

Dick Bond



CIAR Universe=System+Res, =Data+Theory **en-TANGO-ment**

CIFAR Cosmology & Gravity Program: >1985, 22 fellows & scholars (5@UofT), 19 associates + 6 Advisory Board members; CITA: 6+1 faculty, ~26 PDFs & Sr RAs + ~15 grad students; Bond: projects with 3-1 grad students, 5-1 SrRAs, 2 (++) PDFs



Cosmic history: what is U made of? $\Rightarrow \rho_{\text{dm}}/\rho_{\text{b}} = 5.1$

$\Rightarrow \rho_{\text{m}}/\rho_{\text{de}} = .30$ & $\Omega_m = 0.268 \pm 0.012$, $\Omega_\Lambda = 0.736 \pm 0.012$

How Structure in the Universe Arose?: fluctuation generation in curvature from an early **inflaton**: isocurvature, Gravity Wave, non-Gaussianity signatures

(coherence + quantum noise => incoherence via entropy/information generation)

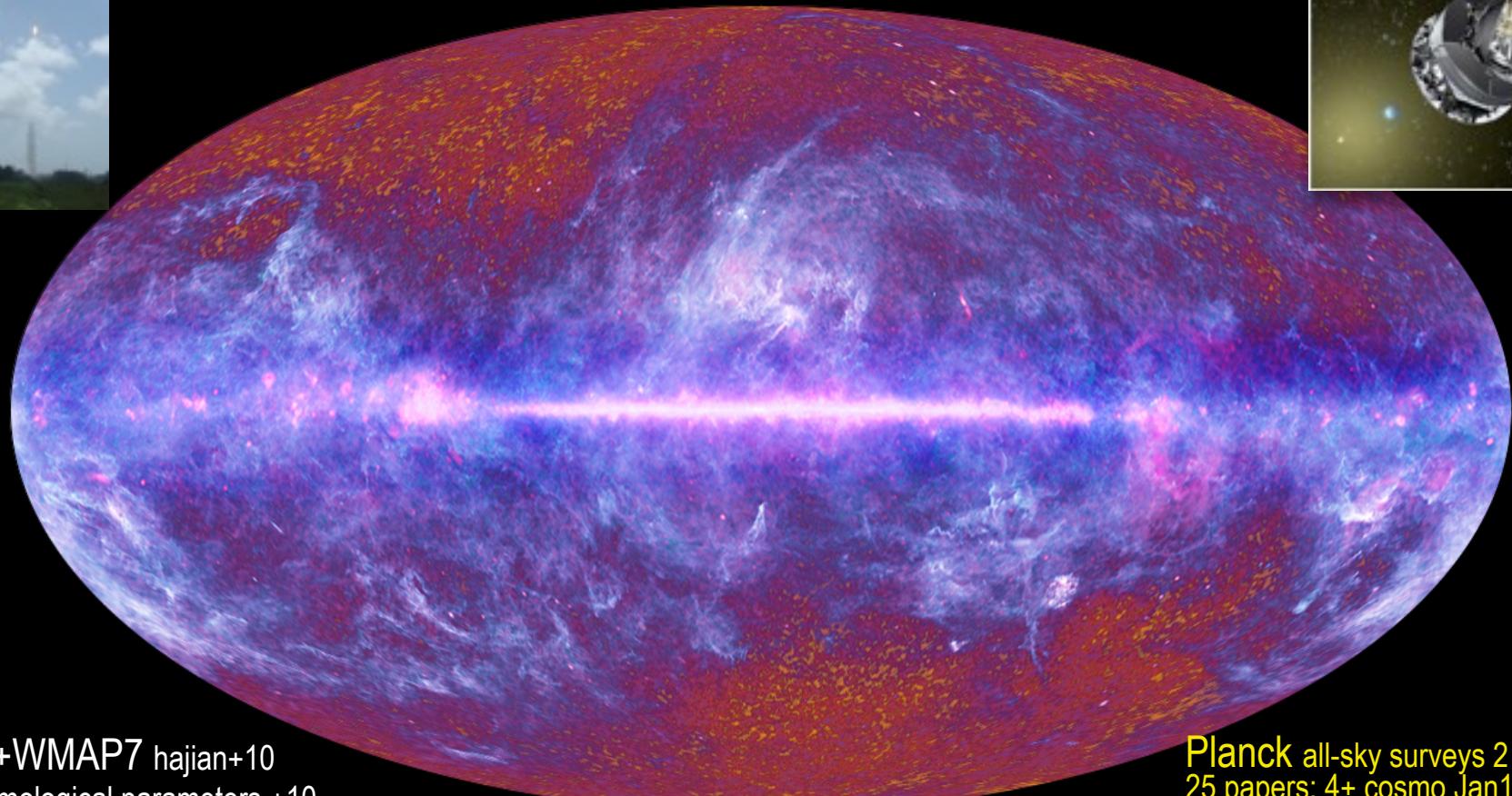
morphs into the nonlinear Cosmic Web: clusters, filaments, voids; galaxies (SZ)

