

Novel LSS/CMB non-Gaussianities from Instabilities & Entropy Generation During and After Inflation



Generate fully-correlated nonGaussian Websky-Ensembles for CMB/LSS probes

Dick Bond @ ICTS 19 01 22

Bond+Braden+Frolov+Huang+Morrison+Stein

varieties of primordial nonG and how to search for them

Simons Modern Inflation group - b2fhms & Eva Silverstein+ & Dan Green+

Origin of the observed entropy in the Universe and SMpp/BSMpp particles

- coarse-grained coherent condensate breaks up into fine-grained fluctuations
- particle creation = (instability => stretch and break via mode-mode coupling)
- episodic stretching (adiabatic) and breaking (non-adiabatic => nonG) during inflation & after

nonlinear multi-field classical coupled system. evolve using lattice simulations. via *pseudo-spectral code & symplectic defrost++ code* => very high accuracy to unveil small nonlinear effects leading to nonG

during inflation (**beyond stochastic inflation. nonlinear k-space burst structure**)

$$\langle \Delta \mathcal{P}_{\phi^A \phi^A}(k) | \Delta V, \Delta m_{eff}^2 \rangle, \quad \langle \Delta \mathcal{P}_{\zeta \zeta}(k) | \Delta V(\phi, \chi) \text{ controls} \rangle \quad \langle \Delta \langle \prod \zeta^N \rangle_{cc} | \Delta V \text{ controls} \rangle \quad \text{SBB89, SB90,91 B95..}$$

& after inflation ends (**modulated heating. marginalize ~50 e-folds of sub-LSS**)

dynamical system Kolmogorov-Sinai entropy cf. true Shannon entropy
nonG ~ "particle" production ~ Shannon entropy generation

$$\Delta S_{\text{flucs,k}} = \text{Trace} \ln[\mathbf{C}_{\phi^A \phi^B} \mathbf{C}_{\Pi_A \Pi_B} - \mathbf{C}_{\phi^A \Pi_B} \mathbf{C}_{\Pi_A \phi^B}] / 2 \approx \ln(\mathbf{n}_{\text{flucs,k}} + 1/2) \text{ cf. old way } \sim \ln[\rho_{Ak}(t) / \hbar \omega_{Ak}(t)]$$

adiabatic flucs encoded in the collective Phonons, fluctuations + condensate = $\zeta_{\mathbf{k}}$

$$\langle \delta_{\mathbf{J}}(\mathbf{x}, \mathbf{t}) | \zeta \rangle = \chi_{\mathbf{J}\zeta}(\mathbf{x}\mathbf{t} | \mathbf{x}_i) * \zeta(\mathbf{x}_i), \quad \chi_{\mathbf{J}\zeta} = \text{linear transfer fn}$$

varieties of primordial nonG and how to search for them

perturbative, nonG part correlated with dominant Gaussian part

see Planck 2015/2018 nonG for exhaustive study and current constraints - 2018 including T+Epol
local f_{nl}^* - current limit cf. f_{nl} std target < 1 . & equilateral orthogonal

relax if uncorrelated quadratic nonG suppressed by at least $\sim \mathcal{E}^2$ Planck2015/2018

nonG 3-point-correlation-pattern measure

f_{nl} : 2.7 ± 5.8 local for Newton potential

$\Rightarrow f_{NL}^* = 0.44 \pm 3.5$ for phonons/3-curvature

$-f_{nl}$: 42.3 ± 75.2 equilateral

-25.3 ± 39.2 orthogonal

other caveats - beyond Planck2015/2018 nonG: some in Planck 2015/2018 Isotropy & Statistics

outside horizon, wide open stochastic inflation huge nonG from feedback.

k-localized nonG. wide open. role of instabilities during inflation to make k-localized zeta-bursts. could even make PBHs

silverstein etal approach. higher N-points.

here numerical pseudo-spectral codes to correct stochastic inflation.

B2FHMS can ensemble-measure everything, N-pt, coherences!

nonG from heating 1 cm comoving scale \Rightarrow to be in observable bands need modulation, but that is natural if there are light fields (heavy fields damp power)

long-lived field-condensates: strings, oscillons, curvaton structures, ...

later phase transition structures - need first order (discontinuity in entropy - latent heat) cf. second order (discontinuity in second derivative aka fluctuations) & smoother higher order, eg adiabatic evolution of particle content

out-of-equilibrium decays



ζ TOPOGRAPHY & CARTOGRAPHY

of our Hubble-patch bit of the early universe: RECONSTRUCT

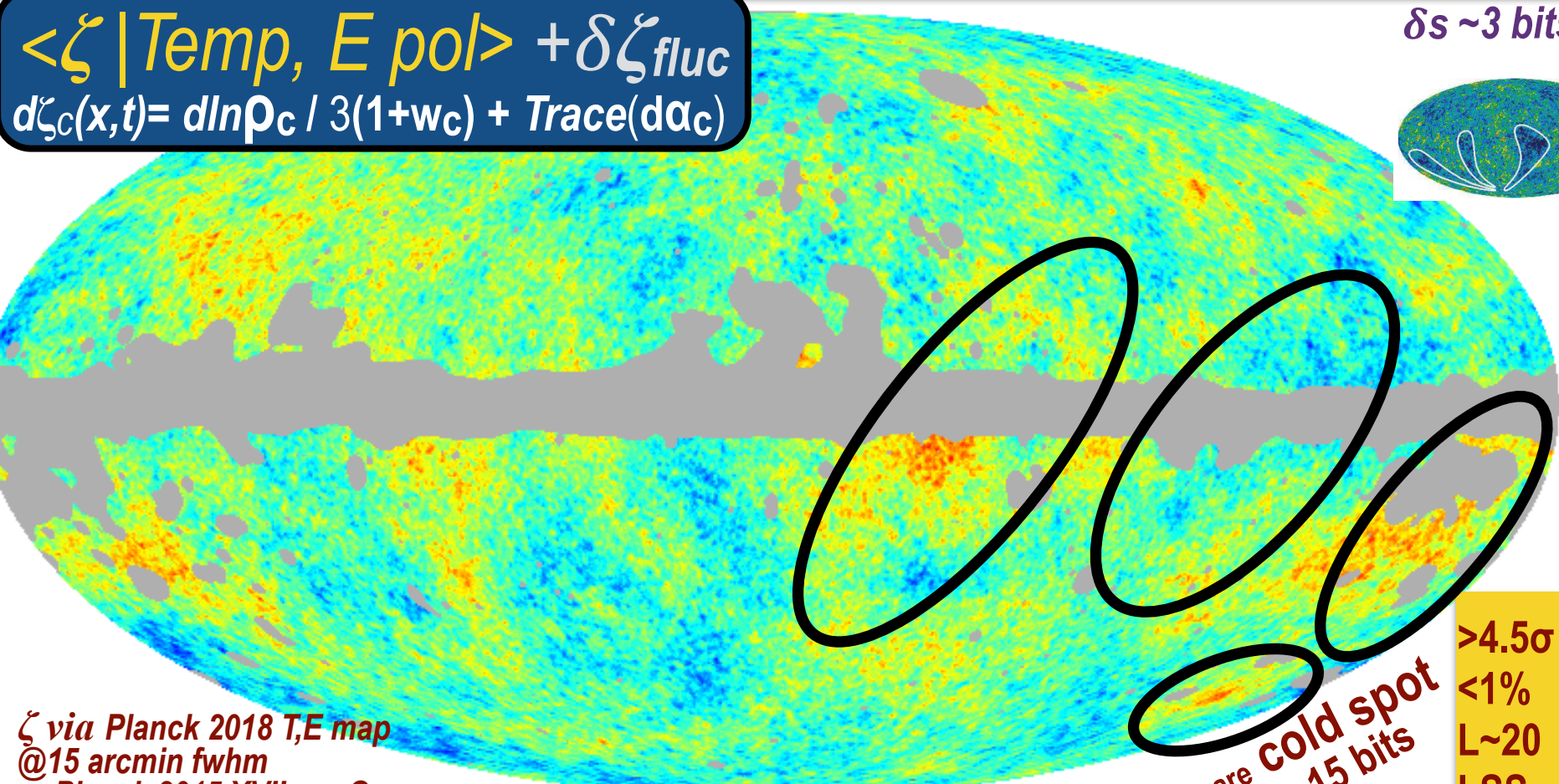
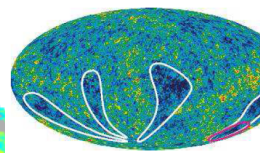
Beyond the Standard Model of cosmology? $SM_c = \text{tilted } \Lambda\text{CDM} + r \text{ via } (\zeta, h_{+x})$

BSM_c = SM_c + primordial anomalies (nonG) in the true ζ -WebSky anomalies @ low L => sample variance limited ~2σ's *CMB TT power L~ 20-30 dip => ζ-Spectrum k-dip*

$$\langle \zeta | \text{Temp}, E \text{ pol} \rangle + \delta \zeta_{\text{fluc}}$$

$$d\zeta_c(x,t) = d \ln \rho_c / 3(1+w_c) + \text{Trace}(d\alpha_c)$$

$\delta s \sim 3 \text{ bits}$



ζ via Planck 2018 T,E map @15 arcmin fwhm \approx Planck 2015 XVII nonG

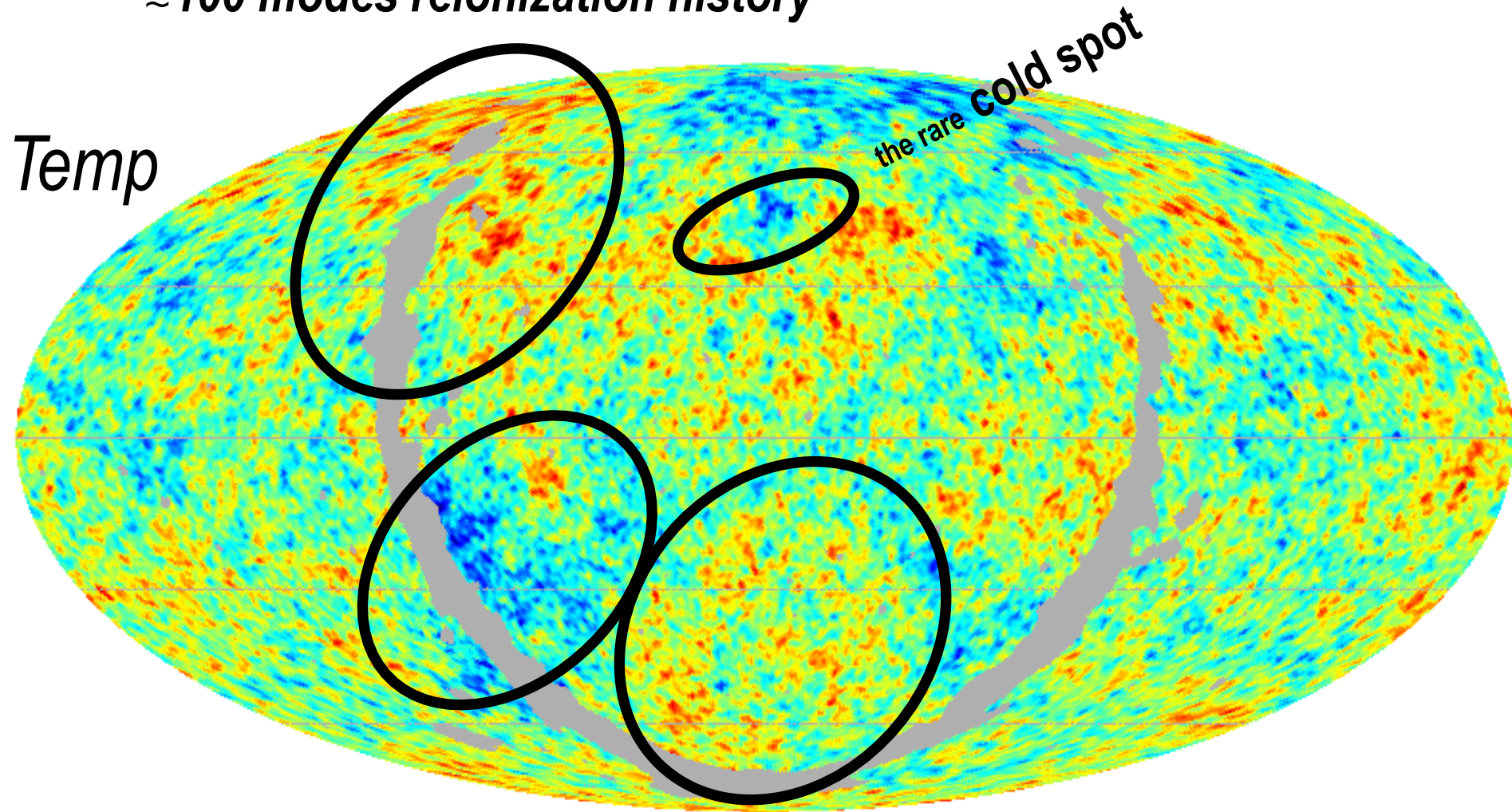
GUTA = Grand Unified Theory of Anomalies?

the rare cold spot $\delta s \sim 15 \text{ bits}$

>4.5σ
<1%
L~20
LSS
void?

