps, e.g., onto 12 bands in k -space (LM projection)

=> a quadratic map, fully includes lensing & BB from BKP

$$\ell_k \equiv k D_{\text{rec}}$$

$$k D_{\text{rec}} \gtrsim L$$

the exploration of the $L=20-30$ anomaly will improve in Planck2017 + BICEP/KECK2017 + Spider 201x



uniform $n_s = 0.968$
P15+LSS best fit

superb 12-knot fit $k \sim .008$ to $.3$

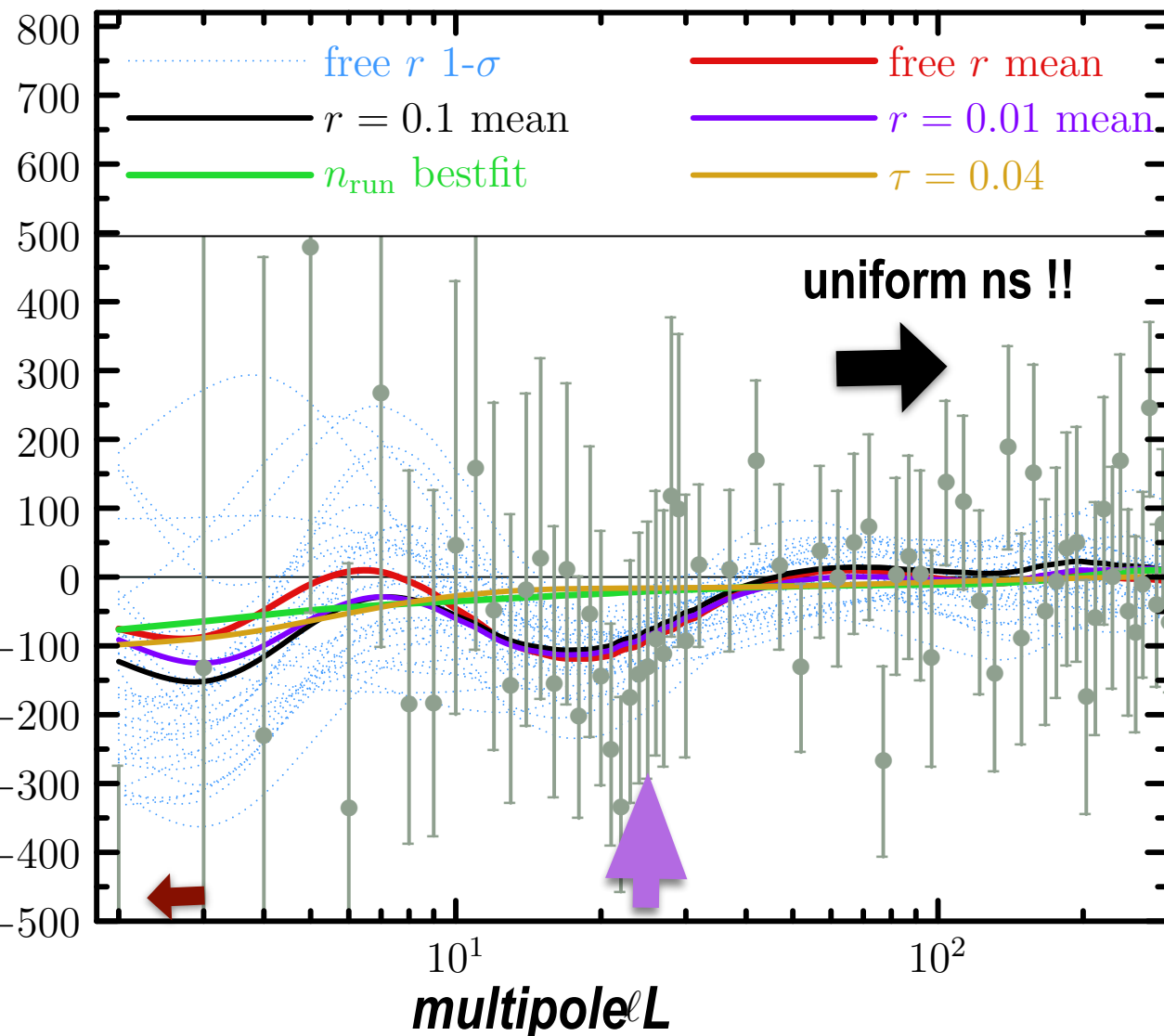
$r < .11$ 95%CL cf.
 $r < 0.09$ uniform n_s

trajectories of $\mathcal{D}_{TT,L}$

cf. Planck 2014 Commander Low L spectrum with Blackwell-Rao errors

12 knots, cubic spline

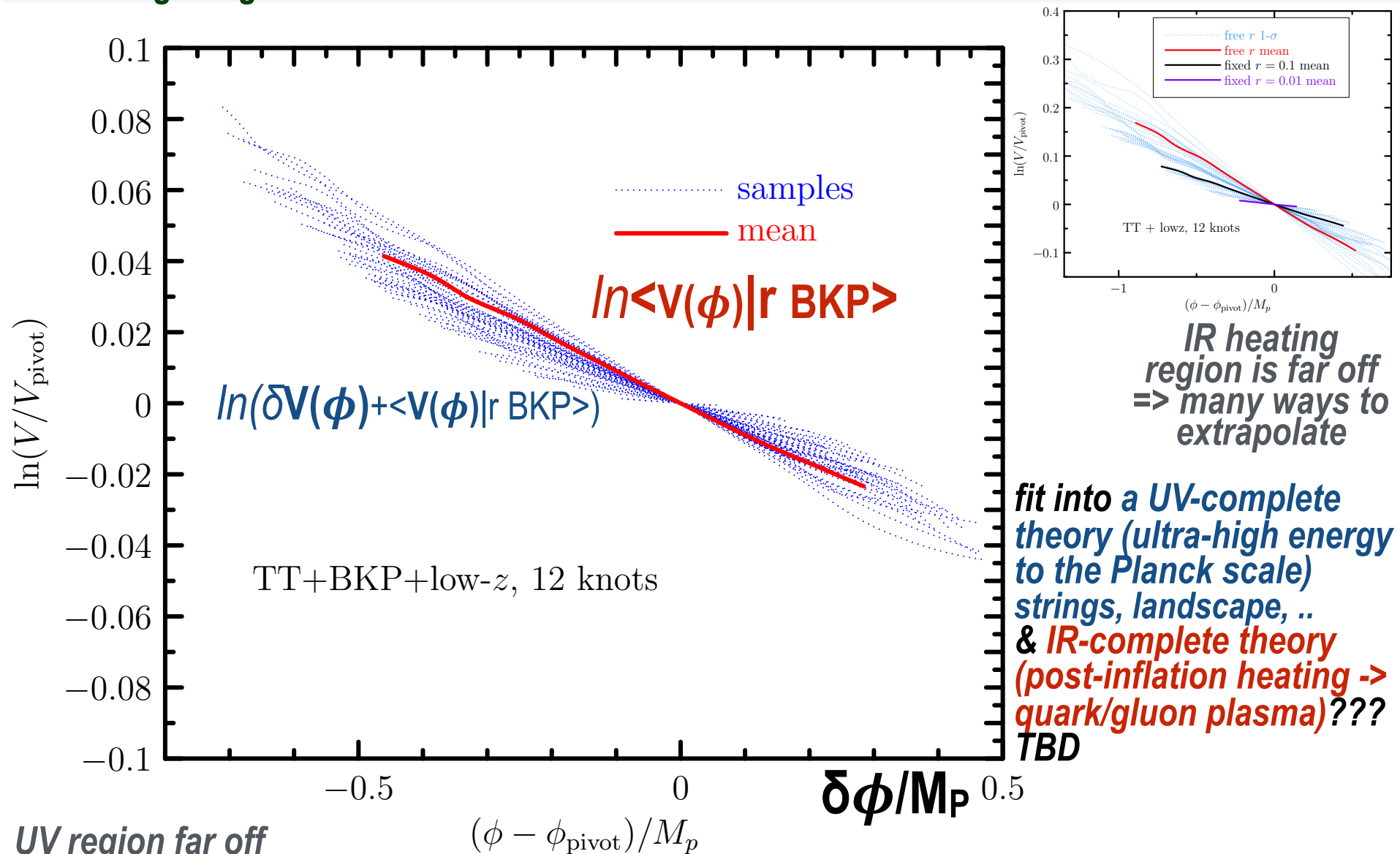
$\Delta\mathcal{D}_{TT,L}$



running of \mathcal{P}_ζ
 \equiv 3 Chebyshev modes
 \Rightarrow very stiff
 \Rightarrow not what the data wants
 Lower $\tau \Rightarrow$ shape similar to
 running at low L
 similar response on $\mathcal{D}_{TT,L}$
 for constrained & free r
 modified by τ freedom

running of \mathcal{P}_ζ
NOT wanted
 the down-up-down
 tendency
 is here to stay,
 2014-2022-...

**inflaton $V(\phi)$ -maps $= 3M_P^2 H^2 (1-\epsilon/3)$ HJ eqn, $d\phi/M_P/d\ln a = \pm \sqrt{2\epsilon}$
 along the gradient / Morse flow**



UV region far off
 \Rightarrow many ways to extrapolate

r to ± 0.02 Spider forecast

r to ± 0.003 AdvACTpol forecast w/ fgnds

early universe **acceleration histories = EOS histories** $3(1+w)/2$

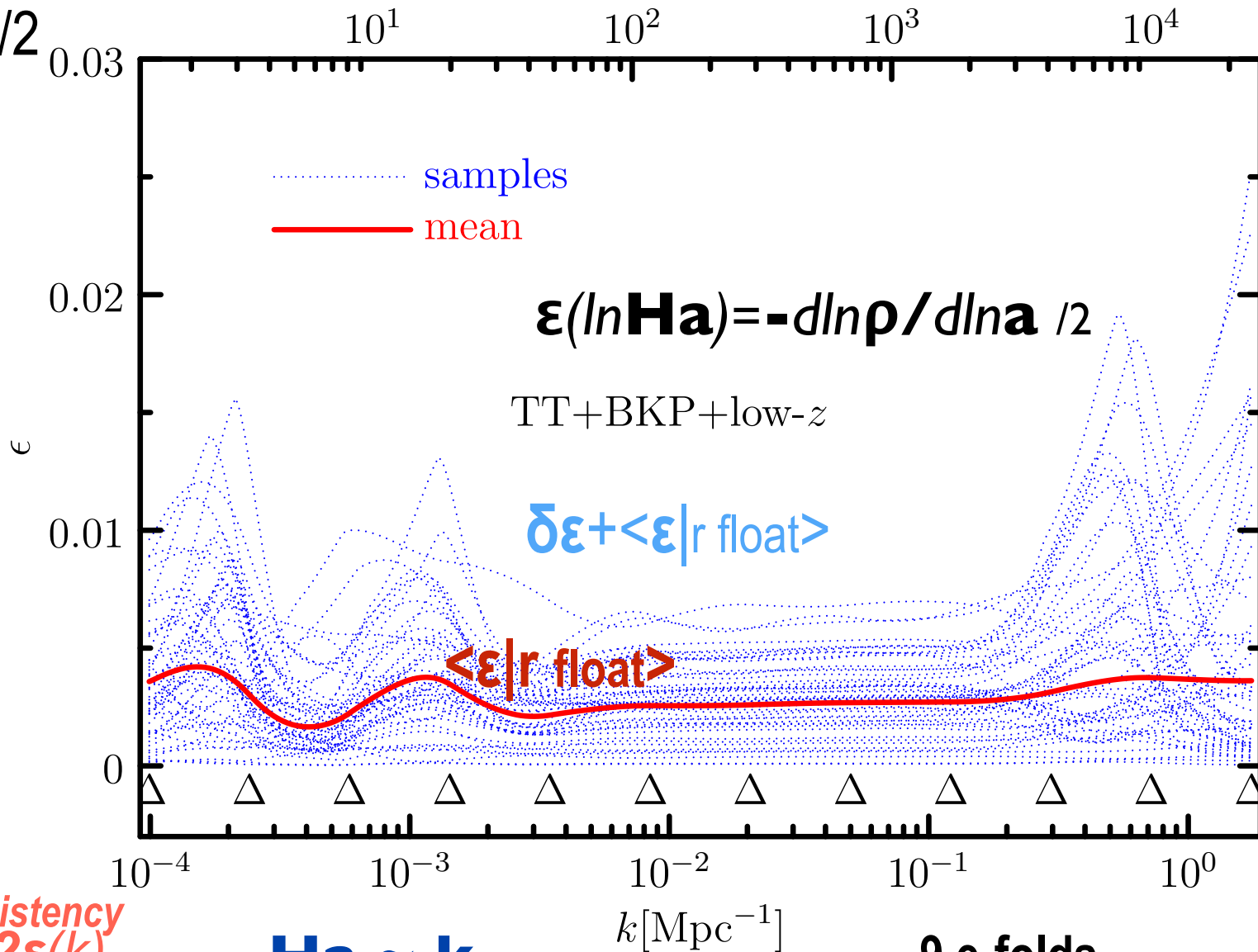
Planck 2015+BKP+LSS

$$\ell_k \equiv kD_{\text{rec}}$$

$$\boldsymbol{\varepsilon} = 3(1+w)/2$$

$$\approx r(k)/16$$

$$\sim \mathcal{P}_{\text{GW}}/\mathcal{P}_{\zeta}$$



inflation consistency
 $-n_t \approx r/8 \approx 2\varepsilon(k)$
 $1-n_s \approx 2\varepsilon + d \ln \varepsilon / d \ln H a$

Ha ~ k

9 e-folds

Will any

Anomalies in the CMB or Tensions with the CMB

turn into
BSMc Subdominant Physics?

