

Extending stochastic inflation to include coherent pulses to explain instability-driven multi-field simulation ensembles of concentrated primordial non-Gaussianity

Dick Bond @simons eU 22 04 26 Tom Morrison (GS), Jonathan Braden, Tomas Galvez



$3\zeta(\mathbf{x}, t) = \int_{\text{field-path}} (dE + p dV) / (E + pV) \sim dS / \beta(E + pV)$ the entropic structure measure

$\zeta(\mathbf{x}, t) \sim \int \Delta \zeta_{nG} \text{-Prominences}(\mathbf{x} - \mathbf{x}_c) dN_c(\mathbf{x}_c, R_c) + \text{Gaussian random } \zeta\text{-flucs}$

$\zeta(\mathbf{x}, t) \sim \sum_p \Delta \zeta_{nG}(\chi_p(\mathbf{x}, \alpha_e)) + \text{Gaussian random } \zeta\text{-flucs}, \chi_p(\mathbf{x}, \alpha_e) = \text{Gaussian random}$

nonG from instabilities during inflation as well as **modulated heating nonG**

- Inflation probes fundamental physics in the CMB/LSS and higher k scales through ζ . Major effort: linear transfer / nonlinear transport $\zeta_G + \zeta_{nG}$ from early U through neutrino decoupling and photon decoupling through the gravitational instability of DM and baryons into the entangled cosmic web (\mathbf{x}, t) and its luminous observables - and component separate / disentangle non-G ζ -forms / templates to measure. We use Webskys. Deep Learning approaches are underway, Seljak, Wandelt + @MIC2022 meeting
- NonG = from nearly G aka **perturbative nonG everywhere distributed - Planck well-constrained, LSS measures just beginning (via MIC+)** - to **space (\mathbf{x}) and resolution (\mathbf{k}) confined nonG extreme peaks** aka concentrations to wall, string, etal ζ -memories => complex, so let simulations be our guide to what's possible, what may be generic, and how to measure the forms.
- Conclude: with m_{AB}^2 temporarily < 0 we find localized concentrations of NonG is a generic feature of breaking/restoring symmetry during inflation for a wide range of potential parameters. Superposed almost statistically independently (?) on a relatively Gaussian base. => **In search of Buried Treasure in LSS/CMB**

varieties of primordial nonG and how to search for them

perturbative: the nonG component is strongly correlated with the dominant Gaussian
see Planck2018 IX (nonG) for an exhaustive study and current constraints - using both T+Epol

Planck2018 IX arxiv 1905.05697 correlated nonG

nonG 3-point-correlation-pattern measure

f_{nl} : -0.9 ± 5.1 local for -Newton potential

$\Rightarrow f_{NL^*} = -1.54 \pm 3.0$ for phonons/V-strain $f_{NL^*} \sigma_{\zeta} < .0006$

f_{nl} : -26 ± 47 equilateral

-38 ± 24 orthogonal

SO to .0003

cf. Planck2015 XVII

nonG 3-point-correlation-pattern measure

$f_{nl} = 0.8 \pm 5.0$ local for -Newton potential

$\Rightarrow f_{NL^*} = -0.52 \pm 3.0$ for phonons/V-strain

$-f_{nl}$: -4 ± 43 equilateral

-26 ± 21 orthogonal

f_{nl} : $to \pm 0.5$! SphereX ~2023 - all-sky near-IR satellite 6 arcsec resolution 96 bands

Planck 2015/2018 VII (Isotropy & Statistics) arxiv 1906.02552. blind nonG stats in scale space \Rightarrow no strong evidence
e.g., Kolmogorov-Sinai test on $n(T, E | Prominences)$ at 2 scales. rare low-L anomalies, **cold spot in T not E**

beyond Planck2018: higher res CMB (SO,S4) $\sim f_{sky} L_{max}^2$ cf. LSS $f_{sky} L_{max}^2 \times k_{max} d_{max}$ z-space, LIM more modes

varieties of primordial nonG and how to search for them

perturbative: the nonG component is strongly correlated with the dominant Gaussian
see Planck2018 IX (nonG) for an exhaustive study and current constraints - using both T+Epol

Planck2018 IX arxiv 1905.05697 correlated nonG

nonG 3-point-correlation-pattern measure

f_{nl} : -0.9 ± 5.1 local for -Newton potential

$\Rightarrow f_{NL^*} = -1.54 \pm 3.0$ for phonons/V-strain $f_{NL^*} \sigma_{\zeta} < .0006$

f_{nl} : -26 ± 47 equilateral

-38 ± 24 orthogonal

SO to .0003

cf. BOSS bispectrum+power spectrum MIC22

nonG 3-point-correlation-pattern measure

$f_{nl} = -30 \pm 29$; -33 ± 28 local

\Rightarrow .. Senatore+; Zaldarriaga, Philcox+

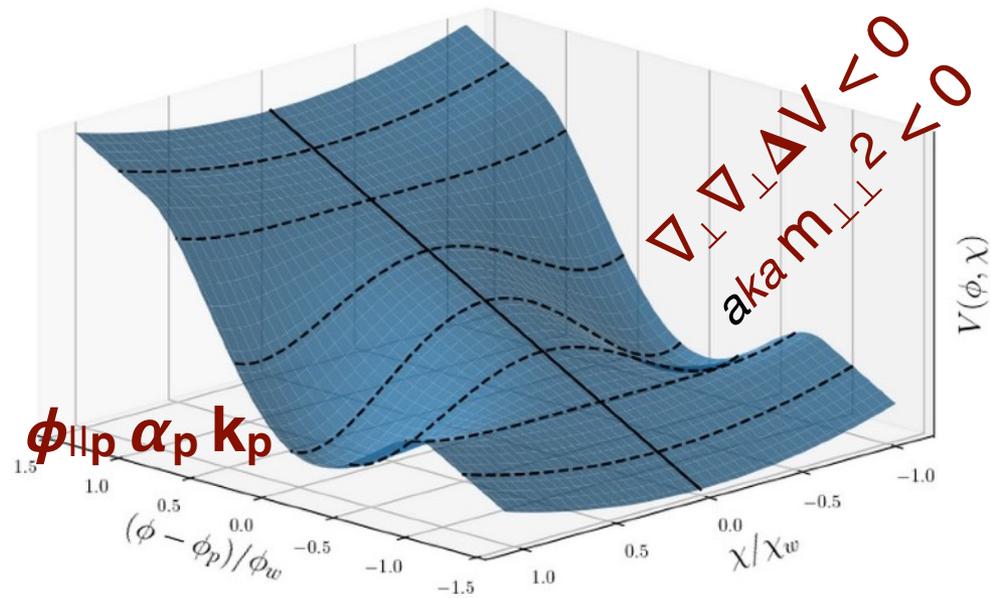
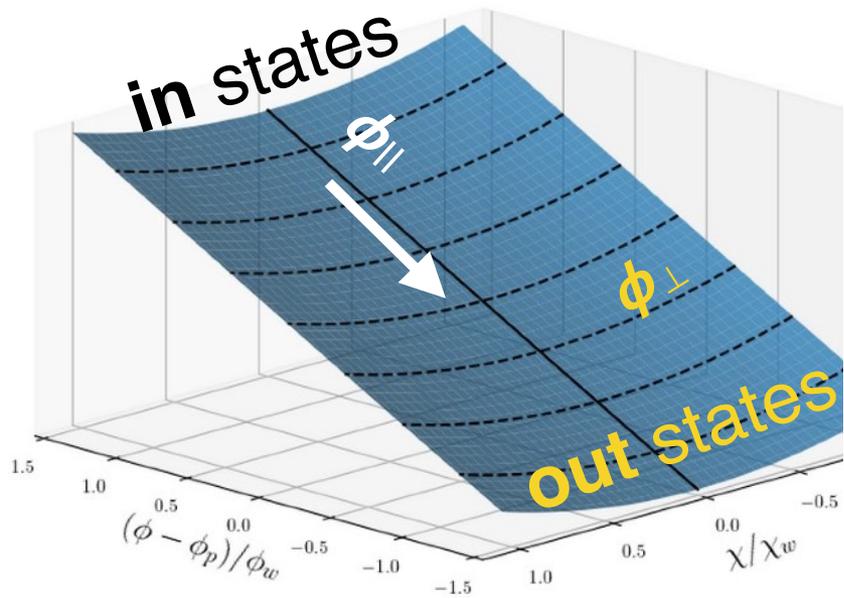
f_{nl} : 2 ± 212 ; 940 ± 600 equilateral

126 ± 72 ; 170 ± 170 orthogonal

f_{nl} : $to \pm 0.5$! SphereX ~2023 - all-sky near-IR satellite 6 arcsec resolution 96 bands

Planck 2015/2018 VII (Isotropy & Statistics) arxiv 1906.02552. blind nonG stats in scale space \Rightarrow no strong evidence
e.g., Kolmogorov-Sinai test on $n(T, E | Prominences)$ at 2 scales. rare low-L anomalies, cold spot in T not E

beyond Planck2018: higher res CMB (SO,S4) $\sim f_{sky} L_{max}^2$ cf. LSS $f_{sky} L_{max}^2 \times k_{max} d_{max}$ z-space, LIM more modes



We get to choose k_p and instability strength & V_0
 k_p may be high and buried or - with heavy constraints
 - in the CMB+LSS range. There could be many k_p
 instability events

experiments:
 χ -light \rightarrow χ -heavy
 ΔV wide \rightarrow ΔV narrow
 weak \rightarrow strong instability

unperturbed inflation flow +:

$$\Delta V = 1/4\lambda(\phi)[(\chi^2 - v^2)^2 - v^4],$$

controls strength & on-off

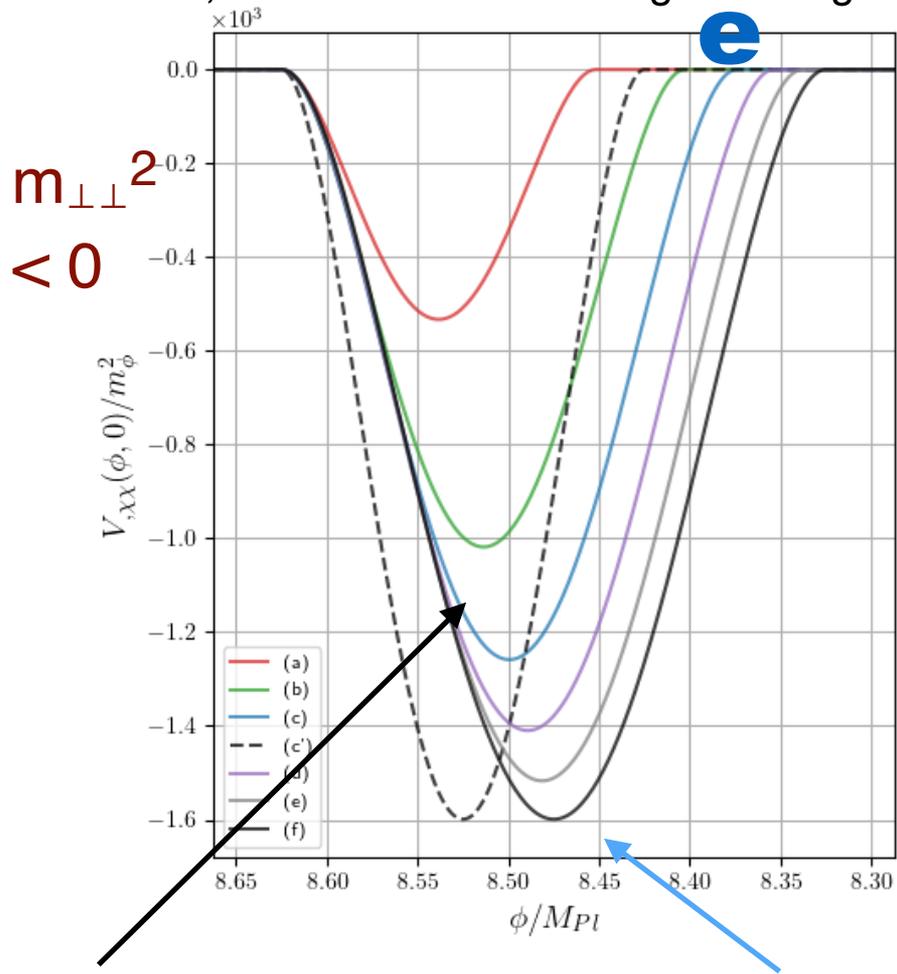
vev

an ensemble of symmetry breaking/restoration potentials

$$\Delta V = 1/4\lambda(\phi)[(\chi^2 - v^2)^2 - v^4],$$

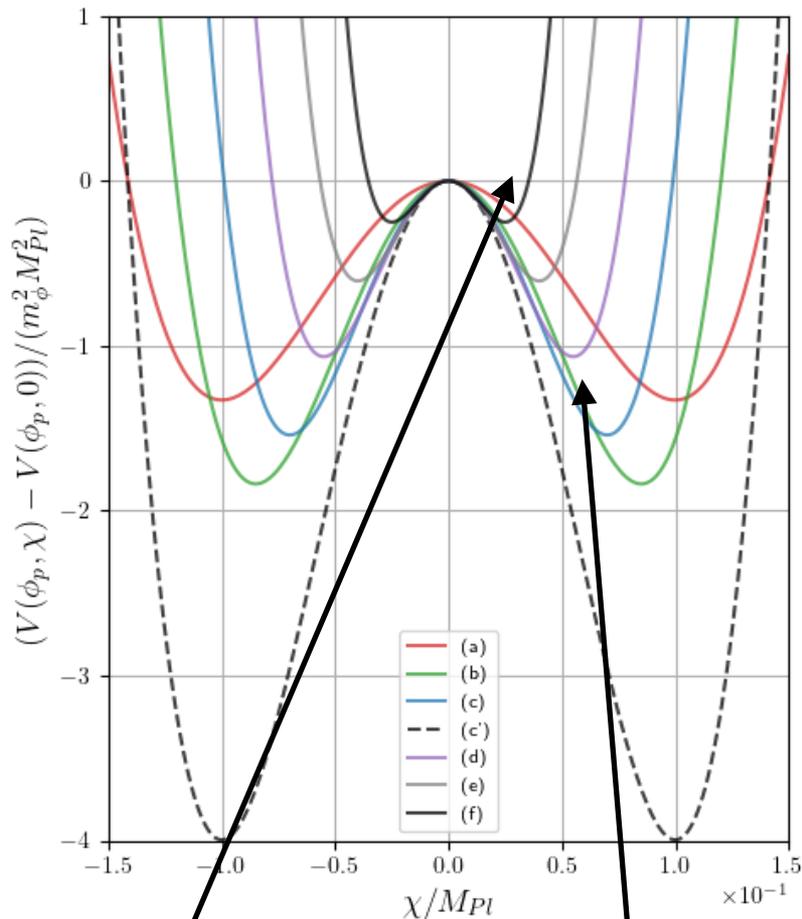
X U, U= 0 to 1 overall strength scaling

- a deep breath in the downward flow
- U makes differential templates easier to compute

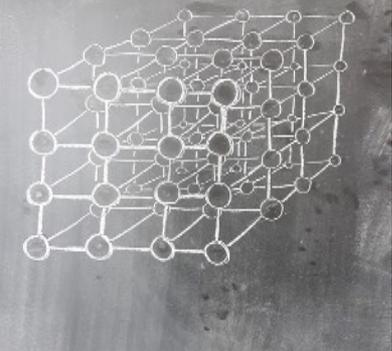


leads to ζ peaks - common

leads to wall memory in ζ



leads to ζ peaks - common

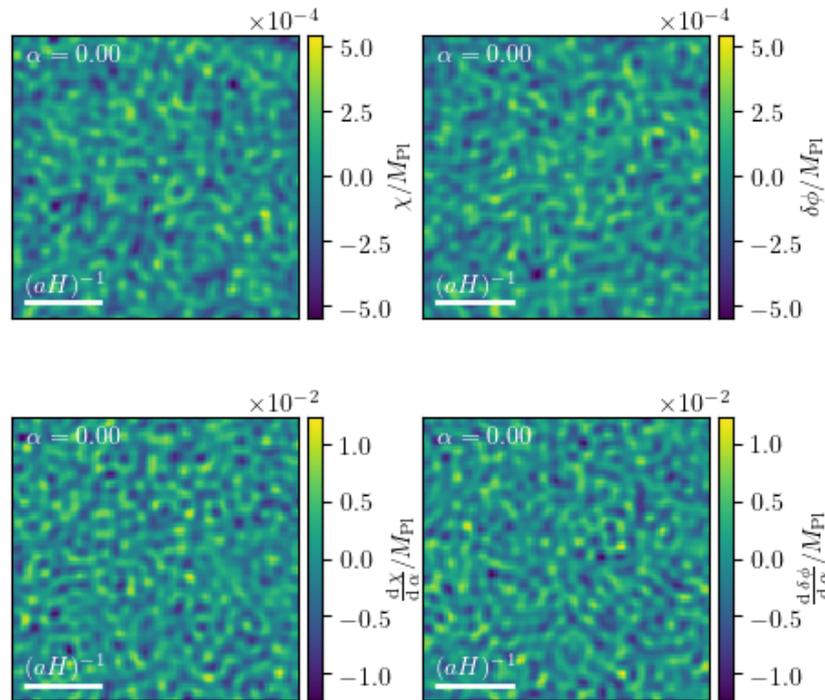


in state in the native phase space variables

Discretizing space onto a lattice recasts PDEs as many coupled ODEs, which can be integrated numerically.