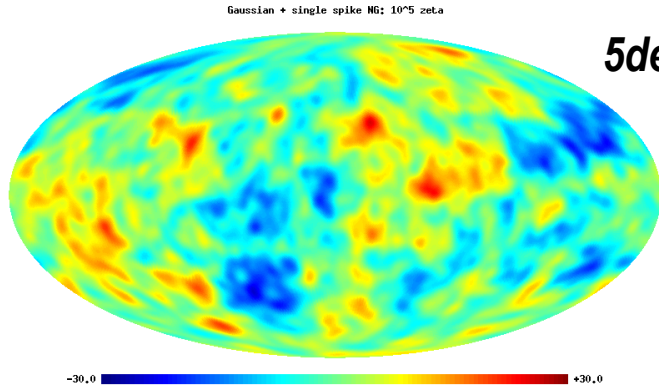


CMB $\sim 10,000,000$ T/E modes of Λ CDM
 ≈ 500 modes of anomaly
 ≈ 100 modes reionization history

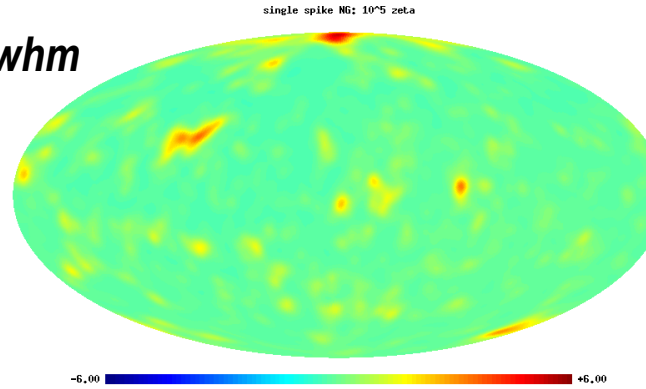


2D intermittency WMAP cold spot

CMB+LSS mocks to test: standard Gaussian inflaton ζ_{inf} + subdominant uncorrelated ζ_{isoc}
e.g., from modulated preheating



5deg fwhm



**scan sims to get
chance intermittent
alignment to get a
WMAP “cold spot”**

**intermittent nG from
early U preheating
lattice sims**

tunable peak model

**also cf. quadratic nG:
correlated fNL**

uncorrelated large fNL_{eff}

uncorrelated nonG ‘wide open’ cf. usual correlated highly constrained nonG

the true quadratic ζ -Websky of the ζ -scape

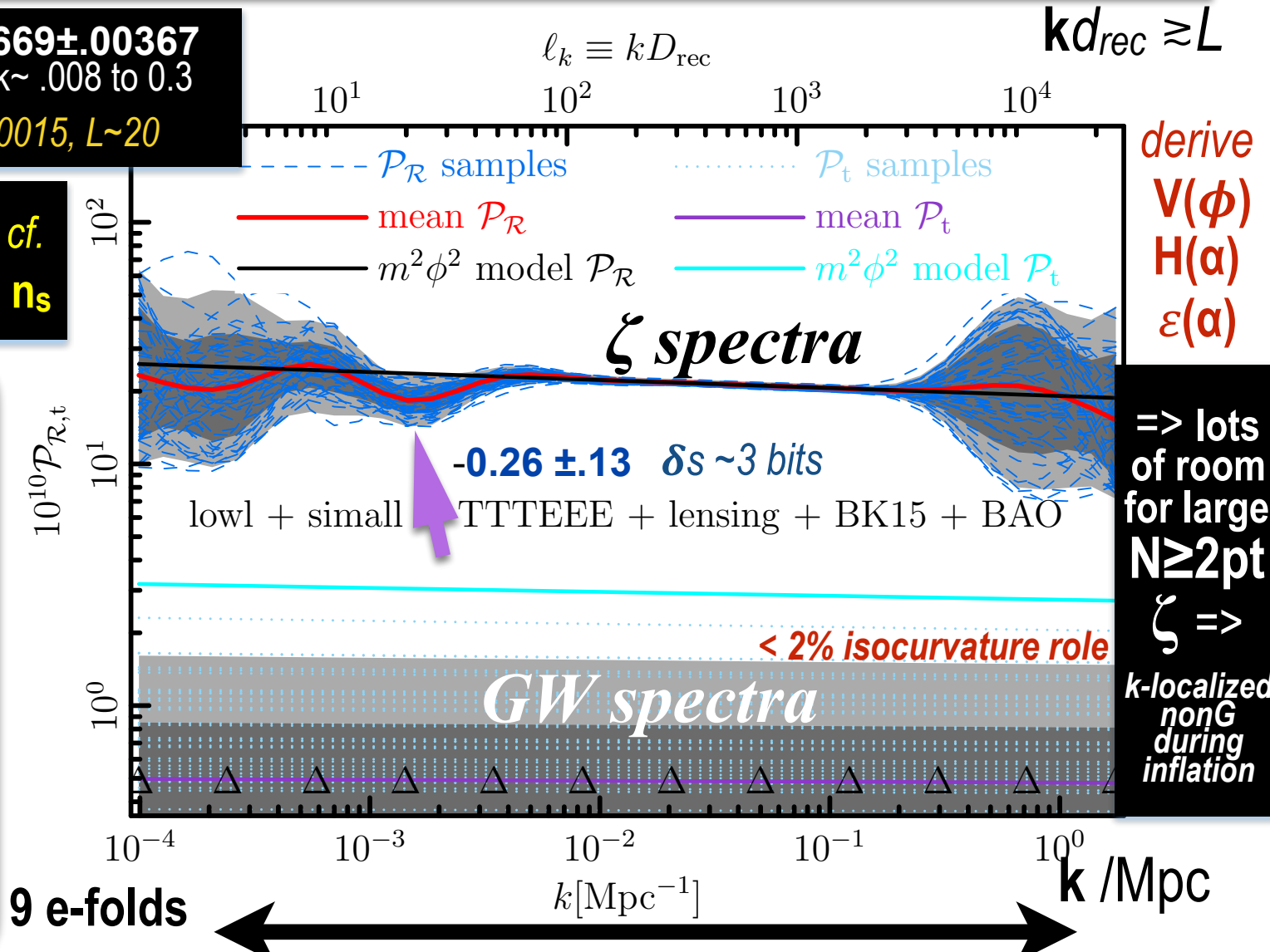
Planck 2018 X inflation: TTTEE lowL Epol + CMB lens + BK15 BB + BAO

CMB TT power L~ 20-30 dip => ζ -Spectrum k-dip; includes CMB lensing, parameter marginalization

uniform $n_s = 0.9669 \pm 0.00367$
 superb 12-knot fit $k \sim .008$ to 0.3
 $\delta s \sim 3$ bits @ $k \sim .0015$, $L \sim 20$

$r < 0.069$ 95%CL cf.
 $r < 0.061$ uniform n_s

stochastic inflation
 $|\phi_{cg}; \phi_{fg}\rangle$
 coherent state picture of coarse-grain condensate + fine-grain bugoliubov fluctuations QM-correct

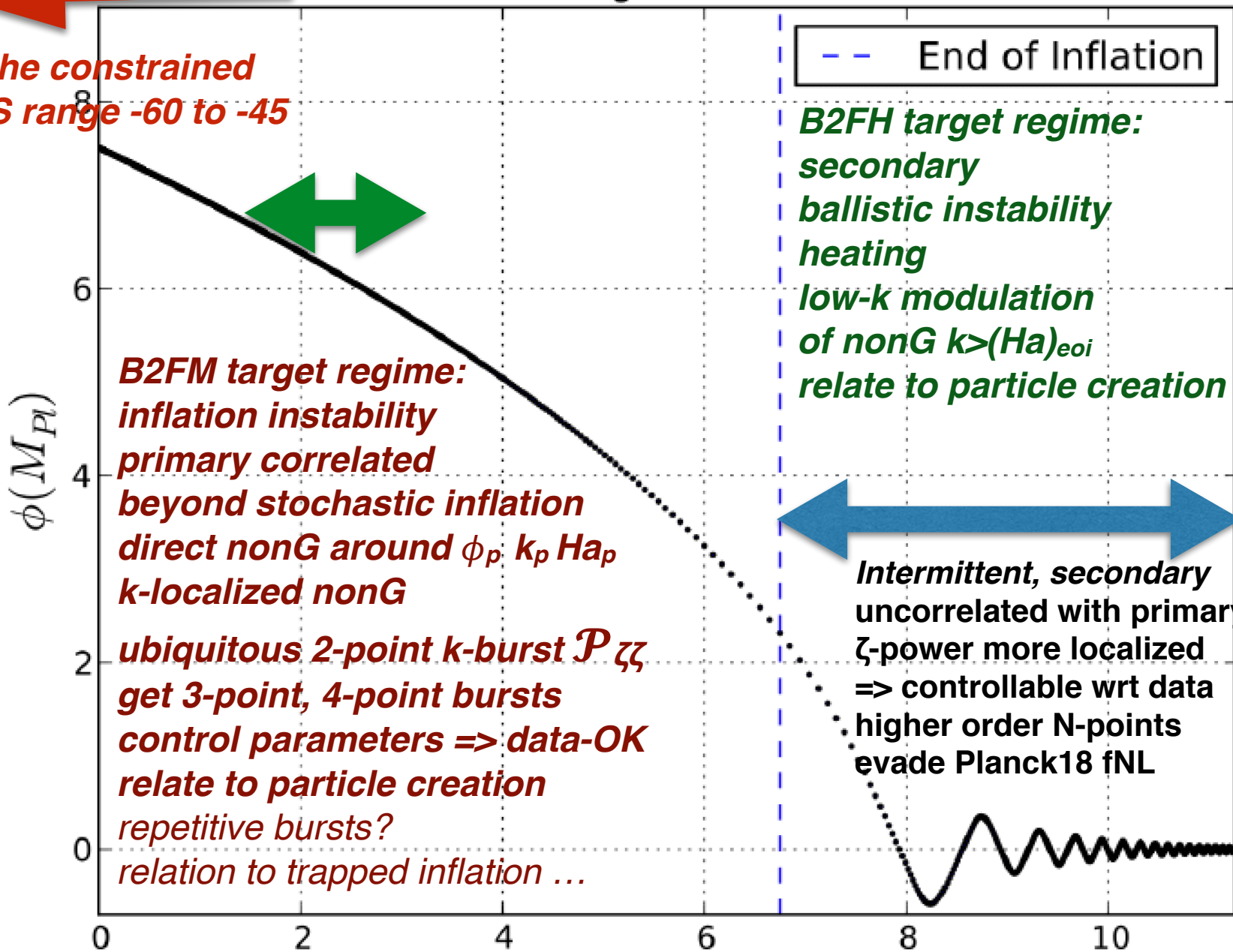


derive $V(\phi)$
 $H(\alpha)$
 $\varepsilon(\alpha)$

=> lots of room for large $N \geq 2$ pt $\zeta \Rightarrow$



to the constrained LSS range -60 to -45



--- End of Inflation

*B2FH target regime:
secondary
ballistic instability
heating
low-k modulation
of nonG $k > (Ha)_{eoi}$
relate to particle creation*

*B2FM target regime:
inflation instability
primary correlated
beyond stochastic inflation
direct nonG around ϕ_p, k_p, Ha_p
k-localized nonG
ubiquitous 2-point k-burst $\mathcal{P}_{\zeta\zeta}$
get 3-point, 4-point bursts
control parameters => data-OK
relate to particle creation
repetitive bursts?
relation to trapped inflation ...*

*Intermittent, secondary
uncorrelated with primary
 ζ -power more localized
=> controllable wrt data
higher order N-points
evade Planck18 fNL*

apply to PBHs etc!!

$\ln(a)$

Webskys: Mocking with *PeakPatches*+*Hydro*+



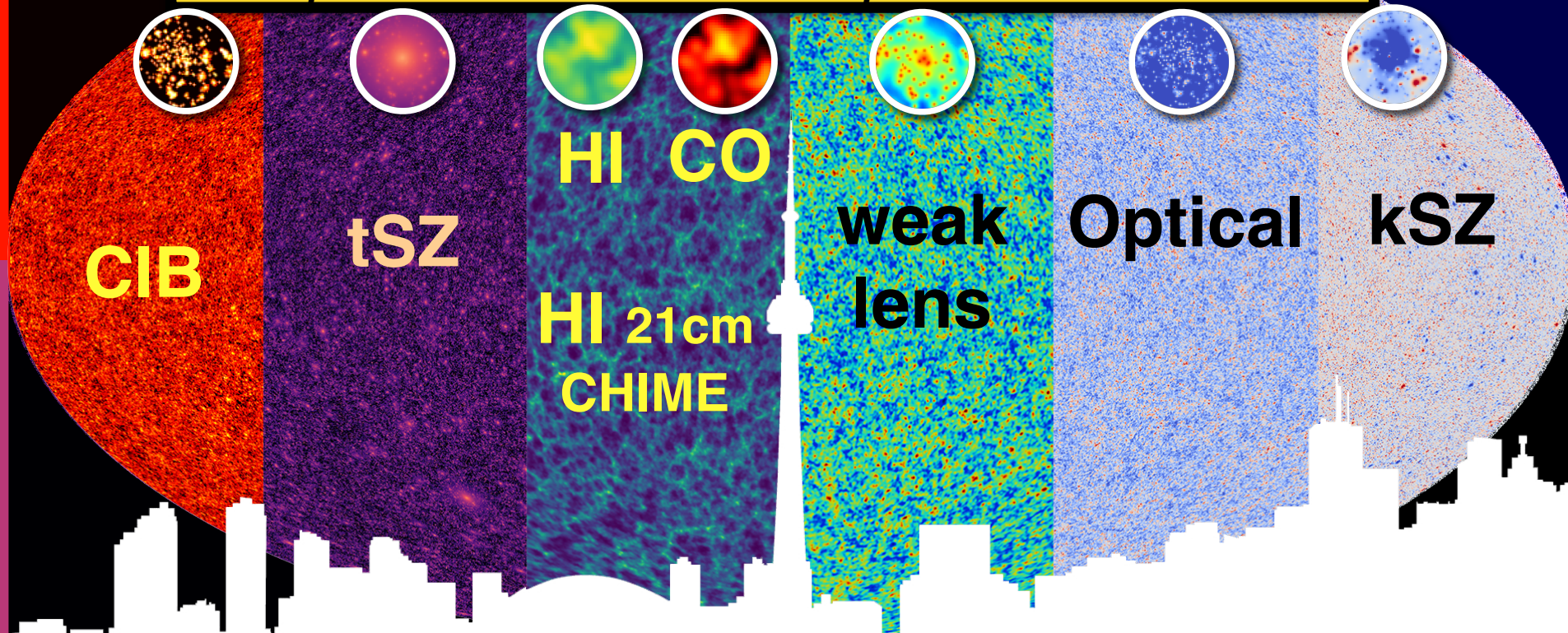
THEN BBKS, BCEK, B+Myers91,93,96, BKP96 web, BW96 importance