What is the fate of the U: dark energy properties driving late inflation

CMBology & xCDM, x=dark energy+tilt: the cosmic standard model
status@Sept11: uses WMAP7+ACT (SPT), past: Boom, CBI, Acbar,.. (QuAD, ...)

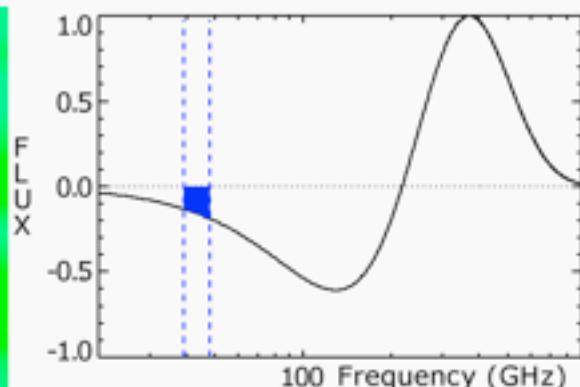
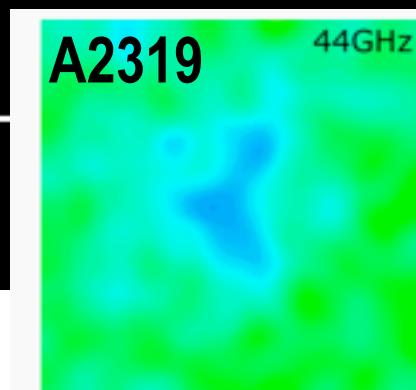
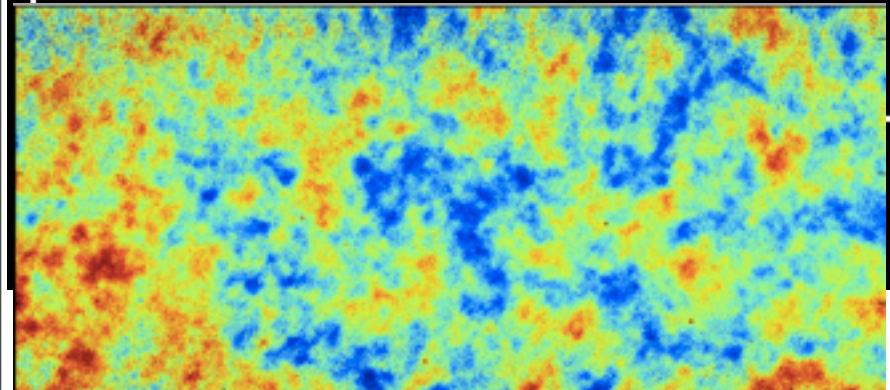
Planck cosmology Jan13 ultra-precision on cosmic parameters Jan11(25p), Feb12 ACTpol, ABS, Spider, Quiet-2, .. ALMA, CARMA, Mustang2 on GBT, CCAT

Cluster Information from Compton Heating of the CMB: ACT, Planck & Theory



ACT+WMAP7 hajian+10
=>cosmological parameters +10

Planck all-sky surveys 2 Jan11
25 papers; 4+ cosmo Jan13



Planck satellite, CMB all-sky, 9 frequencies, & polarization: ESA+NASA+CSA, B since 93, CSA since 02, launch may09, 4 full skies done, 5 in all, great data B+