Very high accuracy
pseudospectral code
with symplectic
integrator (developed by Braden for preheating)
=> extract very small nonG from full nonlinear dynamics, including all stochastic inflation effects

Energy conservation to 10^{-13}



$$\frac{d\phi^A}{d\tau} = a^{-2}\Pi^A,$$

$$\frac{d\Pi^A}{d\tau} = -\nabla^2\phi^A - a^4\frac{\partial V}{\partial\phi^A},$$

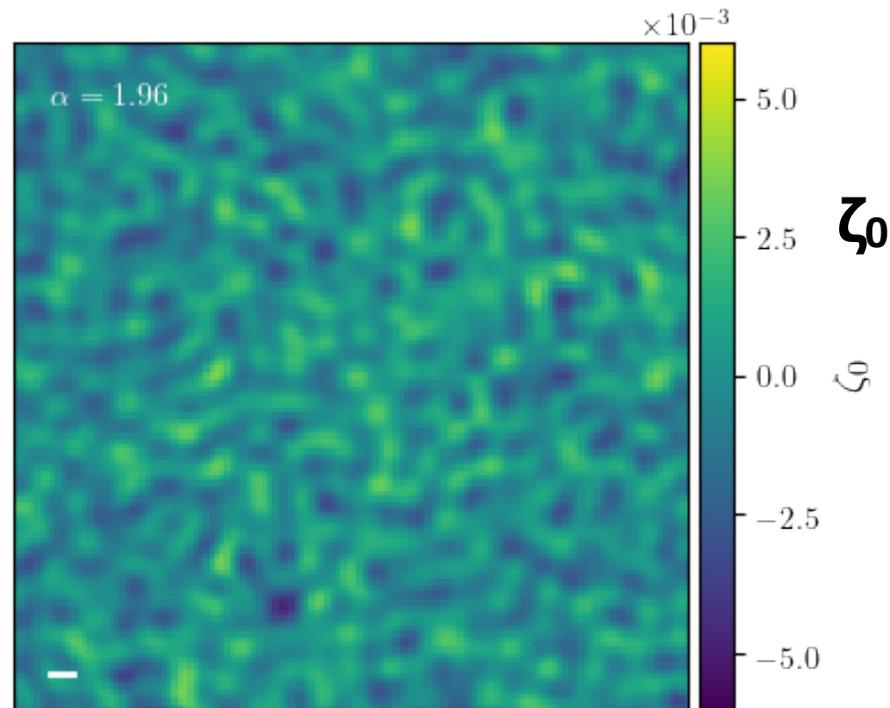
$$\frac{d\alpha}{d\tau} = \frac{-1}{6M_{Pl}}\Pi^a,$$

$$\frac{d\Pi^a}{d\tau} = -\langle -a^{-3}\Pi^A\Pi^A + a\nabla\phi^A \cdot \nabla\phi^A + 4a^3V \rangle.$$

$$\zeta = \frac{\dot{\phi}\nabla^2\phi + \nabla\dot{\phi} \cdot \nabla\phi + \dot{\chi}\nabla^2\chi + \nabla\dot{\chi} \cdot \nabla\chi}{3a^2(\rho + P)}$$

out state

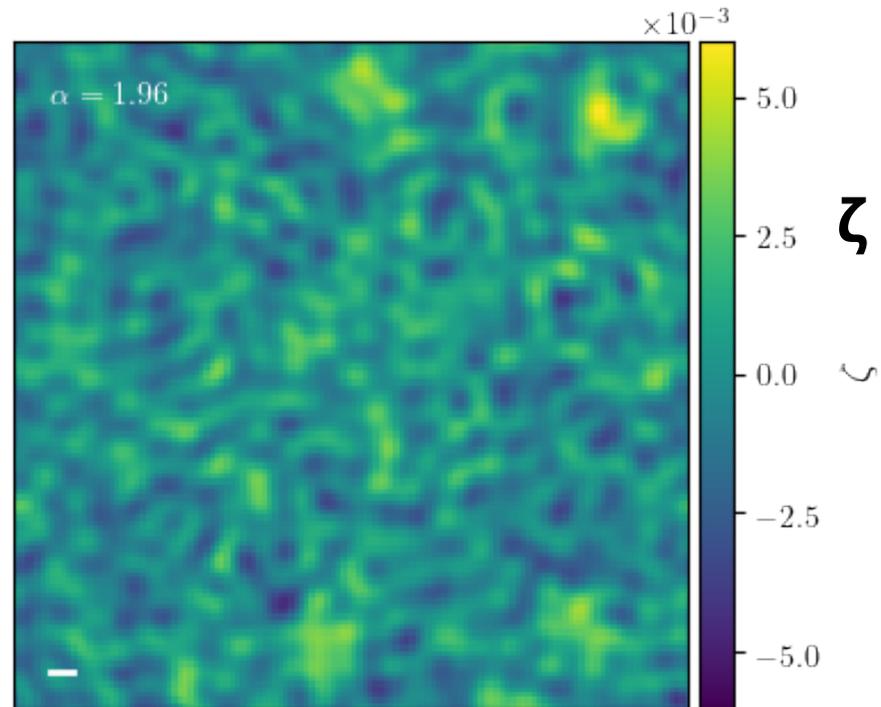
$$\zeta_f(x, t_f | V_0; \text{Bunch-Davies GRF IC})$$



V0 control of amplitude

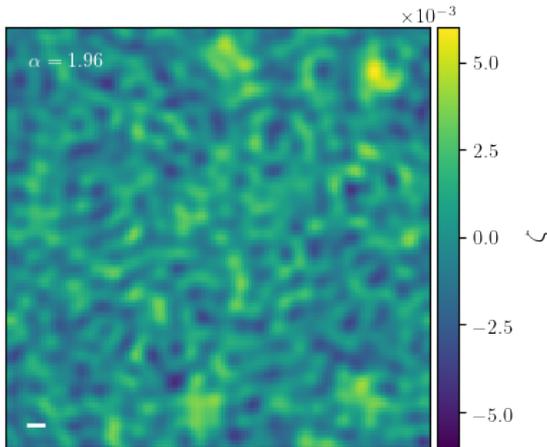
out state

$\zeta_f(x \mid \Delta V + V_0 ; \text{Bunch-Davies GRF IC})$

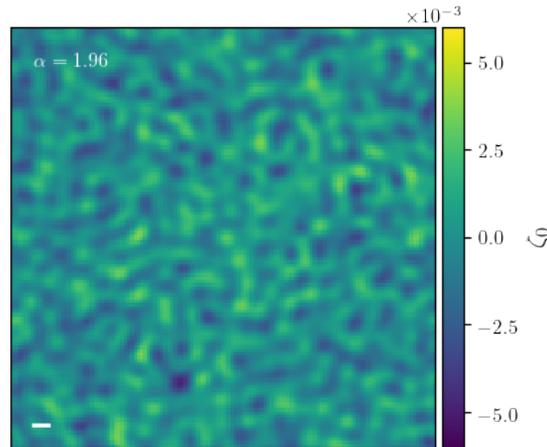


out state

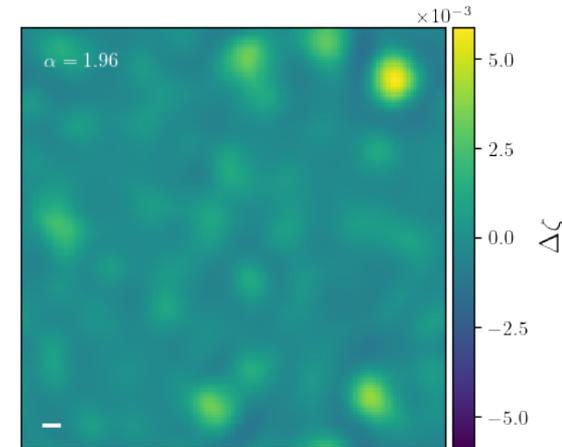
$$\zeta_f(x | \Delta V + V_0)$$



$$\zeta_f(x | V_0)$$



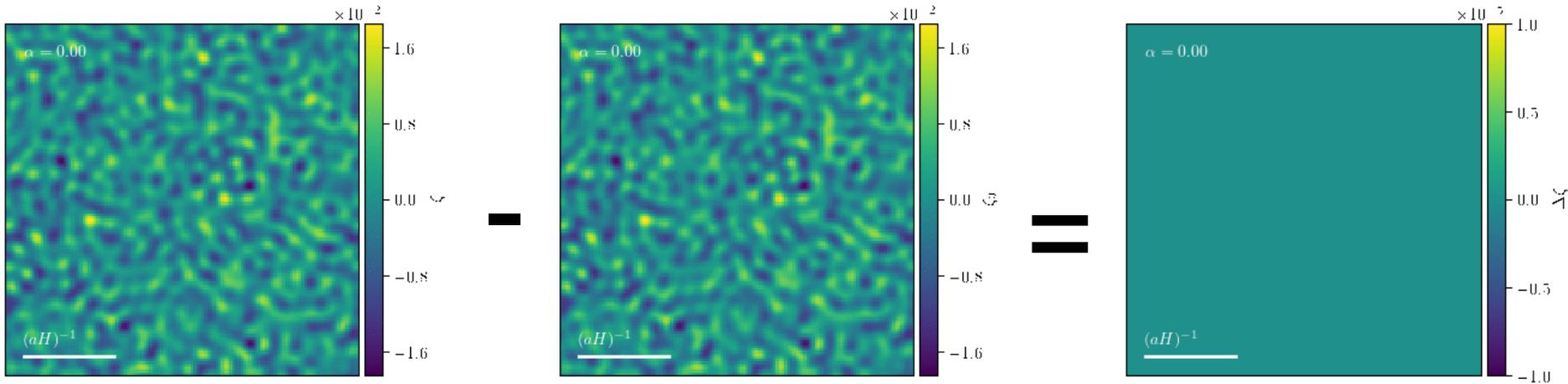
$$\Delta \zeta$$



Subtracting off a baseline sim using V_0 gives $\Delta \zeta = \zeta - \zeta_0$, this removes nearly Gaussian 'noise' V_0 generates from the NonG signal.

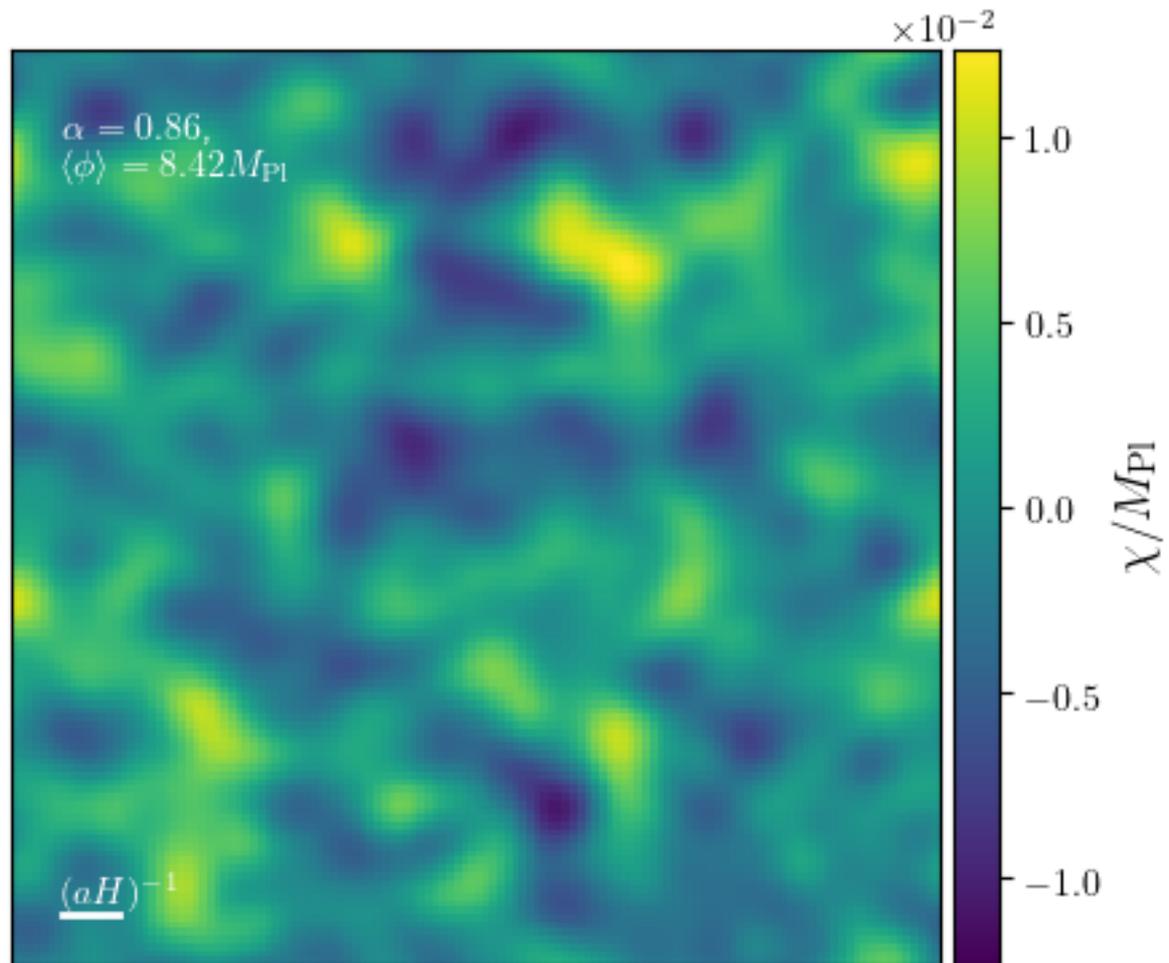
=> $\Delta \zeta$ has strong local NonG concentrations. Convolved with transfer function to get the gravitational potential for N-body/gas, linear to nonlinear cosmic web. How to unravel the $\Delta \zeta$ from the gastrophysical dynamics. Subtle effects in CMB/LSS => Top down is easier than bottom up 'backward' unravelling.