NOW: CITA mini-industry *Alvarez, Bond, George Stein* 2018, .. Validation SAB18 + Euclid 2018 validation a,b,c

Berger, Battaglia, Codis, van Engelen, Motloch, Huang, Frolov,

now 19.1 *Lague, Lokken, Murray, Keating, Lahklani, Breysse, bruno, connor, ronan, furen, remi, jason lee* ++

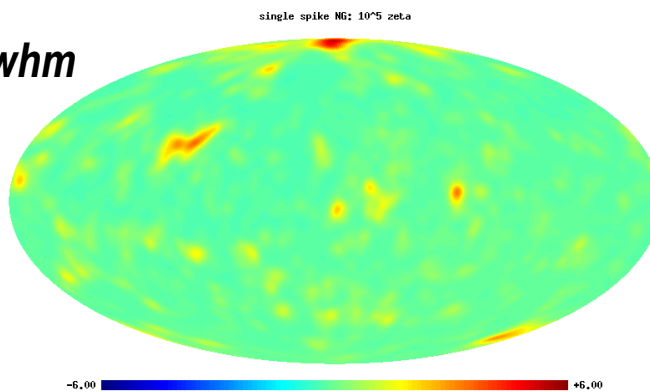
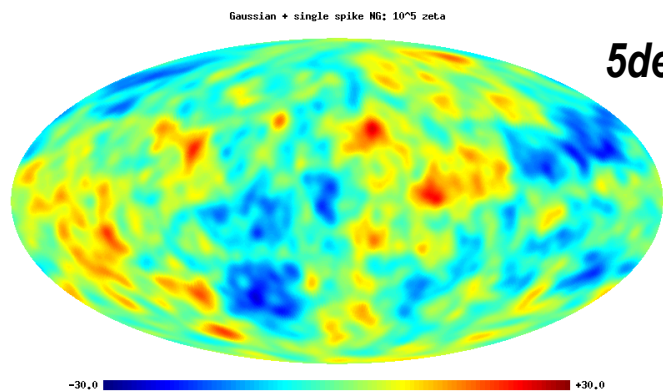
need **End to End** mocks: **BSMc, nonG, DE/modG, Mnu, ...**
need all signals to be correlated, 1, 2, 3, .. Npt
need speed to build ensembles & explore BSMc



Planck, AdvACT, SO, CMB-S4, CCATp, EUCLID, LSST, DES, CHIME, HIRAX, COMAP, ...SKA

2D intermittency WMAP cold spot

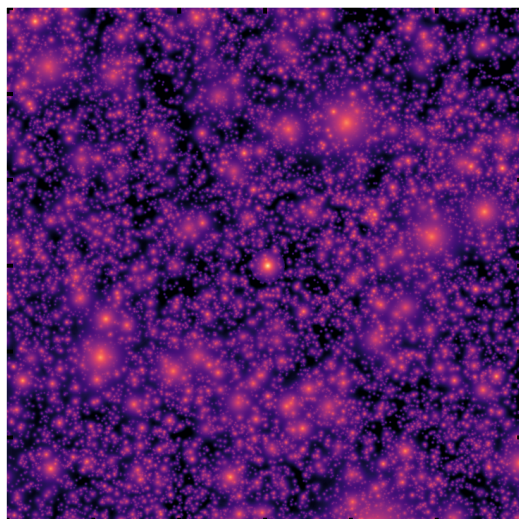
CMB+LSS mocks to test: standard Gaussian inflaton ζ_{inf} + subdominant uncorrelated ζ_{isoc}
e.g., from modulated preheating



scan sims to get
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also cf. quadratic nG:
correlated fNL
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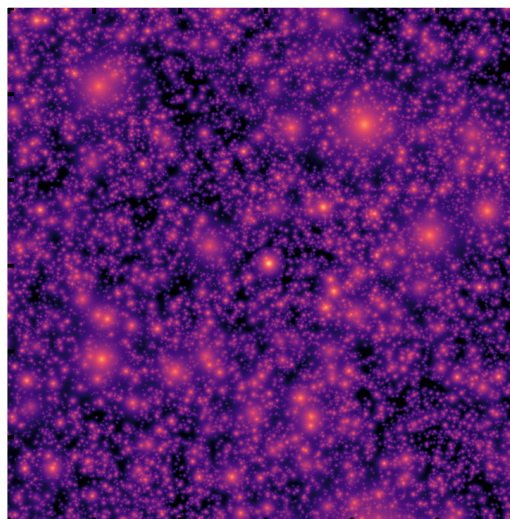
3D intermittency *uncorrelated nonG 'wide open' cf. usual correlated highly constrained nonG*

LSS tSZ: Gaussian std



B2FH, b+braden+frolov+huang

LSS tSZ: Gaussian std +
subdominant uncorrelated ζ



ABSB+FH, alvarez+b+stein+frolov+huang

COMap sims using Li+ Mhalo \Rightarrow L_{CO} cf. CIB *a la* Planck13,15

Danger: correlated stochasticity of bursty star formation etc.



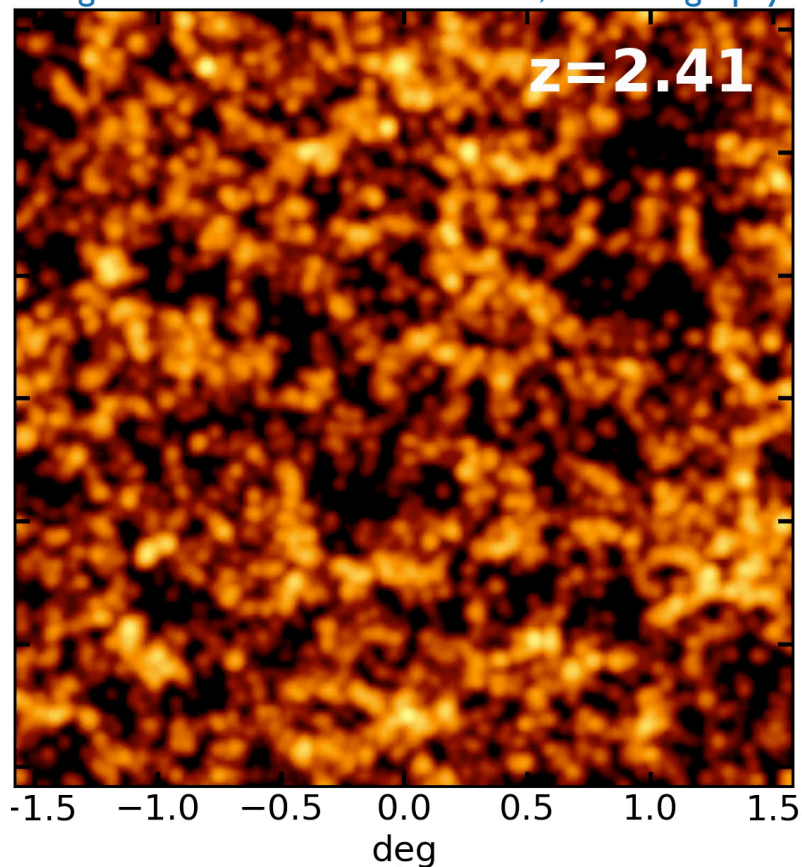
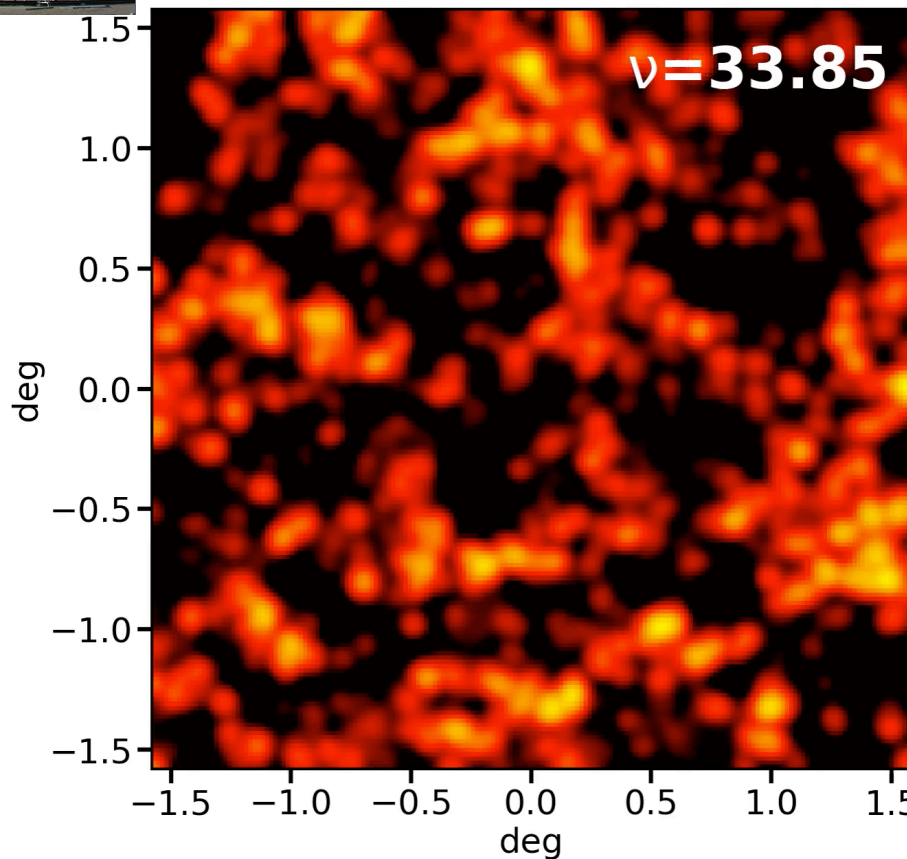
$z=2.4-3.4$ eventually $z=6-8$ cf. CHIME HI $z=.8-2.5$

CO

using Li et al. 2016 Model

CIB

using 217 GHz Planck 2015 Model, no tomography



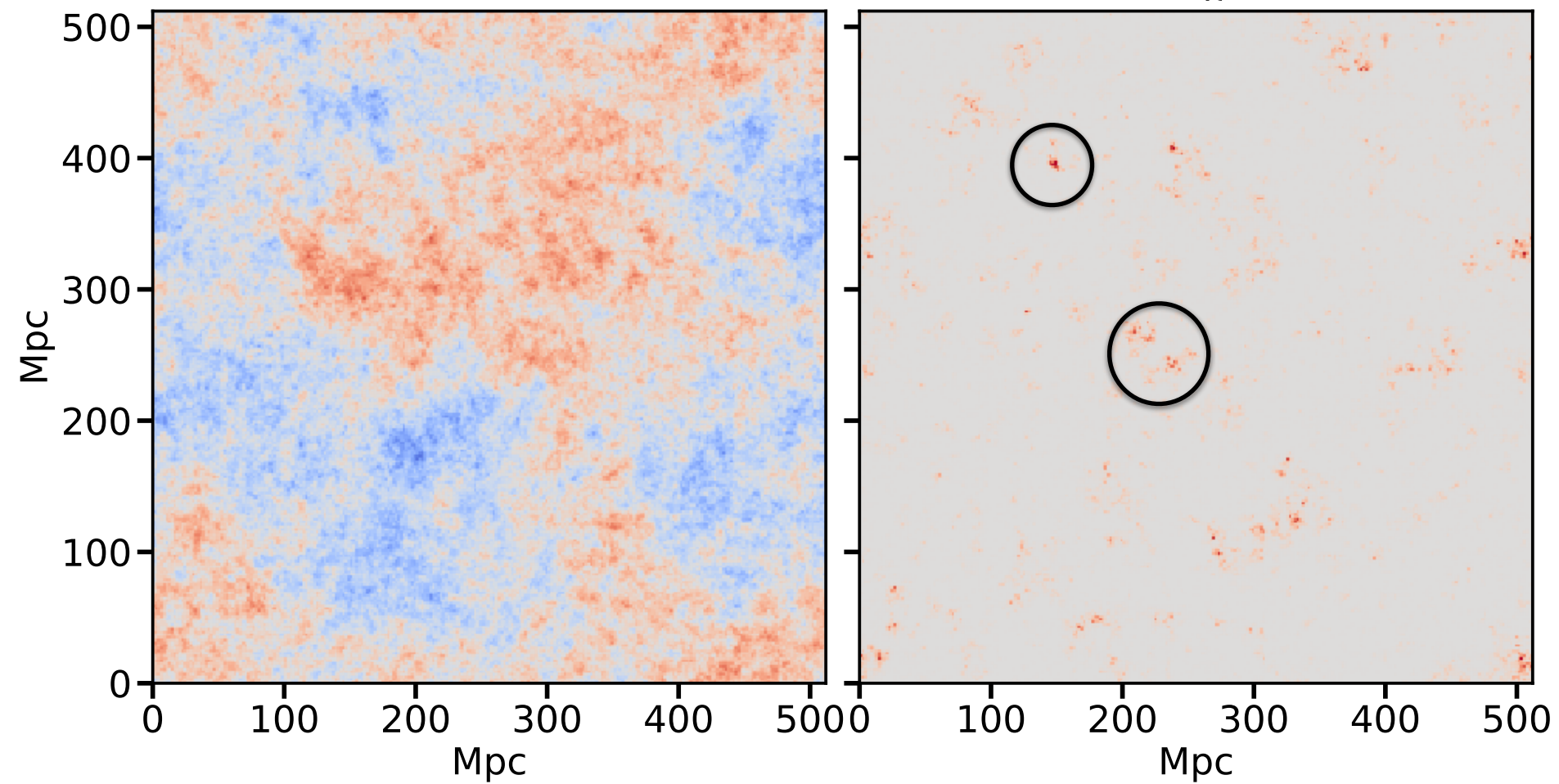
underway: **Lensing of CIB COMap HImaps kSZ tSZ**
nonG sources seen through a nonG lens