Planck2015+LSS some tension released. still H_0 tension but not bad agreement+a bright future

Galaxy Lensing tension persists, systematics? CMB lensing A_L

Cluster σ_{8SZ} cf $\sigma_{8primary}$ tension relaxing, with large $KE_{bulk}/KE_{thermal}$ corrections, hydro expected tho

Beyond the Standard Model of cosmology? $SM_c = \text{tilted } \Lambda\text{CDM} + r(\zeta, h_{+x})$

$BSM_c = SM_c + \text{primordial anomalies}$

$\sim 10,000,000$ T/E modes = $t\Lambda\text{CDM}$, $\lesssim 500$ modes of anomaly

vast unexplored parts of the ζ -scape CMB is 2D

hope to use 3D **LSS** tomography $f_{\text{sky}} L_{\text{max}}^2 k_{\text{max}} d_{\text{max}}$

**CMB TT power $L \sim 20-30$ dip \Rightarrow
Grand Unified ζ -Spectrum k-dip**

$\langle \zeta | T, E\text{-pol} \rangle$

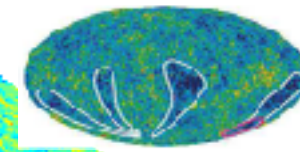
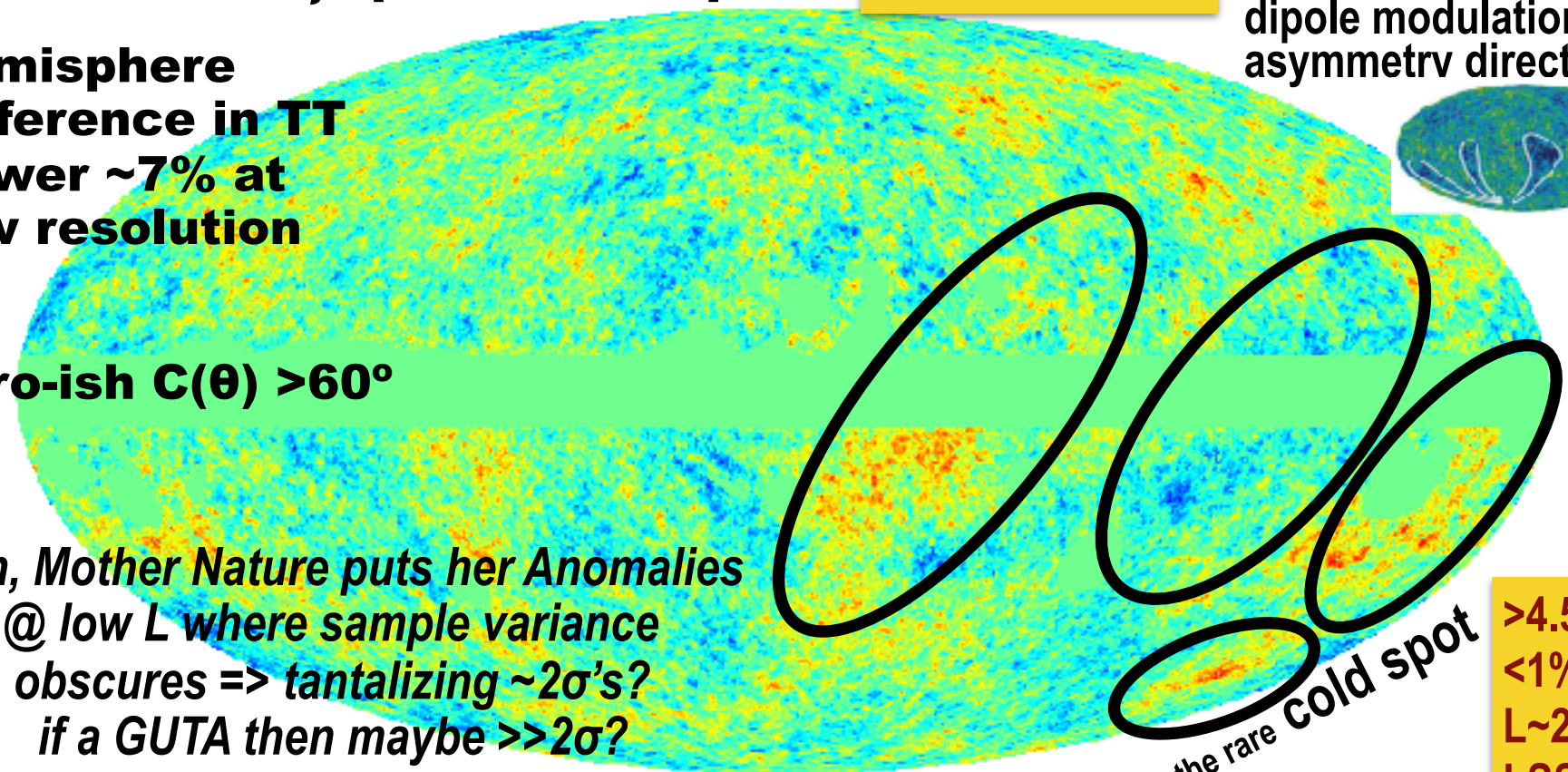
octupole/quadrupole alignment

dipole modulation/
asymmetry direction

**hemisphere
difference in TT
power $\sim 7\%$ at
low resolution**

zero-ish $C(\theta) > 60^\circ$

*sigh, Mother Nature puts her Anomalies
@ low L where sample variance
obscures \Rightarrow tantalizing $\sim 2\sigma$'s?
if a GUTA then maybe $\gg 2\sigma$?*



the rare cold spot

$> 4.5\sigma$
 $< 1\%$
 $L \sim 20$
LSS
void?

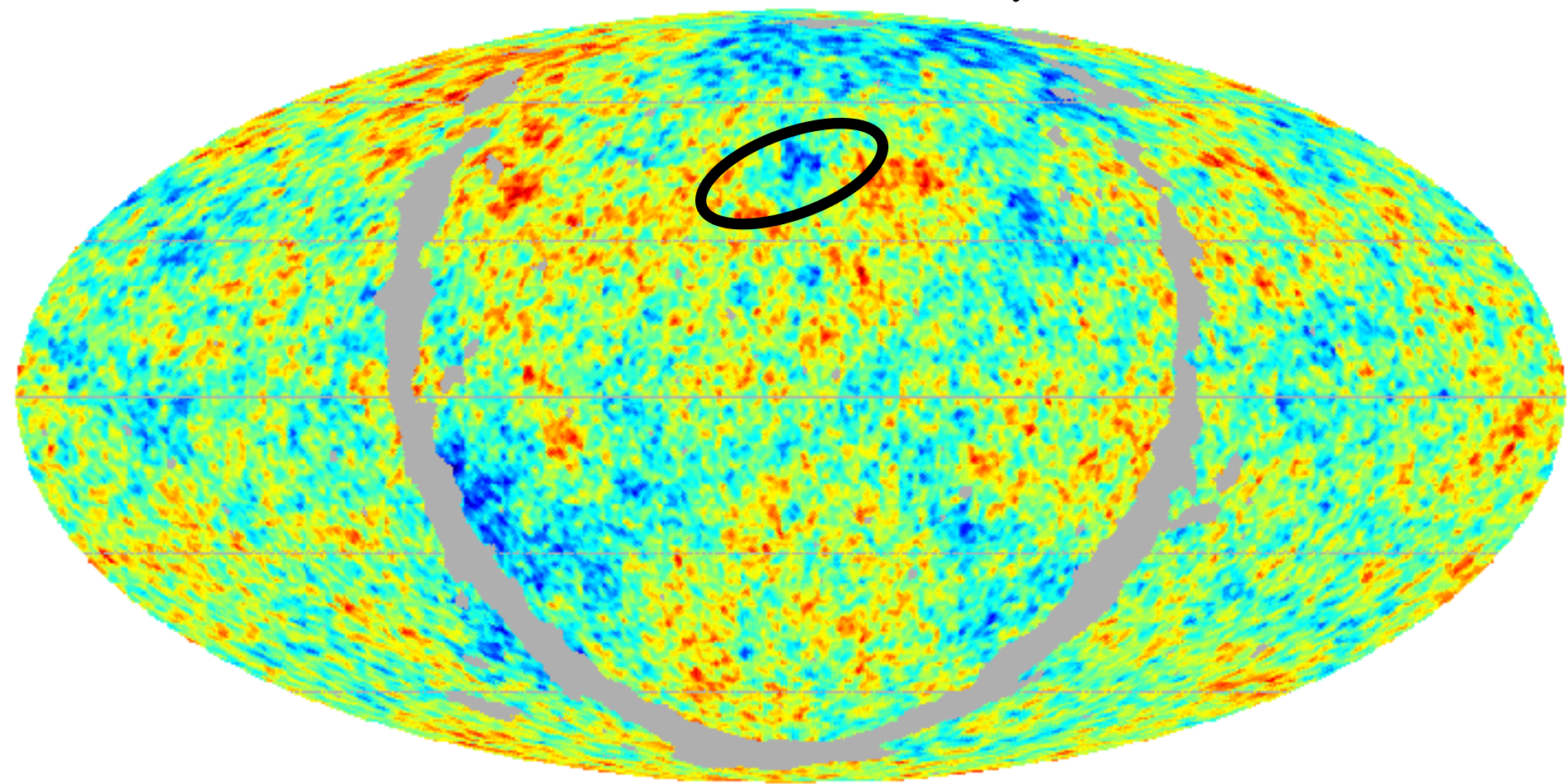
GUTA = Grand Unified Theory of Anomalies? TBD **intermittent?**

looking at the CMB cold spot again as an anomaly example

>4.5 σ <1% L~20 LSS void?

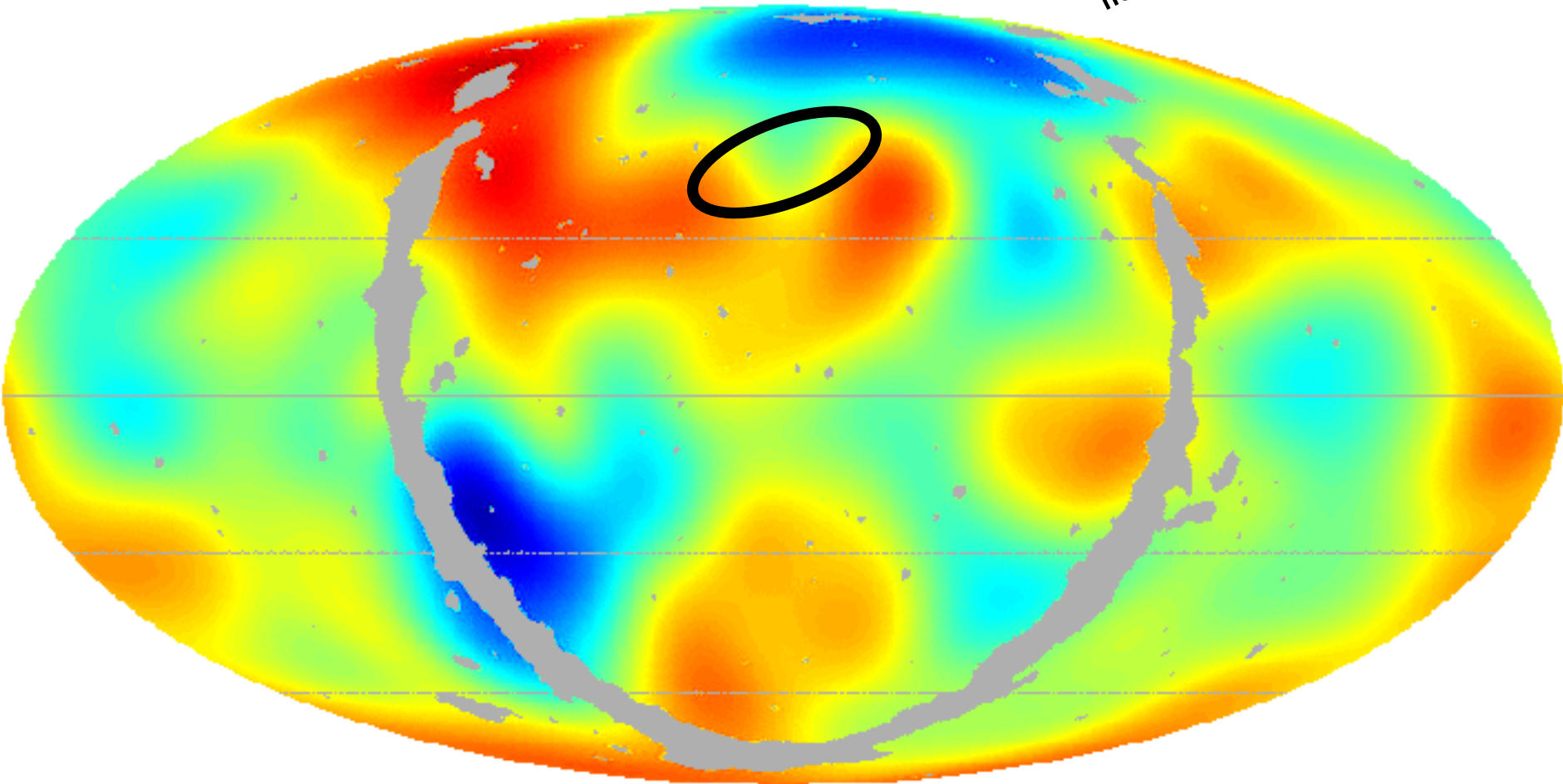
B+Huang tried hard to make a GUTA = Grand Unified Theory of Anomalies? new ways of looking at the anomalies (comparing harmonic and real space in various ways) but no GUTA ... TBD

the rare **cold spot**



Gaussian smoothing $l = 6$ (FWHM 20.8deg)

no cold spot

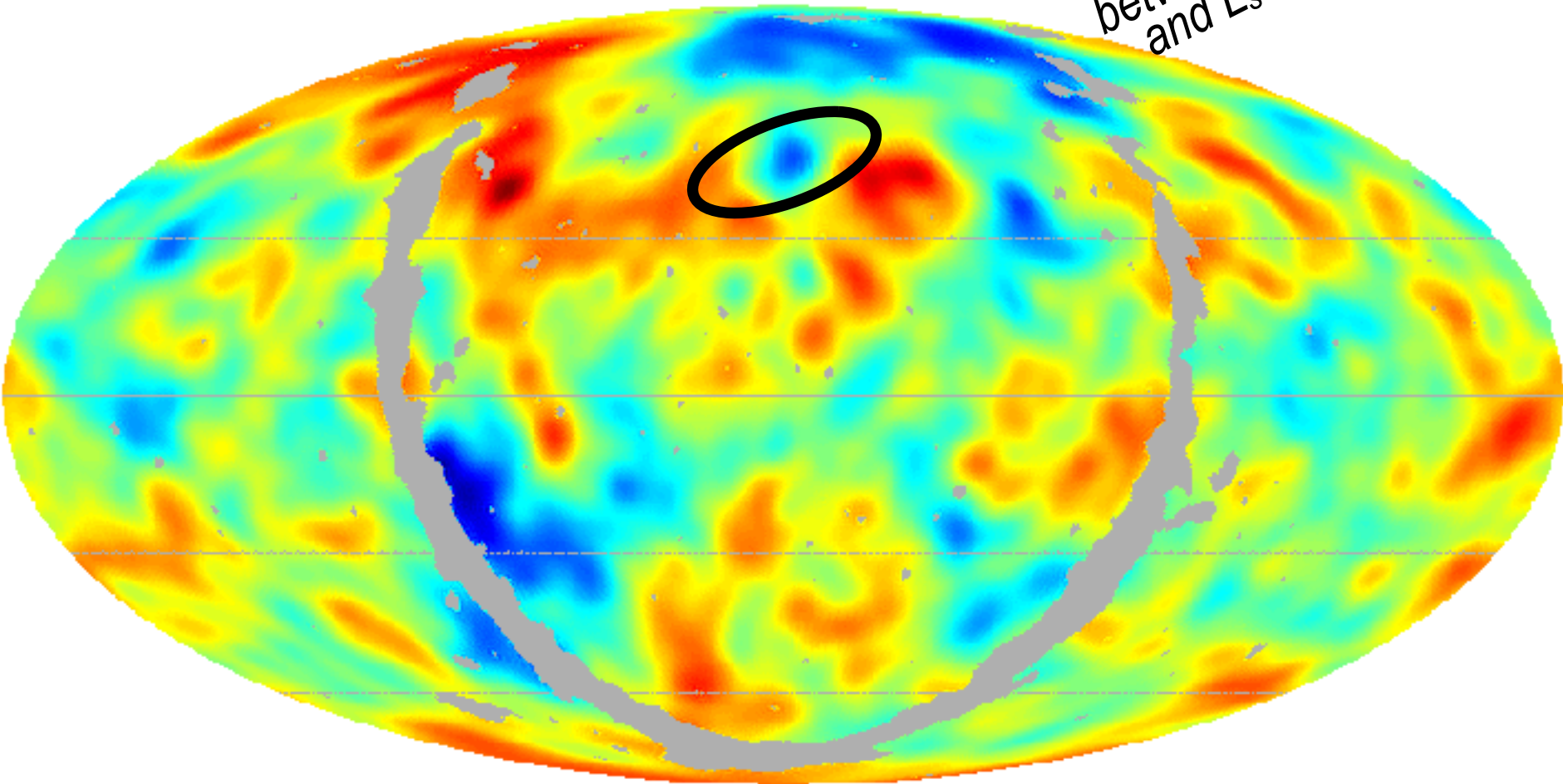


-101.