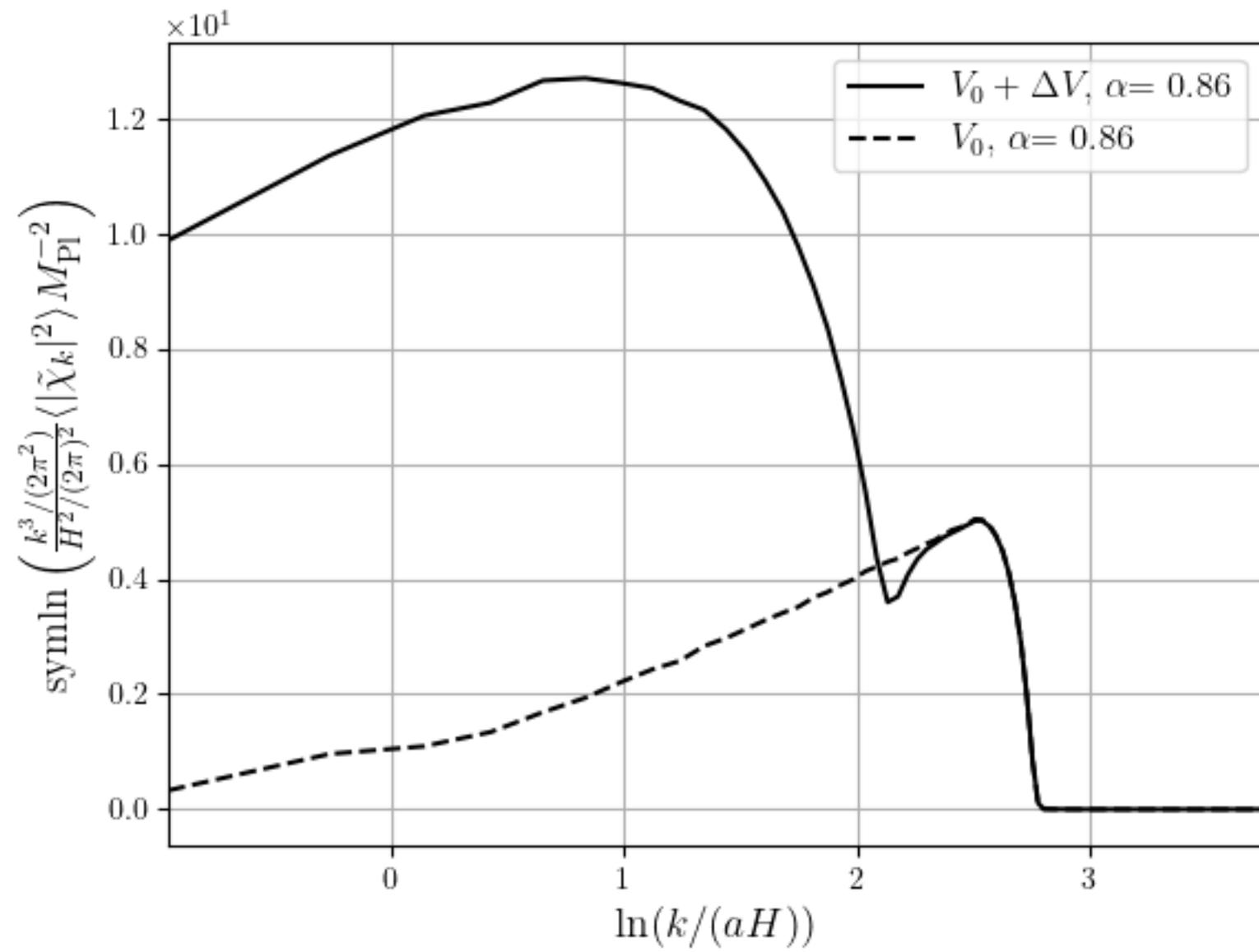
in state \Rightarrow out state, the movie



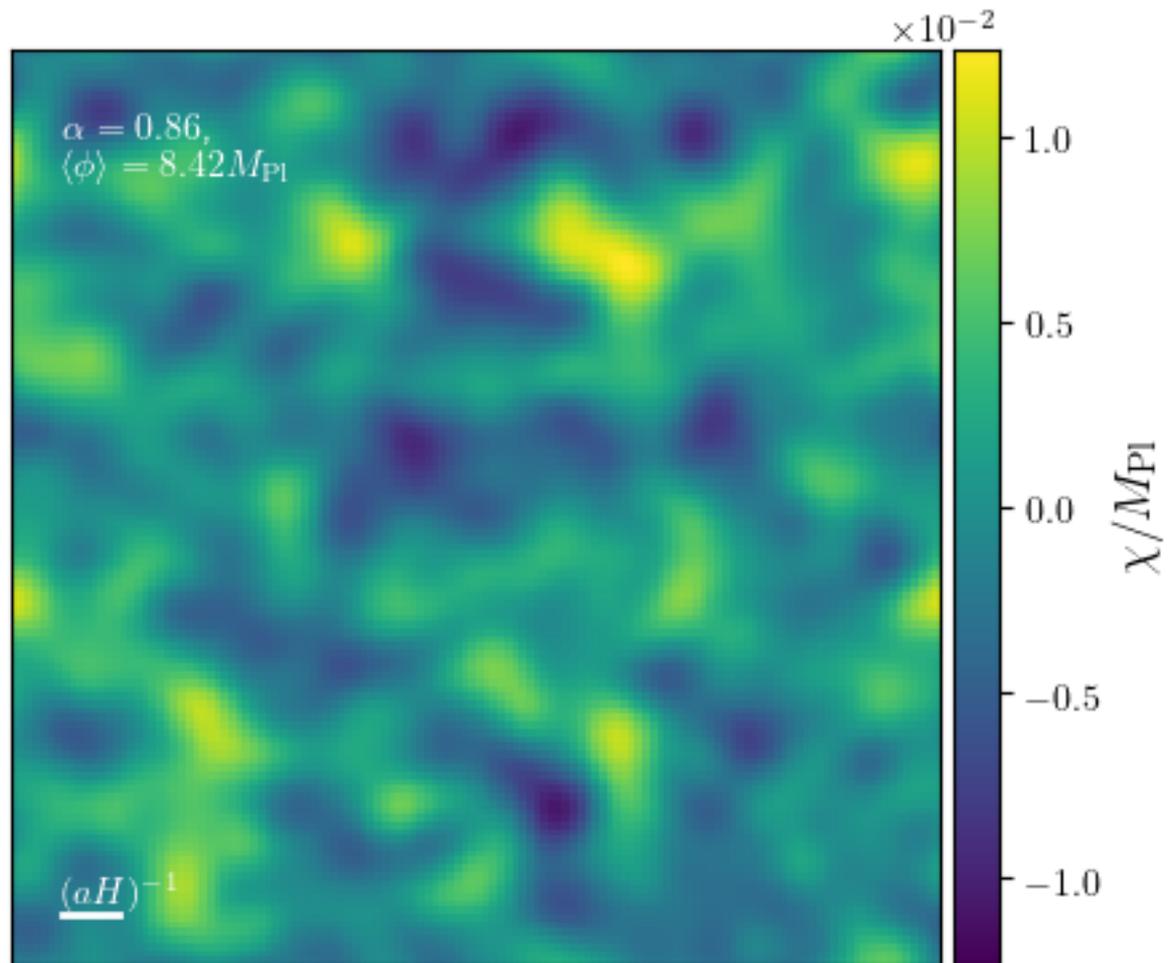
$\Delta\zeta_f > 0$ in the final state. Shorter arrested dynamics can lead to $\Delta\zeta_f < 0$

