

from SuperWeb simplicity to complex Intermittency in the Cosmic Web



Zeldovich 100th:
Moscow June 2014,
Tallin IAU 308 June 2014

the Russian school & how Arnold fits into the evolving cosmic story 79-82 (Arnold, Shandarin & Zeldovich 1982)



Dick Bond &
Sergei Shandarin, U of Kansas



CITA
ICAT

Canadian Institute for
Theoretical Astrophysics

L'institut Canadien
d'astrophysique théorique

our Hubble patch

Initial conditions: a Gaussian random field in energy density & spatial curvature

initial content: scalar fields (inflation) + gravity + ??

post-inflation heating content: the standard model of particle physics, quark/gluon plasma

content now: baryons in a gas, photons, neutrinos, (cold) dark matter, dark energy

linear evolution phase, to reheating $1/a \sim 10+$ when nonlinearity => the 1st stars form nonlinear SCALE grows, defining the evolving cosmic web

ingredients now: clusters/groups = halos of Dark Matter, filaments, membranes, voids; galaxies, stars

IN EVERY teaspoon of air ~ 5 cubic cm

Ordinary Matter $\sim \text{amu}/\text{nm}^3$ 4.8% O₂ N₂; H, He

THE DARK

Dark Matter

$\sim \text{amu}/\text{m}^3$ $26.0 \pm 1\%$ compressed in MilkyWay $\sim 0.1 \text{ amu}/\text{cm}^3$; for LHC@CERN-type relics ~ 1 every 10 cm

Dark Energy

\sim vacuum potential $\sim 3 \text{ amu}/\text{m}^3$ $69.2 \pm 1.0\%$

THE LIGHT

cosmic radiation

the 1st light of the universe $412/\text{cm}^3$ 0.005%

cosmic neutrinos \sim cosmic photons > 0.47%

cosmic gravity waves << cosmic photons

THE VACUUM

Higgs@CERN vacuum origin of mass

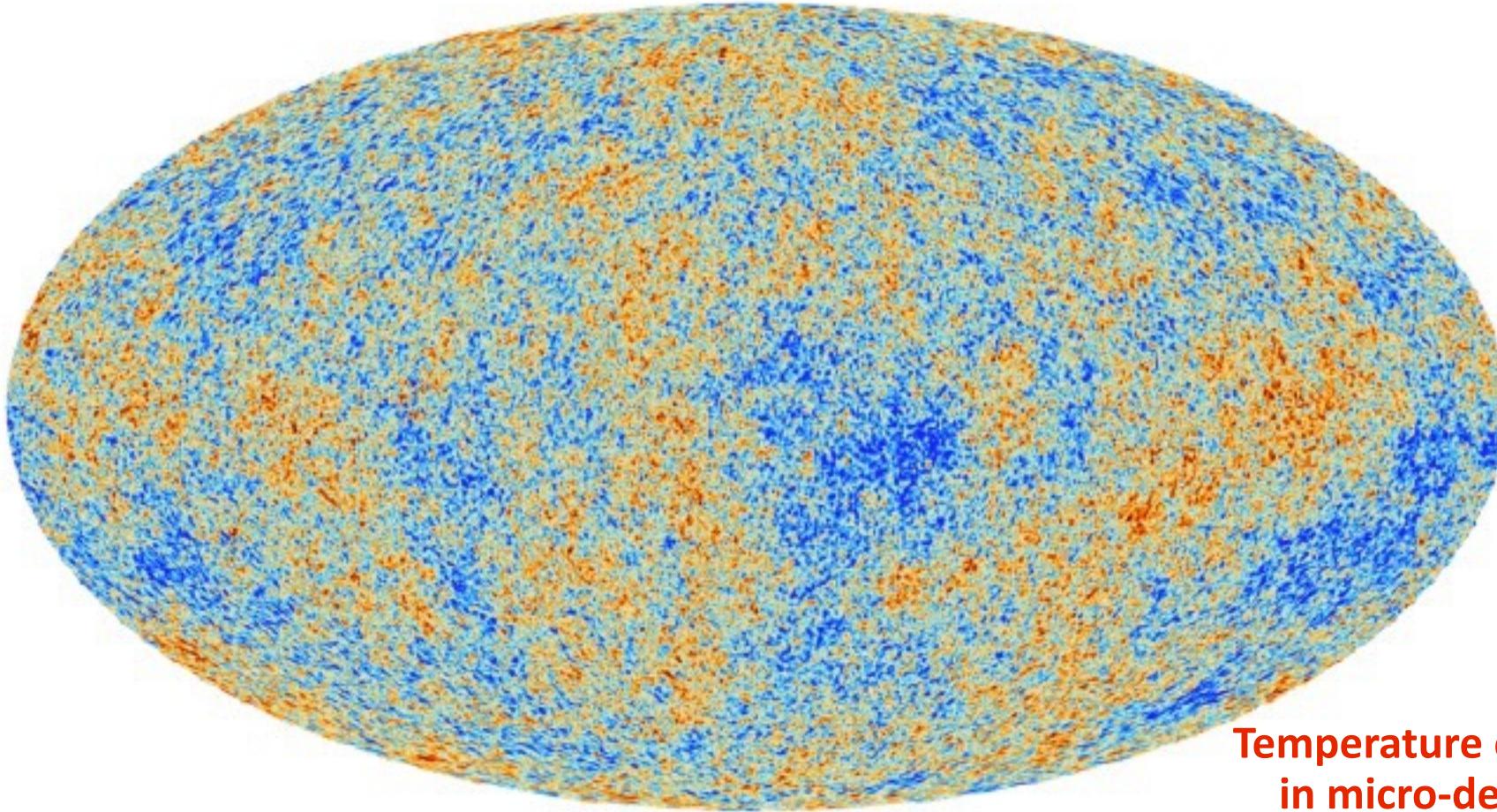
vacuum fluctuations origin of all the cosmic structure we see

Planck's primordial light unveiled, March 21, 2013 (Dec. 1, 2014)

reveals the **SIMPLICITY** of primordial cosmic structure
in 7^+ numbers

=> learn **matter & energy content & structure** at $a \sim e^{-7}$ 380000 yr

=> infer structure far far earlier $a \sim e^{-127} \sim 1/10^{55}$ **in ~ 2 numbers**



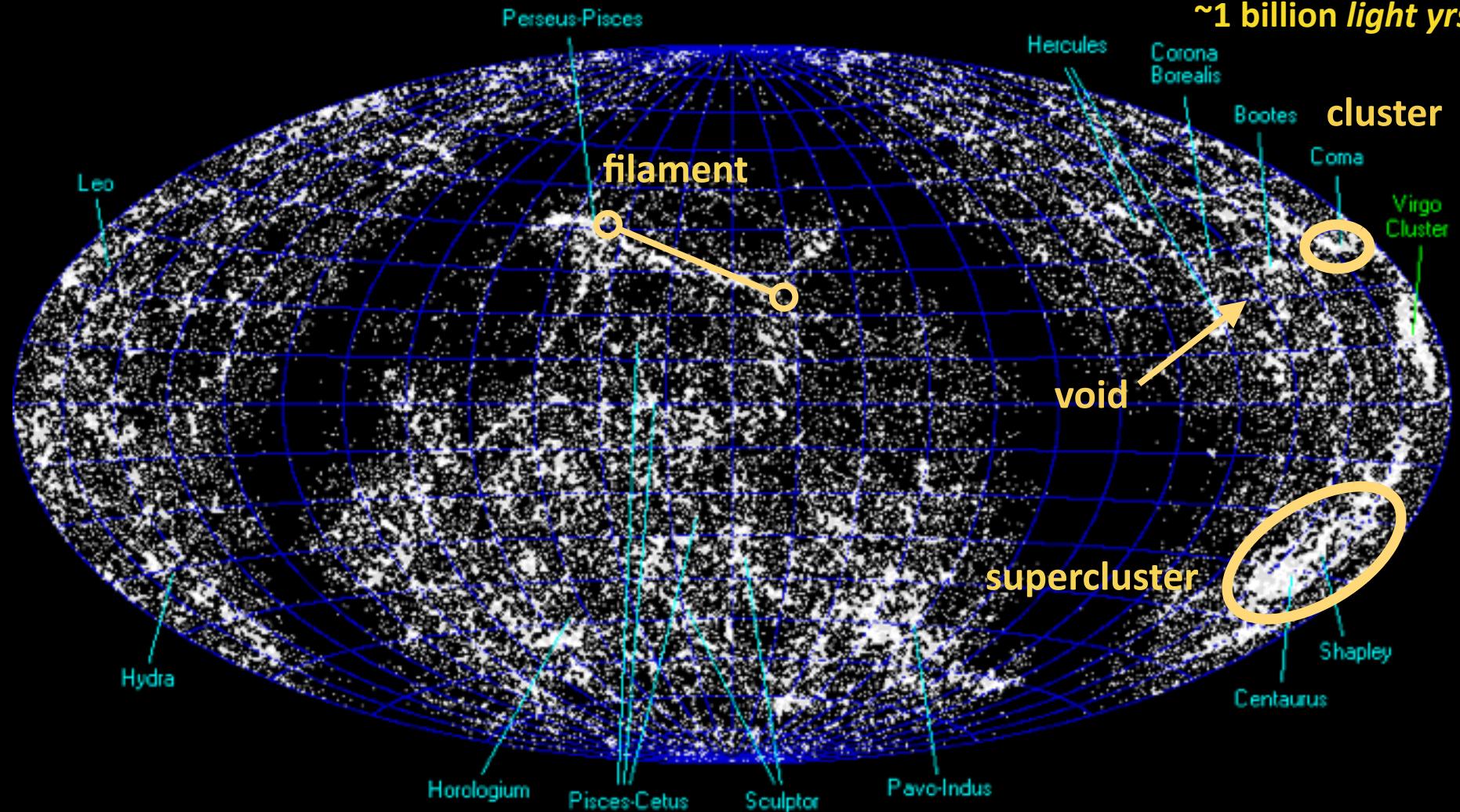
Temperature changes
in micro-degrees

Gaussian 2D temperature field @ 370000 yrs => Gaussian random 3D density field from the ultra-early universe