+72.6

Gaussian smoothing $\lambda = 20$ (FWHM 6.6deg)

cold spot
emerges
between $L_s=6$
and $L_s=20$



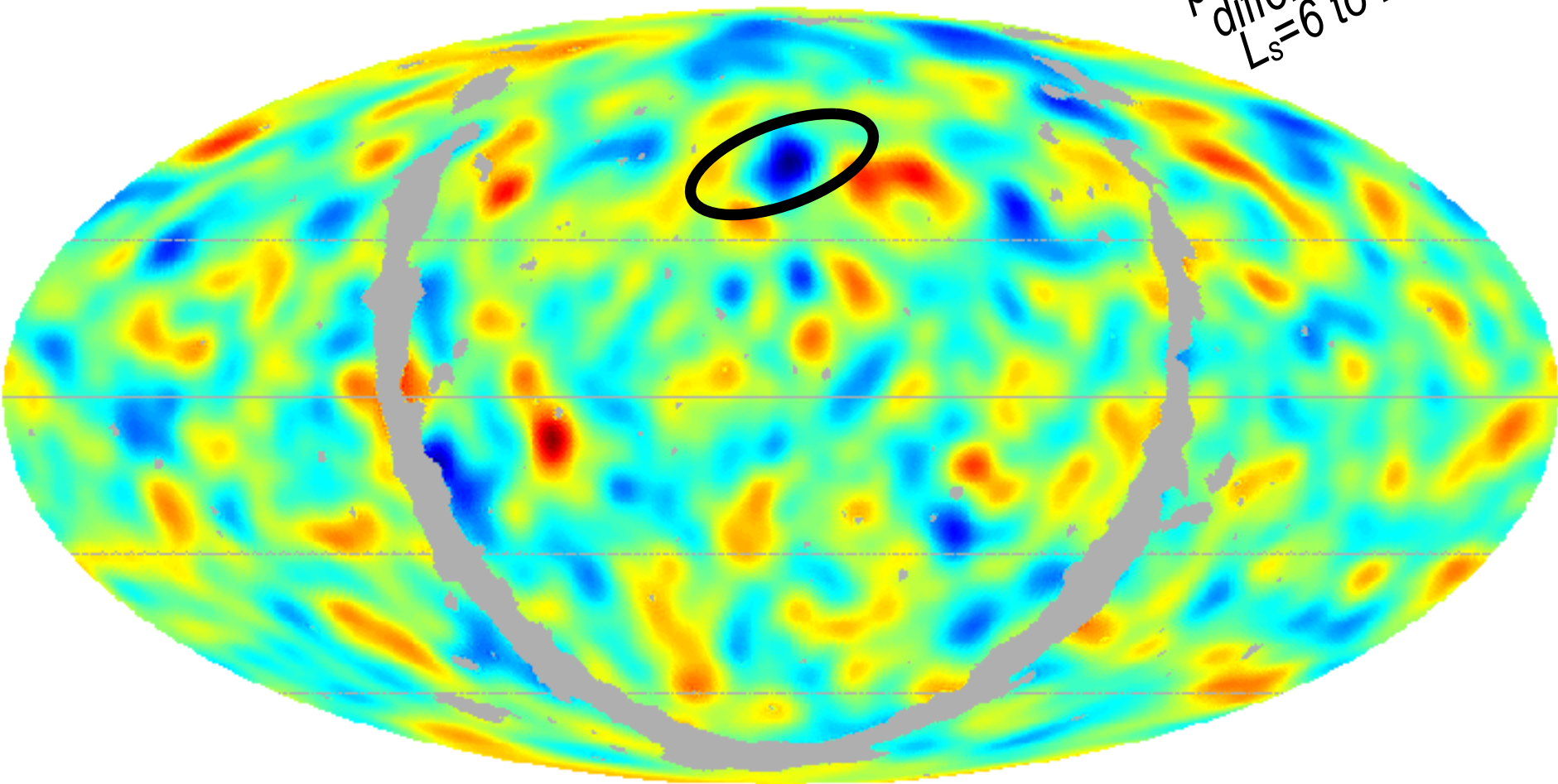
-165.



+125.

Difference map between $l_{\text{smooth}} = 20$ and $l_{\text{smooth}} = 6$

cold spot
prominent in the
difference map
 $l_s=6$ to $l_s=20$



-94.8

+90.4

$$W(\ell) = e^{-\frac{\chi(\ell|1)}{2(\ell_2+1/2)^2}} - e^{-\frac{\chi(\ell|1)}{2(\ell_1+1/2)^2}} \quad (\ell_2 > \ell_1)$$

ℓ_1	ℓ_2	T_{cold}/σ_T	cold-spot p value	T_{hot}/σ_T	hot-spot p value
2	20	-3.5	29.9%	3.2	60.2%
4	20	-4.0	10.1%	3.9	13.9%
6	20	-4.5	2.0%	4.2	4.7%
8	20	-4.5	2.1%	4.3	4.5%
10	20	-4.5	3.0%	4.4	3.9%

tantalizing that the cold spot is the same L-band range as the L p spec dip, but all of our tools have not teased out a relation

B+Huang 2015

*0
e.g. low L constrained fields do not make a nice low-L cavity for the cold spot to be boosted up*

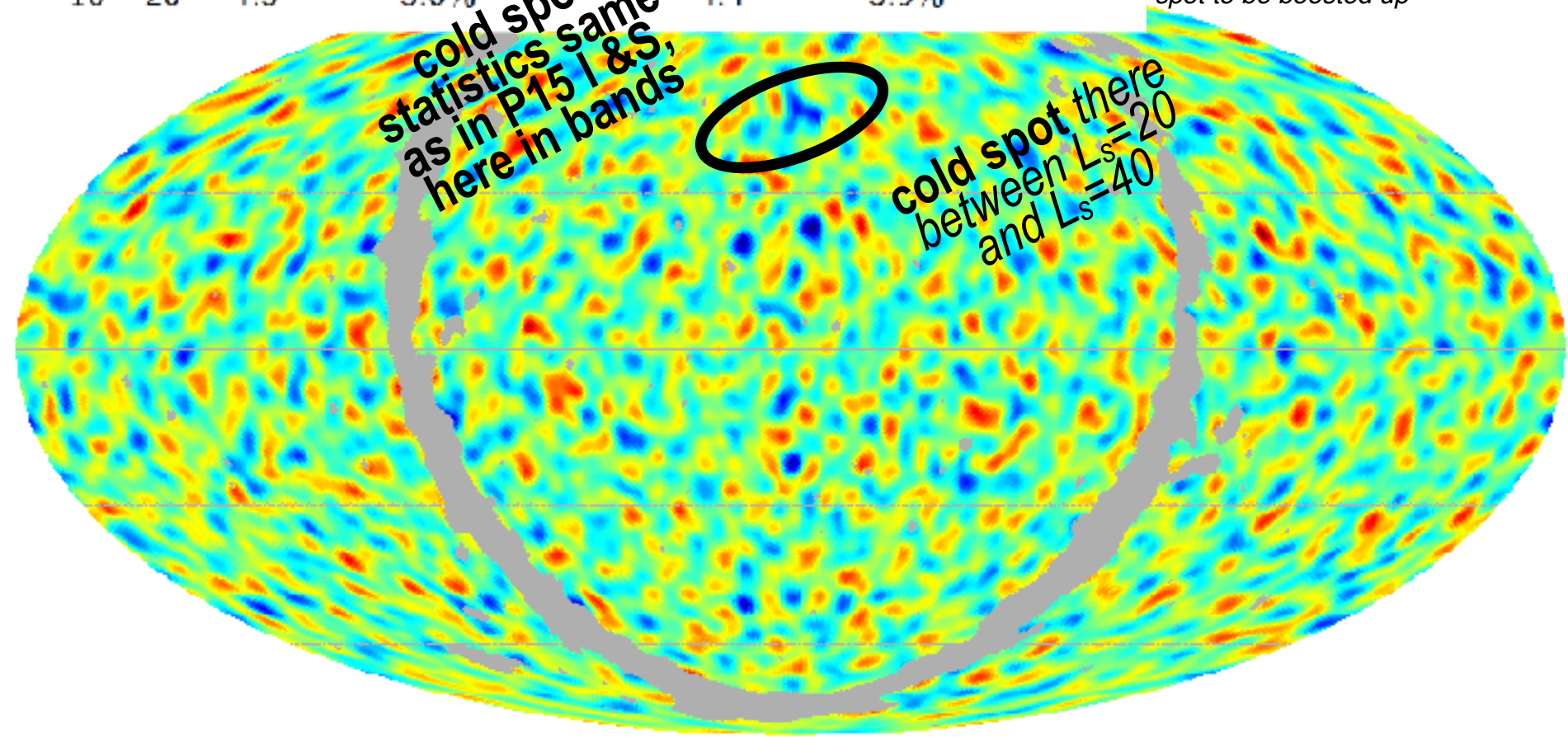
cold spot statistics same as in P15 I & S, here in bands

cold spot there between $L_s=20$ and $L_s=40$

-65.7



+59.7



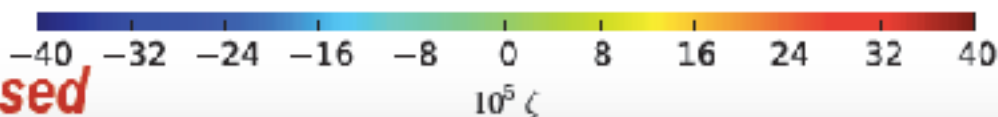
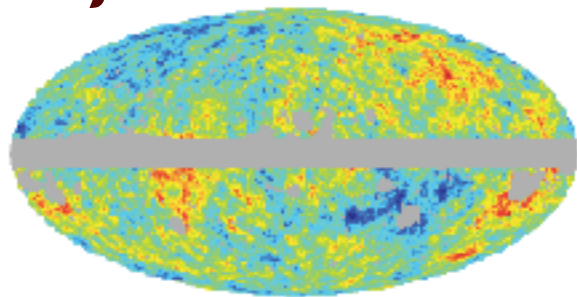
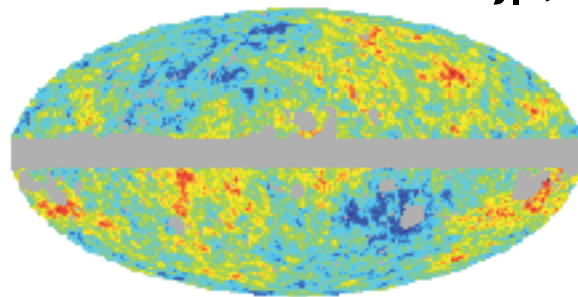
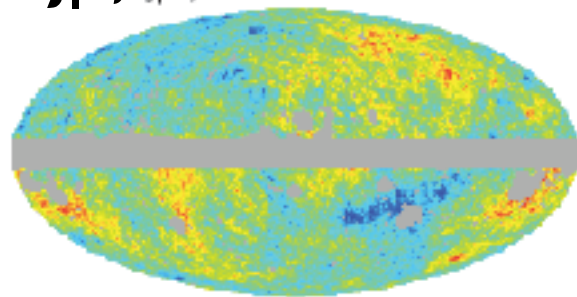
how intermittency could amplify the cold spot to statistical correctness

from $>4.5\sigma$ Gaussian random field anomaly

$\langle \zeta | T, E \rangle$

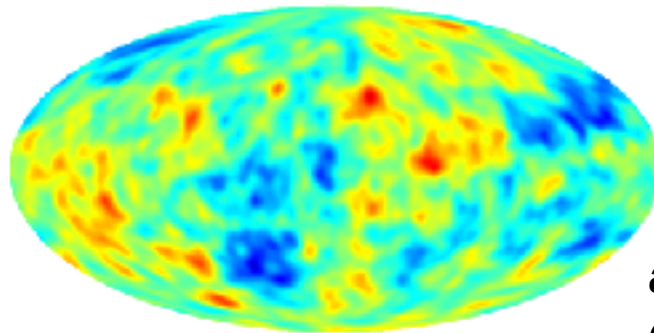
Planck 2015 XVII nonG

$\langle \zeta | T, E \rangle + \delta \zeta$

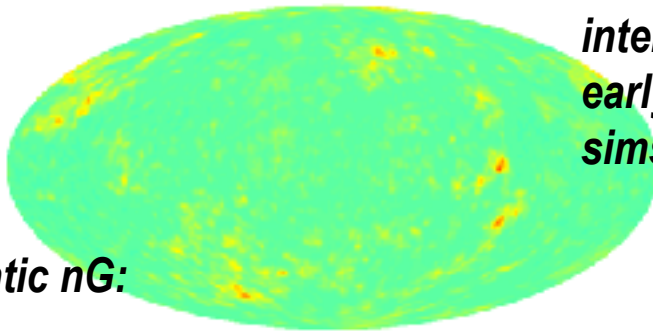


caution: not de-lensed

visibility mask

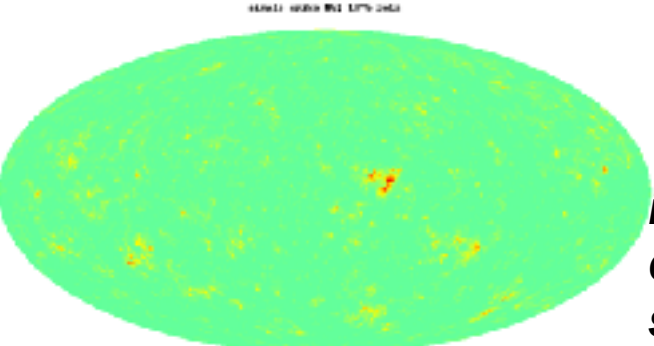


5deg fwhm



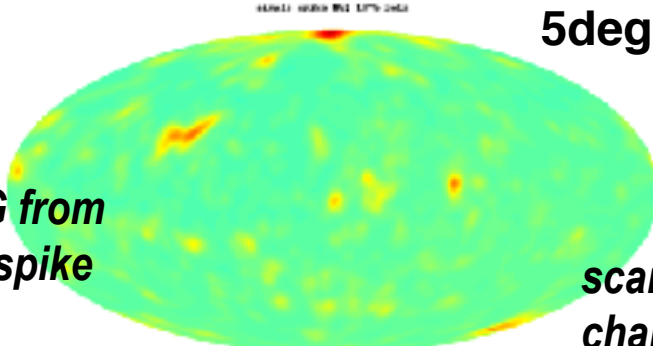
intermittent nG from early U preheating sims - too small

also cf. quadratic nG: correlated fNL uncorrelated large fNL_{eff}



40 arcmin fwhm

intermittent nG from early U single spike sims - tunable amplitude, get the "cold spot" etal