slice_chi_e_cp

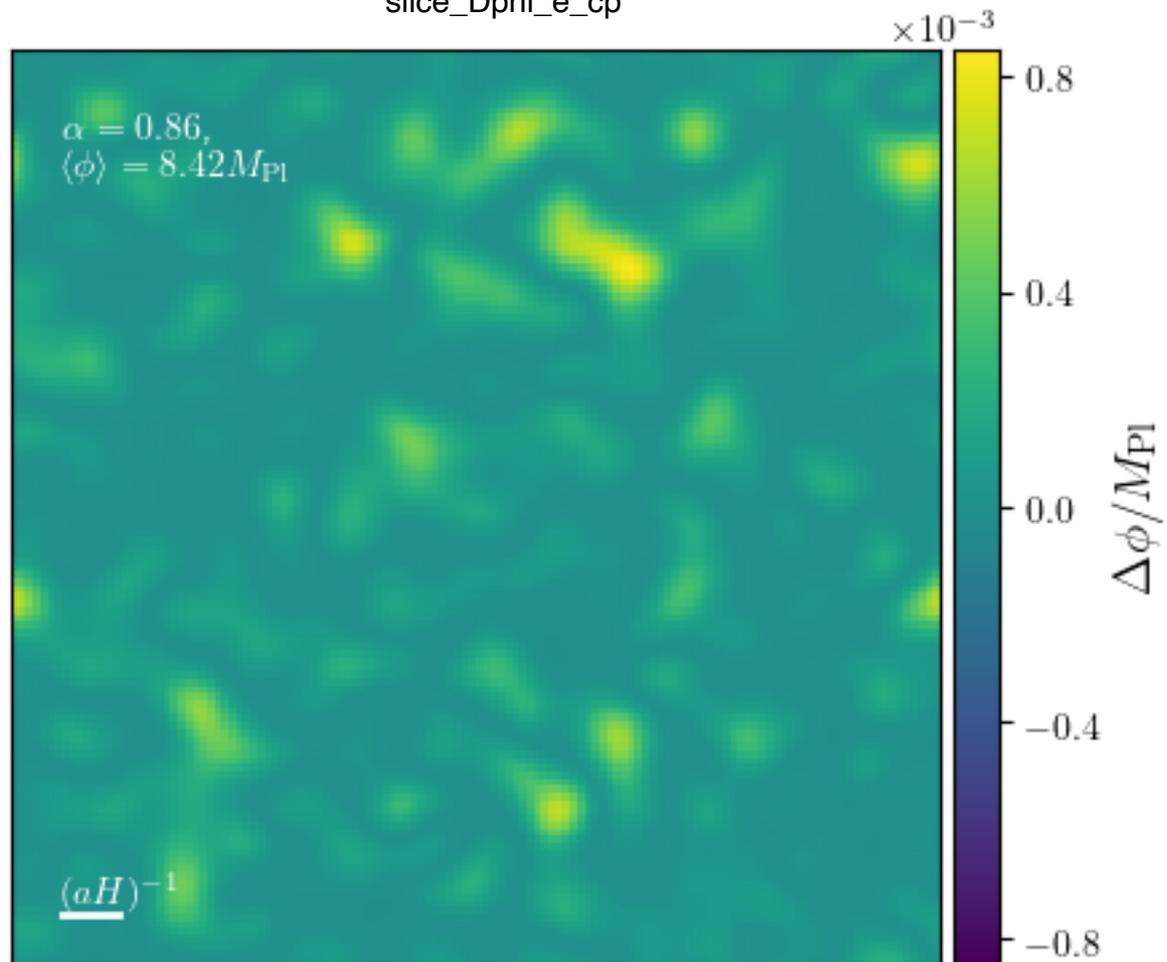




slice_chi_e_cp

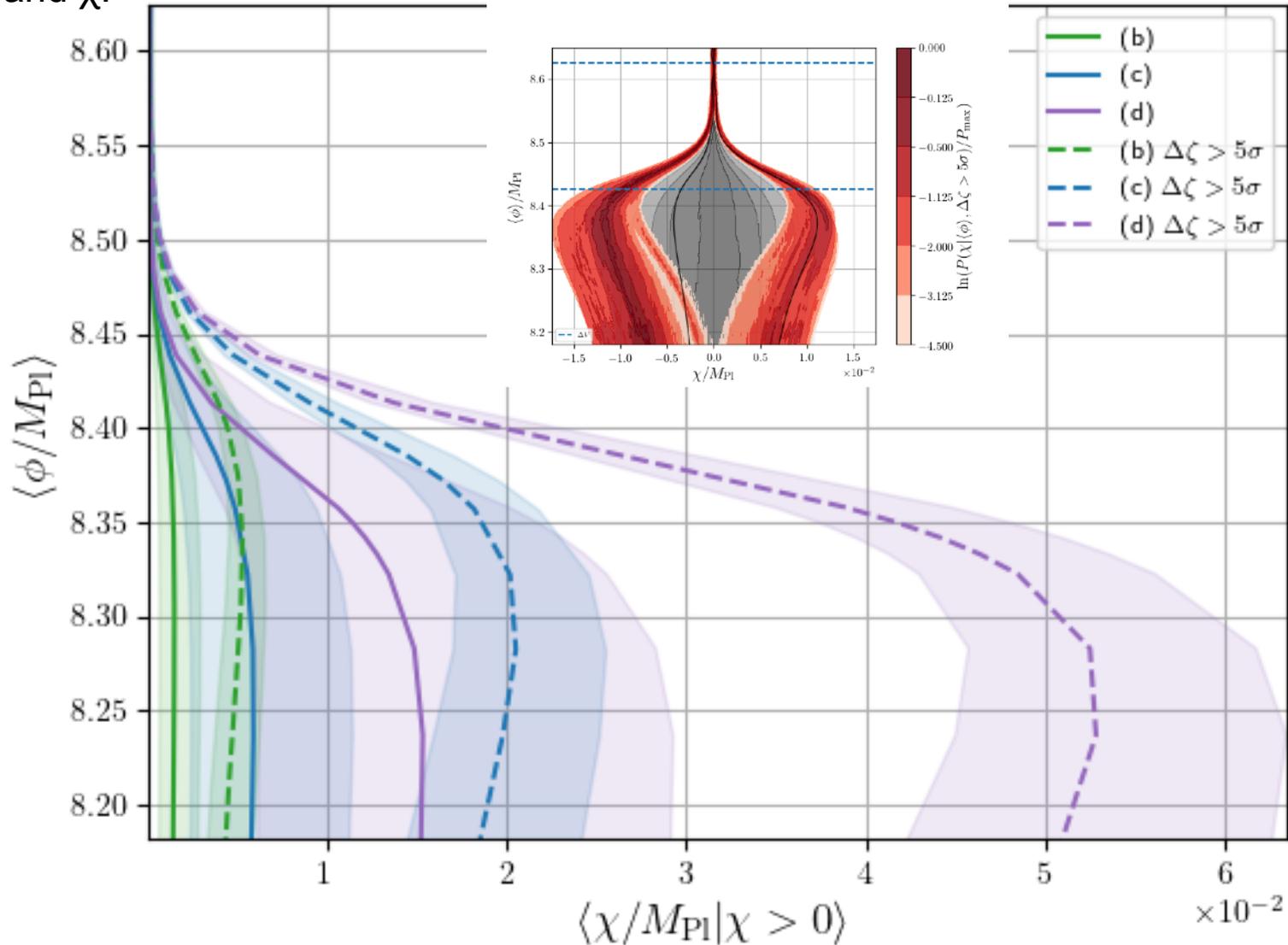


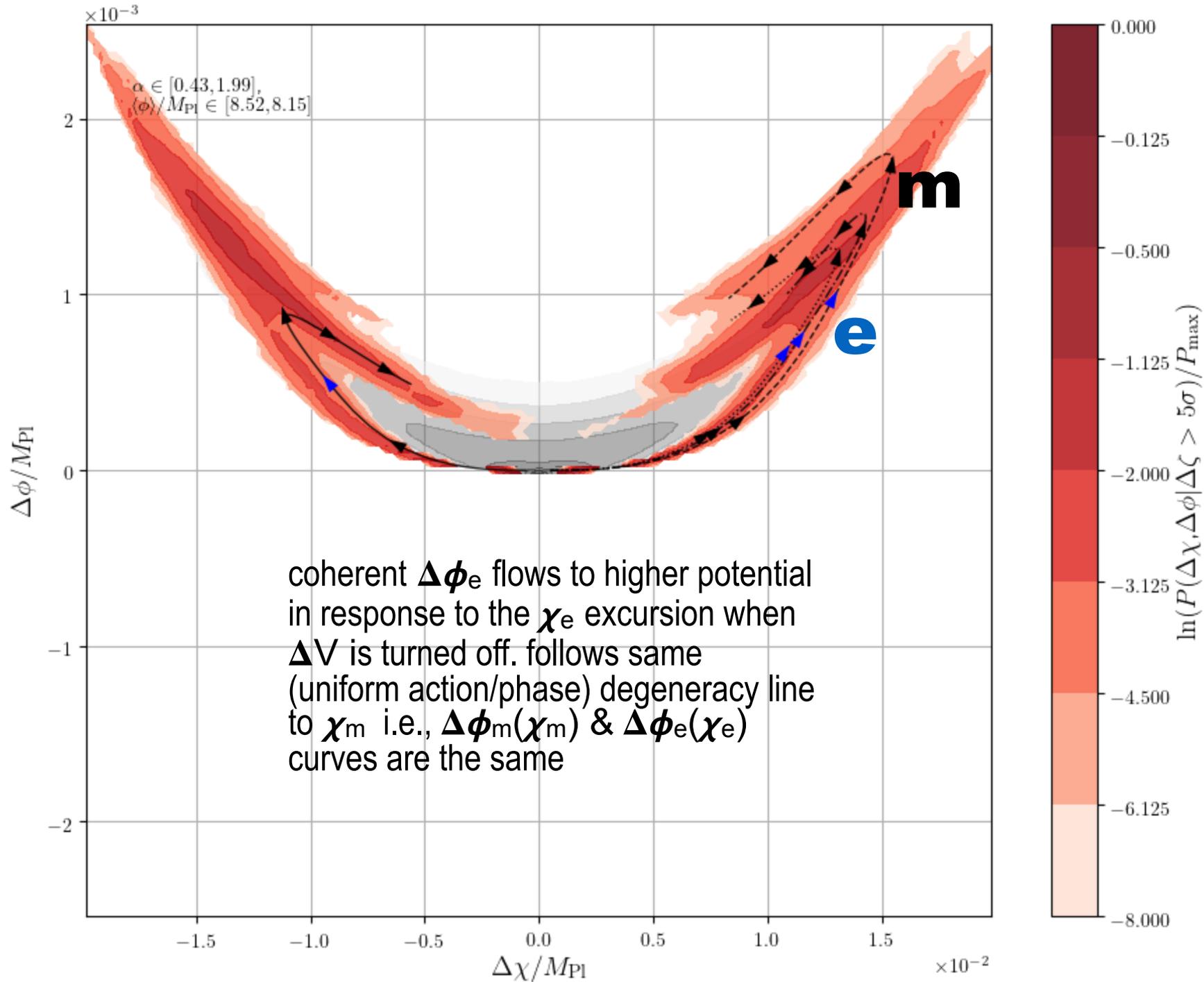
slice_Dphi_e_cp

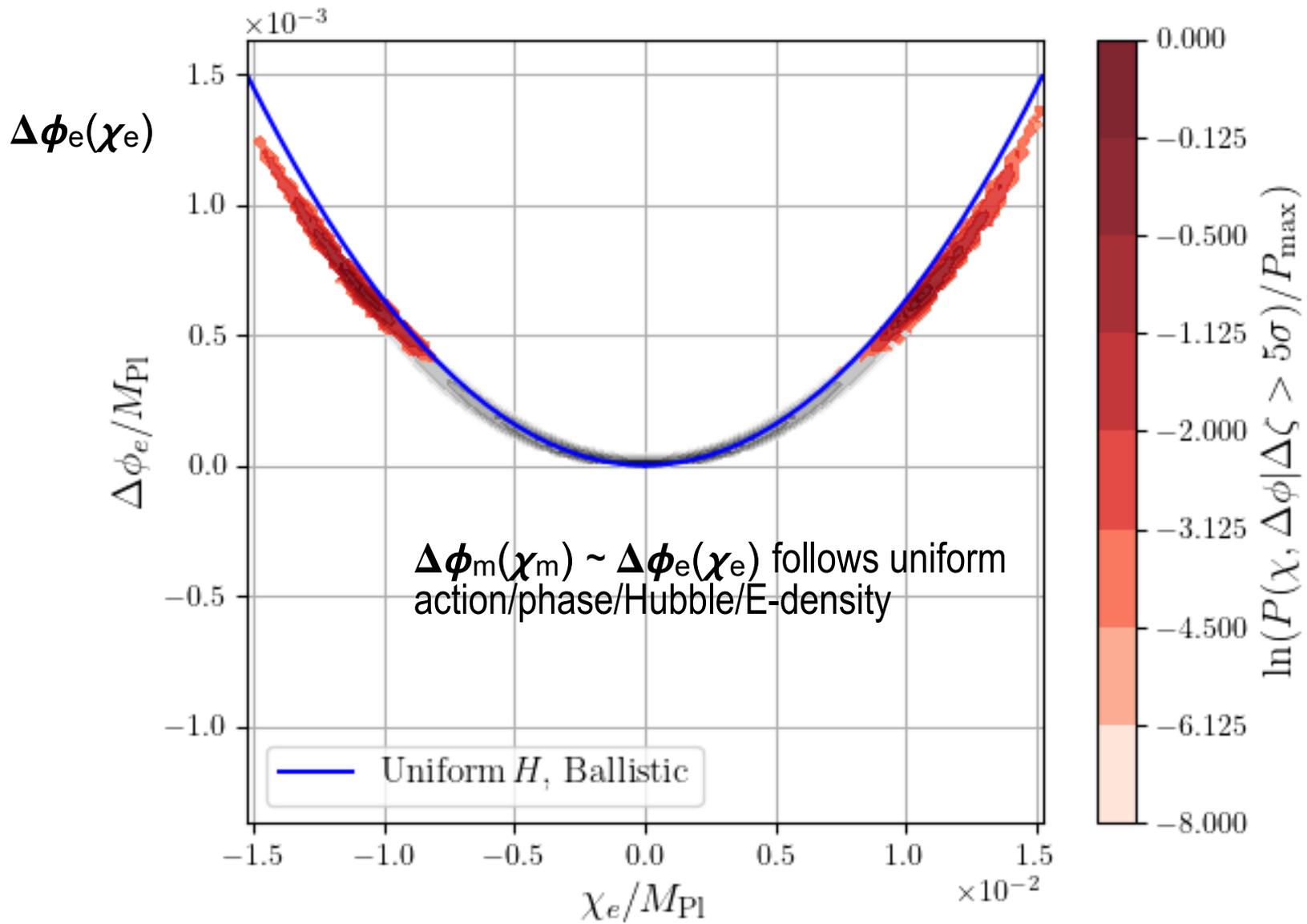


post select trajectories with $\Delta\zeta > 5\sigma$.

Trajectories leading to $\Delta\zeta$ concentrations are those which undergo the maximum excursion during the ΔV instability, leading to strongest nonlinear interactions between ϕ and χ .



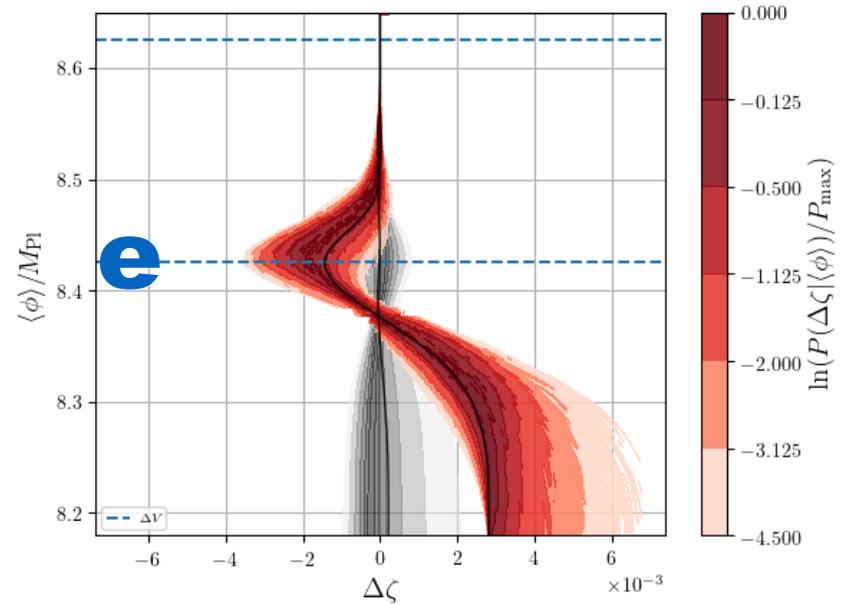
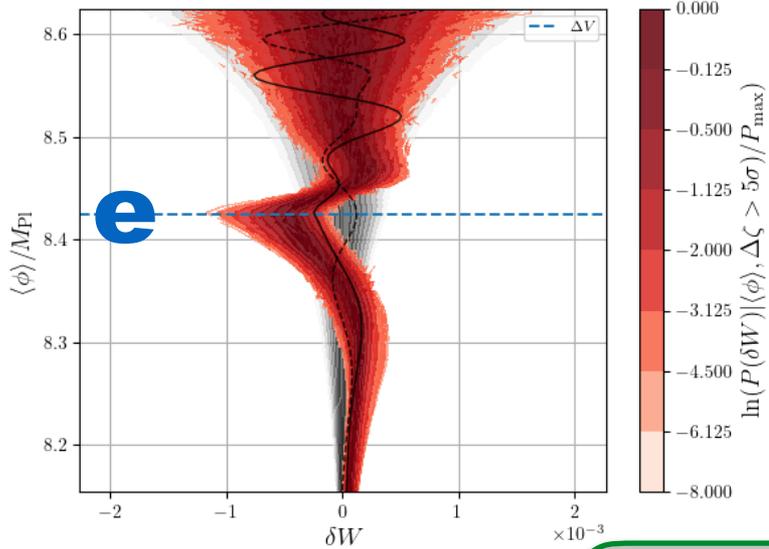




post – burst final : $\Delta\zeta_p \approx \Delta\phi_p(\chi_e)/\sqrt{2M_P\epsilon_e}$ or $\Delta\zeta_p \approx \Delta\phi_p(\chi_m)/\sqrt{2M_P\epsilon_m}$, i.e., function of χ_e

post select trajectories with $\Delta\zeta > 5\sigma$.

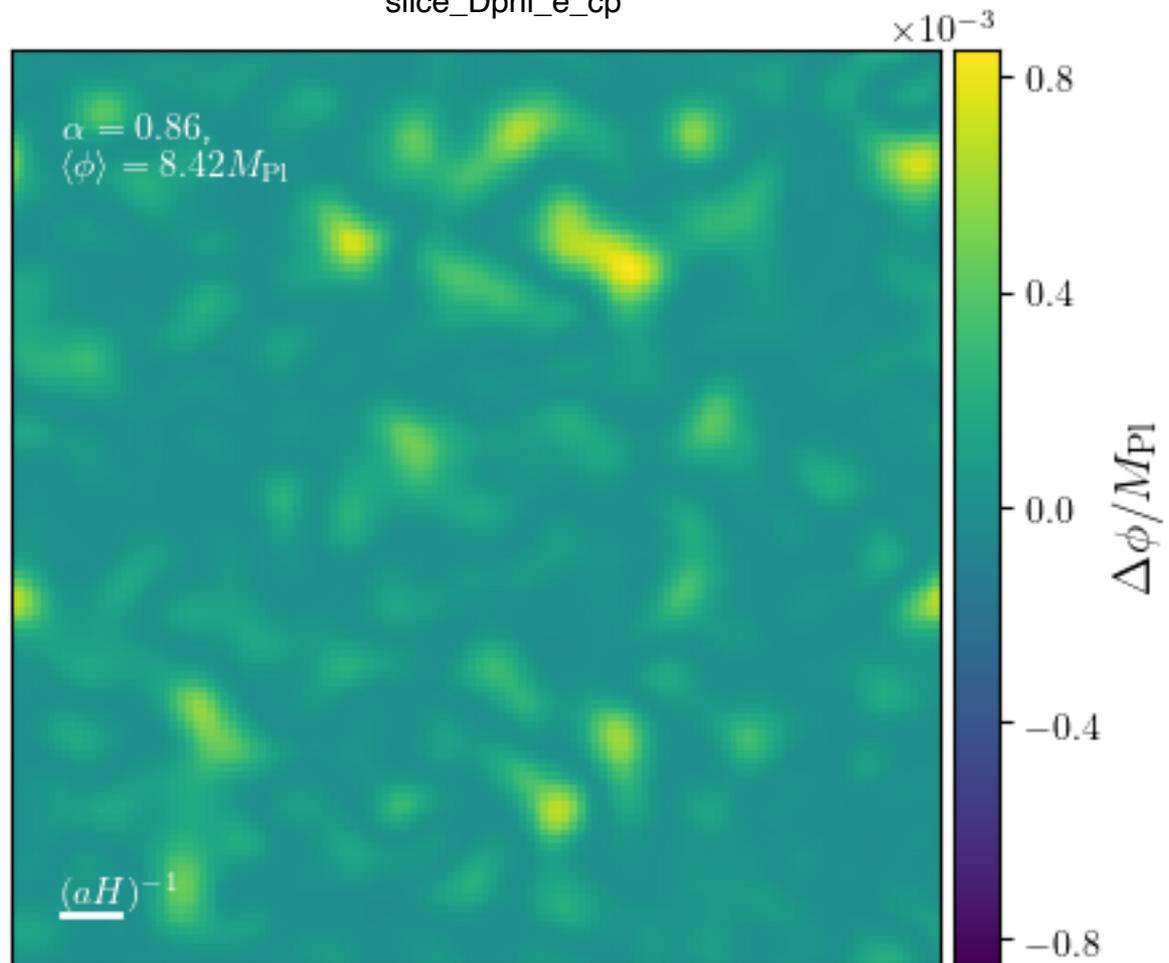
Trajectories leading to $\Delta\zeta$ concentration are those which undergo the maximum excursion during the ΔV instabilities due to nonlinear interactions between ϕ and χ .