Primordial Non-Gaussianity in observable Webskys constructed with the mass-Peak Patch method
+ gas-halo response functions/susceptibilities

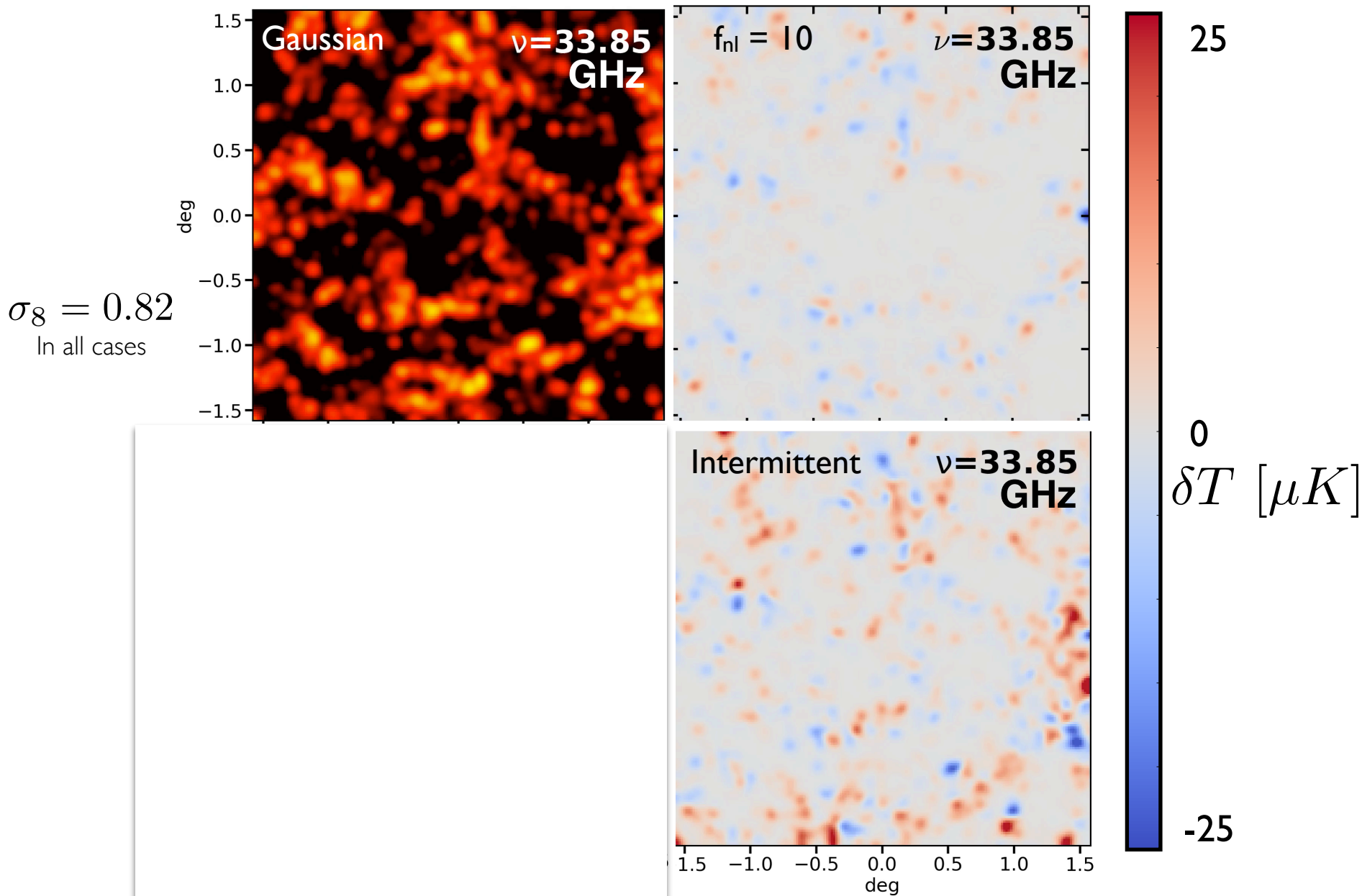
Intermittent Non-Gaussian case
uncorrelated ζ [GRF]

ζ_G

$\zeta_{F(\chi)}$



Primordial Non-Gaussianity in CO example: the LCDM signal and 2 nonG difference maps - a movie



large scale => CHIME much larger volume is better

after inflation - instabilities => entropy => nonG

$$dS/dt(t, g) =>$$

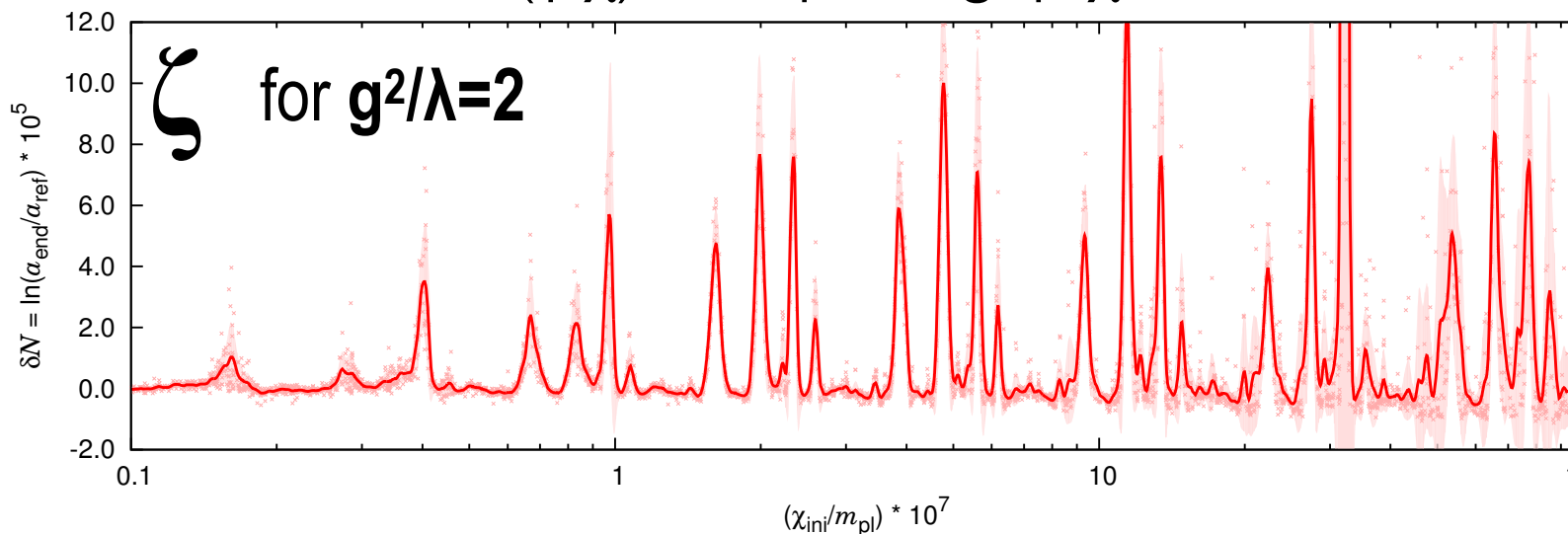
the Shock-in-time: entropy production rate

$$\zeta_{\text{shock}}(\chi_{c, \text{eoi}}(x) | g^2/\lambda) => \text{Chaotic Billiards: NonG from Parametric Resonance in Preheating}$$

B+Frolov, Huang, Kofman 09

B+Braden, Frolov, Huang 19

$$V(\phi, \chi) = 1/4 \lambda \phi^4 + 1/2 g^2 \phi^2 \chi^2$$

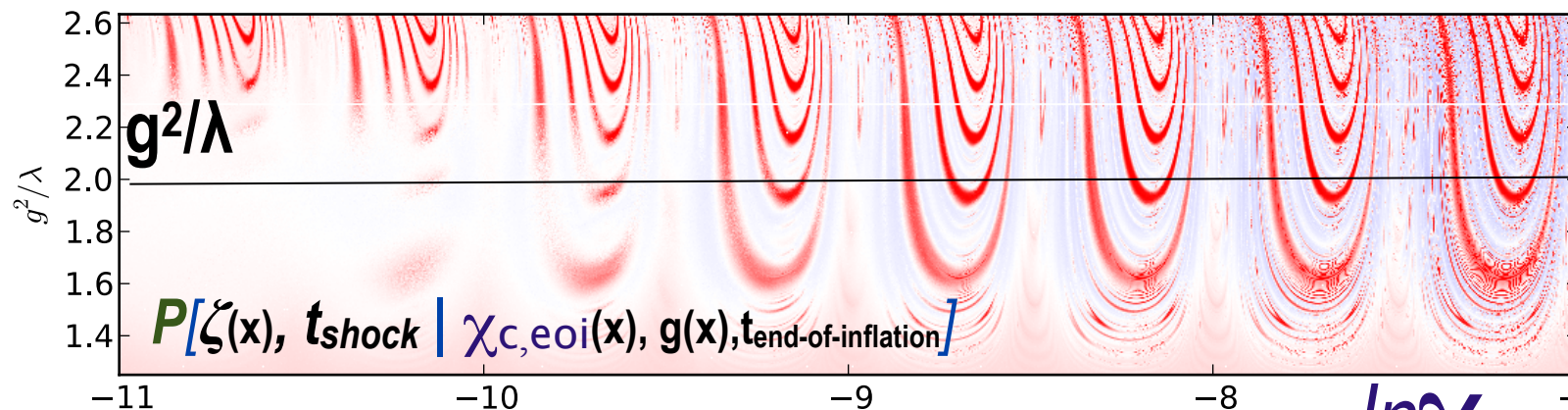


smooth over
~50 e-folds of
HF structure
 $\chi_c \Rightarrow \zeta_c$
 ζ -bias cf.
late-time
density-bias
 $\chi_{>h}$ control
parameter

computational
tour de force

huge number of
 64^3 sims to
show the
wondrous
complexity of

$\zeta(\chi_c, g^2/\lambda)$



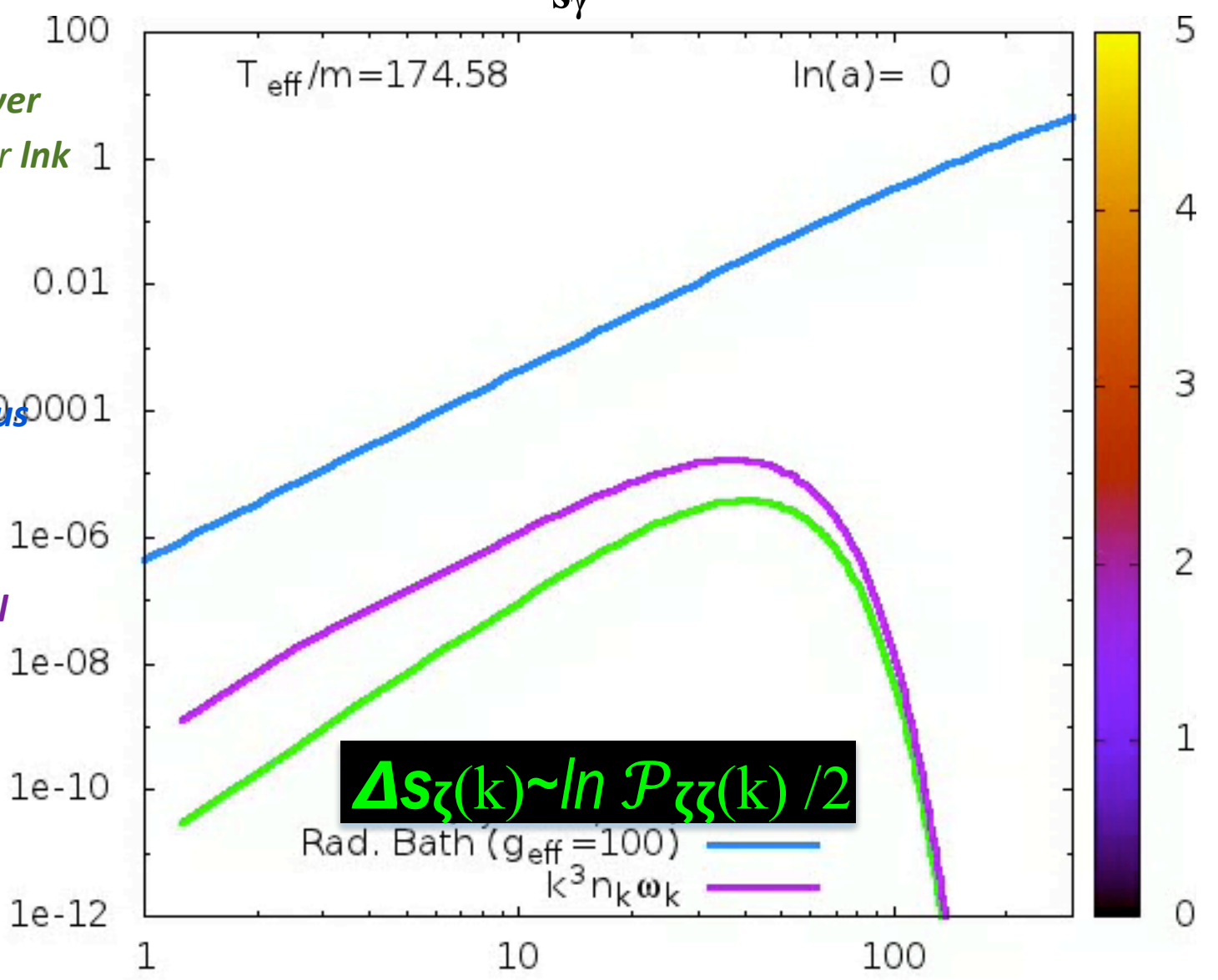
gigafigure of lattice simulations $\ln(\chi_0/\phi_0)/\mu_0 T$