Cosmic Web of 60,000 nearby galaxies: exhibits “local” COMPLEXITY

$$a \sim e^{-0.1} = 1/1.1$$

~1 billion *light yrs*



hard won observational emergence of the web. 79-81 sparse info, e.g., of Coma supercluster. So what Arnold, Shandarin and Zeldovich knew was very very much less, ie speculative theory

to $a \sim 0.9$ via 3D maps
cosmic web of nearby superclusters < Gigayr

$a = e^0 = 1$ now

$a \sim e^{-0.1} = 1/1.1$

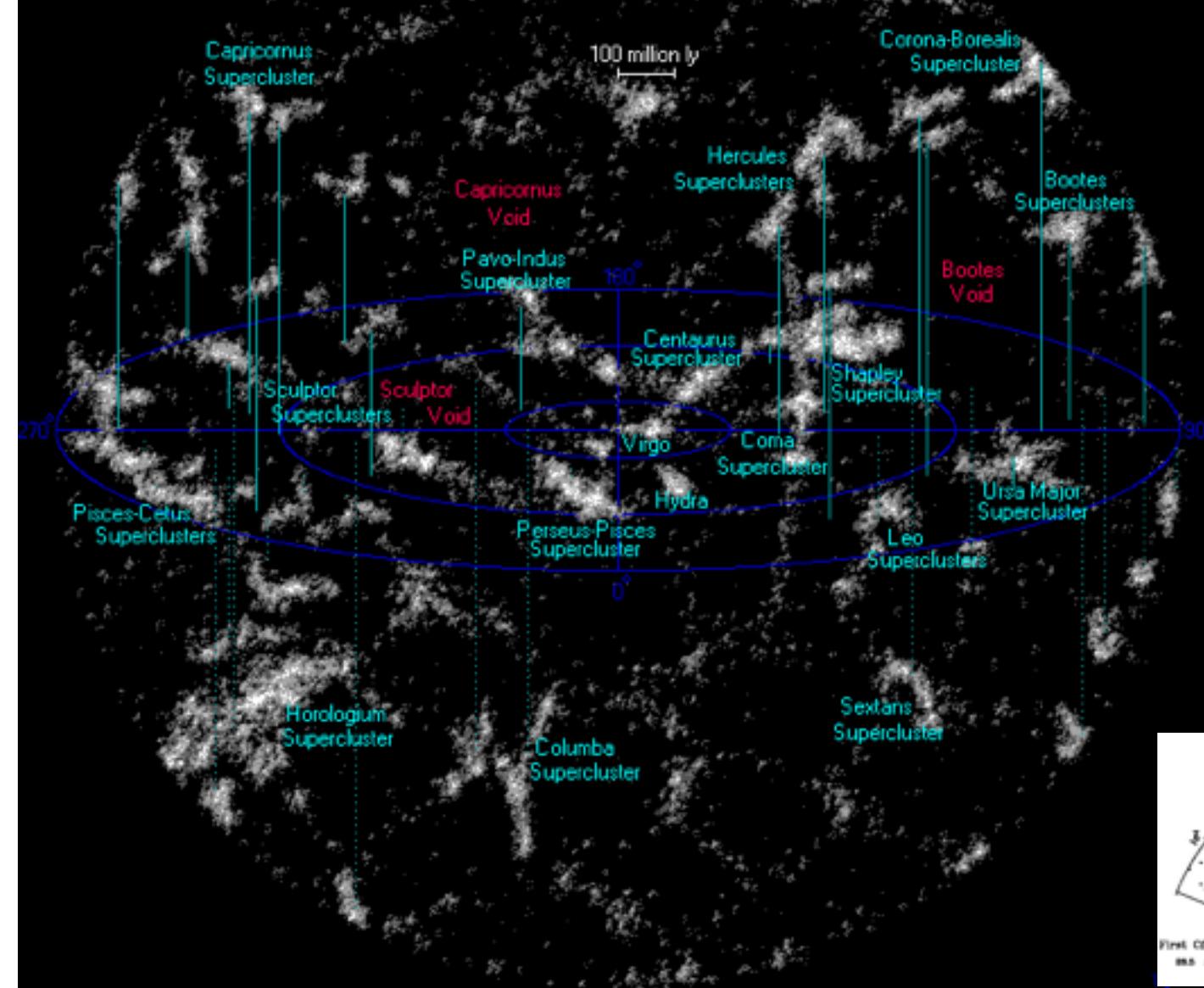
70s adiabatic
pancake
(physical filter)
Doroshkevich, Zeldovich

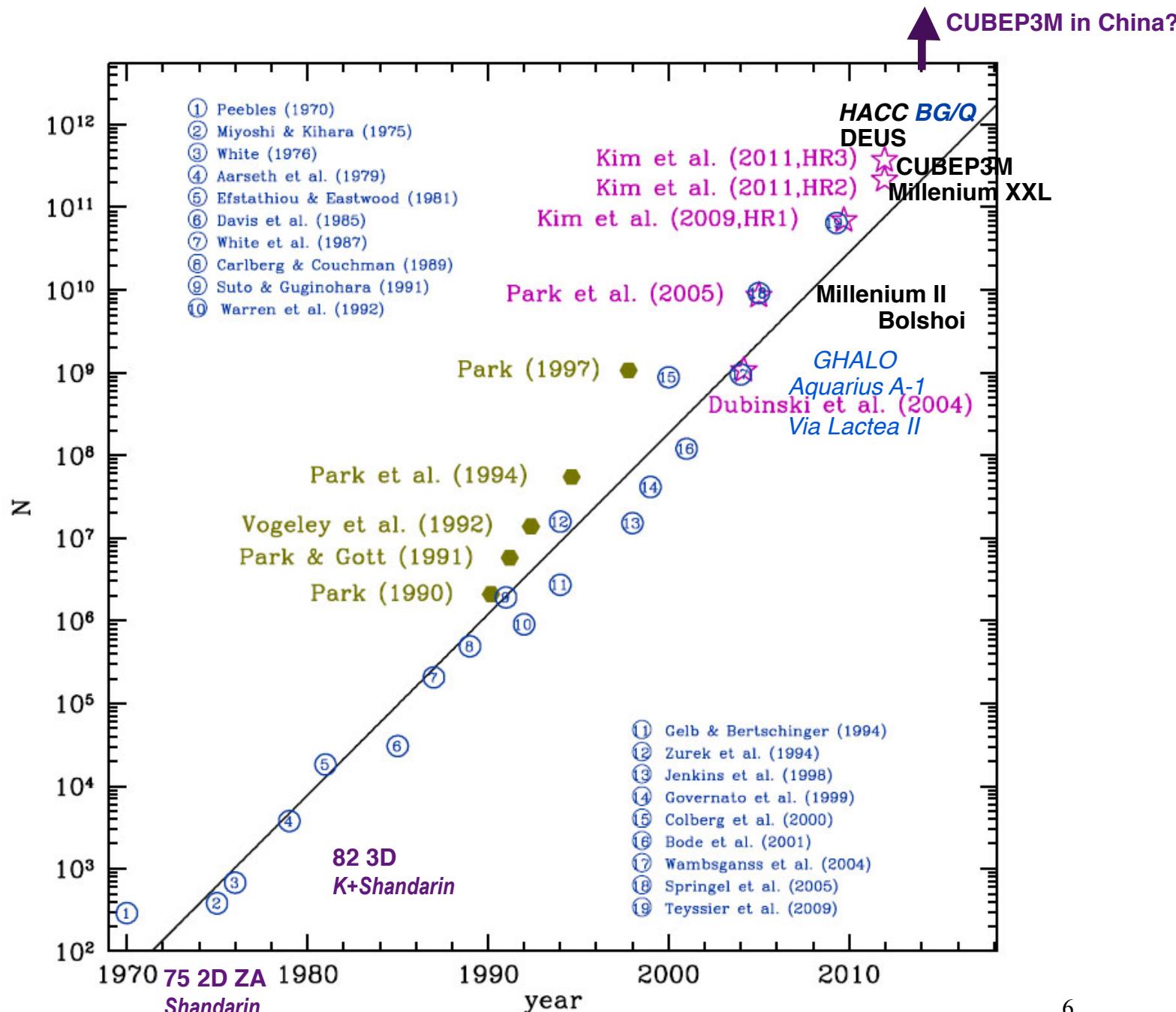
cf.

70s isoc B/BH
(power law CorrFn)
Basko

**miracle of
CDM = grand
unification
of east & west
ideas
with \sim HSZ
spectrum
emergence of
superclusters**

Peebles vs.
70s Einasto+..
80 + Oort +

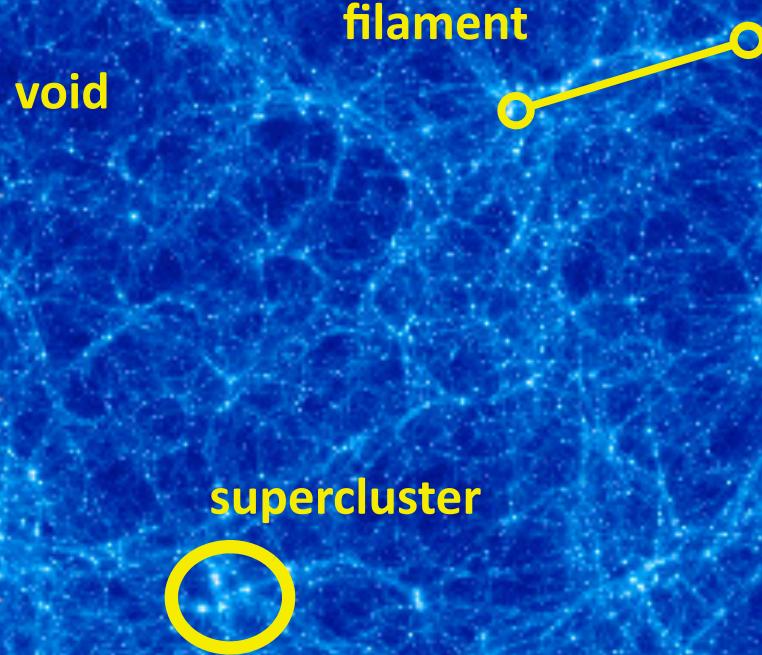




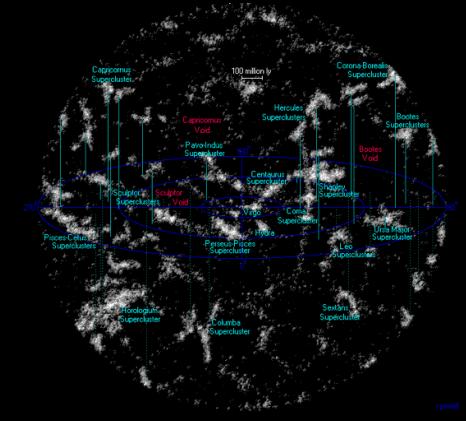
Simulation of the 7⁺ numbers

begets the **Cosmic Web** of clusters now $a \sim 1$ & galaxies then $a \sim 1/4$