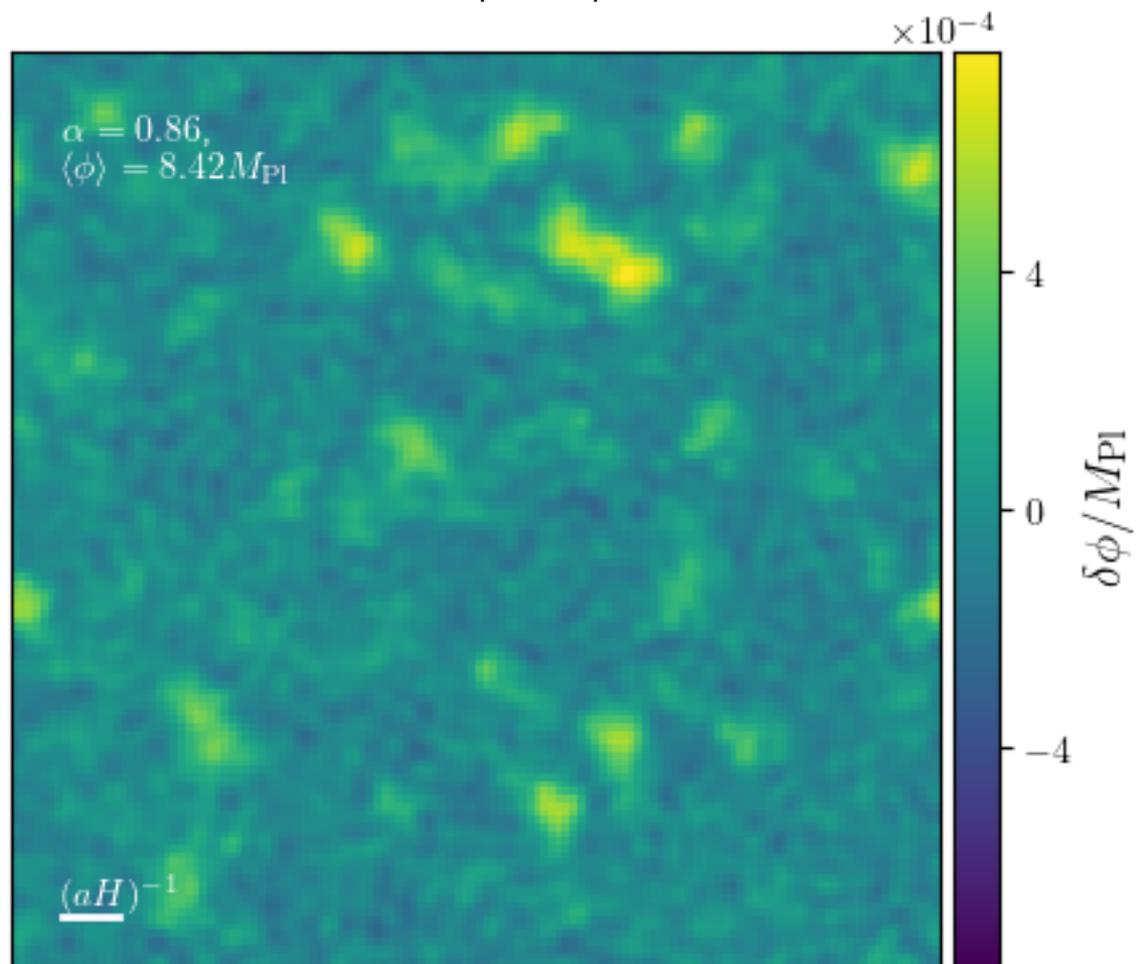
$$d\zeta = \left\{ \nabla^2 \delta\phi_A \cdot d\bar{\phi}^A + \nabla \delta\phi_A \cdot \nabla d\delta\phi^A + \nabla^2 \delta\phi_A \cdot d\delta\phi^A \right\} / 3(\rho + p)_{tot}$$

$$\text{Stochastic framework : } d\zeta_{cg} = d\zeta_{fg \rightarrow cg, Linear} + d\zeta_{cg, NL} + d\zeta_{fg \rightarrow cg, NL} = -d \ln \epsilon$$