$\ln \chi_{c, \text{eoi}}$

coherent inflaton => incoherent mode cascade of fields thru a shock-in-time to thermal equilibrium

$S_{U_i} \sim 0$; $S_{U_{tot,m+r}}/n_b \sim 1.66 \times 10^{10}$ bits/b; $s_\gamma/n_\gamma = 5.2$ bits/ $\Upsilon = 2130/411$; $s_v = 21/22$

In $p/\langle \rho \rangle$ power spectrum per $\ln k$ aka entropy aka phonon aka ζ cf. instantaneous full thermal spectrum cf. conventional energy spectrum using a pseudo particle occupation number

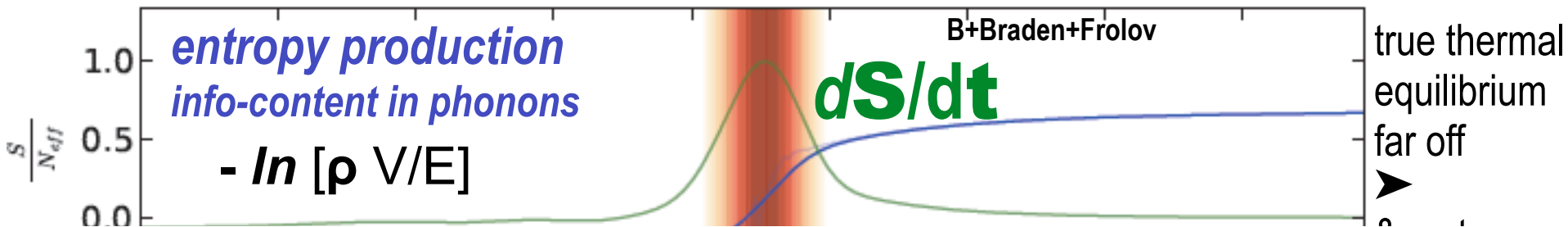


$$V(\phi, \chi) = 1/2 m^2 \phi^2 + 1/2 g^2 \phi^2$$

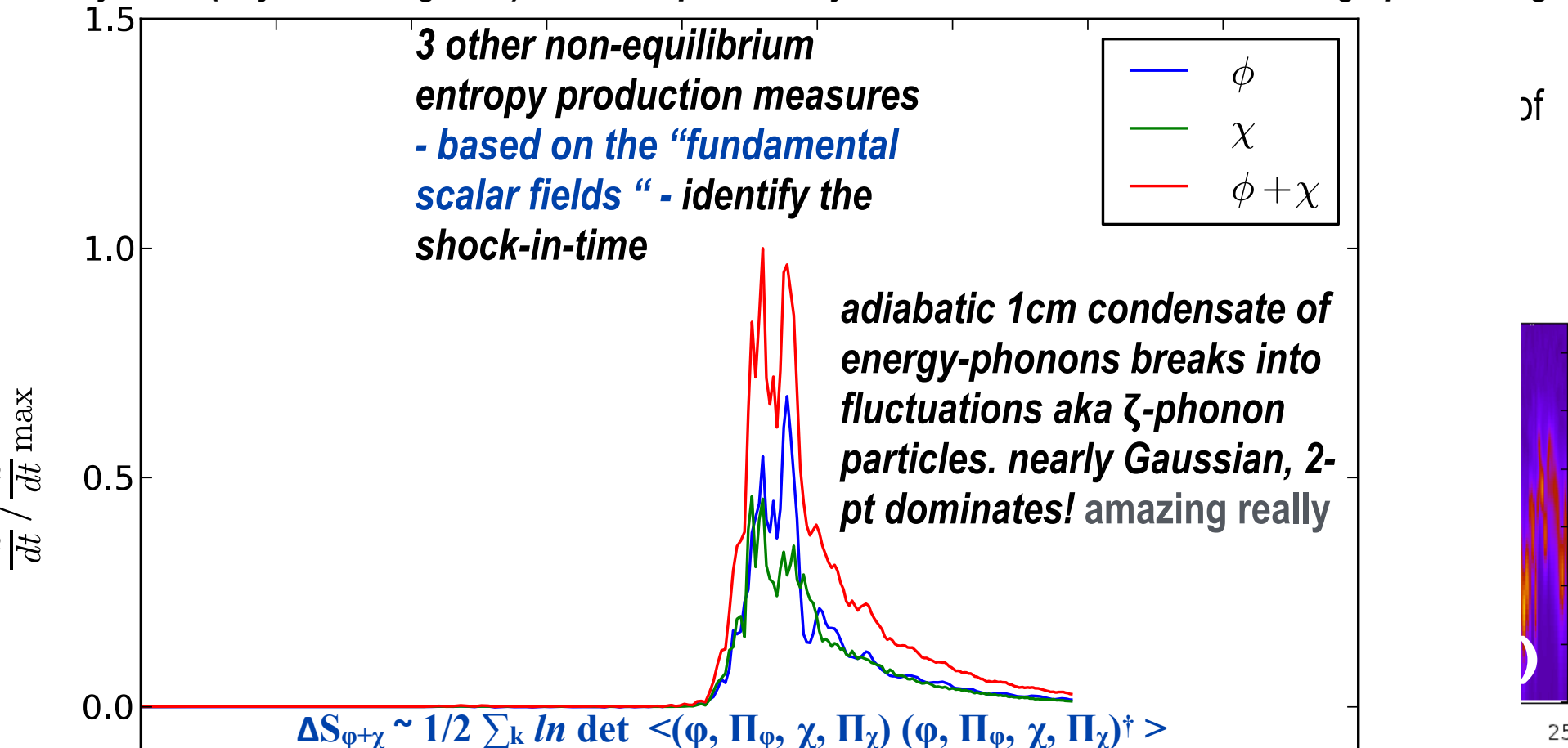
k/m momentum

B+Braden+Frolov

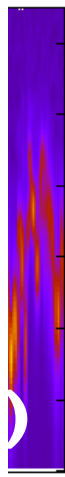
nonG from large-scale modulations of the shock-in-times of preheating



decay rates (Feynman diagrams) and transport theory difficult to make accurate through preheating



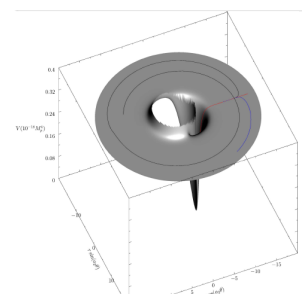
of



stochastic inflation:
 the battle of classical drift V_C &
 diffusion of quantum fluctuations V_D
 $V_C \sim \nabla S_R$ $V_D = D_H \nabla S_I$
 eternal inflation $\Rightarrow V_D$ dominates
 emergence $\Rightarrow V_C$ dominates

inflaton+isocons potential $V(\varphi, \chi, \dots) = ?$

$\varphi_C, \chi_C, \dots, \alpha_C, H_C$



$V(\varphi, \chi, \dots)$

Preheating After
 Roulette Inflation

$\langle \tau \rangle =$

$\hbar H_C$
 quantum
 diffusion
 spatial jitter

entropy
 generation in
 preheating
 from the
 coherent
 inflaton
 (origin of all
 matter)

$-M_{Pl} \nabla \ln H_C$
 drift

isocon directions χ
 e.g., axion

let there be
 heat

SEMIFUNCTIONAL INFLATION-NO
 ENZEMER

conformal potential-flattening eg Higgs inflation SBB89 etc

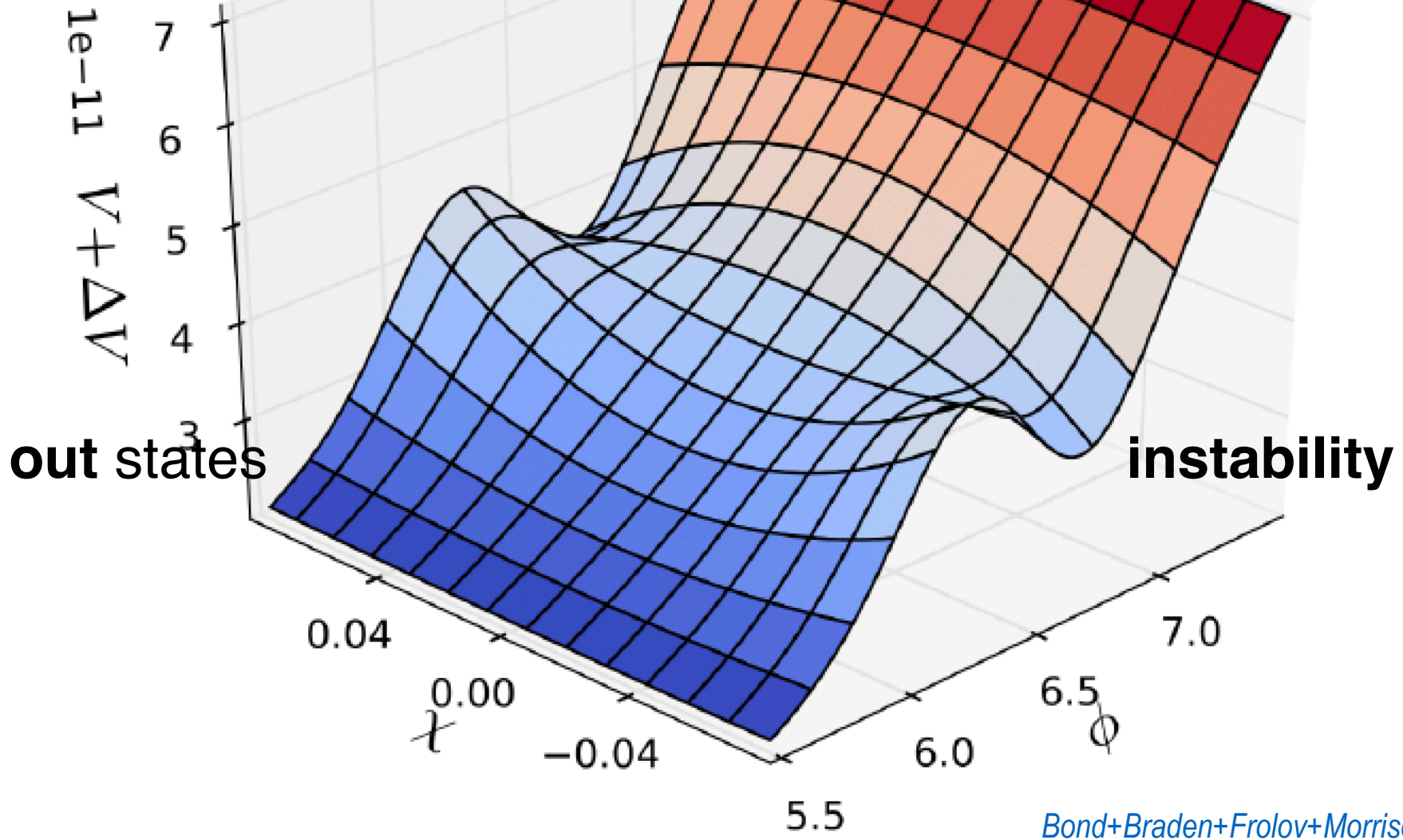
B2FH, b+braden+frolov+huang

during inflation - instabilities => entropy => nonG

*numerical experiments
of in-out states through
localized ΔV .
chain together .. oscillating*