SIMPLICITY to COMPLEXITY under Gravity



1st light simplicity
 $a \sim e^{-7} \sim 1/1100$

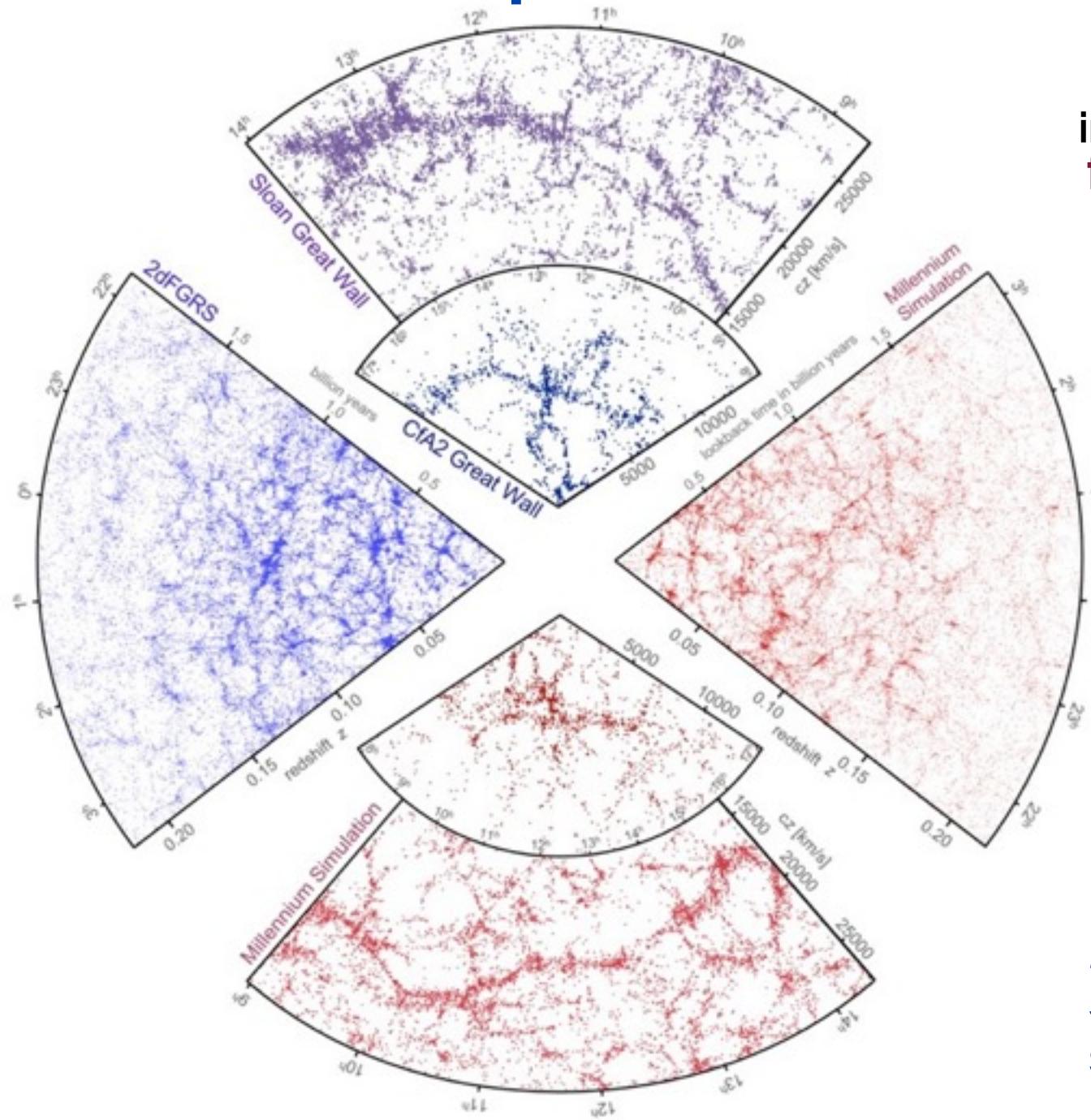


~ billion light years

state of the art simulation
 $a \sim 1$ to $1/1.1$

ordinary matter
dark matter
dark energy

to $a \sim 0.8$ via 3D maps



Collisionless matter
Simulation of the
initial Gaussian random
field characterized by
7⁺ numbers
does indeed beget the
Cosmic Web

Millennium simulation web
site “propaganda” on
sims cf. z^8 -space data

and to **a ~ 0.7 to 0.5 via 3D maps**

VIPERS using VIMOS@VLT release Oct 4, 2013, 57K redshifts, z=0.45 to z=0.95, $6e7 (h^{-1}\text{Mpc})^3$, higher sampling than LRG BAO surveys Guzzo+13 cover CFHTLS wide fields, 64% done, 24 sq deg

Field W1

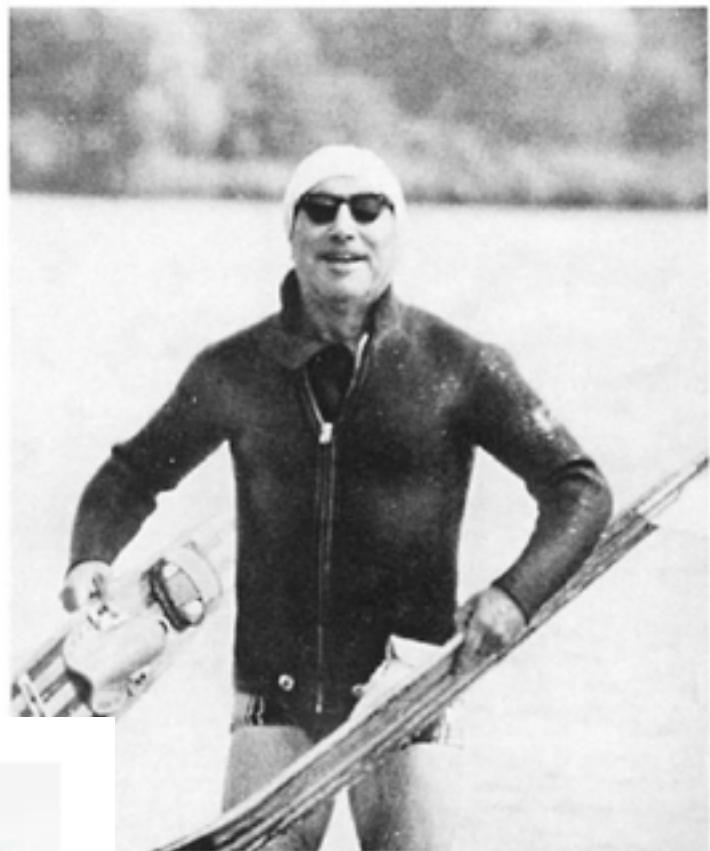
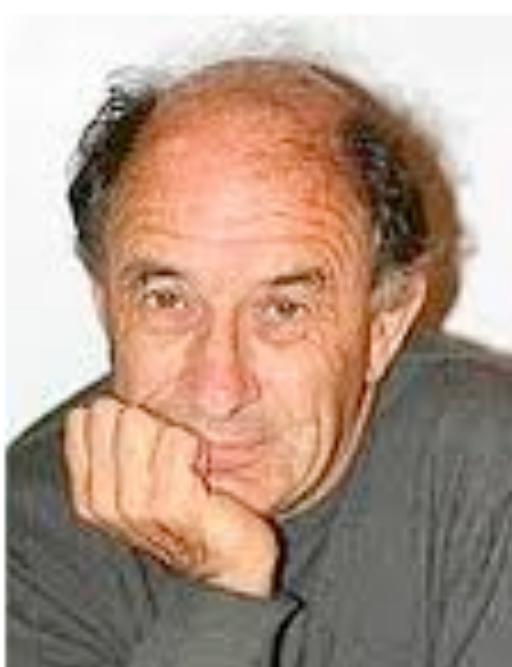


Field W4

Yakov Borisovich
Zel'dovich
1914 - 1987



Vladimir Igorevich
Arnold
1937 - 2010



На здоровье

Terviseks

Z70,ZES82
AZS82
eJ



Zeldovich 1970 approximation $\mathbf{X}(\mathbf{r},t) = a(t) (\mathbf{r} + \mathbf{s}(\mathbf{r},t))$
general map of a cold medium, an onto multi-stream map

The first numerical
simulation of the
Zel'dovich
Approximation
in 2D

Shandarin 1975

published in review
by Doroshkevich
Zeldovich
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(in Russian)

Later in
Dorshkevich,
Shandarin 1978

influential for Arnold,
Shandarin Zeldovich 1982



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brief history of understanding objects and their distribution in the cosmic web

50s+ HALOS hierarchy, small round objects => large round objects

70s: Doroshkevich, Shandarin, Zeldovich: 1st order Lagrangian dynamics, statistics of 1D collapsing entities (caustics & pancakes) in a GRF; 80s: Arnold, Shandarin & Zeldovich: influential picture of 1st order catastrophes; $1D \Rightarrow 2D \Rightarrow 3D$ pancake \Rightarrow filament \Rightarrow cluster flows