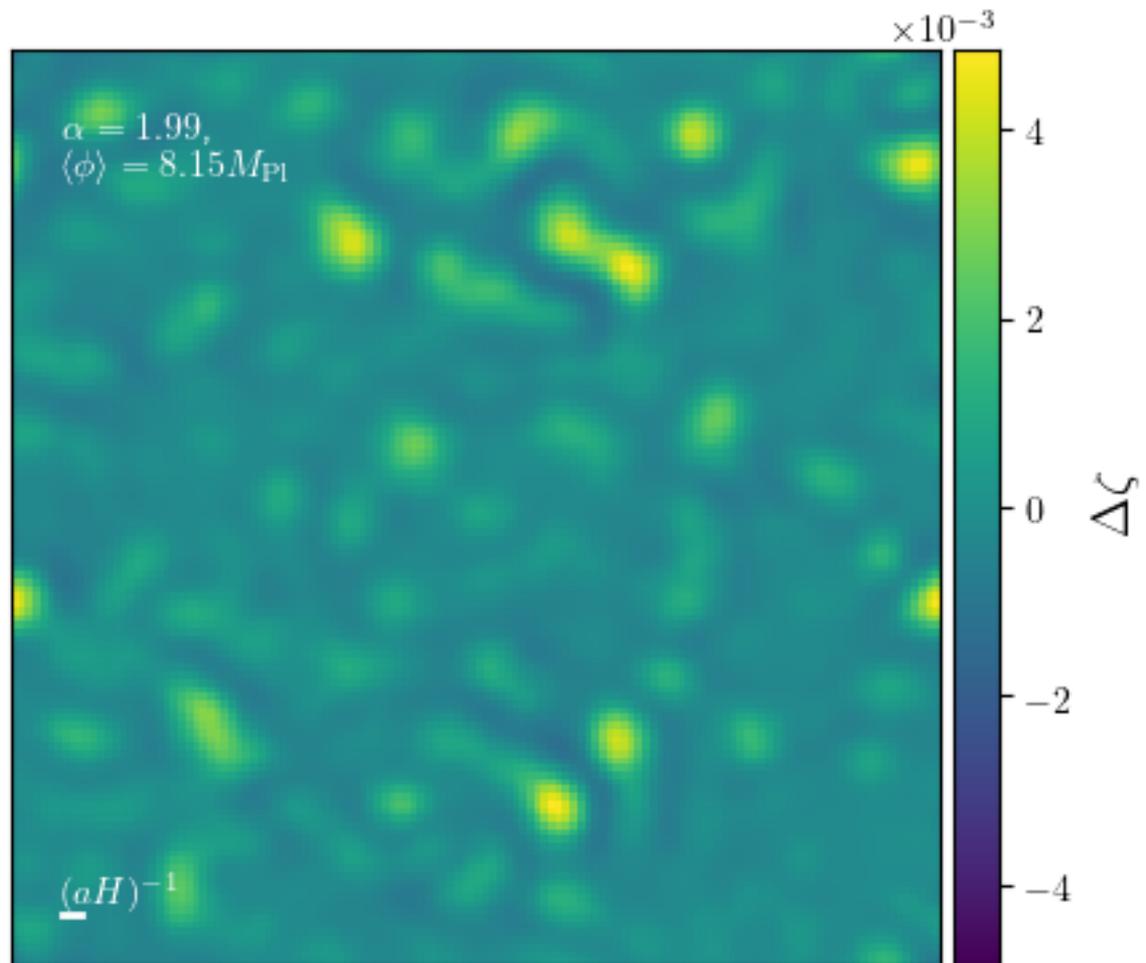
slice_Dphi_e_cp



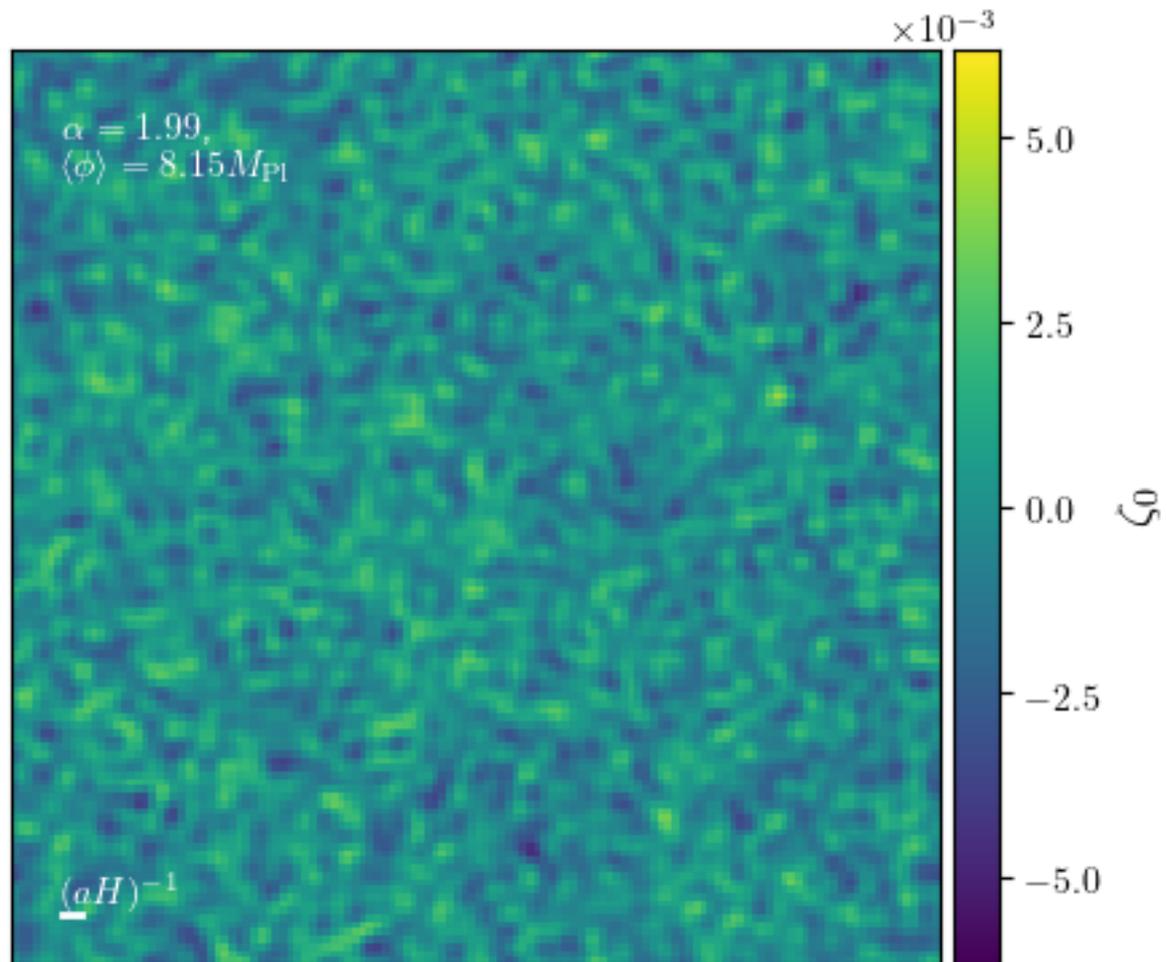
slice_phi_e_cp



slice_Dzeta_f_cp

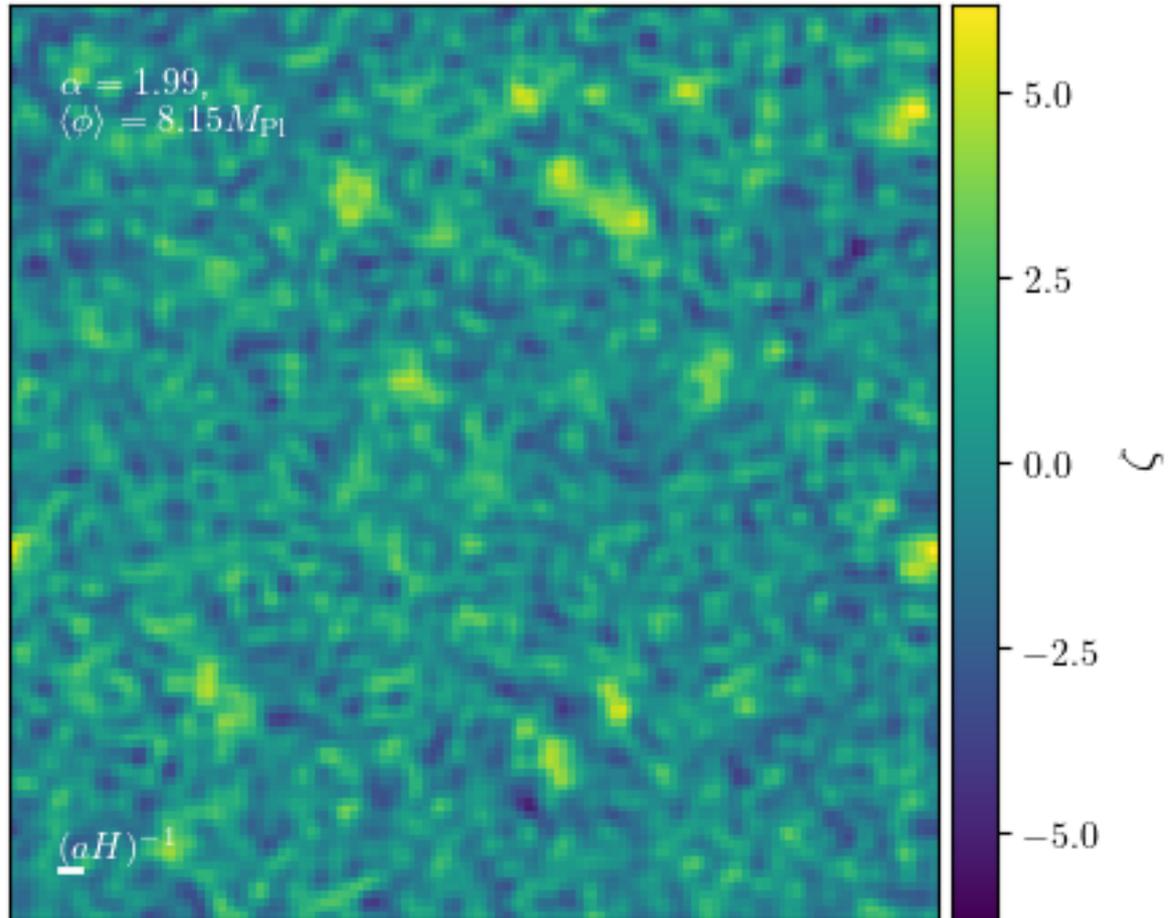


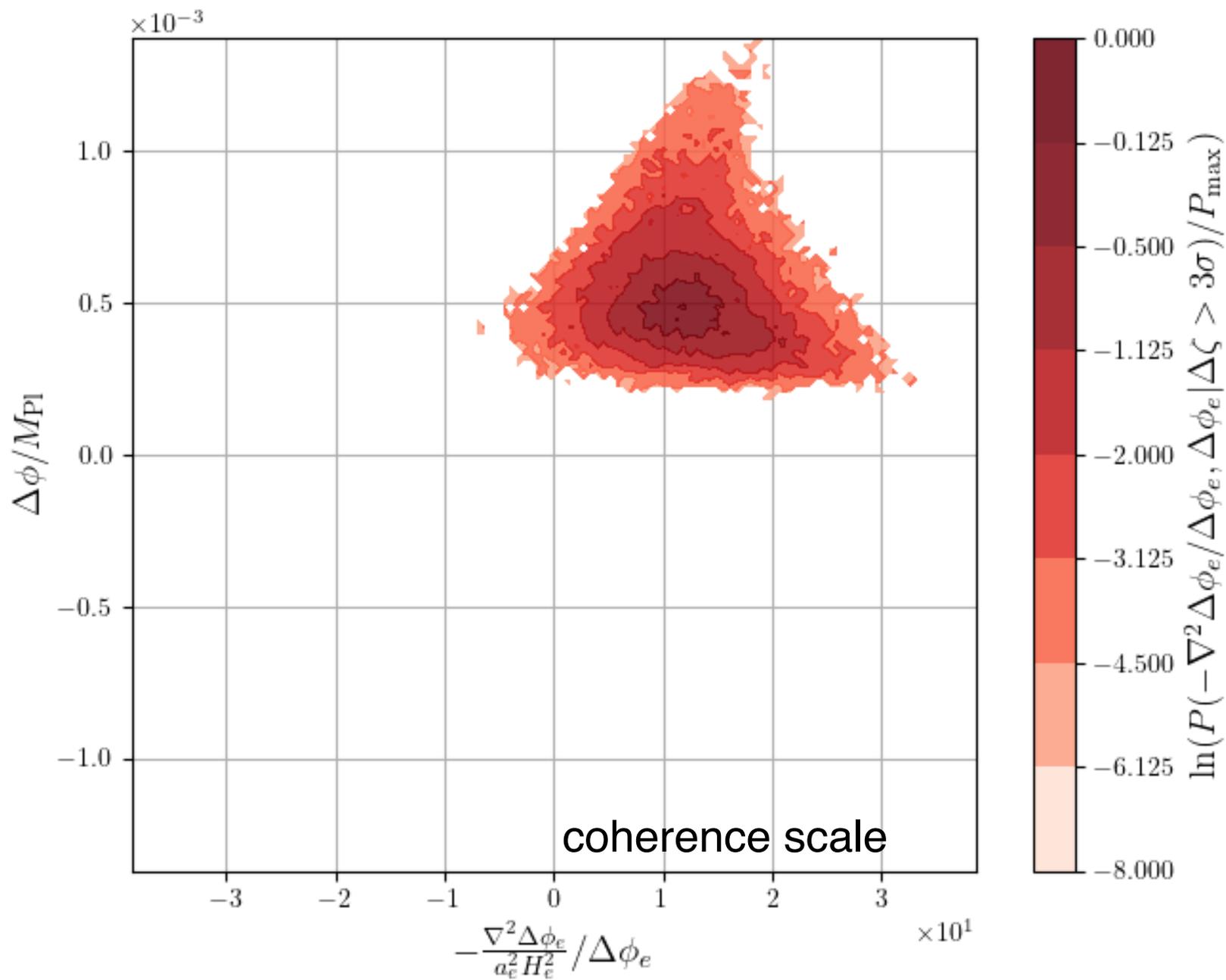
slice_zeta0_f_cp

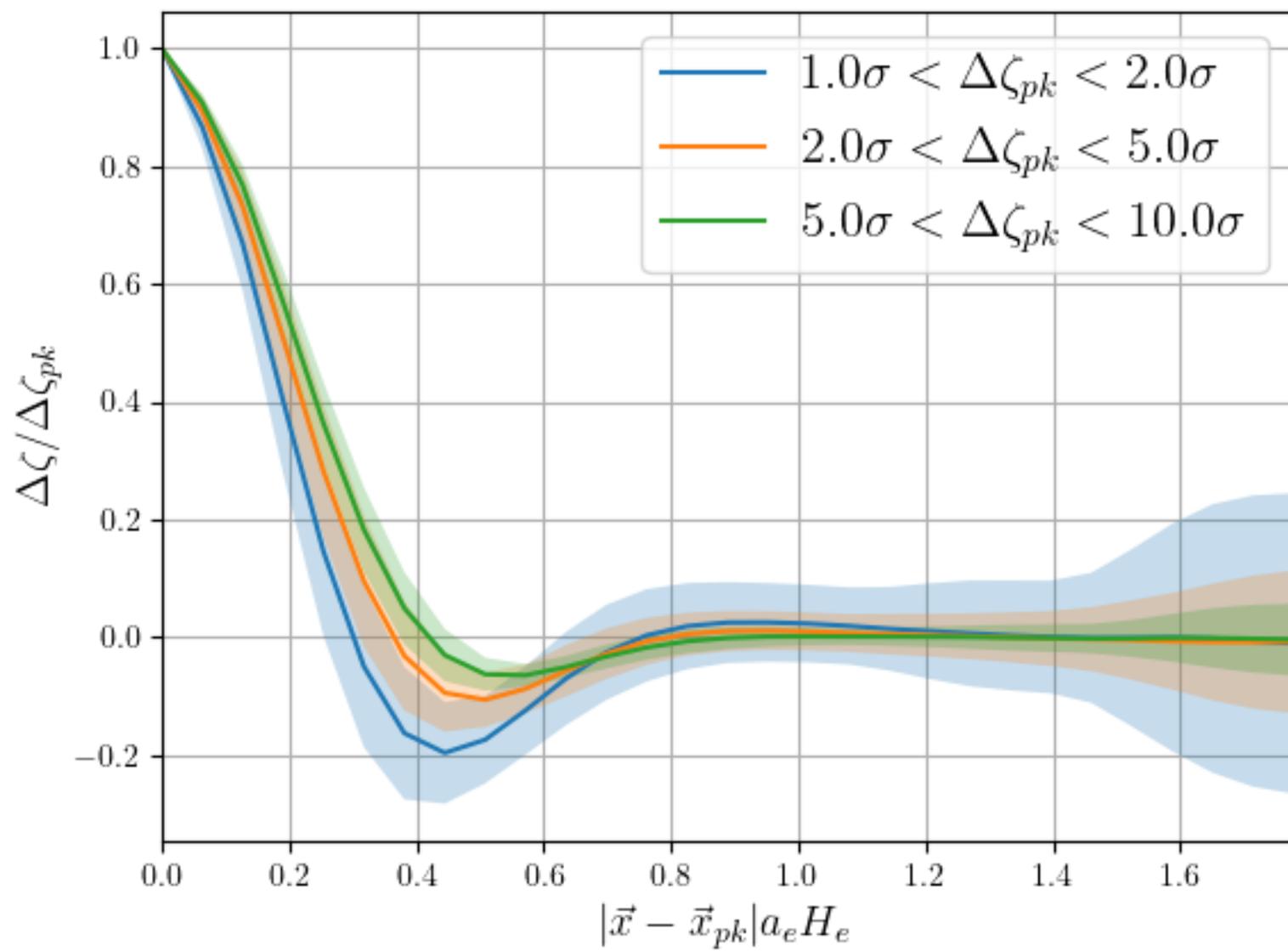


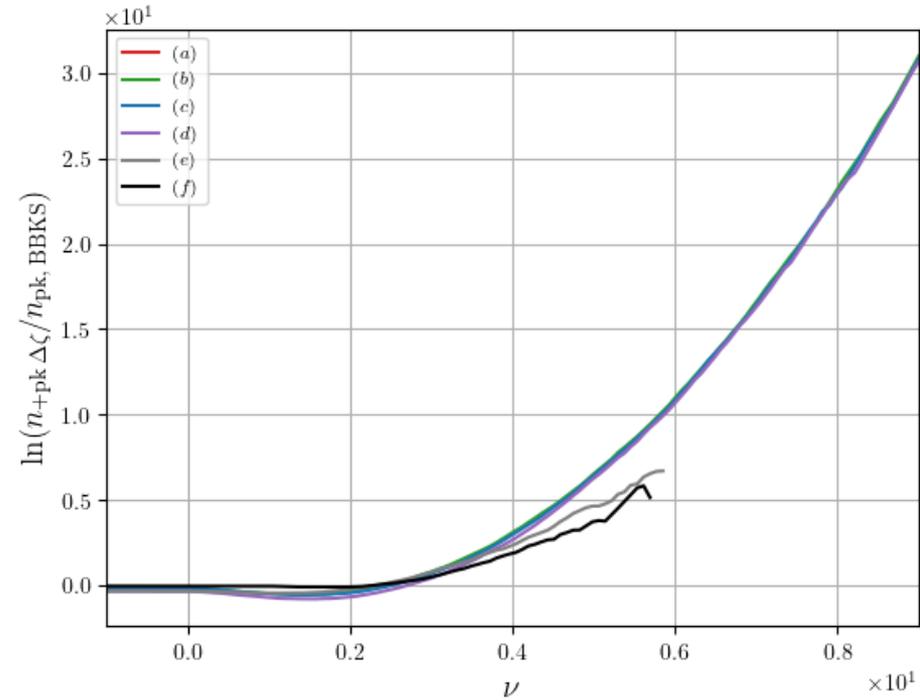
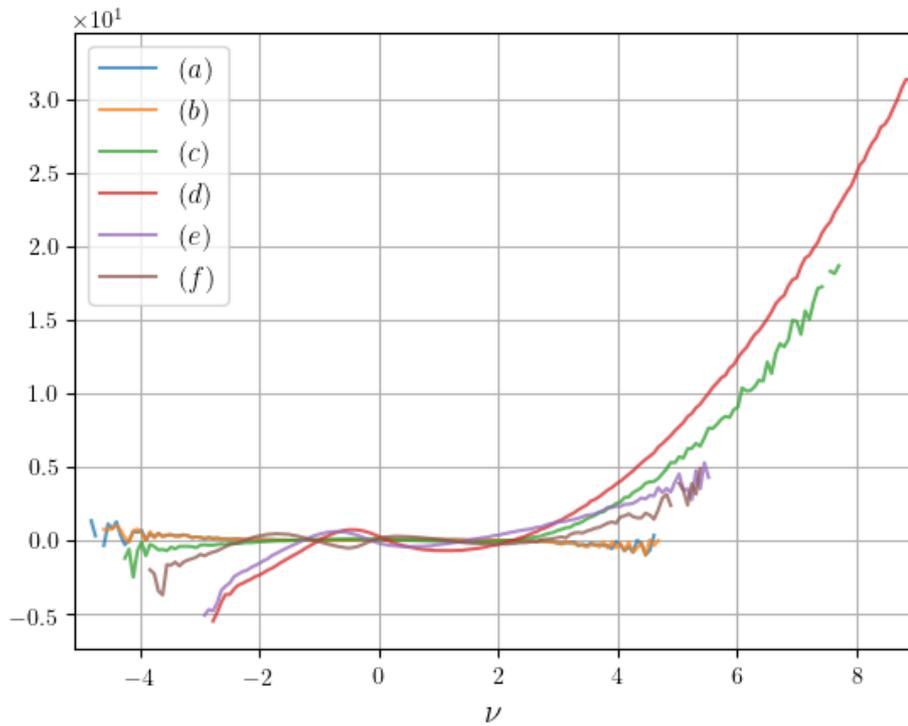
slice_zeta_f_cp

$\times 10^{-3}$







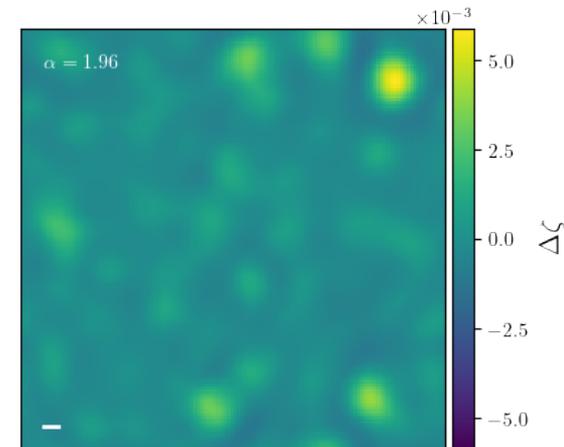


1-point prob dist function measures
NonG induced by ΔV .

NonG shows up as a long tail of high
excursions in excess of the expectation
for a Gaussian field, which is close to
what V_0 gives.

peak density relative to BBKS

$\nu = \text{field}/\text{sigma_field}$

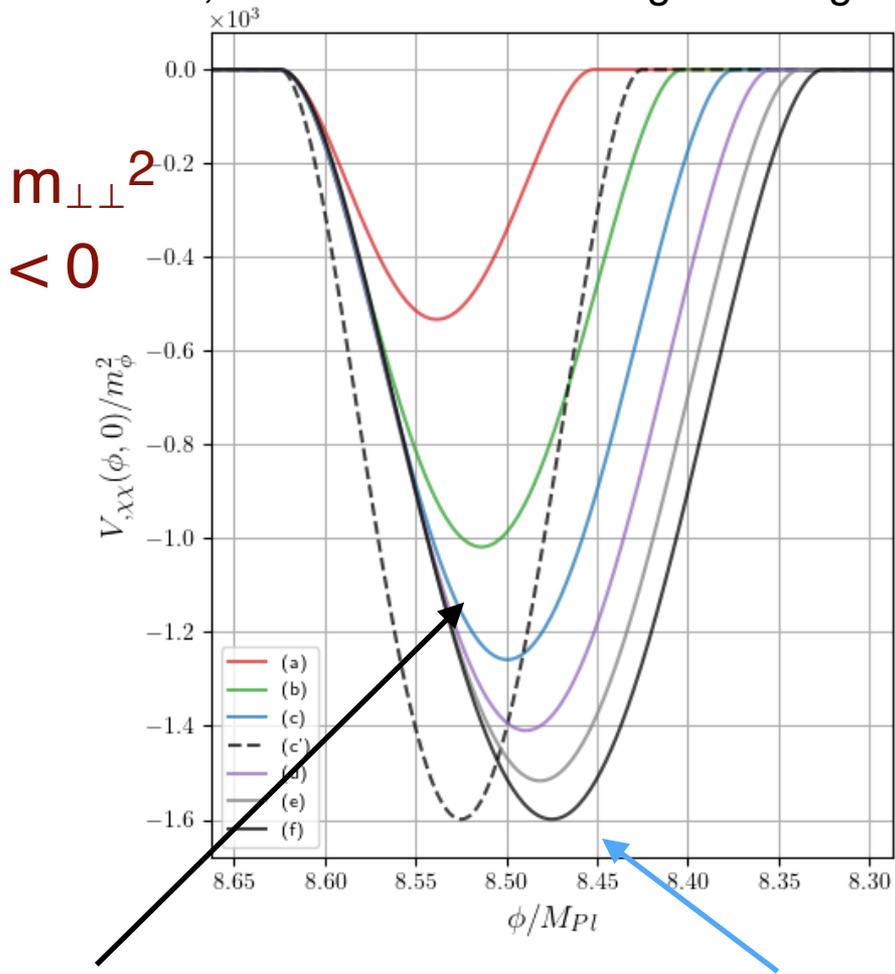


an ensemble of symmetry breaking/restoration potentials

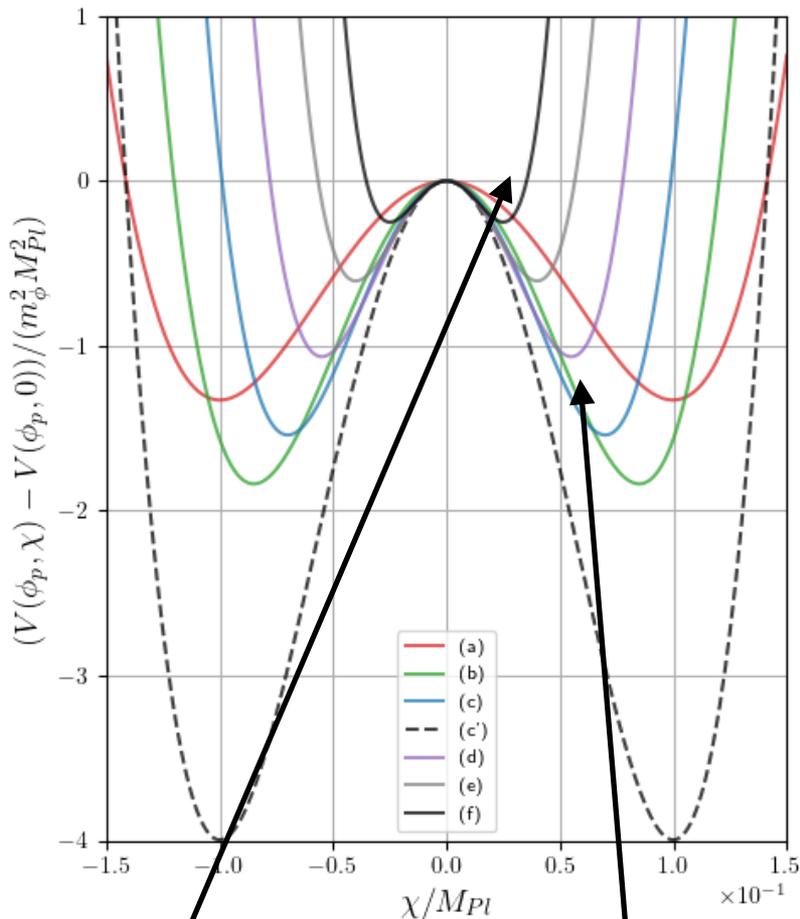
$$\Delta V = 1/4\lambda(\phi)[(\chi^2 - v^2)^2 - v^4],$$