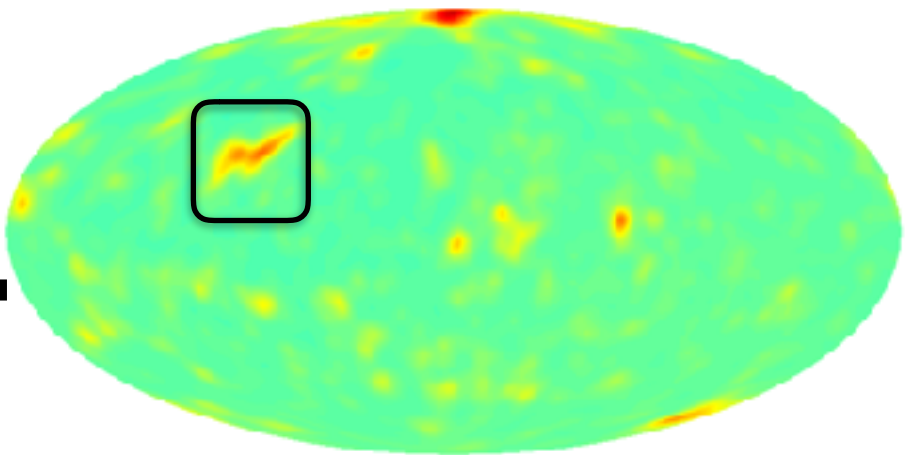
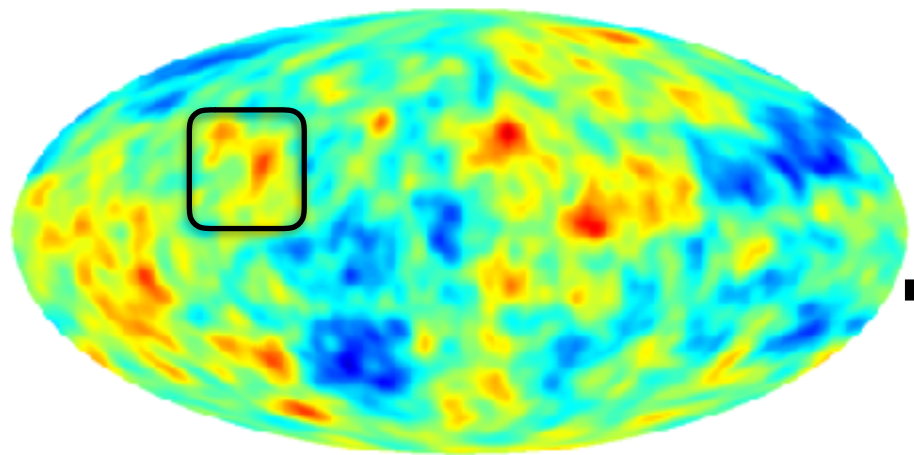


5deg fwhm

scan sims to get chance intermittent alignment to get a "cold spot"

Gaussian; 10°S zeta

single spike 4G; 10°S zeta

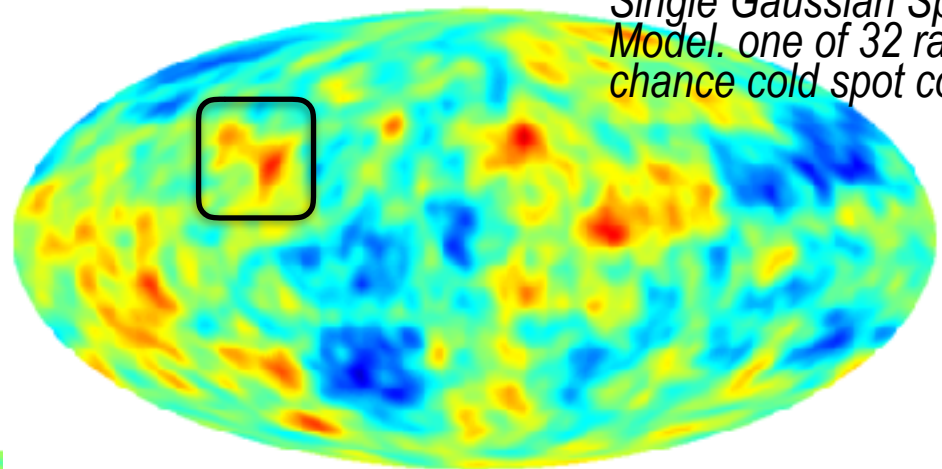


+



Gaussian + single spike 4G; 10°S zeta

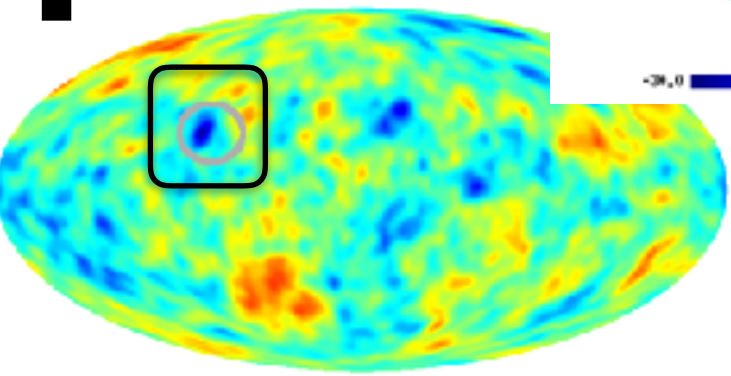
=



Single Gaussian Spike intermittent nonG Model. one of 32 random choices to yield chance cold spot constructive interference.

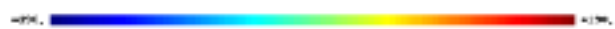
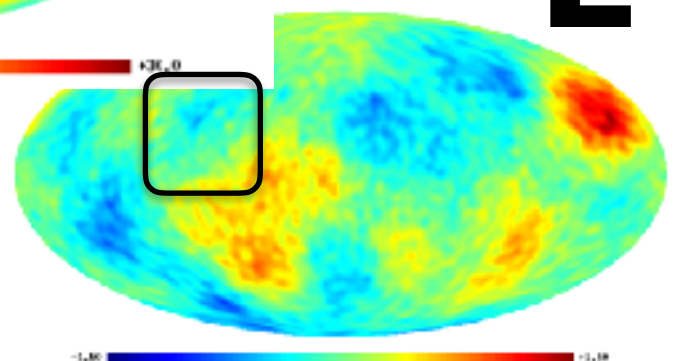
T

Gaussian + single spike 4G; T 6x80



E

single spike 4G; E 10x20

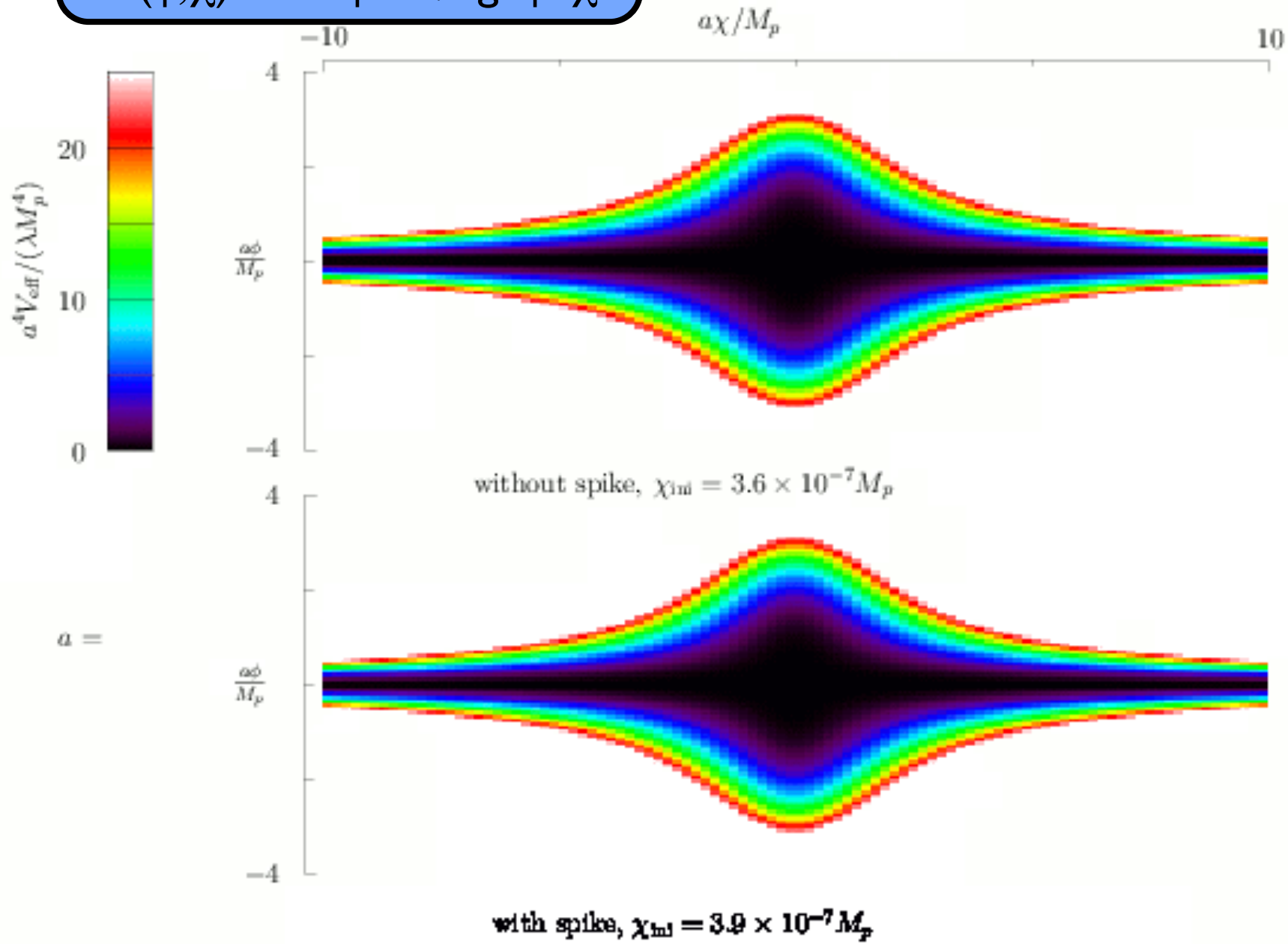


caustics in ballistic orbits

caustics are ubiquitous: **LSS/cosmic web & preheating**

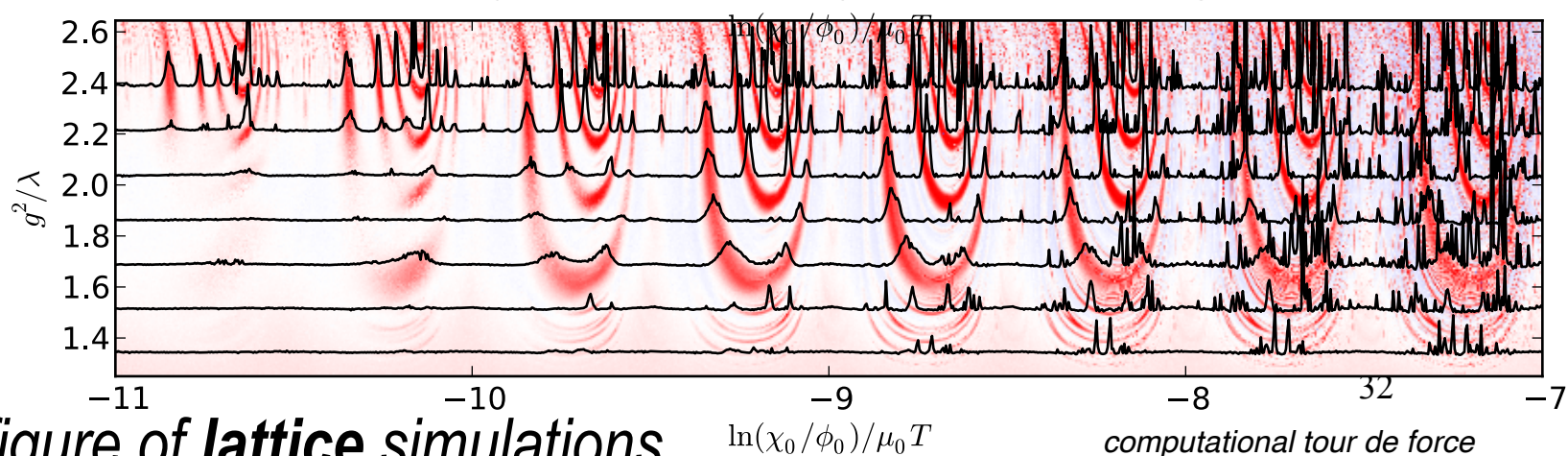
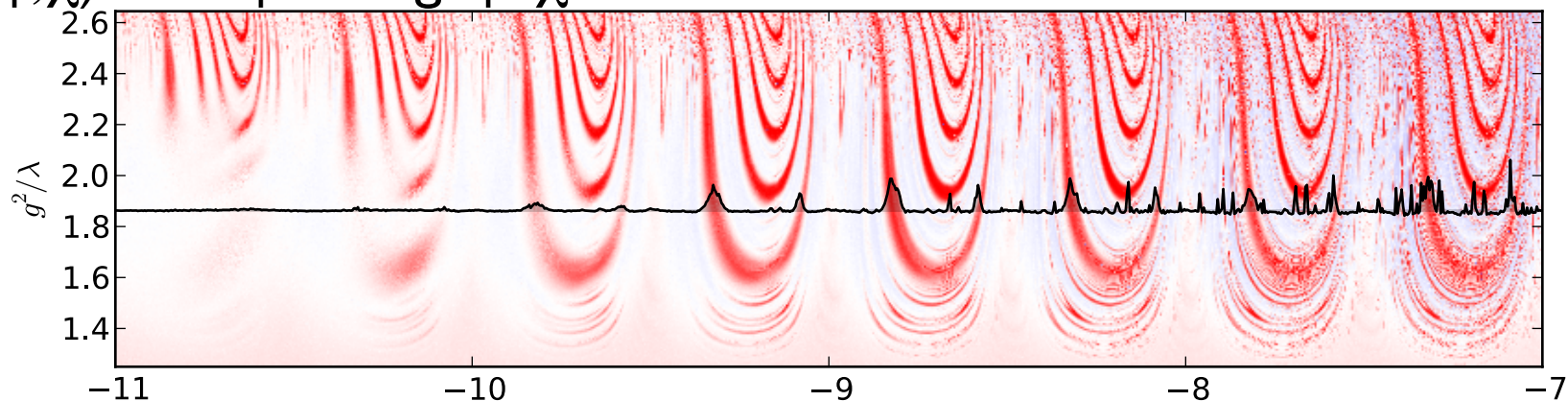
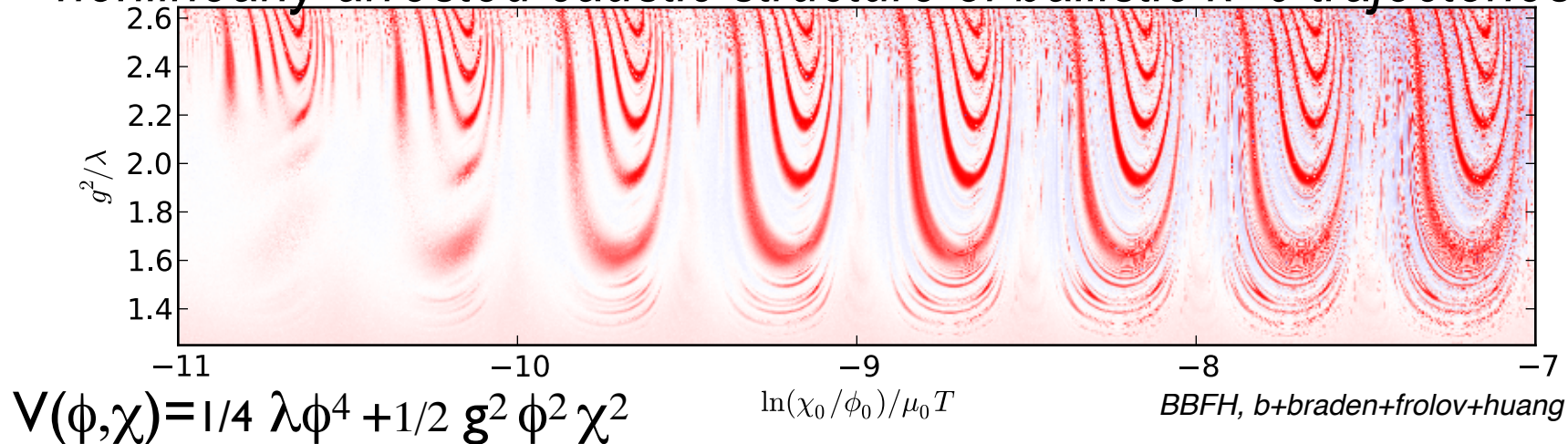