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80s: hot, warm & **cold** collisionless **dark matter** paradigm $\Rightarrow \text{LCDM}$
87: $X =$ includes vacuum energy

90s-00s: data settled on $X = \Lambda + \text{tilt} \Rightarrow \text{dark-energy + tilt}$

$X(r,t) = a(t)(r + s(r,t))$ general map of a cold medium, an onto multi-stream map;
 $dX^i/a = (V^i - HX^i)/a dt + e^i_l(r,t)dr^l = v_{\text{pec}}^i dt + (\delta^i_l + \epsilon^i_l(r,t))dr^l$, where
 $\rho_m/\langle\rho_m\rangle = 1 + \delta_m = 1/\det(1 + \epsilon) \Rightarrow \ln \rho/\langle\rho\rangle = -\text{Trace } \ln(1 + \epsilon)$ **$\epsilon = \text{strain tensor}$**

Lagrangian 1st order linear $s(r,t) = D(t)s(r) = -D(t)\nabla\Psi_s(r)$ separable 1-1 & onto \Rightarrow caustics,
 $\Delta\Psi_s = \delta_L = -\text{Tr } \epsilon = \Phi_P(a/D)/4\pi G \langle\rho_m\rangle a^3$ $\epsilon \sim \text{tidal tensor}$: velocity potential $\Psi_v = -dD/dt\Psi_s$, $d\epsilon/dt \sim \text{shear}$

the celebrated Zeldovich 1970 approximation, with caustics aka singularities $\rho \rightarrow 0$ when $1 + \epsilon \rightarrow 0$ pancakes, but
Lagrangian to Eulerian map in 3D \Rightarrow ASZ82 BUT nonlinear tidal corrections and “hot” regions of collapsed space, not good for halos

The first numerical simulation of the Zel'dovich Approximation in 2D

Shandarin 1975

published in review
by Doroshkevich
Zeldovich
Sunyaev 1975
(in Russian)

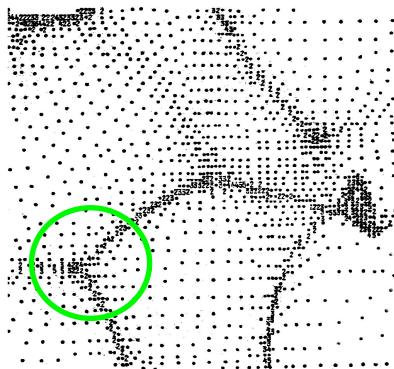
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Shandarin Zeldovich 1982

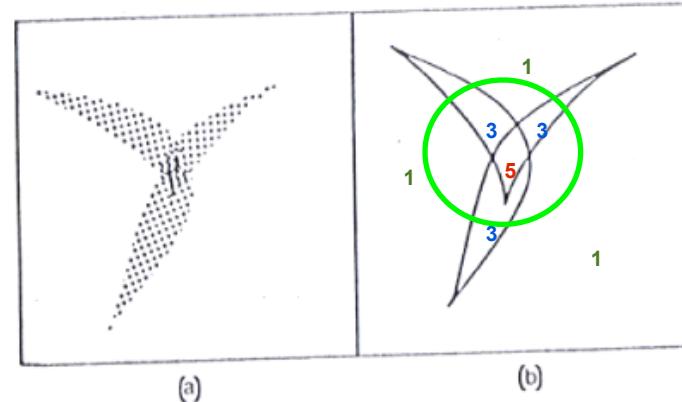


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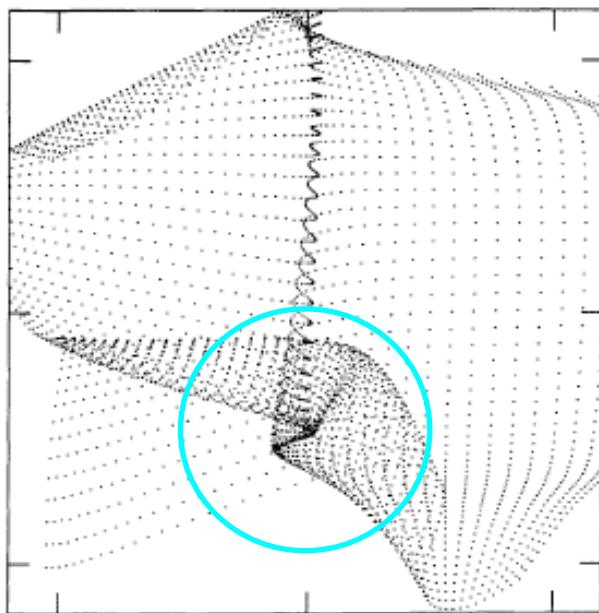
2D example: A4 (“swallow tail”) singularity