X U, U= 0 to 1 overall strength scaling

- a deep breath in the downward flow
- U makes differential templates easier to compute



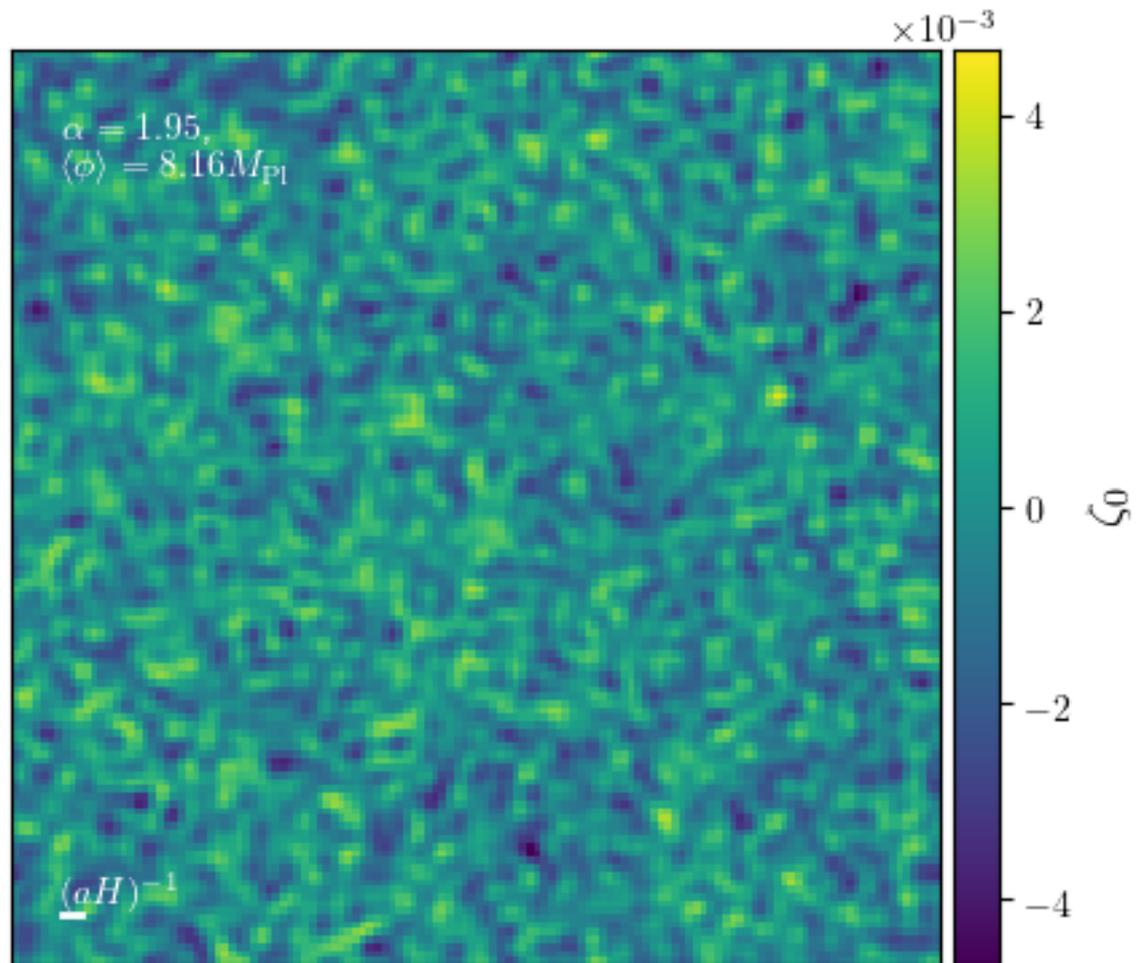
$m_{\perp\perp}^2 < 0$



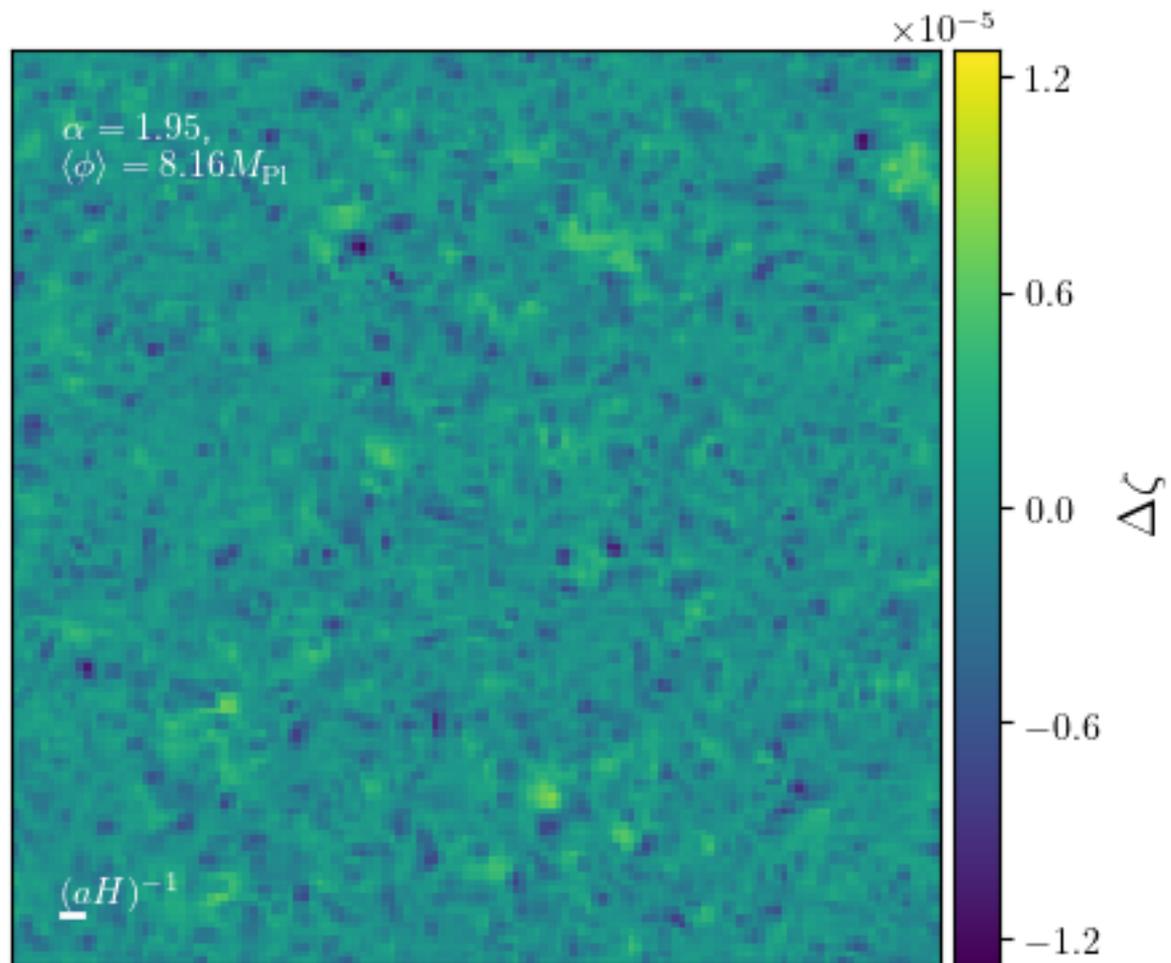
leads to ζ peaks - common

leads to wall memory in ζ

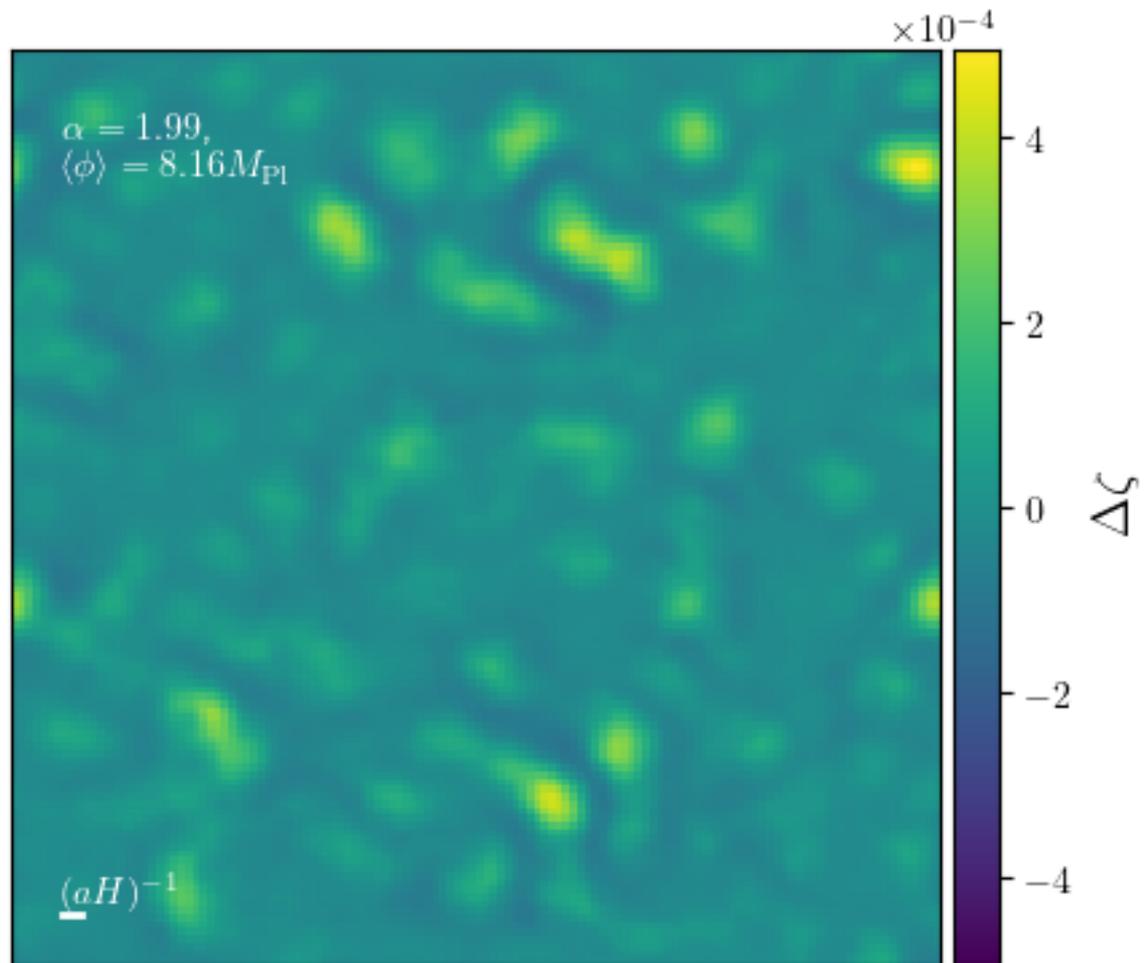
leads to ζ peaks - common



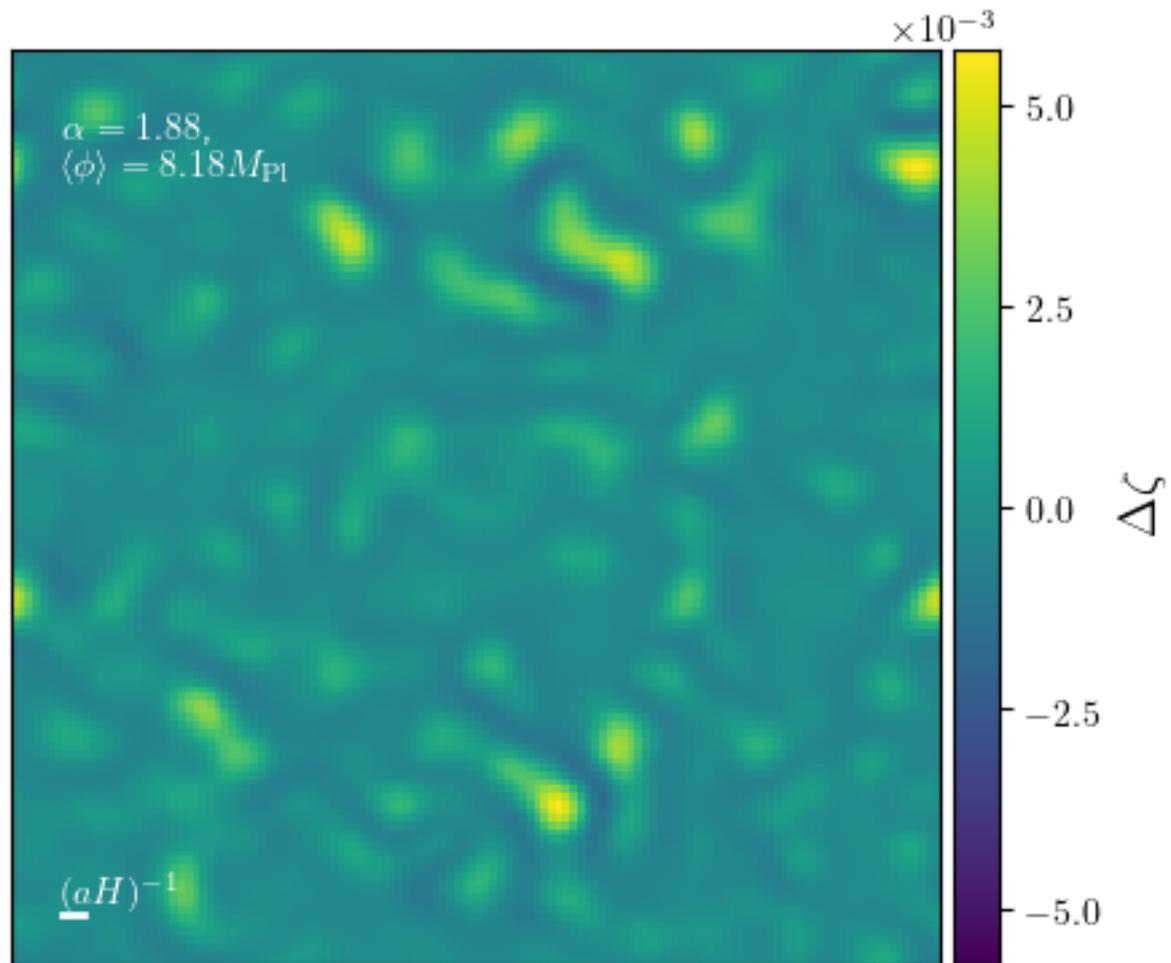
Dzeta_slice_a



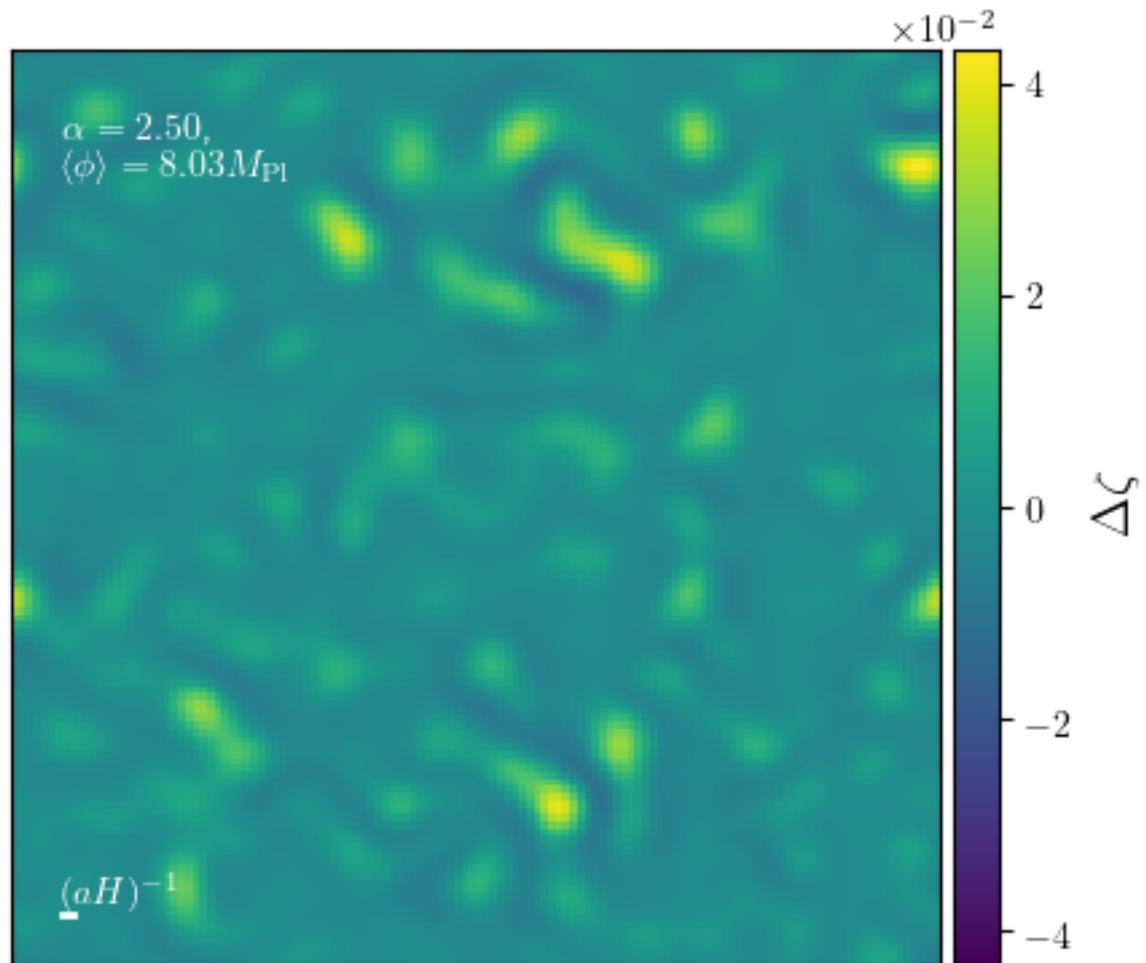
Dzeta_slice_b



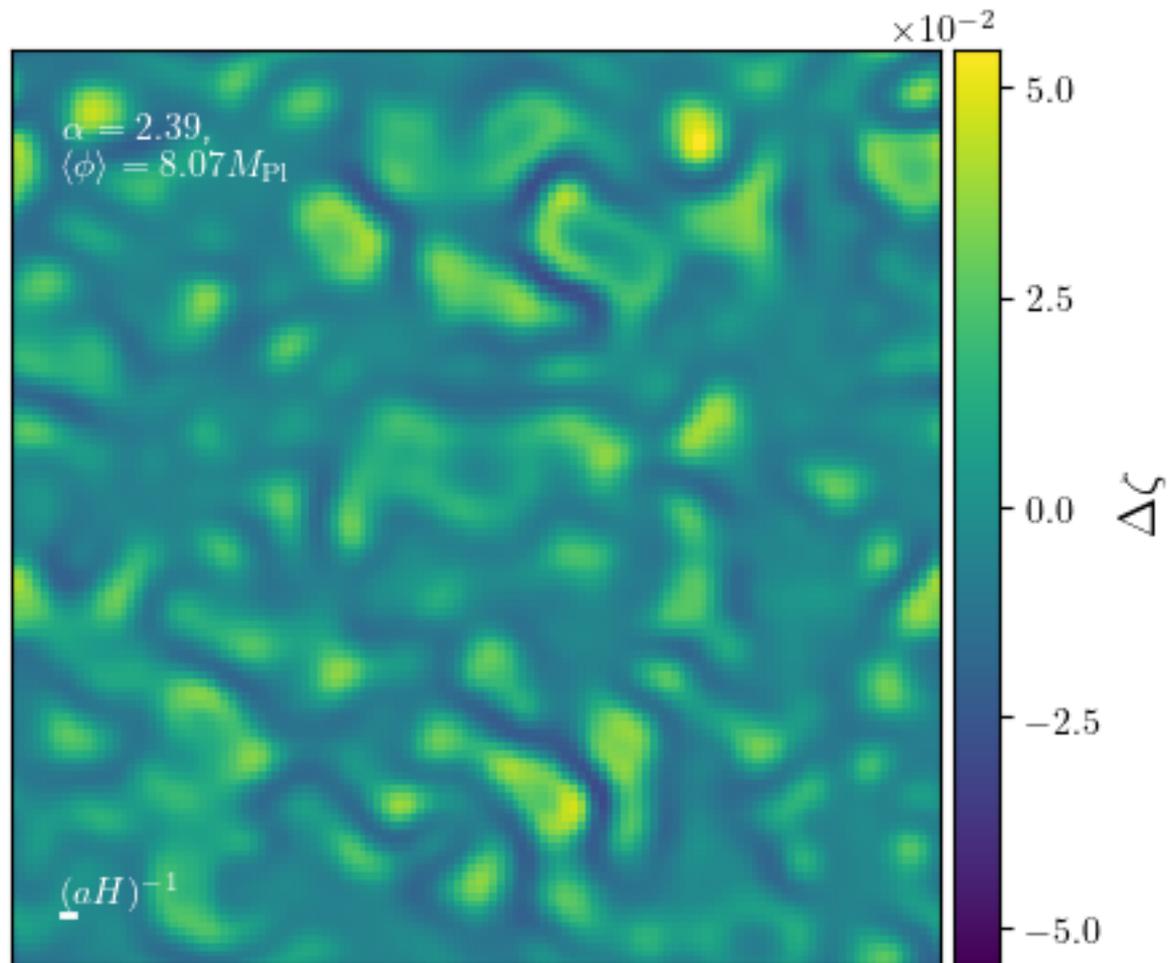
Dzeta_slice_c



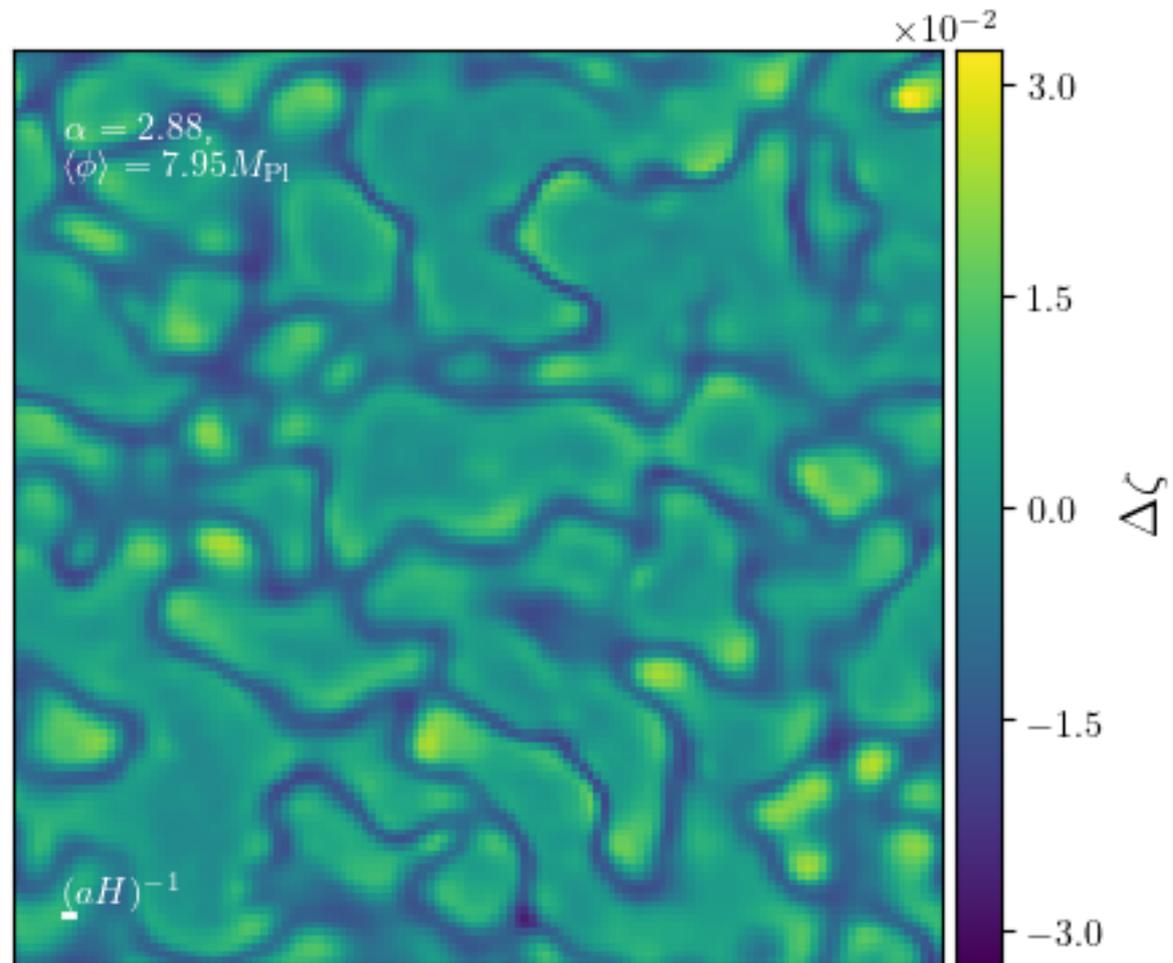
Dzeta_slice_d



Dzeta_slice_e



Dzeta_slice_f



Extending stochastic inflation to include coherent pulses to explain instability-driven multi-field simulation ensembles of concentrated primordial non-Gaussianity

Dick Bond @simons eU 22 04 26 Tom Morrison (GS), Jonathan Braden, Tomas Galvez



$3\zeta(x,t) = \int_{\text{field-path}} (dE+pdV)/(E+pV) \sim dS/\beta(E+pV)$ the entropic structure measure

$\zeta(x,t) \sim \int \Delta\zeta_{nG}\text{-Prominences}(x-x_c) dN_c(x_c R_c) + \text{Gaussian random } \zeta\text{-flucs}$

$\zeta(x,t) \sim \sum_p \Delta\zeta_{nG}(\chi(x, \alpha_e)) + \text{Gaussian random } \zeta\text{-flucs}, \chi(x, \alpha_e) = \text{Gaussian random}$

nonG from instabilities during inflation as well as **modulated heating nonG**

$m^2 > 0$ Spin spectrum for cosmic collider of Arkani-Hamad & Maldacena but small response in power spectrum & nonG

$m^2 < 0$ why not, m_{AB}^2 eigenspectrum, instability & Spontaneous symmetry breaking, landscape etc. more spectacular nonG

$$s_p(\zeta | \chi_e) = -\ln P(\zeta | \chi_e) = \frac{1}{2}(\zeta - \Delta\zeta_p)^T C_{\zeta\zeta}^{-1}(\zeta - \Delta\zeta_p) + \frac{1}{2} \text{Tr} \ln C_{\zeta\zeta} + \frac{1}{2} \ln(2\pi)$$

$$s_p(\chi_e) = \frac{1}{2} \chi_e^T C_{\chi\chi}^{-1} \chi_e + \frac{1}{2} \text{Tr} \ln C_{\chi\chi} + \frac{1}{2} \ln(2\pi)$$

$$P_p(\zeta) = \int \exp[-s_p(\zeta | \chi_e) - s_p(\chi_e)] d\chi_e \sim \langle \exp[\zeta C^{-1} \Delta\zeta_p] \rangle \sim \text{biased } \zeta$$

nG-pulses during inflation

$$\Delta\zeta_p(\chi_e) \approx \Delta\phi_p(\chi_e)/\sqrt{2M_P\epsilon_e} \text{ or } \Delta\zeta_p(\chi_m) \approx \Delta\phi_p(\chi_m)/\sqrt{2M_P\epsilon_m}$$

Concentrated nonG ζ -Prominences summary + future