$$V(\phi, \chi) = 1/4 \lambda \phi^4 + 1/2 g^2 \phi^2 \chi^2$$



V_{eff} is trajectory dependent

nonlinearly-arrested caustic structure of ballistic $k=0$ trajectories



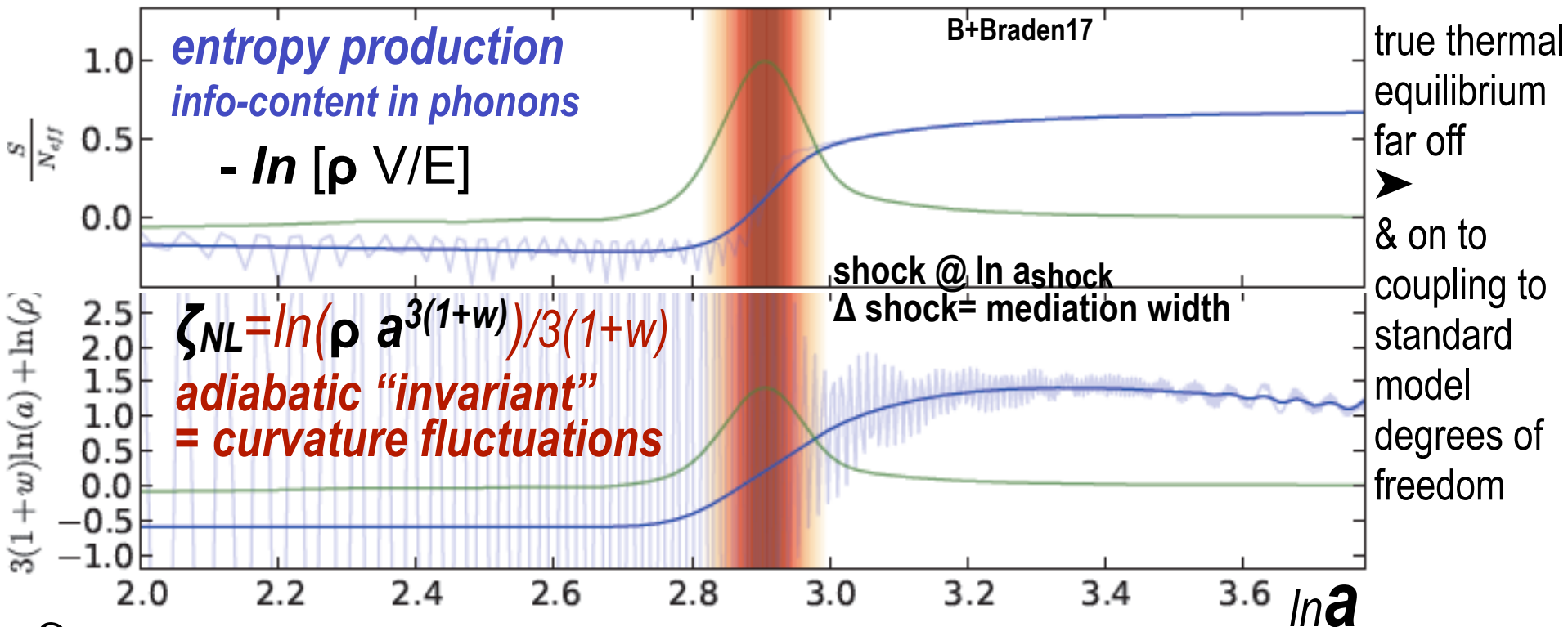
gigafigure of lattice simulations

computational tour de force

**arresting the caustic
orbits via a shock in time,
incoherent cf. coherent
trajectory bundles**

*understanding the ζ -spike structure,
qualitatively YES and quantitatively MAYBE*

nonG from large-scale modulations of the shock-in-times of preheating



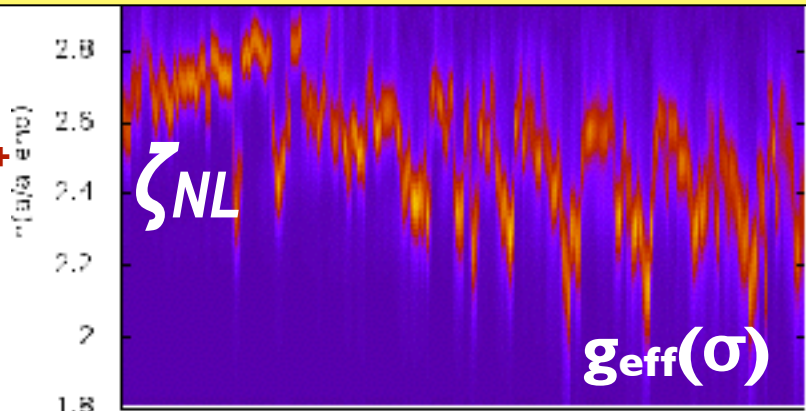
$\delta \zeta_{NLshock}(\mathbf{g}(\sigma(\mathbf{x}))) \Rightarrow$ modulated non-G

$g_0 + g_1 \sigma/M_P, g_0 \exp[\gamma_1 \sigma/M_P], ..$

$V(\phi, \chi) = 1/2 m^2 \phi^2 + 1/2 g_{eff}(\sigma)^2 \phi^2 \chi^2$

$\delta \zeta_{NLshock}(\chi_i(\mathbf{x}) | g^2/\lambda) \Rightarrow$ NonG cold spots ++

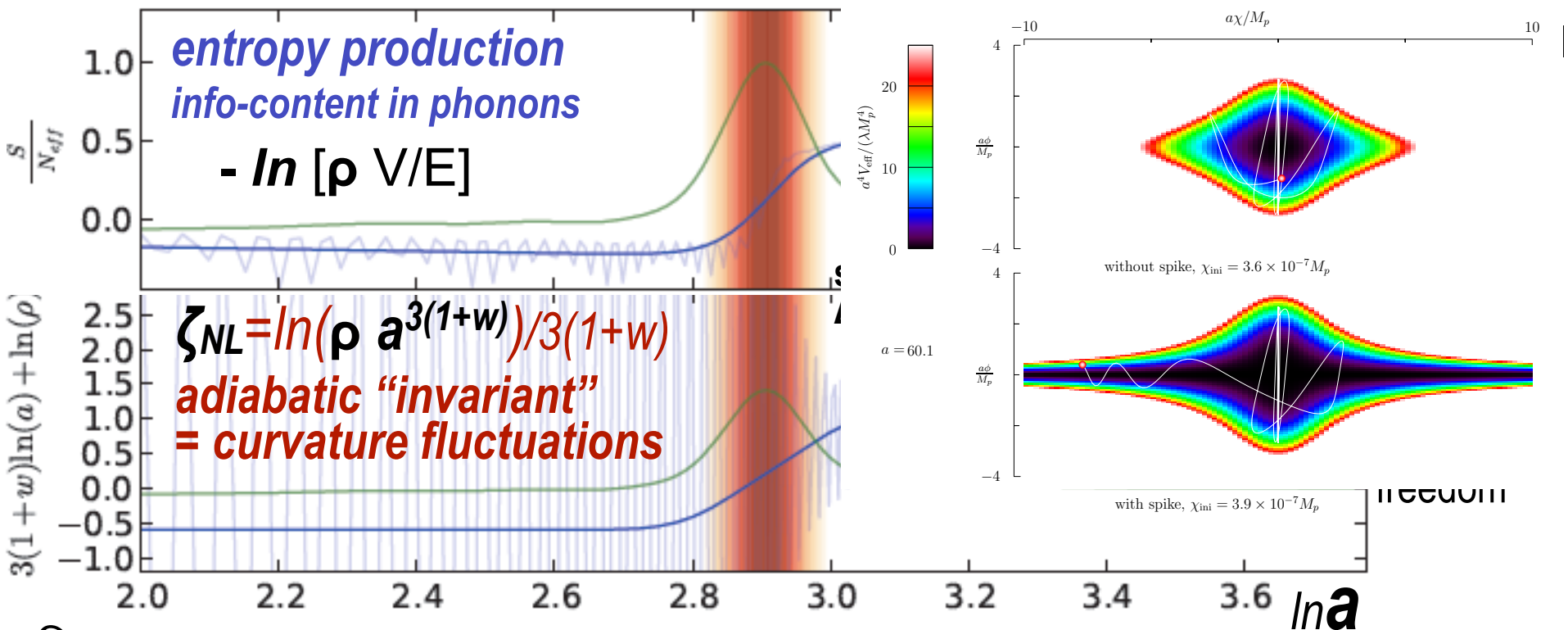
$V(\phi, \chi) = 1/4 \lambda \phi^4 + 1/2 g^2 \phi^2 \chi^2$



V_{eff} is dynamical Bond, Braden, Frolov, Huang17

*unconventional local non-G: no scale built into V;
 perturbative isocon-based f_{NL} ; rare event cold spots*

nonG from large-scale modulations of the shock-in-times of preheating



$\delta \zeta_{NL\text{shock}}(\mathbf{g}(\sigma(\mathbf{x}))) \Rightarrow$ modulated non-G

$$V(\phi, \chi) = 1/2 m^2 \phi^2 + 1/2 g_{\text{eff}}(\sigma)^2 \phi^2 \chi^2$$

$\delta \zeta_{NL\text{shock}}(\chi_i(\mathbf{x}) | g^2/\lambda) \Rightarrow$ NonG cold spots ++

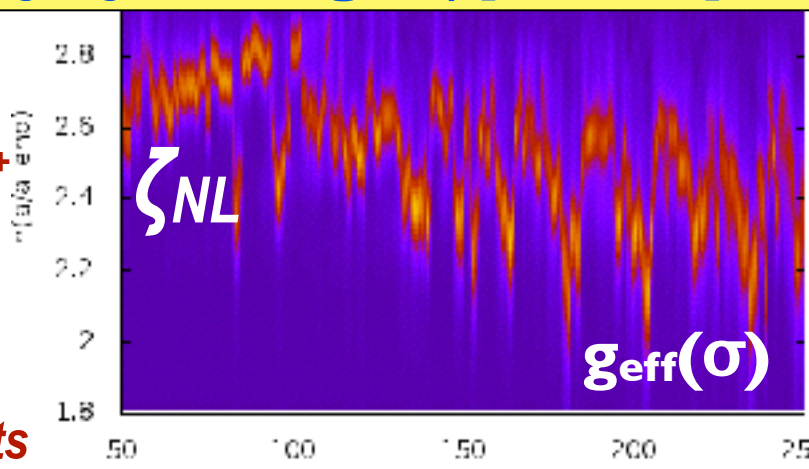
$$V(\phi, \chi) = 1/4 \lambda \phi^4 + 1/2 g^2 \phi^2 \chi^2$$

$$g_0 + g_1 \sigma/M_p, g_0 \exp[\gamma_1 \sigma/M_p], ..$$

V_{eff} is dynamical Bond, Braden, Frolov, Huang17

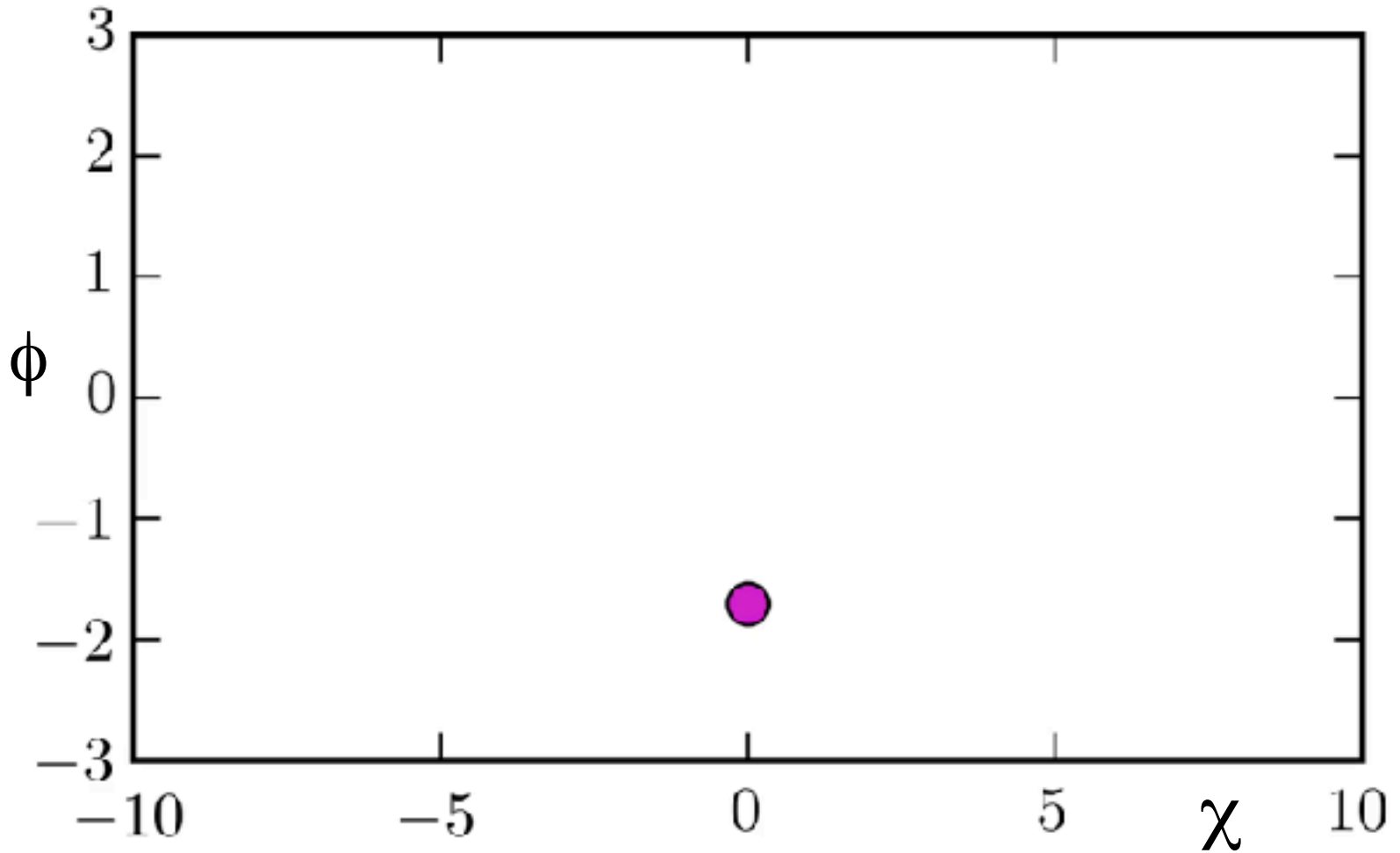
unconventional local non-G: no scale built into V ;

perturbative isocon-based f_{NL} ; rare event cold spots



ballistic billiards $k=0$ mode phase space string evolution.