Miville-Deschenes (SrRA), Nolta (SrRA), Netterfield (Prof), Marleau (DAA), Chluba (SrRA), Martin (Prof) + Contaldi, MacTavish, Crill, Dore (ex-CITAzens)+hundreds more - Europe+US P11 (ISM,SZ,CIB,sources,...), P12, cosmic: P13, P14

WMAP1,3,5,7 all sky 5 frequencies Nolta (SrRA). stopped taking data Aug10 **WMAP9** winter12

s **Early Universe non-Gaussianity:** B+ Braden (GS), Huang (GS=>PDF), Frolov ex-CITAzen

probing CMB non-Gaussianity: B+ Frolov, Nolta, Huang. Cold/Hot Spots; quadratic nonlinearities

CMB@ hi res: CBI2 finished data in 09 - new cluster results, B+ Sievers (SrRA=>Princeton)

Atacama Cosmology Telescope@very hires: B+ Hajian (PDF), Nolta, Sievers, Switzer (SrRA),

Battaglia (GS), Hincks (SJ) finished taking data Jan11 => **ACTpol Fall11** neutrino masses, inflation, cluster physics, dark energy properties, lens ACTpol+Planck +... very powerful!

s **Clusters & Cosmic Web Gasdynamical Simulations, & the Intracluster Radio Web:**

B+ Nick Battaglia (GS, PhD Aug11=>CMU PDF), Pfrommer (ex-SrRA), Sievers (SrRA=>Princeton) ACT,Planck, CBI (SPT,..,SKA)

ABS; Quiet2; Chile-based **Spider** balloon-borne **CMB expts targeting primordial gravitons via CMB pol**
B+ Nolta, Sievers B+ Sievers B+ Netterfield, Farhang (GS) +,Barth-team, ++ GW/scalar $r \approx 0.01V/(10^{16}\text{Gev})^4$

Probing Recombination epoch with the CMB: B+ Chluba (SrRA), Farhang

s **GW & Acceleration Trajectories:** B+ Huang, Vaudrevange, Contaldi; B+ Farhang, Dore, Netterfield

s **Post-Inflation Preheating:** B+ Braden (GS). **How Entropy Arose in U via a Shock-in-Time**

s **Late-time Inflaton Trajectories & Dark Energy cf. CMB,LSS,SN now/future:** B+ Huang

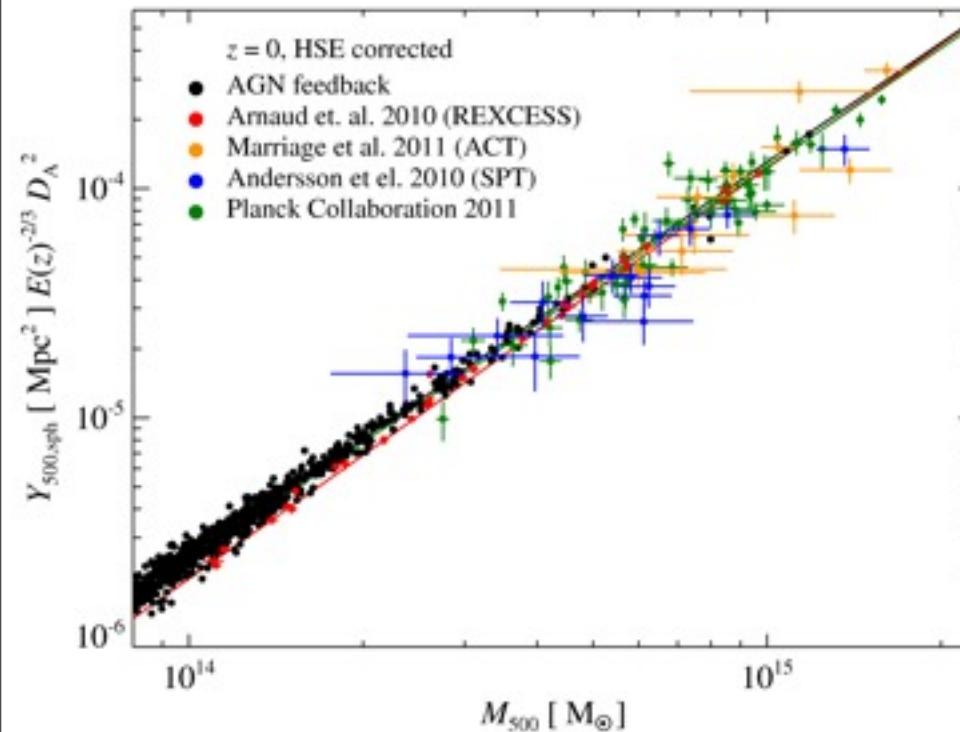
s **Magnetohydrodynamical Turbulence Simulations & Polarized Dust,** B+ Fahrang, Miville-Deschenes, +..., ..