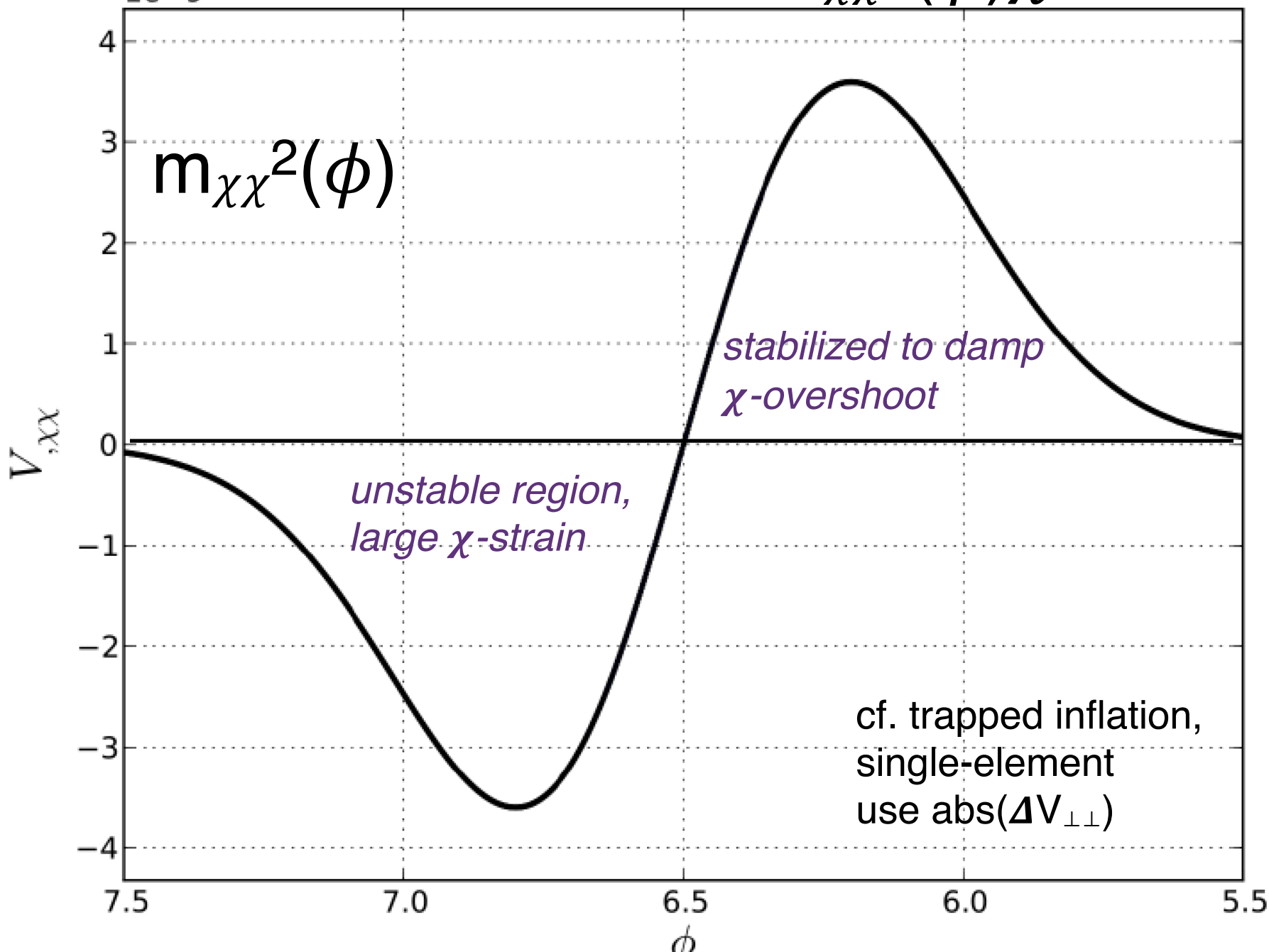
experiment χ -light

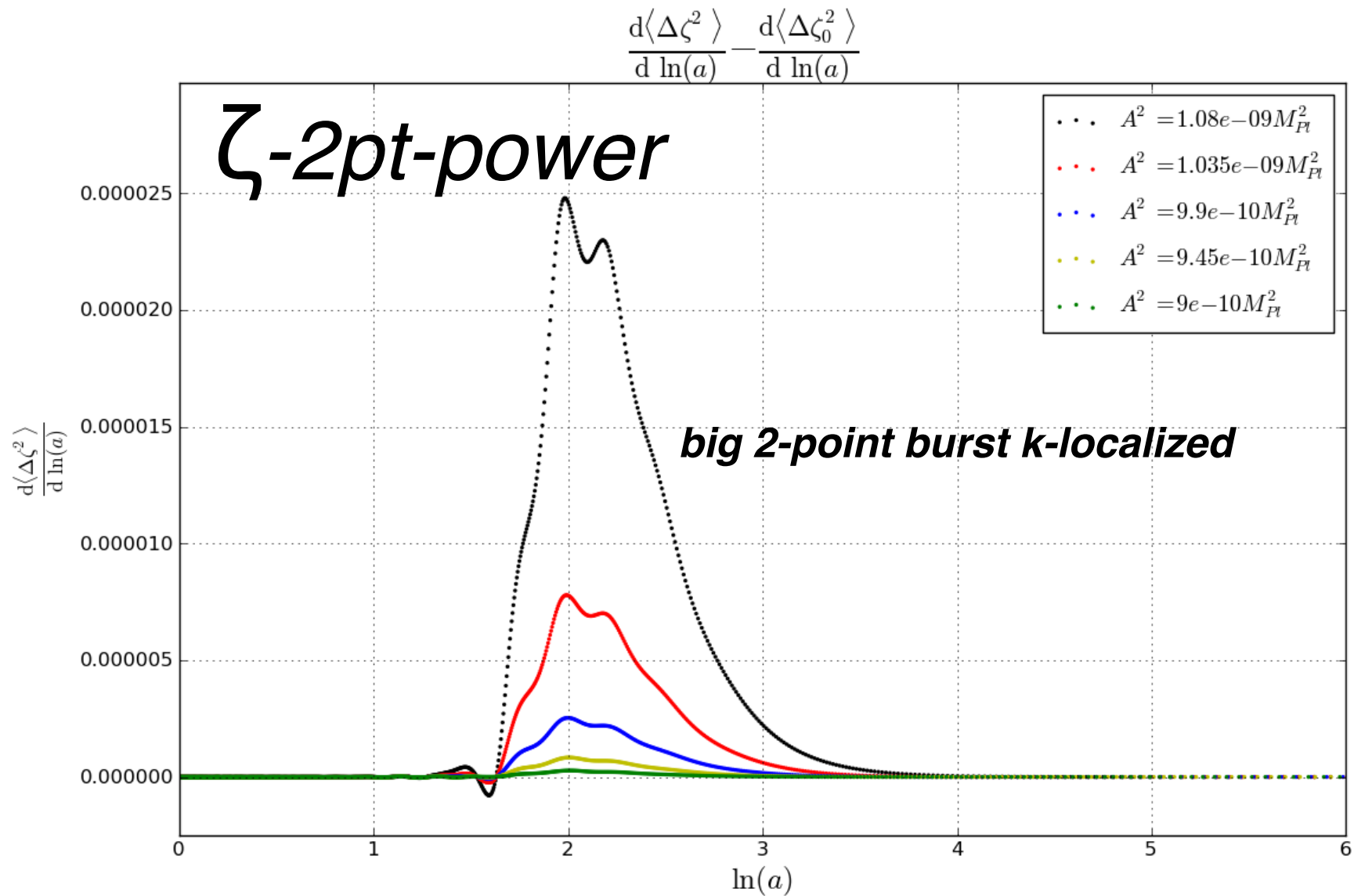
in states

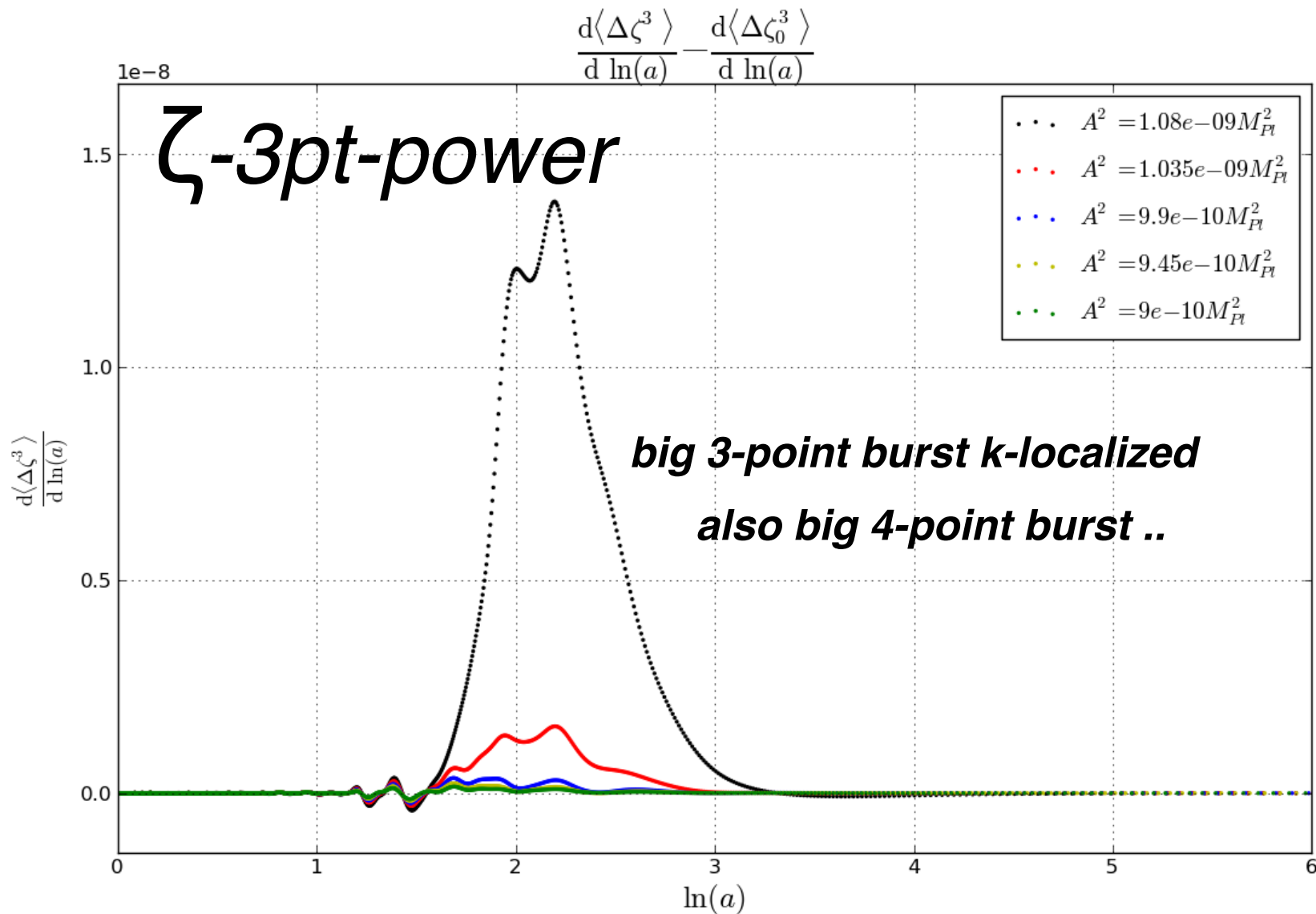


experiment χ -light
1e-9

$$\Delta V_{\perp\perp} = m_{\chi\chi}^2(\phi) \chi^2 / 2$$





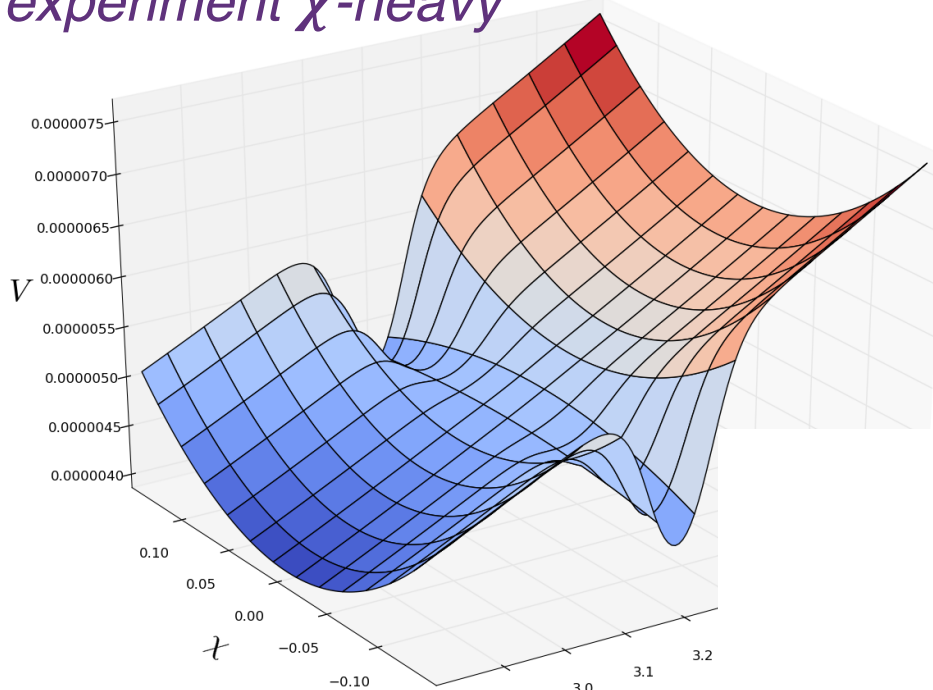


TBD coherence of N -point bursts in N -space ..