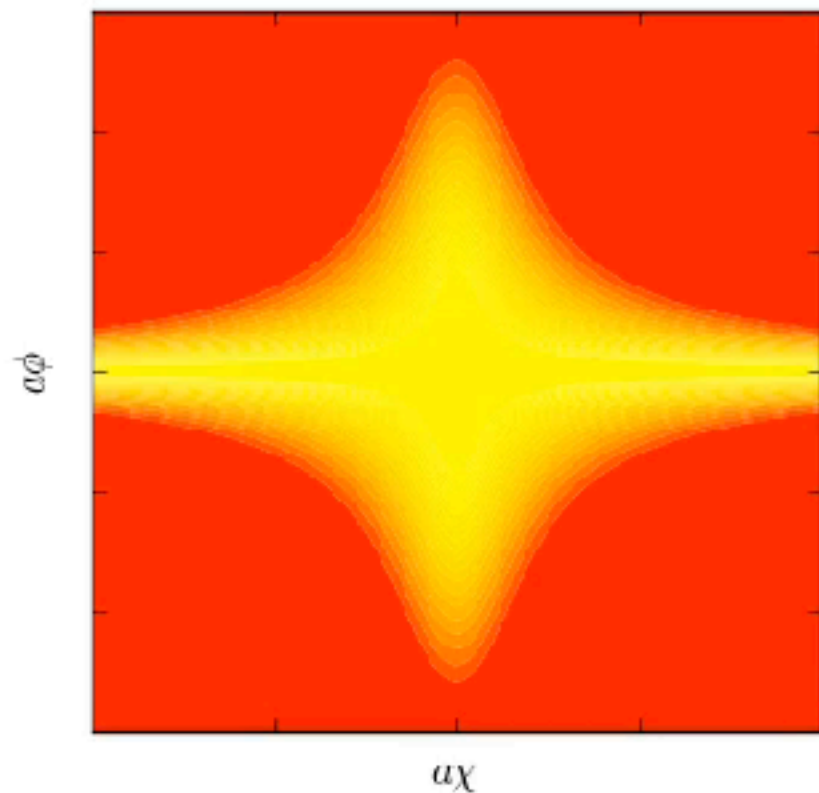
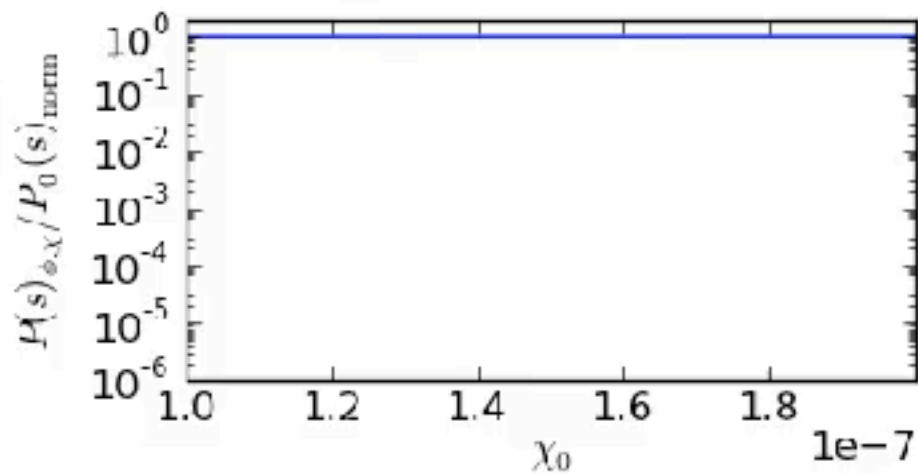
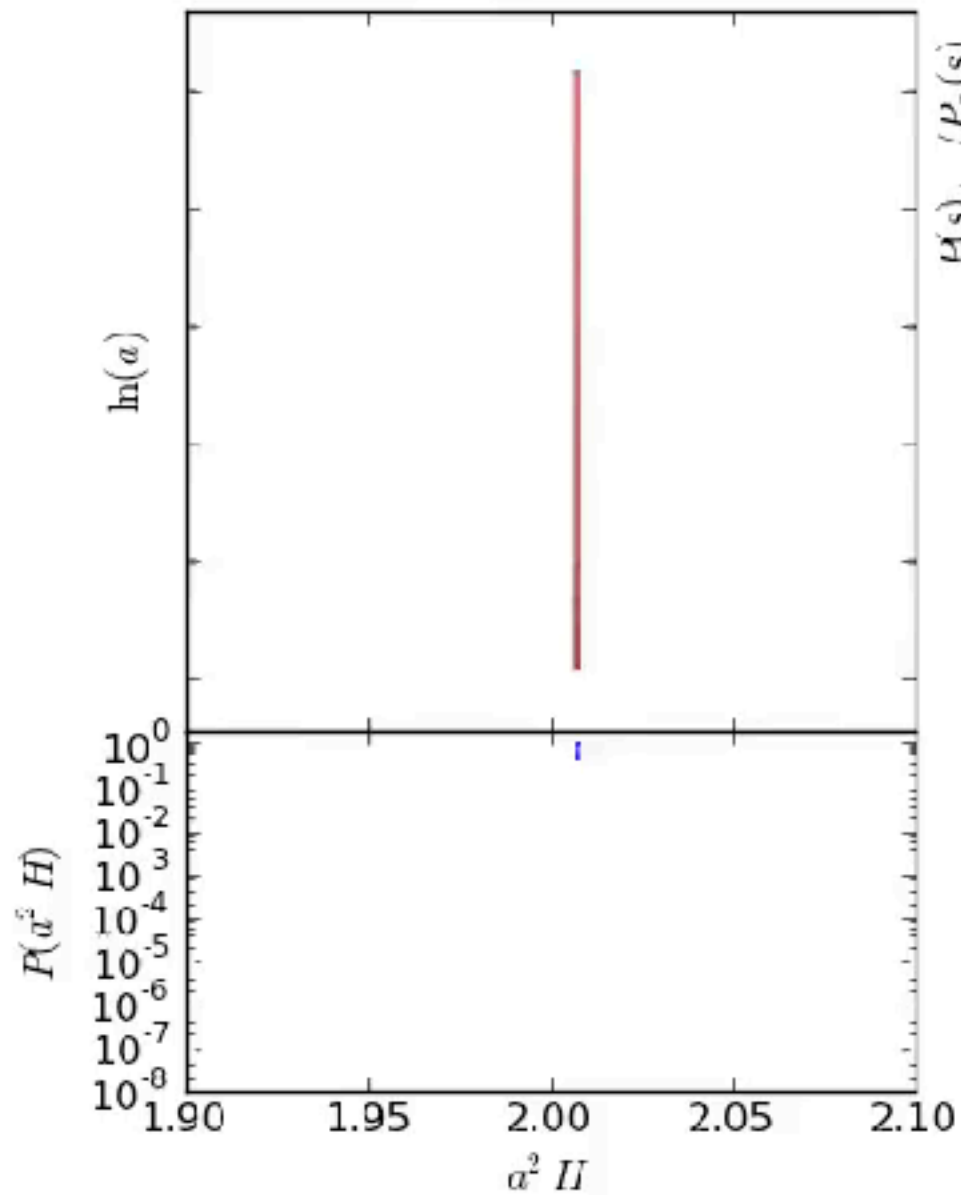
stopping criterion when coarse-grained entropy of field variables rises



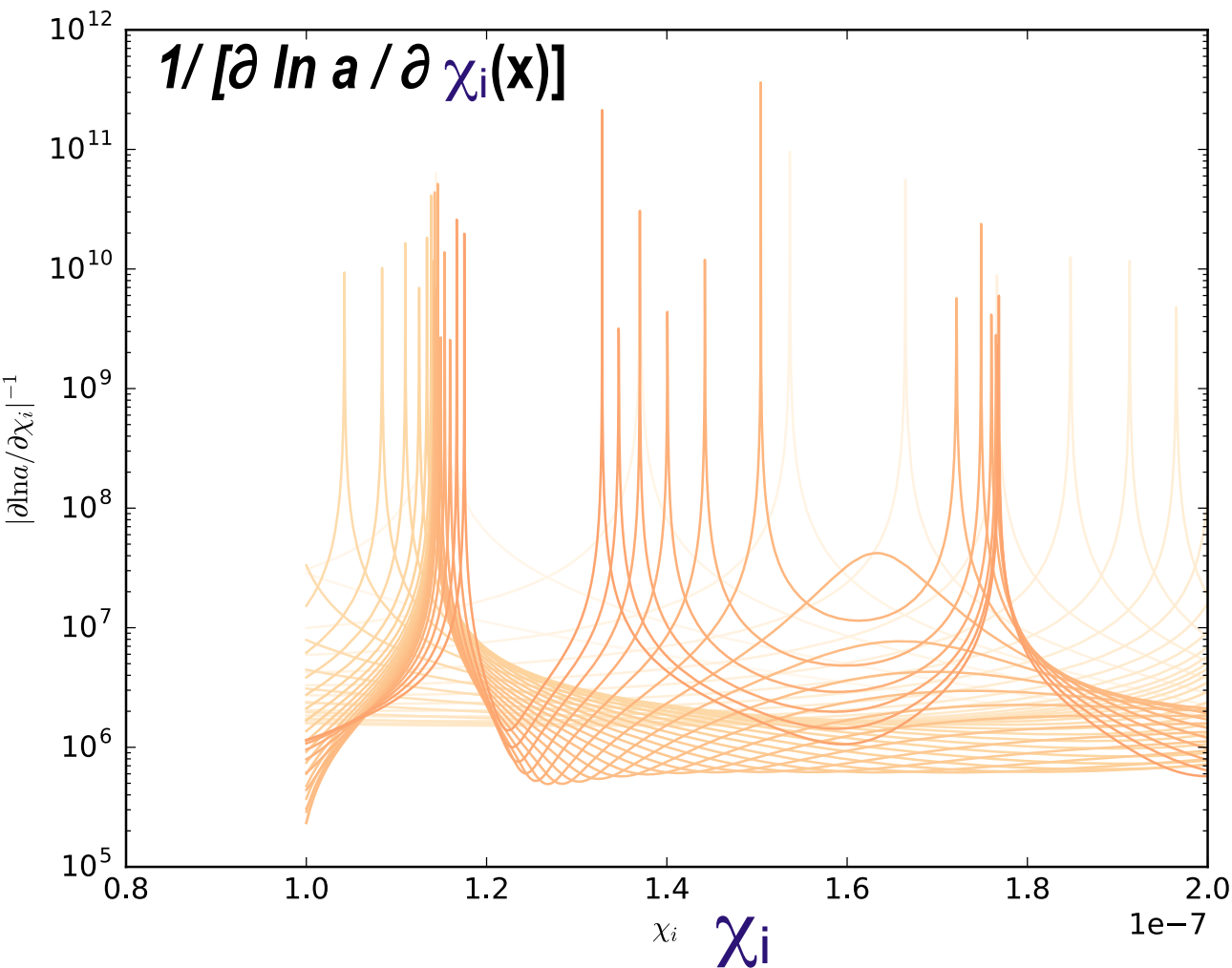
$$V = \frac{1}{4} \lambda \phi^4 + \frac{1}{2} g^2 \phi^2 \chi^2$$



caustics are ubiquitous



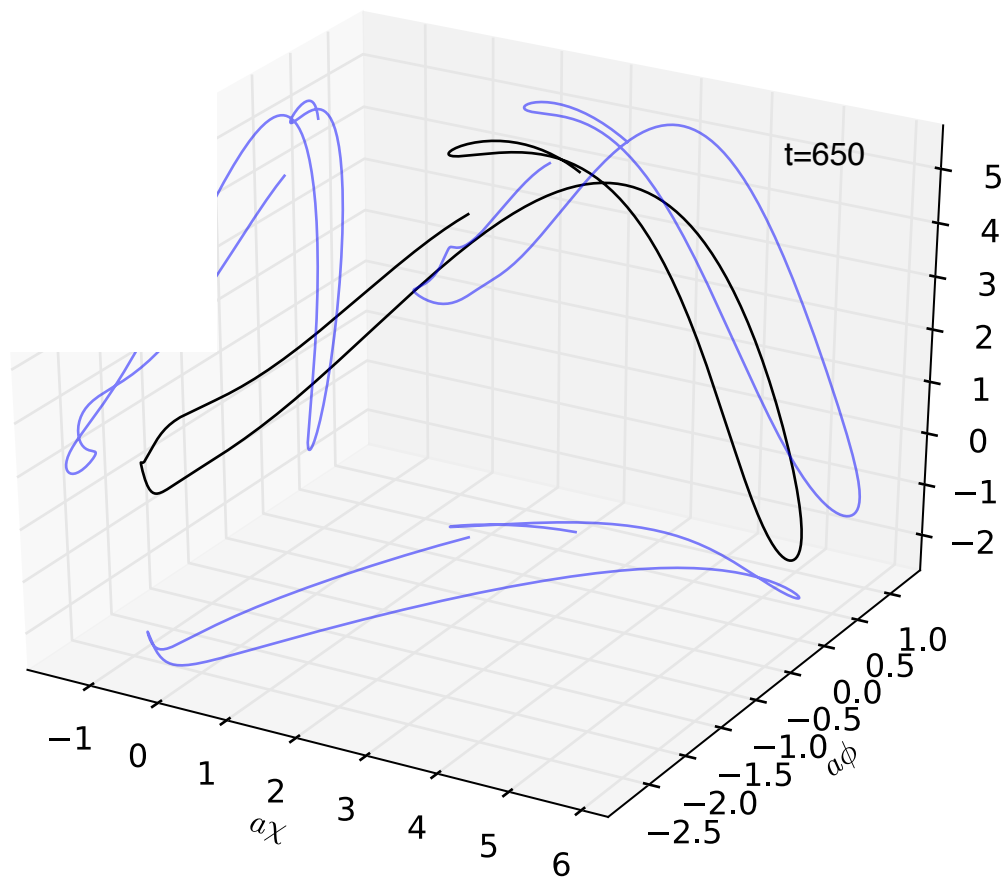
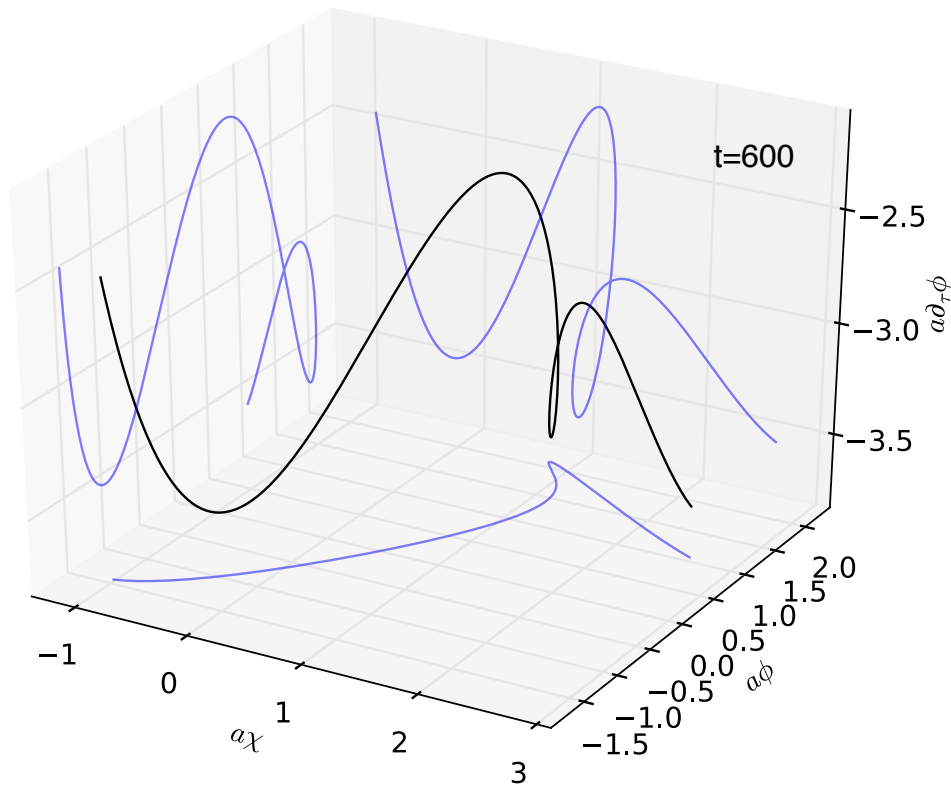
**calculating ballistic evolution to caustics
gives the spikes in good agreement with
full nonlinear lattice simulations**