Shandarin 1975

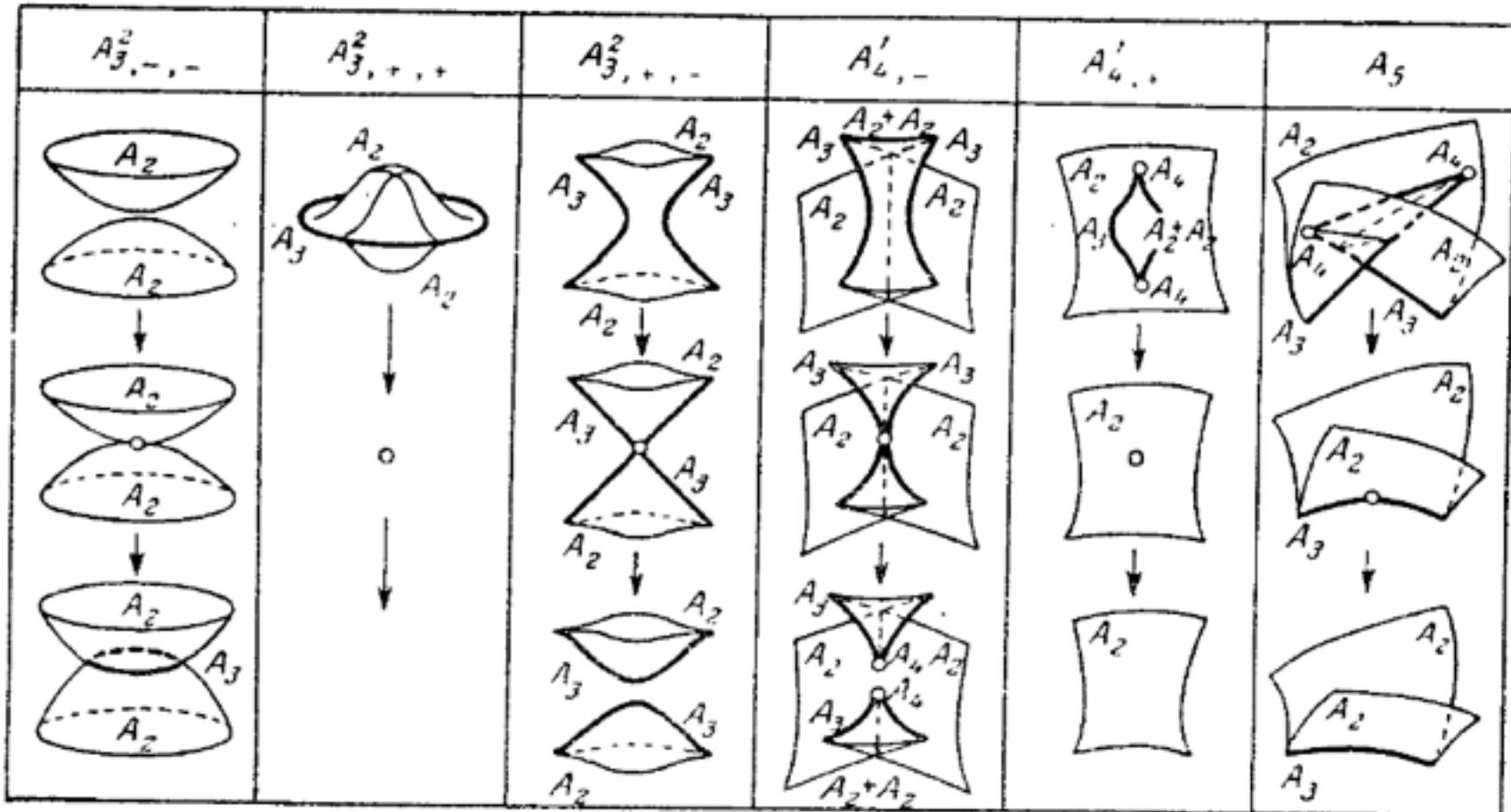


Arnold, Shandarin, Zel'dovich 1982



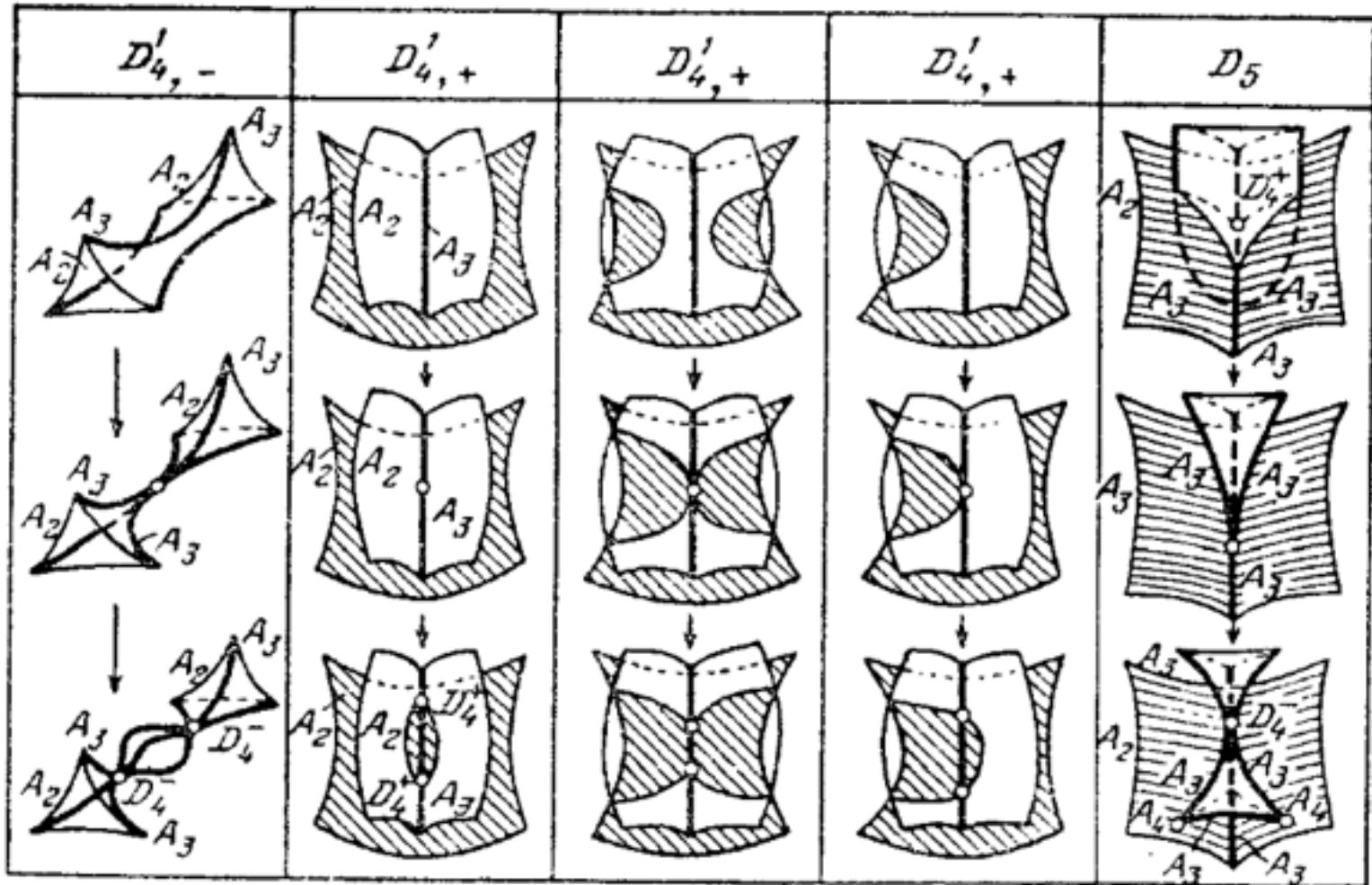
Nusser, Dekel 1990

A - caustics in 3D



Arnold, Shandarin & Zeldovich 1982

D - caustics in 3D



Arnold, Shandarin & Zeldovich 1982

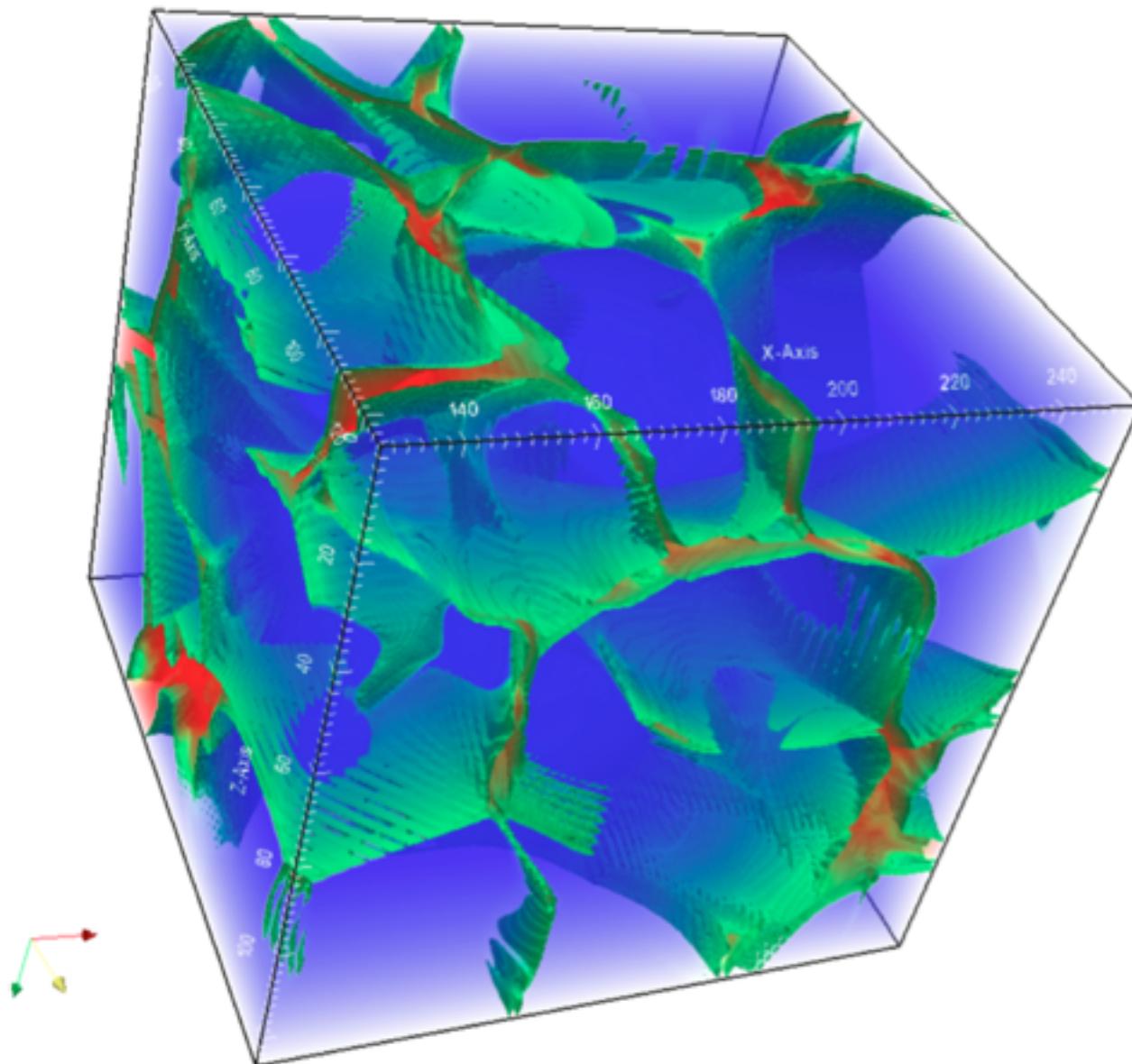
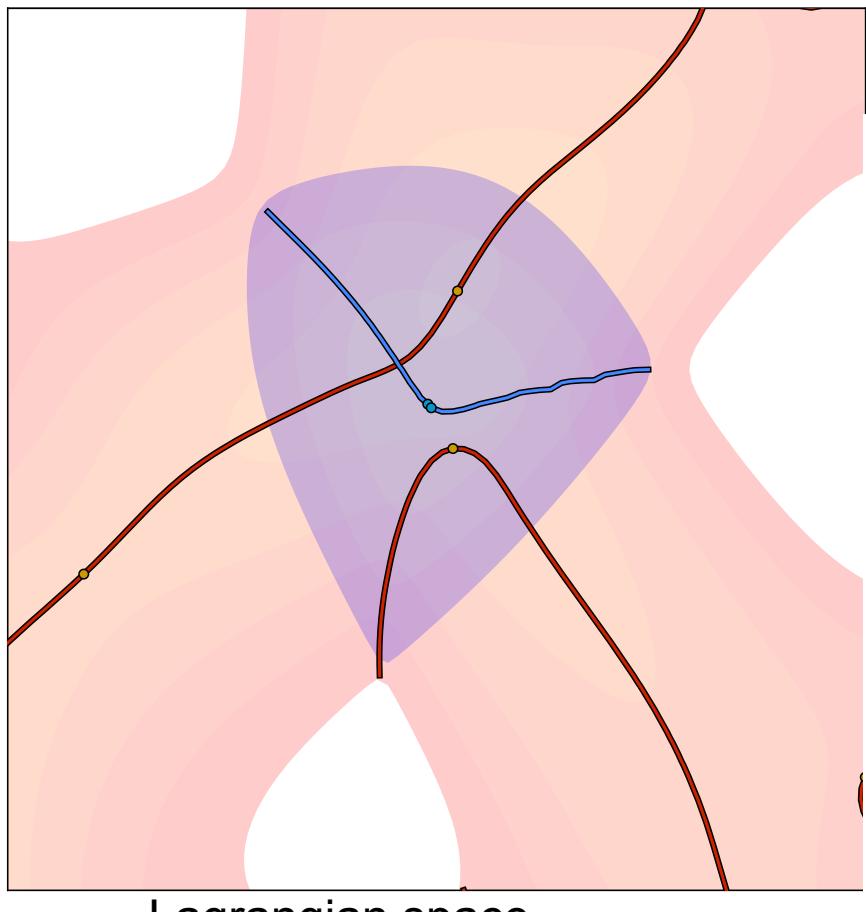


Figure 18. Caustic structure of ZA-evolved mass distribution in 3D. This is a three-dimensional analogy of the top right-panel in Fig. 1. The α -caustics are shown in green and β -caustics in red. The β -caustics are seen only through the openings in the surface of α -caustics by the box faces. This is the reason why one cannot see γ -caustics.