CMB+LSS+ = one single vast entangled multi-messenger experiment probing the underlying BSMc
MassPeakPatches+Hydro+eUsims .. A toolkit for CMB and LSS experiments creating top down websky-ensembles to test BSMc theories (nonG, DM, DE,..) on the Universe: structure is coarse-grain halo+field & fine-grain response functions
make early universe ζ_{tot} maps and characterize most prominent nonG structures: rank-ordered localized modes identification
embed these ζ_{tot} modes in phenomenological ζ_{tot} models to see how to unearth in LSS data - In search of Buried Treasure:
Independent ζ_{tot} components are entangled in gravitational collapse, emissions
10e-fold LSS (Planck, ACT, SO, S4, .., DES, DESI, Euclid, LSST, .., CIB, LIM HI, CII, CO) + 50e-fold & $k \gg k_{LSS}$ (CMB distort, PBH,..)

- We simulated symmetry breaking/restoration during inflation.
- Comparing between two sims using potentials V and V_0 isolates the response $\Delta\zeta$ to the potential feature ΔV , suppressing Gaussian subhorizon and normal (V_0) stochastic 'noise' from the NG signal. **Δ deformed vacua: relative Casimir effect. Schwinger-like instability**
- We find the NonG of $\Delta\zeta$ shows strong local concentrations, **ζ -condensates considered as sourcy with scale related to k_p . Final state post-selection shows the concentrations arise from phase-space trajectories undergoing large excursions during the instability leading to nonlinear interactions between ϕ and χ .**
- Nonlinear mode development is \sim particle creation but not a simple particle creation
- Strongest instability leaves **domain wall memory in ζ - a rather different nonG pattern that the sourcy patterns of less instability strength. No domain wall energy problem, just ghostly memory**
- Conjecture: $\Delta\zeta$ could have hidden patterns reflecting other phase transition configurations, e.g., **fuzzy string memory, texture memory**, maybe oscillon memory. This can occur as well in preheating nonG.
- **Simple model of nonG as bias functions of hilly-spectrum GRFs allows for cosmic websky propagation to test cf. data**

$$\zeta(x,t) \sim \sum_p \Delta\zeta_{nG}(\chi_p(x, \alpha_e)) + \text{Gaussian random } \zeta\text{-flucs}, \chi_p(x, \alpha_e) = \text{Gaussian random}$$