B2FH, b+braden+frolov+huang

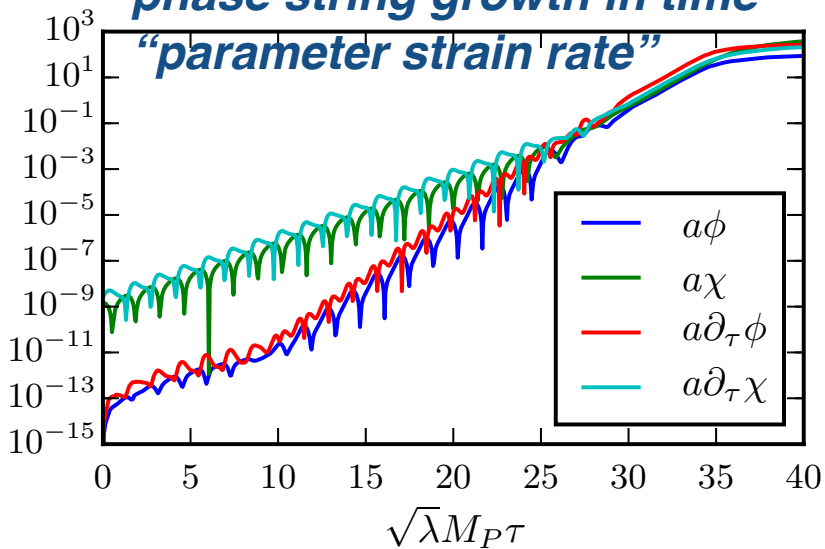
nonG from post-inflation but pre-entropy generation ballistic trajectories can lead to pre-shock-in-time caustics and other phase space convergences in the deformations (!) Zel'dovich map-ish

eg $\partial \ln a / \partial \chi_i(x)$, $\partial \ln a / \partial g(x) \Rightarrow P[\ln a(x), t_{\text{shock}} \mid \chi_i(x), g(x), t_{\text{end-of-inflation}}]$



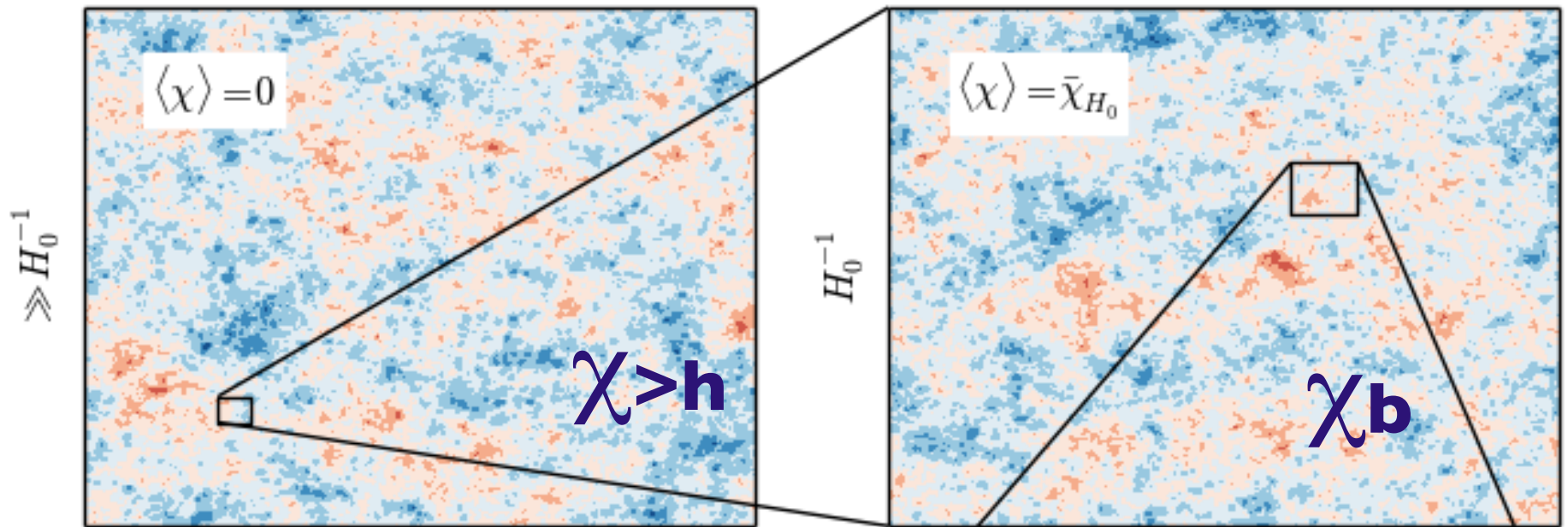
phase string growth in time

"parameter strain rate"



**how modulated caustics in
preheating could give
observable intermittency**

**modulating the caustics
on large scales & super-
horizon scales via isocons**

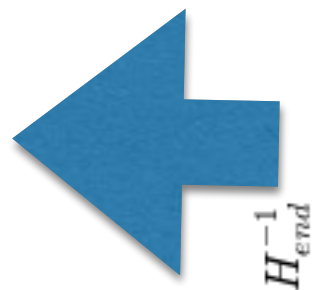


ULSS modulation beyond our Hubble patch

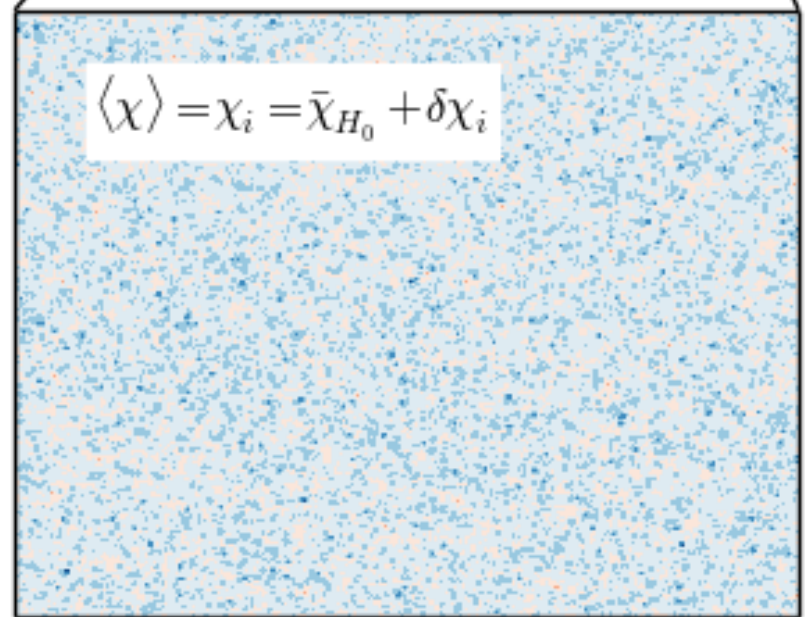
LSS modulation within our Hubble patch

$$\delta \zeta_{NL\text{shock}}(\chi_i(\mathbf{x}) | g^2/\lambda)$$

\Rightarrow *NonG cold spots ++*



H_{end}^{-1}



preheating horizon scale < comoving cm

**how generic will caustic
preheating be? structure
around minima:**

filamentary potentials

define channels

multi-filaments may lead to caustics

modulating post-inflation entropy generation shocks via long range fields

isocon

$\chi(\mathbf{x})$

or

$g(\sigma(\mathbf{x}))$

or..

ϕ

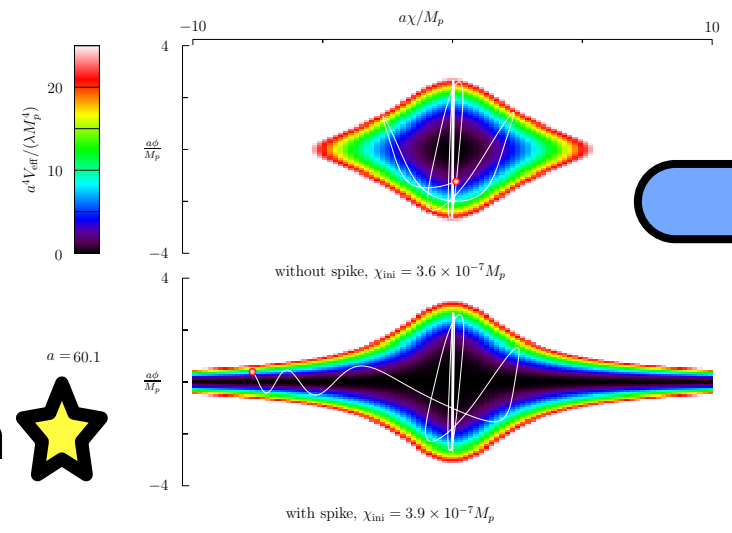
inflaton

pre-heating patch (~1cm)

$S_{U,m+r}$

$\sim 10^{88.6}$

How generic is the intermittent caustic phenomenon? Holds for many basin potentials at the end of inflation. but not if rapid heating

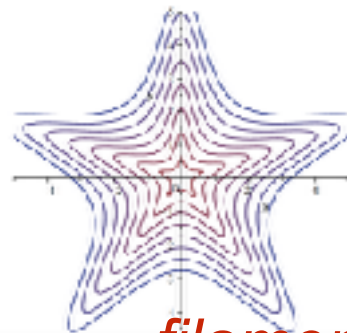


$$1/4 \lambda \phi^4 + 1/2 g^2 \phi^2 \chi^2,$$

$$1/2 m^2 \phi^2 + 1/2 g^2(\sigma) \phi^2 \chi^2$$

$$V(\mathbf{r})U(\cos\theta), r^2 = \phi^2 + \chi^2$$

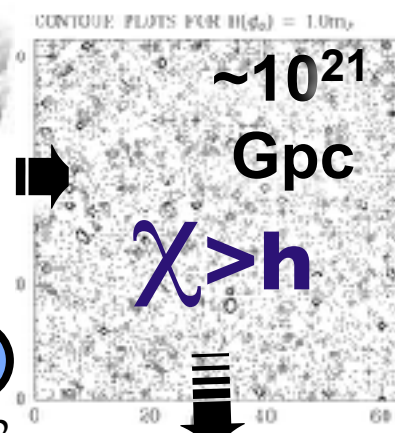
angular variables pNGB natural inflation, racetrack, monodromy, ..
 $V(\mathbf{r}, \theta) = \sum_M V_M(\mathbf{r}) \cos(m\theta)$ pNGB, Roulette r-hole size
 3D $\phi \chi \sigma$ fields $V(\mathbf{r}, \mathbf{n}) = \sum_{LM} V_{LM}(\mathbf{r}) Y_{LM}(\mathbf{n})$
 $V(\phi, \chi) = 1/4 \lambda \phi^4 - 1/2 \xi \phi^2 R + 1/2 g^2 \phi^2 \chi^2$
 conformally transformed potentials a la Higgs/ R^2 , modified kinetic terms, flattened potentials of all sorts B2FH, b+braden+frolov+huang



filaments! 



$S_{U, \text{UUUULSS}}$



1000 Gpc

$\chi > h$

10 Gpc

χ_b

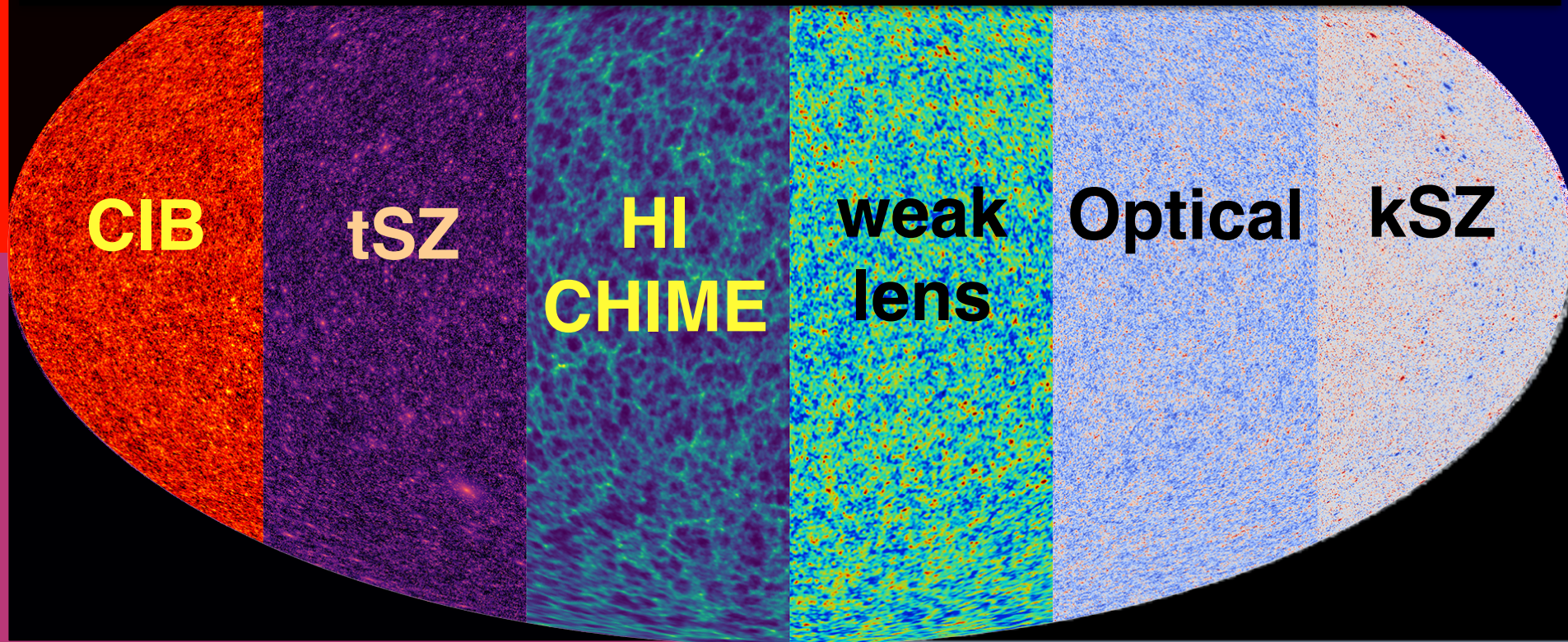
mocking heaven to
explore 3D intermittency
from modulating
preheating, bubble
collisions, etc
a quest for the apparent
breakdown of LSS
homogeneity - but NOT

Mocking Heaven @ CITA Alvarez Bond Stein Battaglia ..



Peak Patch Full Sky Models for Planck, AdvACT, SO, CMB-S4, CCATp, CHIME, HIRAX, SKA, COMAP, EUCLID, LSST, ...