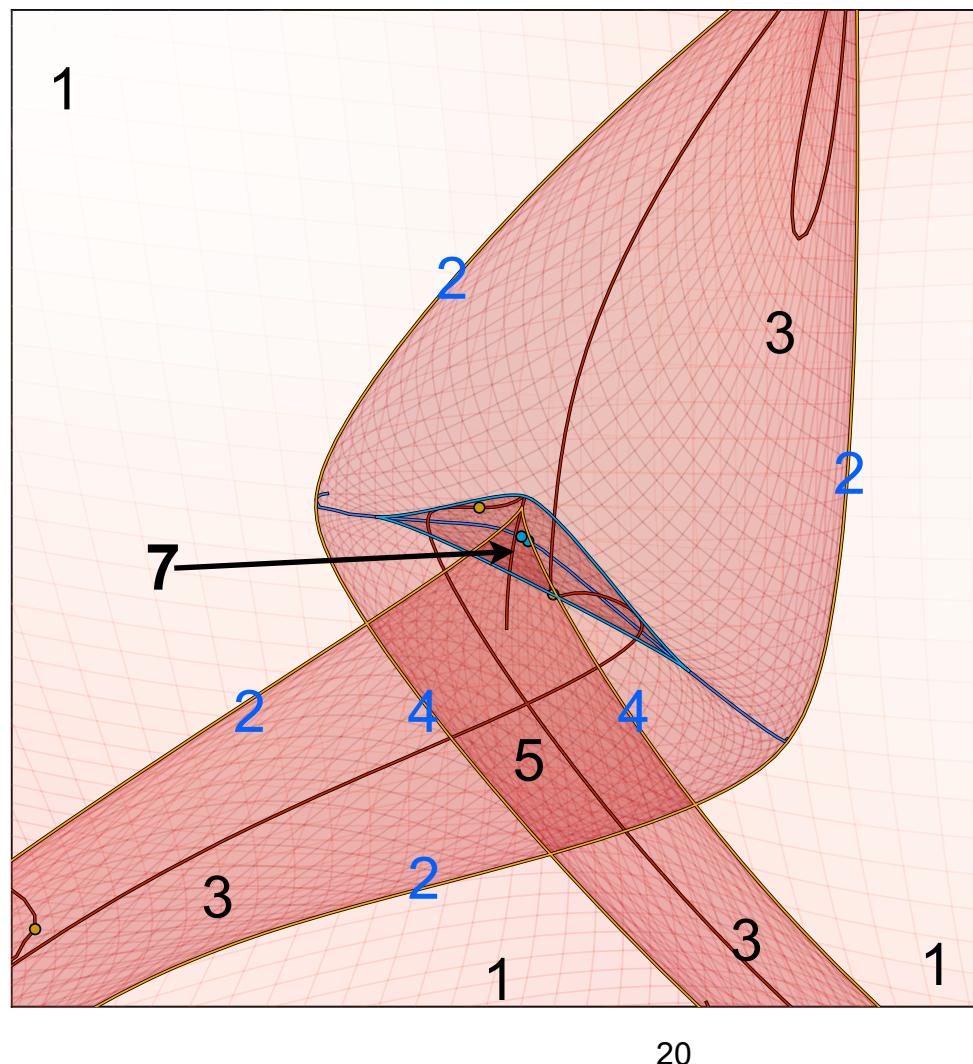
Complexity of caustics



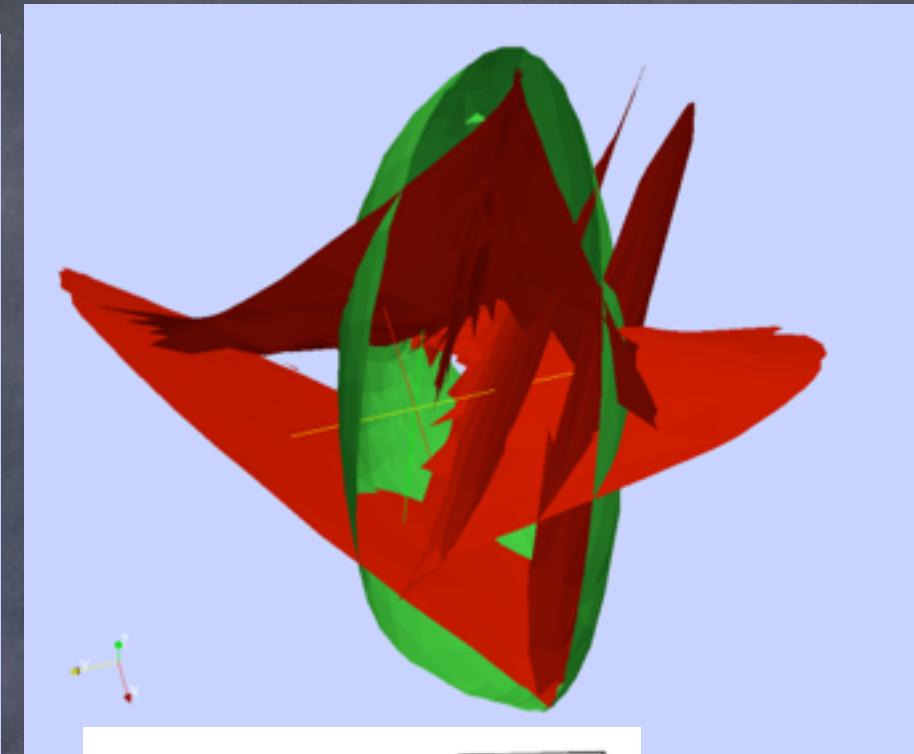
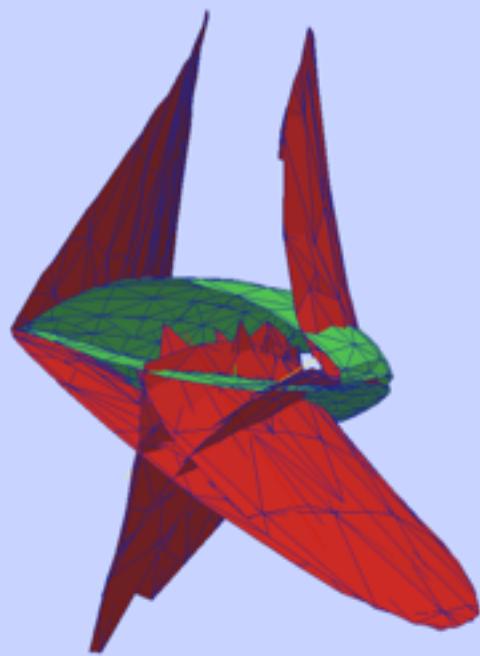
Lagrangian space