during inflation - instabilities => entropy => nonG

numerical experiments
of in-out states through
localized ΔV .
chain together .. oscillating

experiment χ -heavy



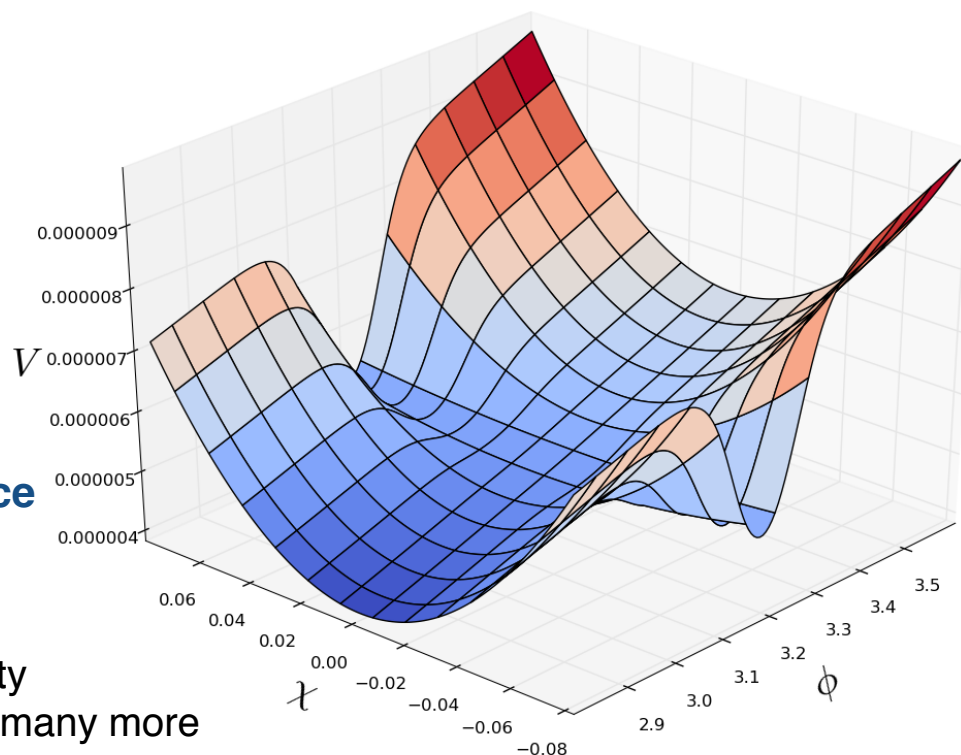
instability potential surface

Bond+Braden+Frolov+Morrison

trapped potential surface

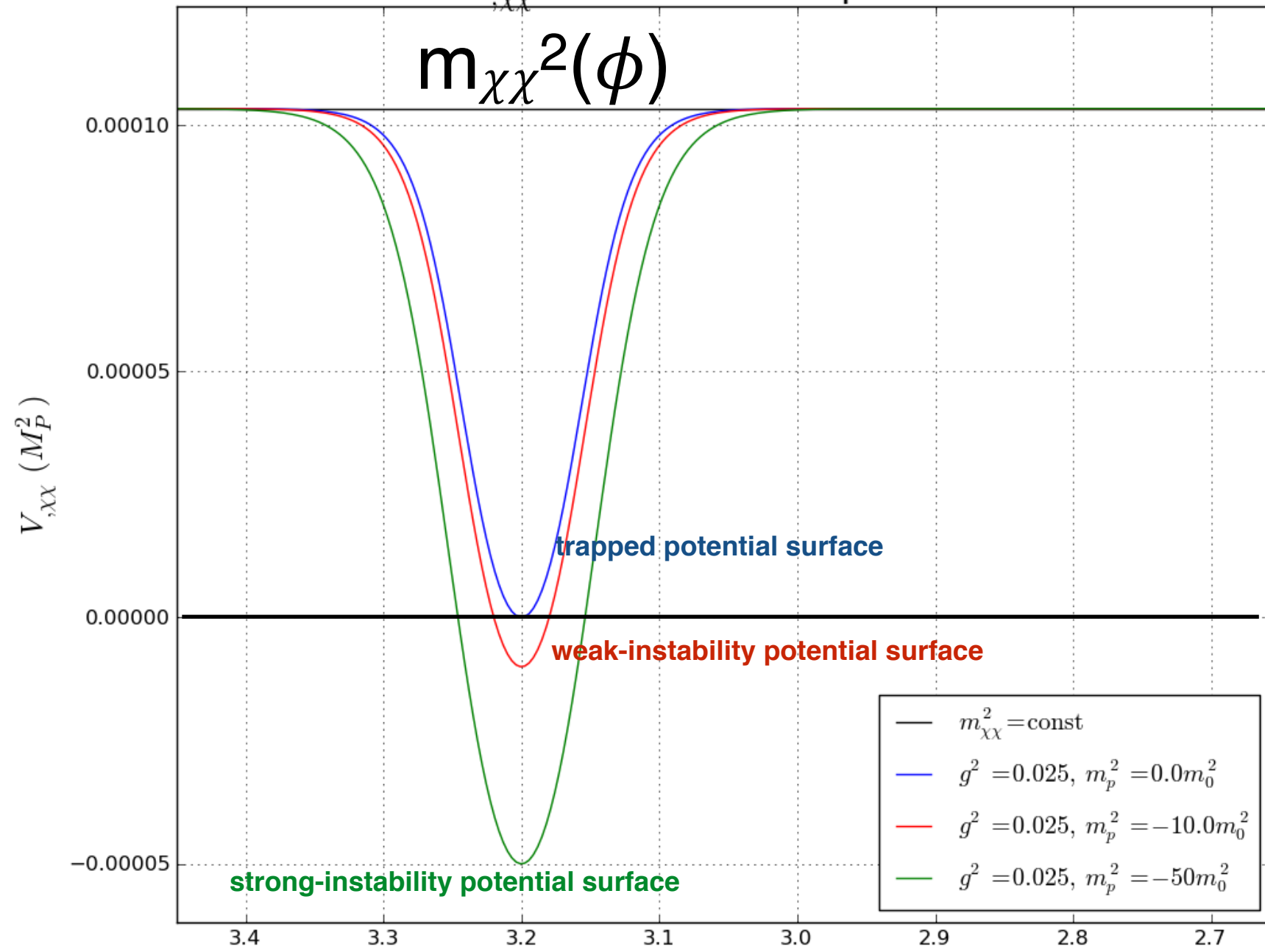
trapped inflation: same parameters, no instability

.. Kofman, Silverstein, Green, Barnaby, Huang, many more



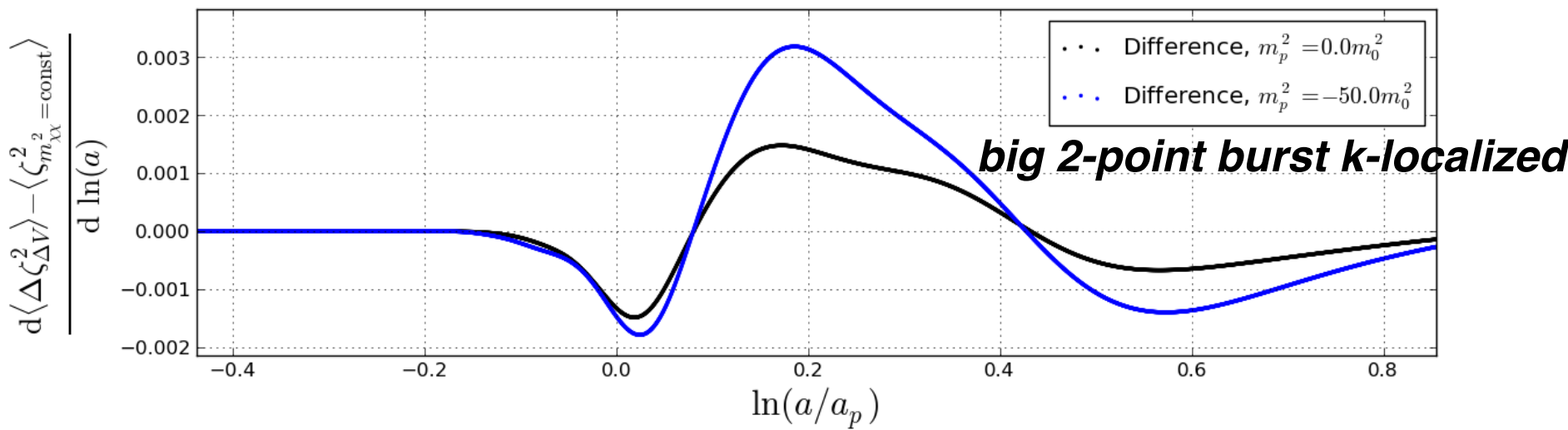
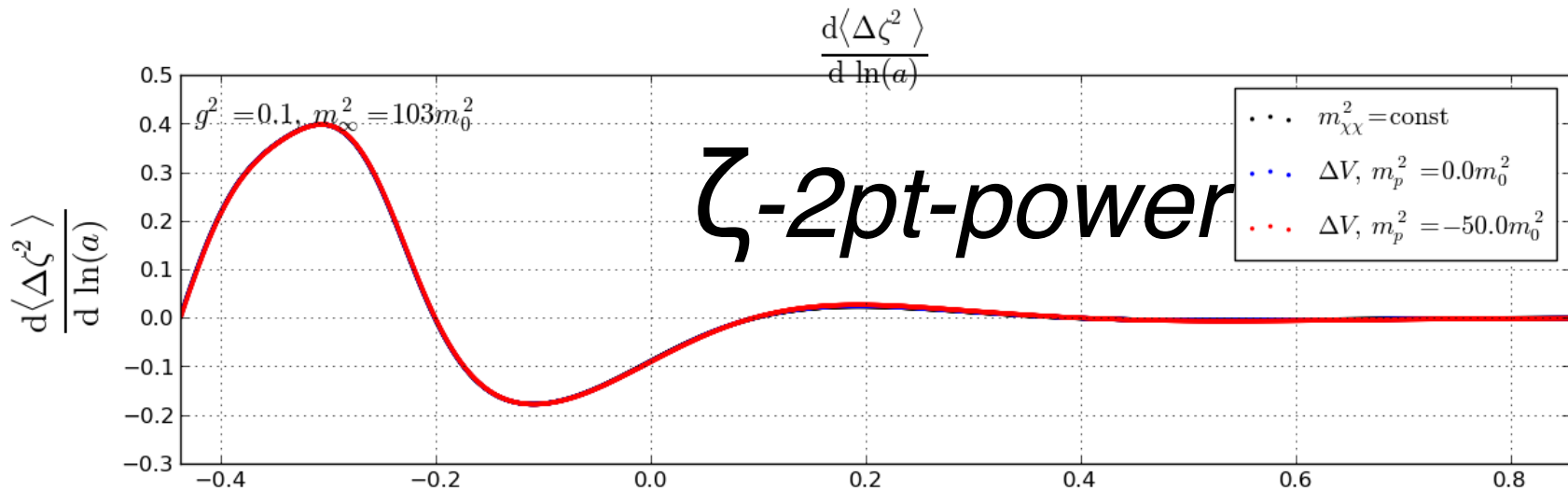
$V_{,\chi\chi}$ Parameter Comparison

$$m_{\chi\chi}^2(\phi)$$



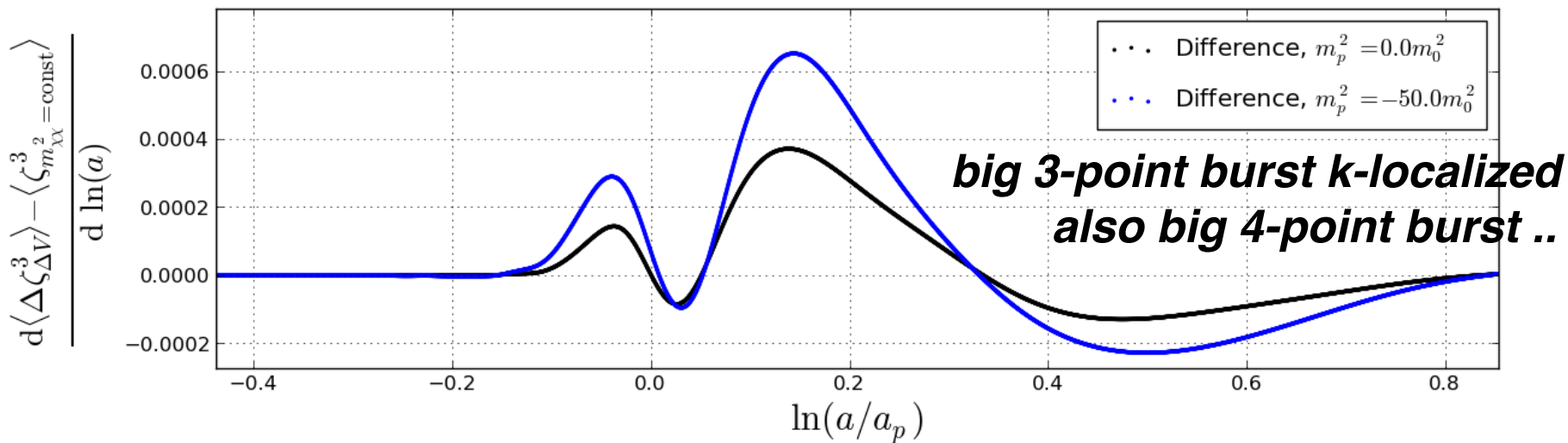
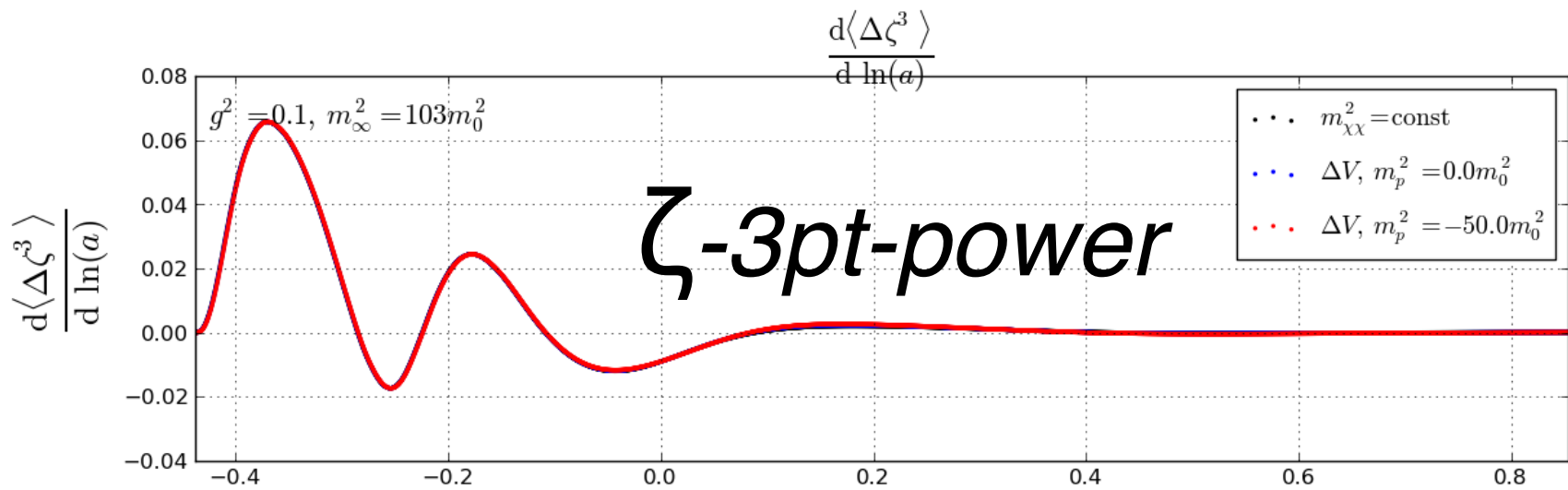
experiment χ -heavy