CHIME, neutral hydrogen via redshifted 21 cm **Baryon Acoustic Oscillations & Dark Energy:** Pen (Prof)++ B+ Canadian Consortium: site DRAO in BC

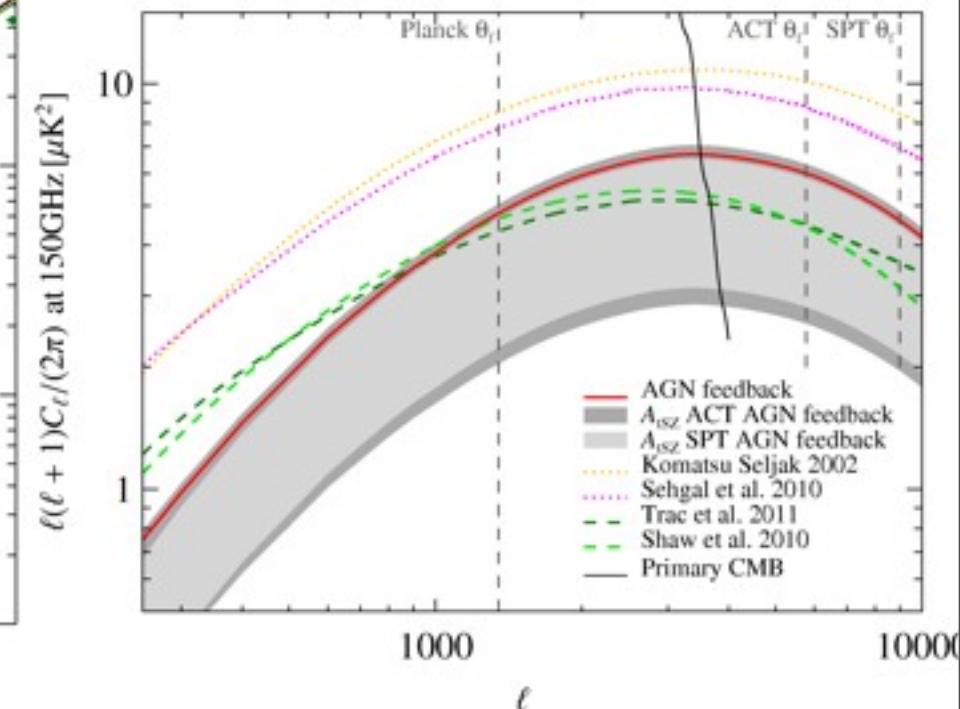
Cluster Coarse-Grained Feedback Sims cf. SZ data ACT, SPT, Planck