Hidding,
Shandarin,
van de Weygaert
2014

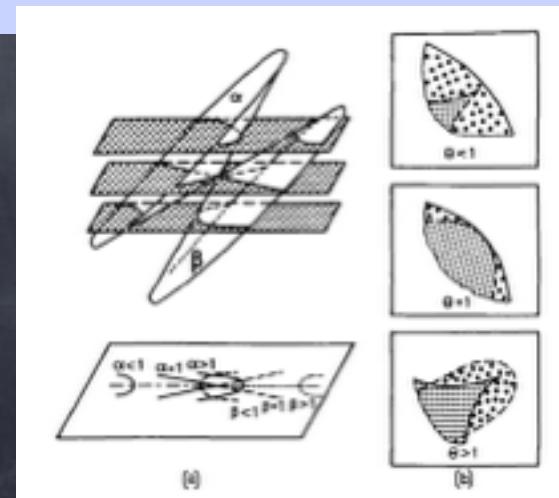
Number of streams
in Eulerian space



D_4 singularity in 3D space

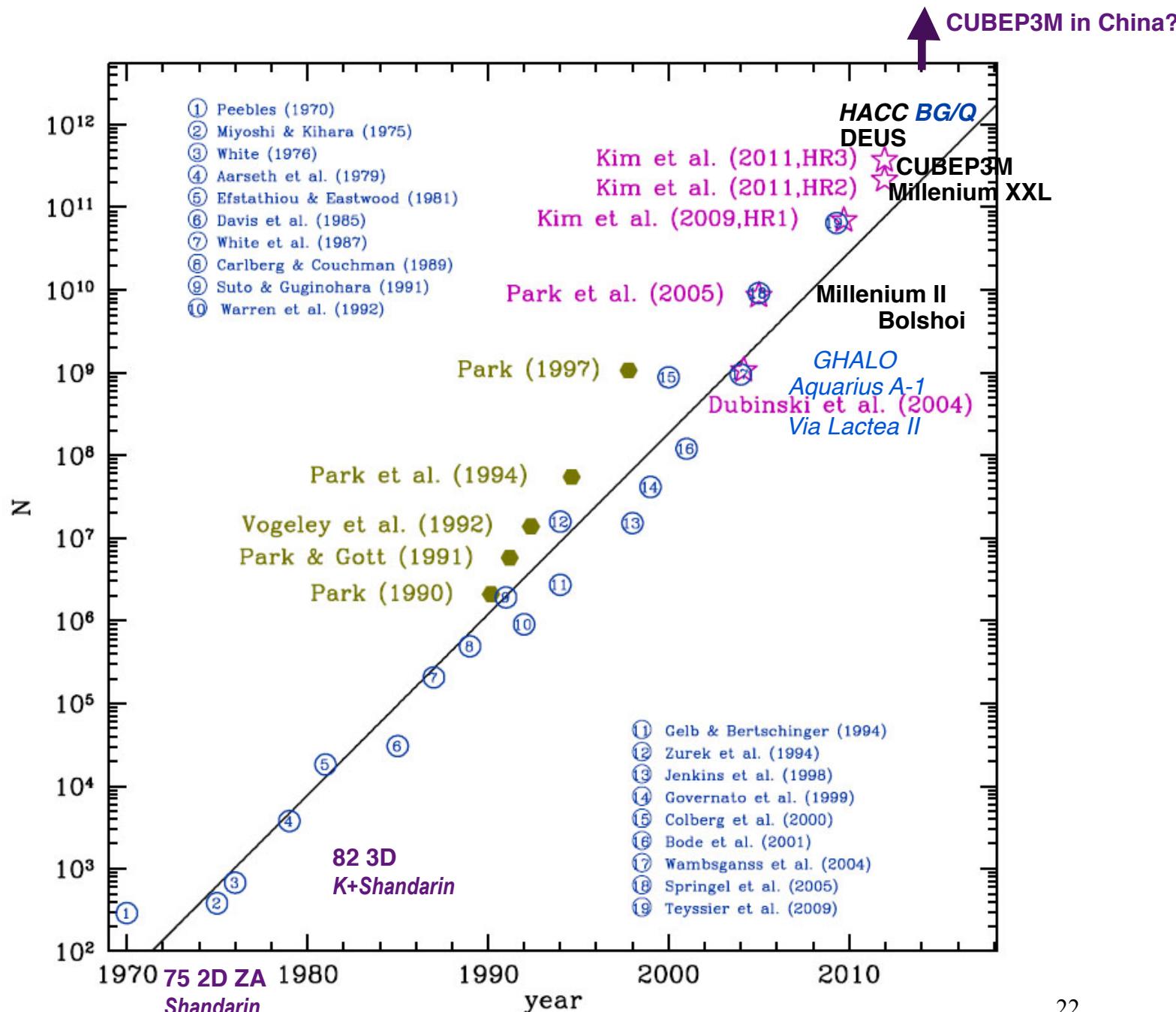


Hidding, Shandarin,
van de Weygaert
in preparation



Arnold,
Shandarin,
Zeldovich 1982

2D



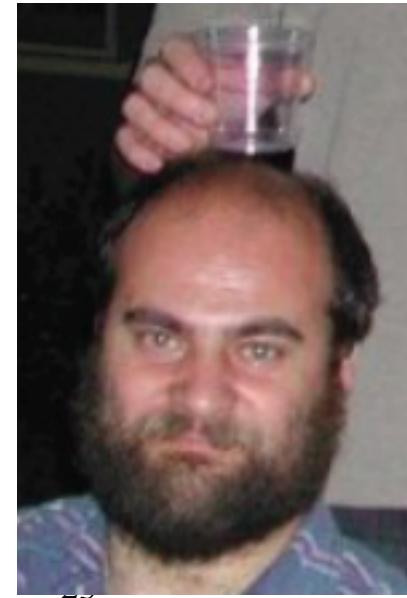
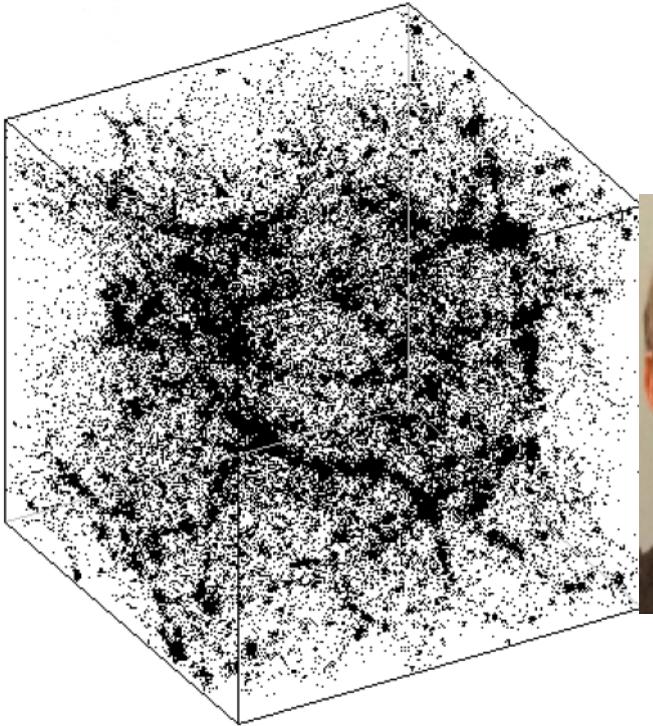
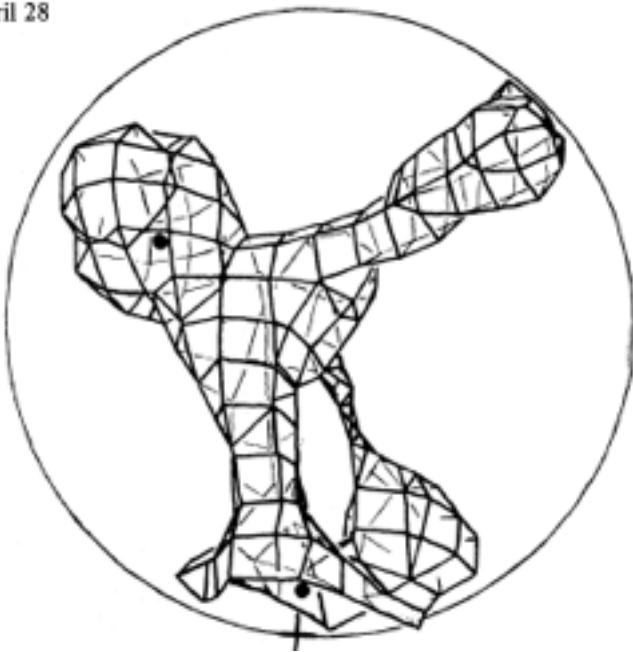
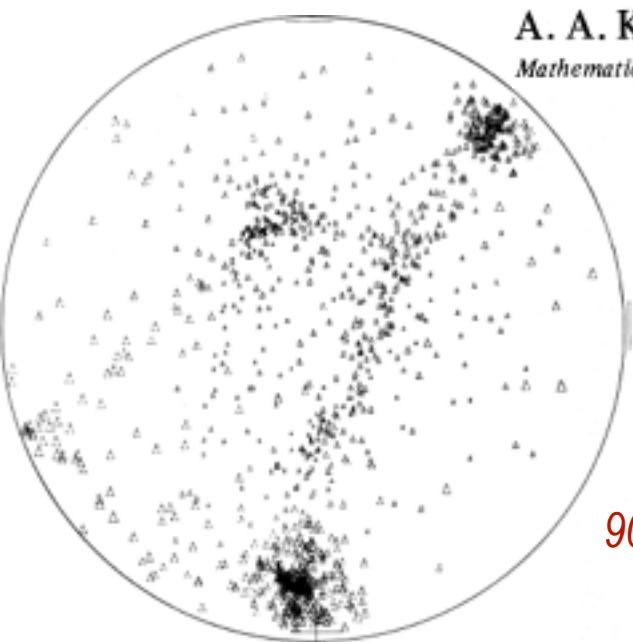
A. A. Klypin and S. F. Shandarin *The Keldysh Institute of Applied Mathematics, Academy of Sciences of USSR, Miusskaja Sq. 4, Moscow 125047, USSR*
Received 1982 November 15; in original form 1982 April 28

3D numerical model of the Universe

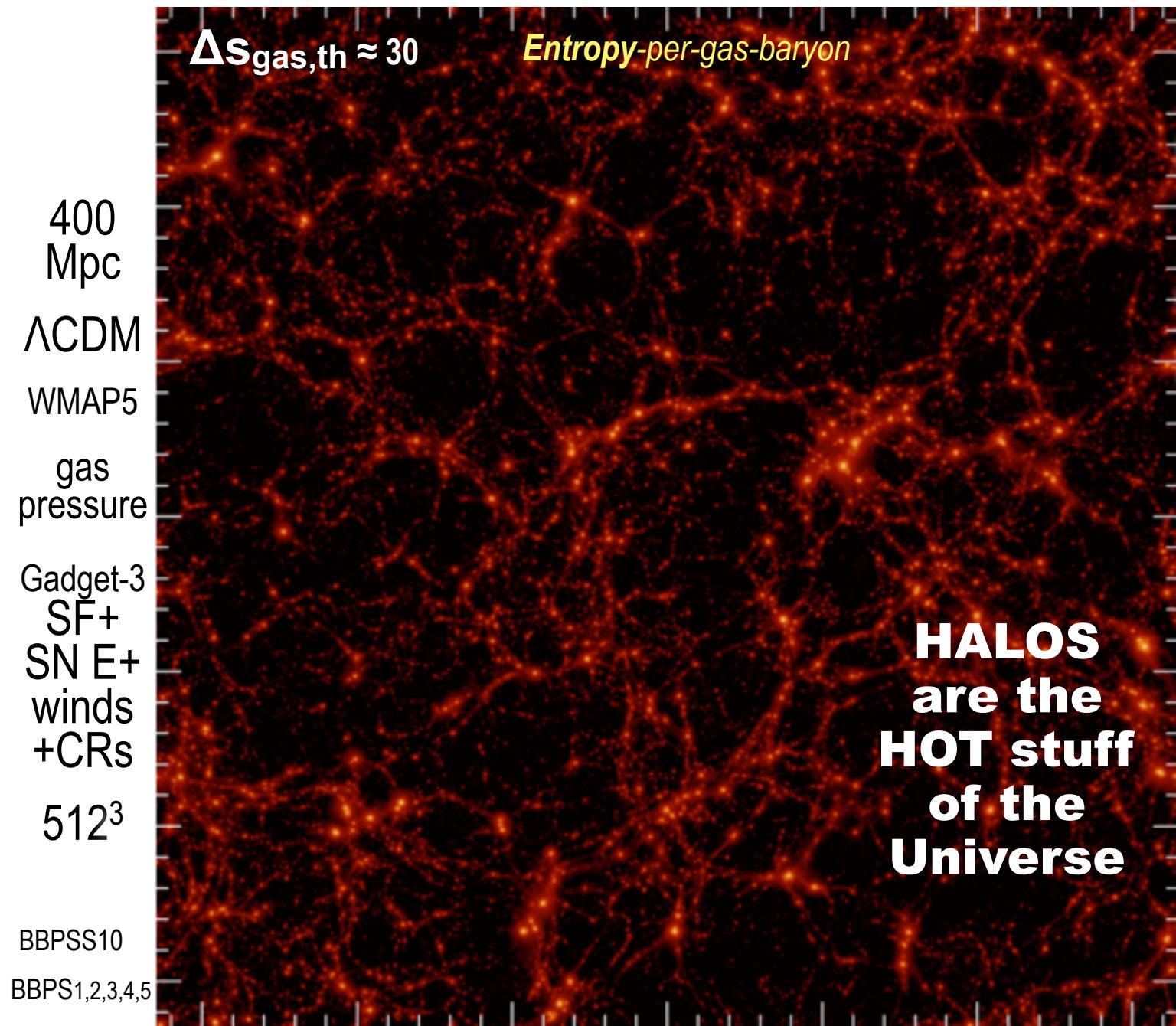
Klypin's vintage 82 $160 h^{-1} \text{Mpc}$ box $32^3 h\text{DM}$