unstable χ cf. trapped



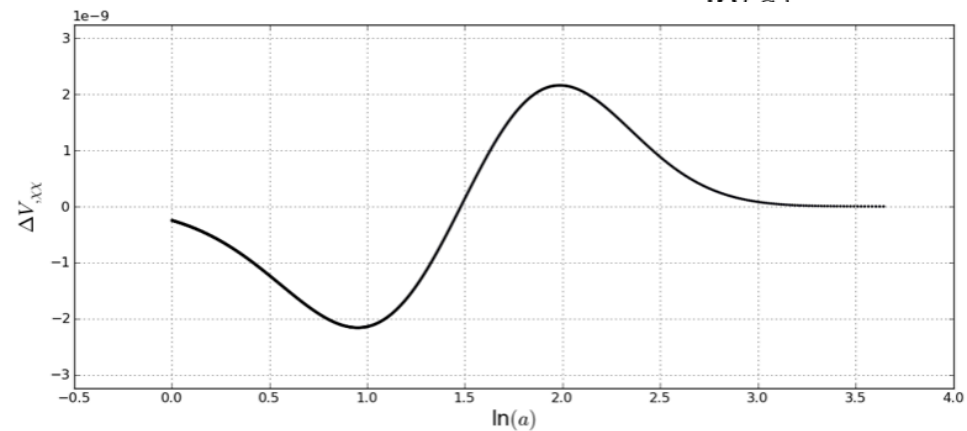
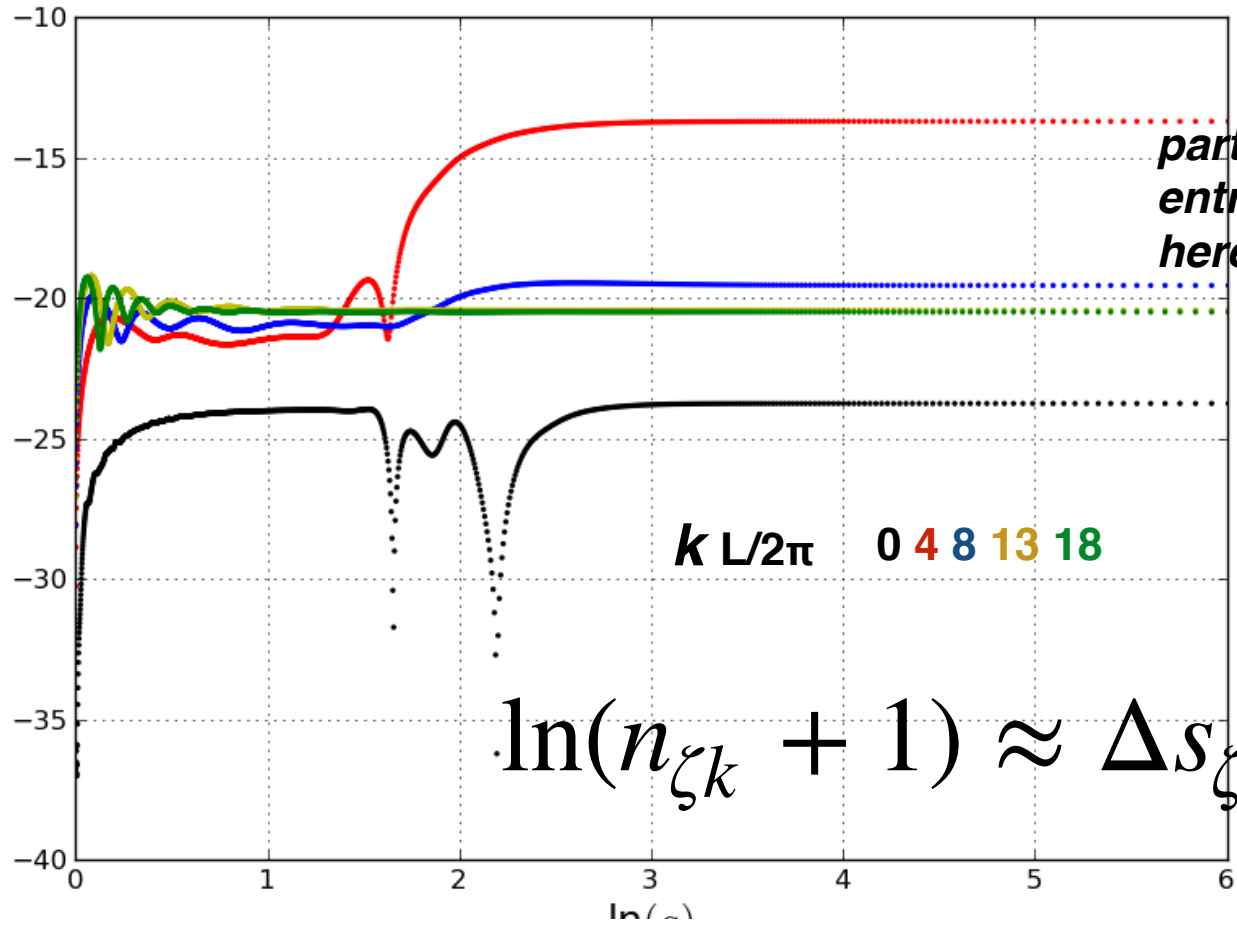
experiment χ -heavy

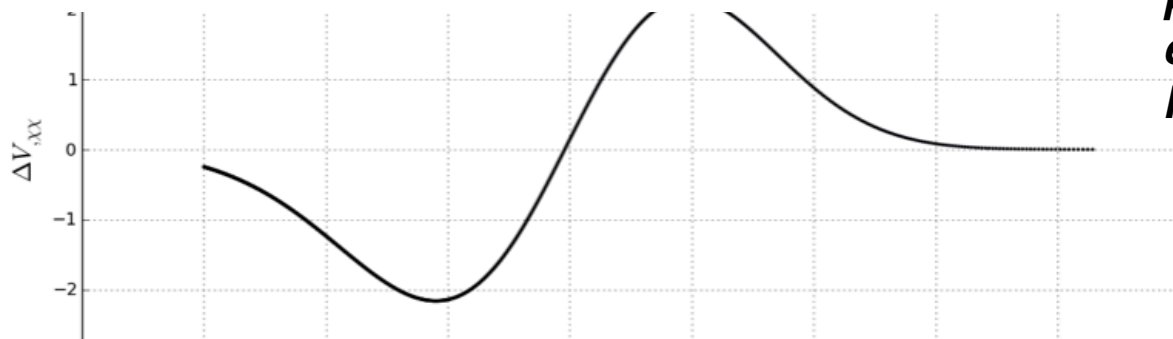
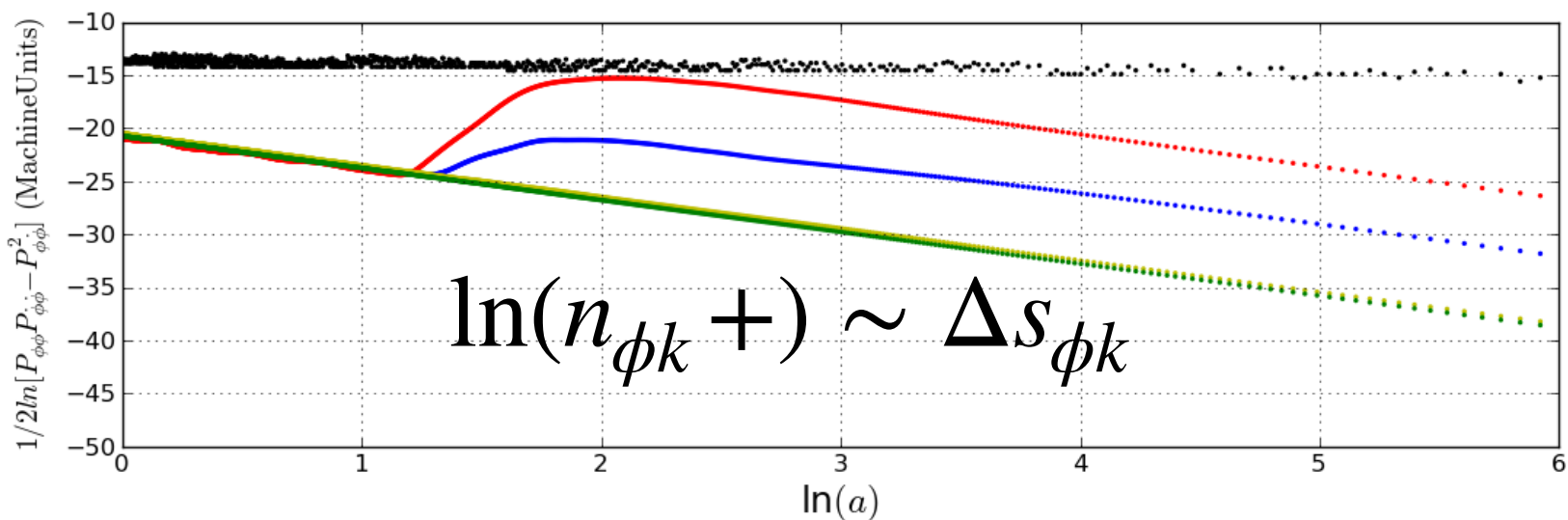
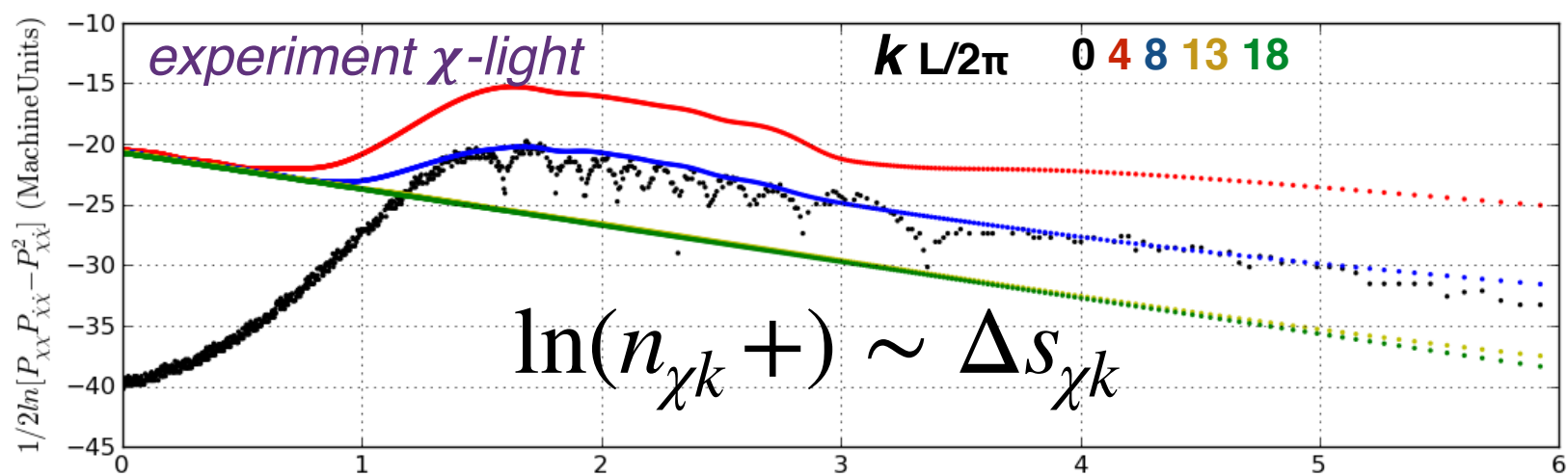
unstable χ cf. trapped



TBD coherence of N -point bursts in N -space ..

experiment χ -light $\ln(P_{\zeta\zeta})$





**particle production picture aka
entropy production picture
here in isocons & inflatons**

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what are the degrees of freedom / parameters of the ultra early Universe? TBD

begin-inflate => inflate => end-inflate => preheat => non-equilibrium heat+entropy
=> *Standard Model particle physics* QG plasma radiation dominated
=> dark matter dominated *structure via gravitational instability* => dark energy now

$$d\zeta(x,t) = (dE+pdV)/3(E+pV) = d \ln \rho_c / 3(1+w_c) + \text{Trace } d\alpha^i_j$$

fit into a *UV-complete theory (ultra-high energy to the Planck scale)* strings, landscape, ..
& *IR-complete theory (post-inflation heating -> quark/gluon plasma)???* TBD

role of (1) *instabilities after inflation*

entropy generation via the breakup of the coherent low-k inflaton condensate into incoherent high-k fluctuations at a “shock-in-time” => **nonGaussianity**

role of (2) *instabilities during inflation*

phenomenology of in-states propagating through localized unstable potential structures to out-states, like scattering theory => **nonGaussianity**

(3) $|cg\rangle \Leftrightarrow |fg\rangle$ condensate/fluctuation framework, for both using coherent states classical-like approach with \hbar .

includes **Bogoliubov** transformations for fluctuations as condensate evolves => **particle creation** interpretation in both heating and inflating regimes.

END