Cluster counts $n_{\text{cl}}(M(Y))dM + t\text{SZ}/k\text{SZ}$ Power spectrum

Battaglia, Bond, Pfrommer, Sievers 2011: I,II,III,IV; BBPS+Sijacki 2010

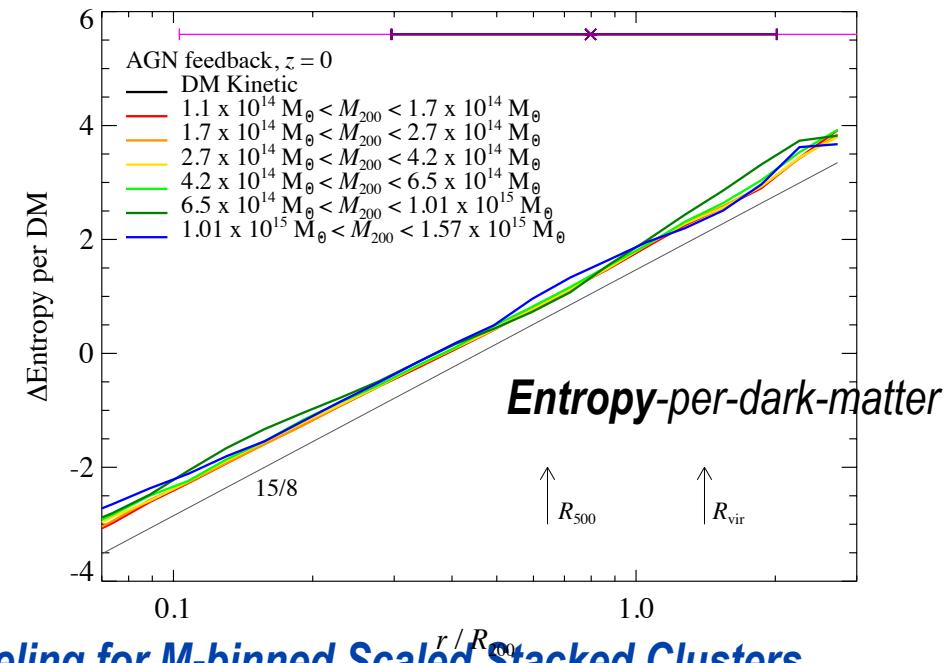
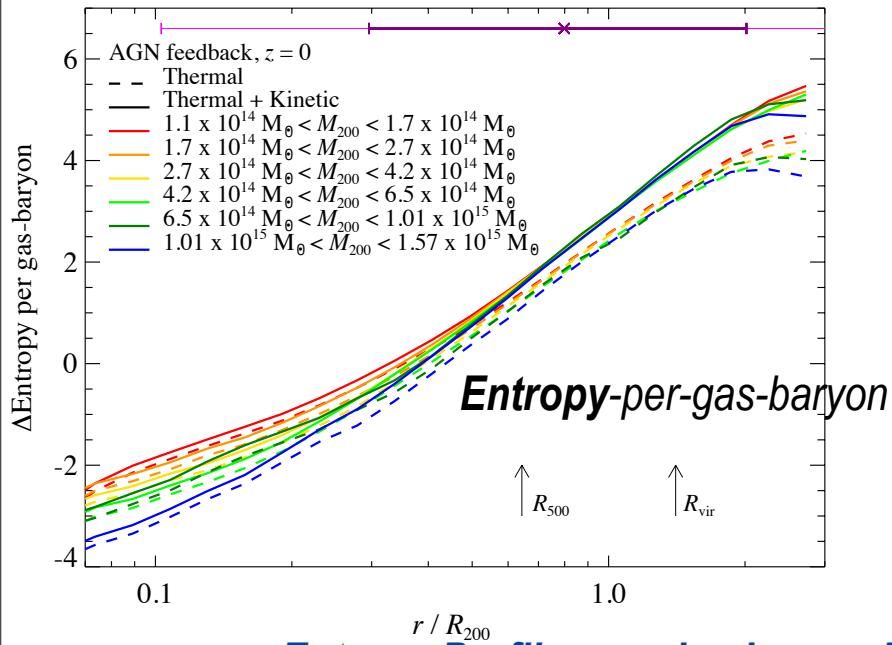


“turbulence” $p_{\text{kin}}/p_{\text{th}} \sim 20\%$ effect
asymmetry long/short $<20\%$
effect; cf. spherical $\sim 30\%$
 Δ input physics $\sim 30\%$ effect



both are sensitive to gastrophysics:
resolution, feedbacks(M, z), kinetic $\langle \delta V \delta V^\dagger \rangle$ cf. thermal pressure, $\langle \delta X \delta X^\dagger \rangle$ anisotropy, p&p-clumping, non-equilibrium cluster-outskirts

Entropy Profiles ($M/z=0$) for M-binned Scaled Stacked Clusters



Entropy Profiles vs physics modeling for M-binned Scaled Stacked Clusters