*need End to End mocks, fully correlated to draw out:
BSMc, DE/modG, Mnu, nonG (correlated, uncorrelated, intermittent),...*



Planck 2015 XII: Full Focal Plane Sims (Nov): FFP8 ensemble of 10K EndtoEnd mission realizations in 1M maps. instrument noise + CMB + PSM + .. (25M NERSC CPU hrs)

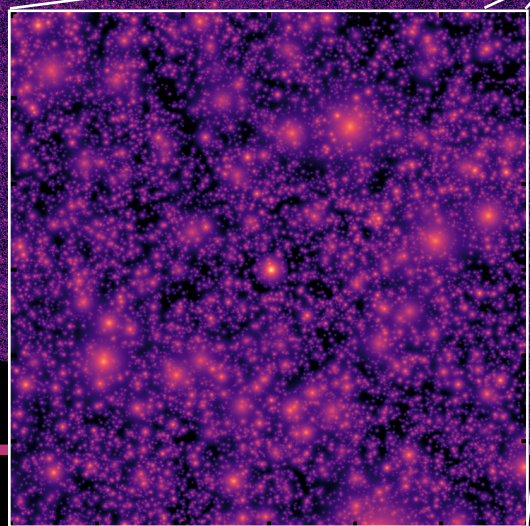
Compton Scattering (Sunyaev-Zeldovich)
Simulations for ACT, Planck, Simons Obs
& CMB Stage 4 Cluster Observations
Using high res Gas Hydro Sims

HI Intensity Mapping
simulations of CHIME / HIRAX ..
 $z=0.8-2.5, \sim(8 \text{ Gpc})^3$

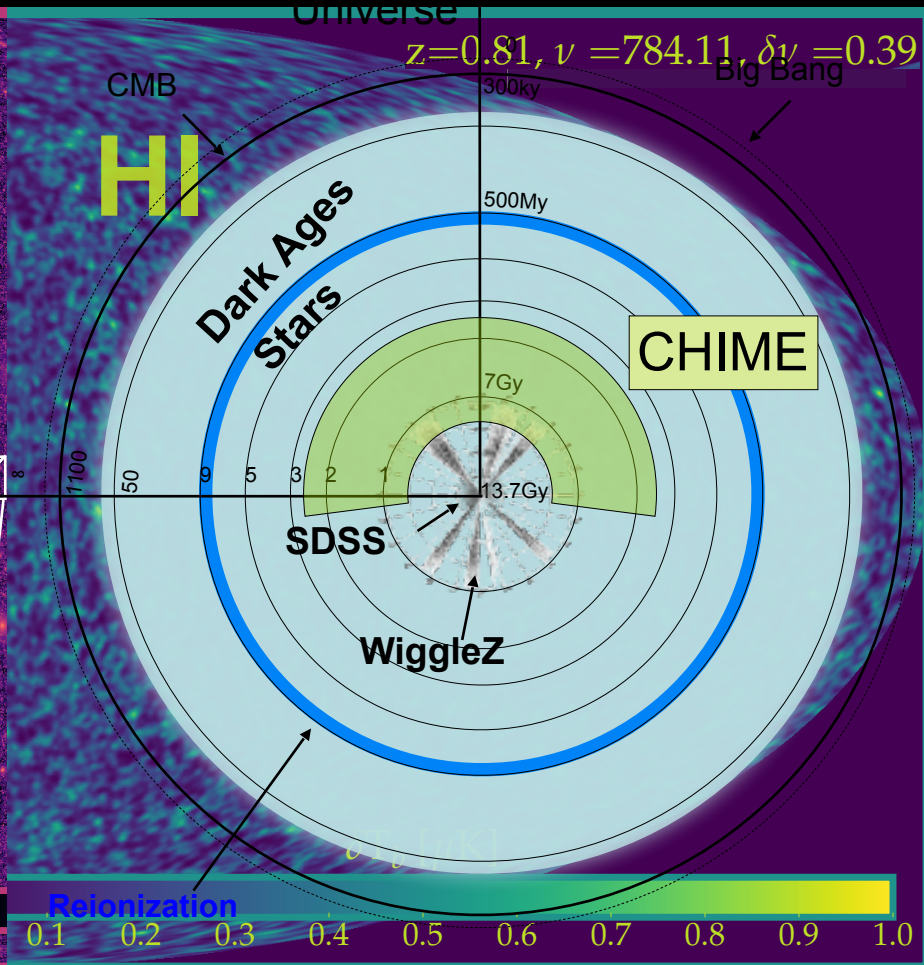
$0.00 < z < 1.25$
 8Gpc, 4096^3 Box

tSZ

Gaussian



6 deg



Compton Scattering (Sunyaev-Zeldovich)
Simulations for ACT, Planck, Simons Obs
& CMB Stage 4 Cluster Observations
Using high res Gas Hydro Sims

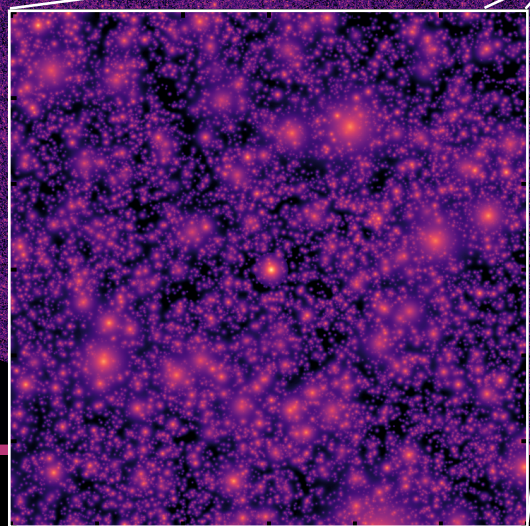
HI Intensity Mapping
simulations of CHIME / HIRAX ..
 $z=0.8-2.5, \sim(8 \text{ Gpc})^3$

$0.00 < z < 1.25$
8Gpc, 4096^3 Box

$z=0.81, \nu = 784.11, \delta\nu = 0.39$

tSZ

HI



Gaussian

$\delta T_b [\mu\text{K}]$



6 deg

Compton Scattering (Sunyaev-Zeldovich)
Simulations for ACT, Planck, Simons Obs
& CMB Stage 4 Cluster Observations
Using high res Gas Hydro Sims

HI Intensity Mapping
simulations of CHIME / HIRAX ..
 $z=0.8-2.5, \sim(8 \text{ Gpc})^3$

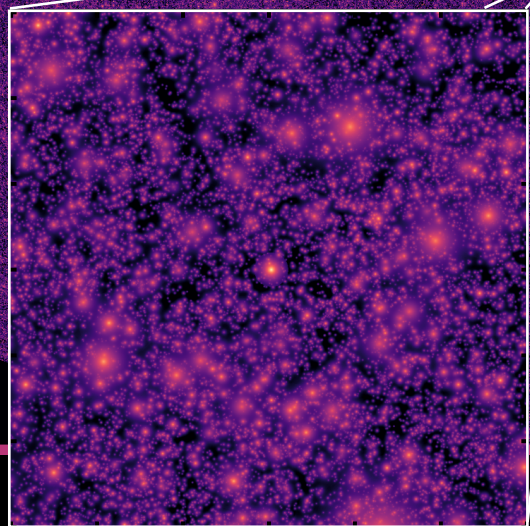
$0.00 < z < 1.25$
8Gpc, 4096^3 Box

tSZ

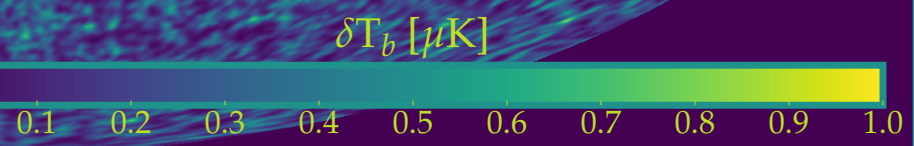
HI

$z=0.81, \nu = 784.11, \delta\nu = 0.39$

correlated
Quadratic
nonG
 $f_{NL} = 25$



6 deg



Compton Scattering (Sunyaev-Zeldovich)
Simulations for ACT, Planck, Simons Obs
& CMB Stage 4 Cluster Observations
Using high res Gas Hydro Sims

HI Intensity Mapping
simulations of CHIME / HIRAX ..
 $z=0.8-2.5, \sim(8 \text{ Gpc})^3$

$0.00 < z < 1.25$
8Gpc, 4096^3 Box

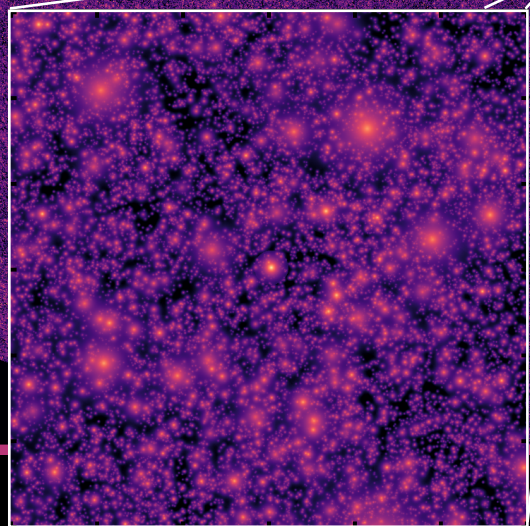
tSZ

HI

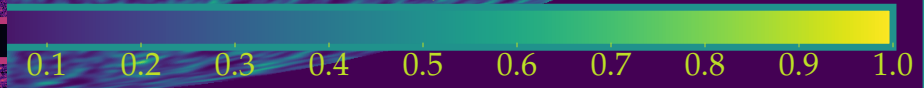
$z=0.81, \nu = 784.11, \delta\nu = 0.39$

uncorrelated
modulated
preheating
intermittent
nonG

Gaussian Spike



6 deg



$\delta T_b [\mu\text{K}]$



this is a very quantitative exercise

e.g., **response of BAO and**

biasing of halos to forms

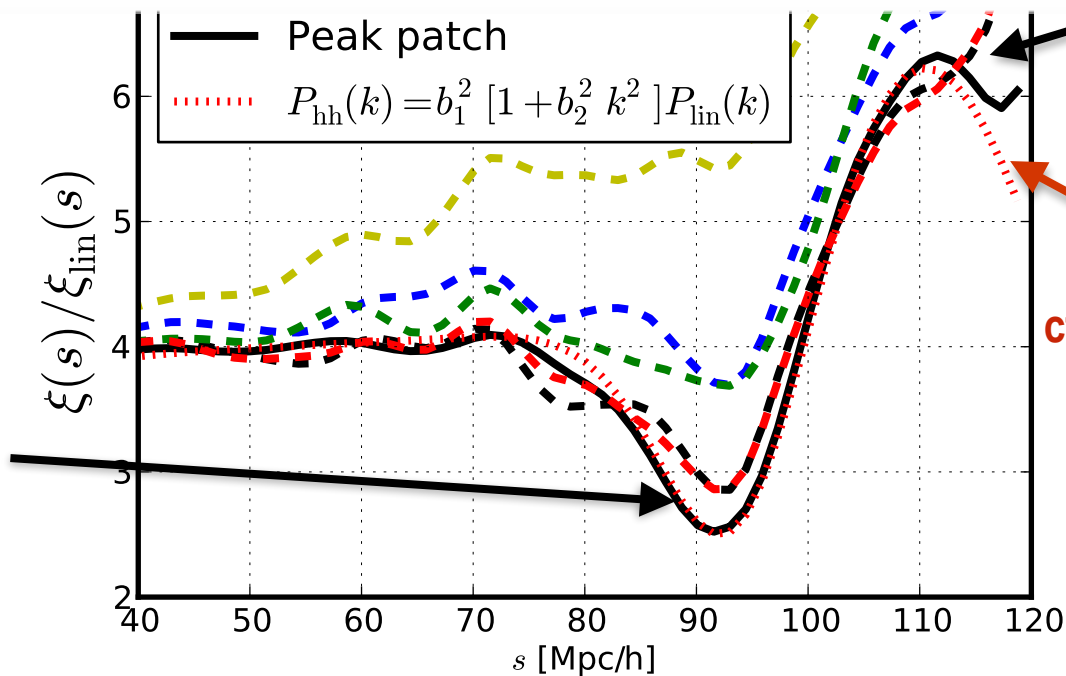
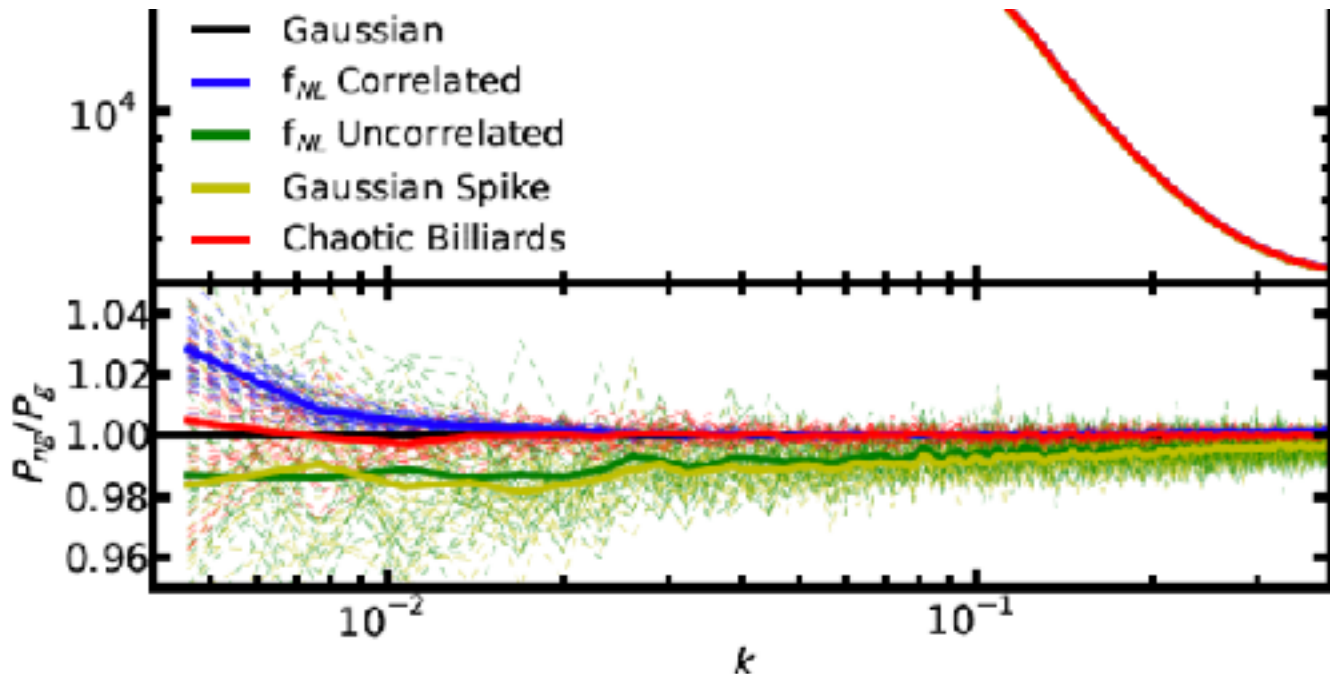
of nonG - correlated cf.

uncorrelated, intermittent

cf. perturbative

e.g., **search for rare superBIAS**

events > supercluster-scale



Positive Curvature:
Overabundance of
Negative Extrema

cf. fit to N-body results:
without any tuning it
comes out from peak
patches naturally

*intermittent nG from
early U single spike*

Positive Curvature:
Overabundance of
Negative Extrema

**highly nonlinear field evolutions happened
(Eol, bubble collisions)!**

amusing subdominant patterns do arise!

lead to observable rare-event CMB/LSS anomalies?

***light isocons cf. heavy isocons, the heavy can lighten up = original SBB nG
isocon modulators, coupling(isocon) modulators, isocon tunneling, isocon
oscillons, isocon short-lived fuzzy-strings, + very long-lived strings***

**or just to weak constraints on multifield potentials,
>horizon fields, nucleation rates, etc.**

alas a 2-number A_s - n_s early universe so far

intermittency frustration: statistical variance is large cf. 2-3 parameter search

***CMB restricts us to a projected 2D ζ -scape to reconstruct
phonon/isotropic-strain power, the future may look much the
same as now for $\zeta \Rightarrow$ potential $V(\phi) \Rightarrow$ acceleration $\epsilon(a)$***

we mock the LSS future *end-to-end* to probe the mode-rich 3D ζ -scape