*It is possible to recognize
some webs connecting
these 'clusters of galaxies'*

90s Klypin to CITA, 'the west is best'



Klypin's vintage 93 $50 h^{-1} \text{Mpc}$ box $128^3 s\text{CDM} = \text{BKP98}$ web workhorse, Couchman's 128^3 for BM91-96



Secondary Anisotropies
(tSZ, kSZ, WL, reion, CIB; hydro)

$s_{b,\text{th}}(x,t)$

baryons get entangled in the cosmic web

Let there be HEAT

$\rho_g(x,t)$

$\rho_{\text{stars}}(x,t)$

$p_e(x,t)$

$I_v(x,t)$

$n_{\text{dust}}(x,t)$

non-Gaussian CDM

entanglement
 $\rho_{\text{dm}}(x,t)$

Peak-patches = “hot” halos

B+Myers 91-96; BBKS 83-86

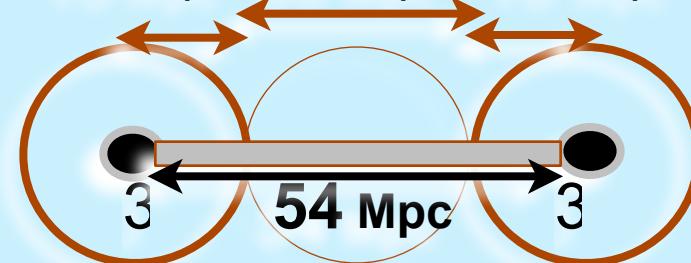
The Cosmic Web

B+Kofman+Pogosyan 96-99

“Molecular” Picture of LSS Filaments & Membranes

HALOS are dynamically HOT, the hierarchical standard model, Λ CDM,
=> scale space (3+1D => 4+1D)
adaptive coarse-grain Zeldovich flows of Lagrangian peak-patches agree with N-body Eulerian halo simulations => fast mock surveys

15 Mpc 30 Mpc 15 Mpc



clusters
 $z \sim 0-1+$
 $\sim 10^{15} M_{\text{sun}}$

1 Mpc

2 Mpc
3.6 Mpc

1 Mpc

galaxies
 $z \sim 2-5$
 $\sim 10^{11.5} M_{\text{sun}}$

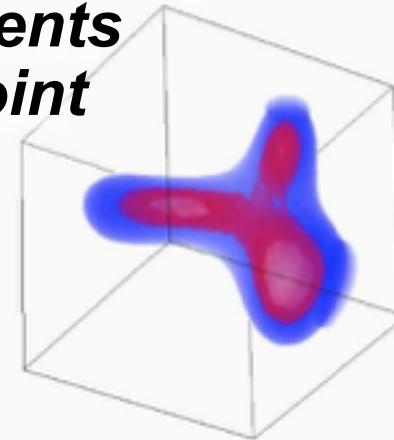
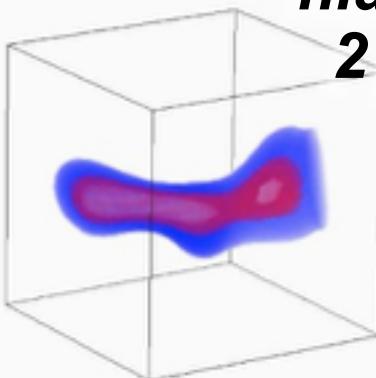
marriage of halos & Zeldovich
hot dynamics => $e^i_j(r_{pk}, t, R_{pk})$



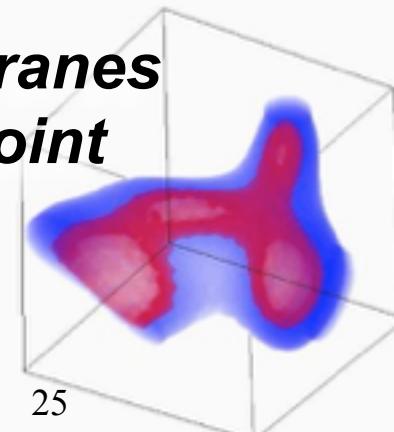
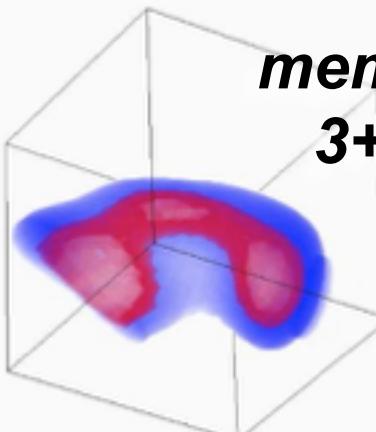
cool dynamics => $s^i(r_{pk}, t, R_{pk})$



stacked (constrained) density fields
filaments
2 point



membranes
3+ point



25

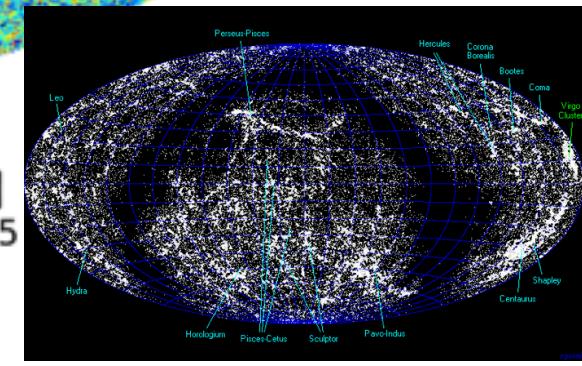
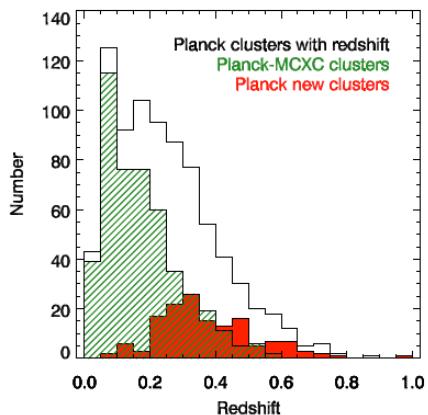
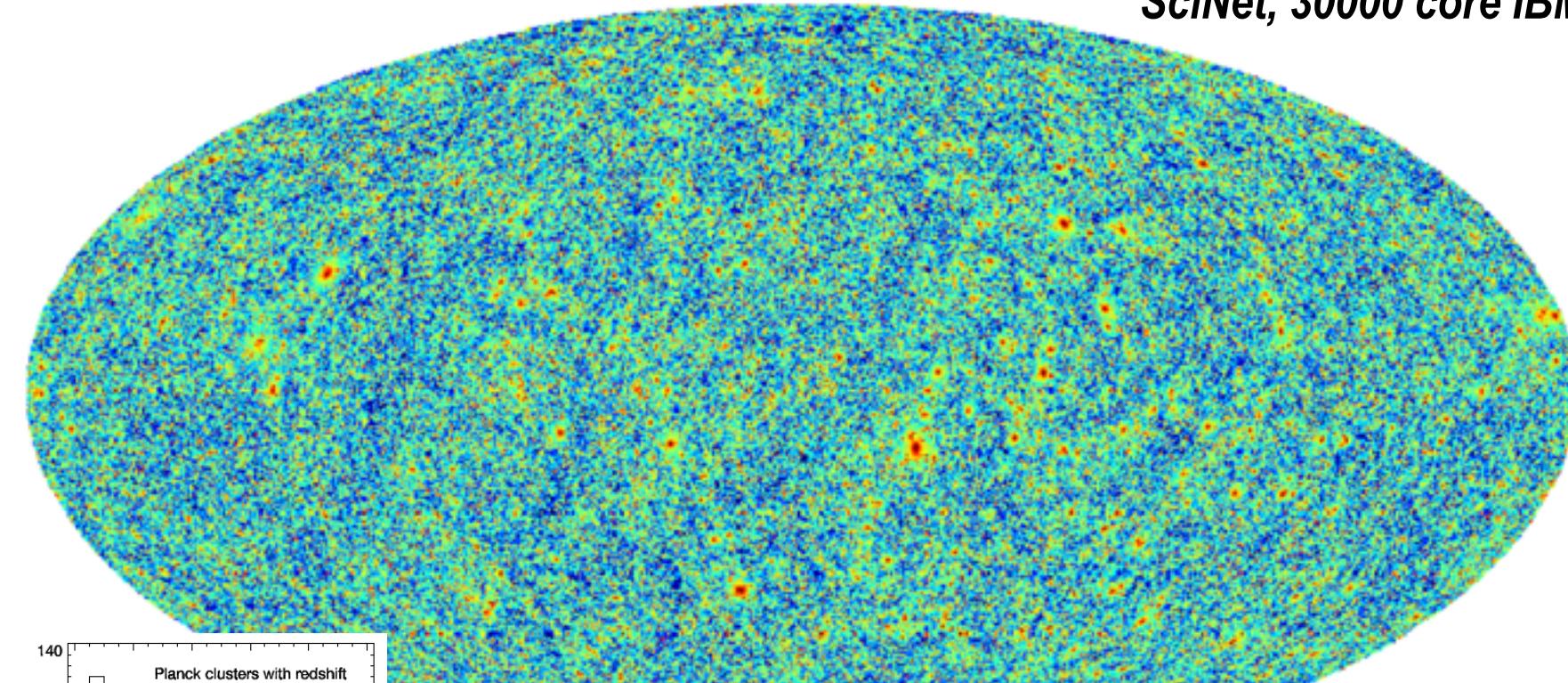
the Cosmic Web of Clusters, seen thru Compton cooling of high pressure electrons by the CMB

tSZ effect

Lightcone Simulation of 35000 Clusters > 1.5×10^{13} M_{sun} to z=0.5 in projected pressure

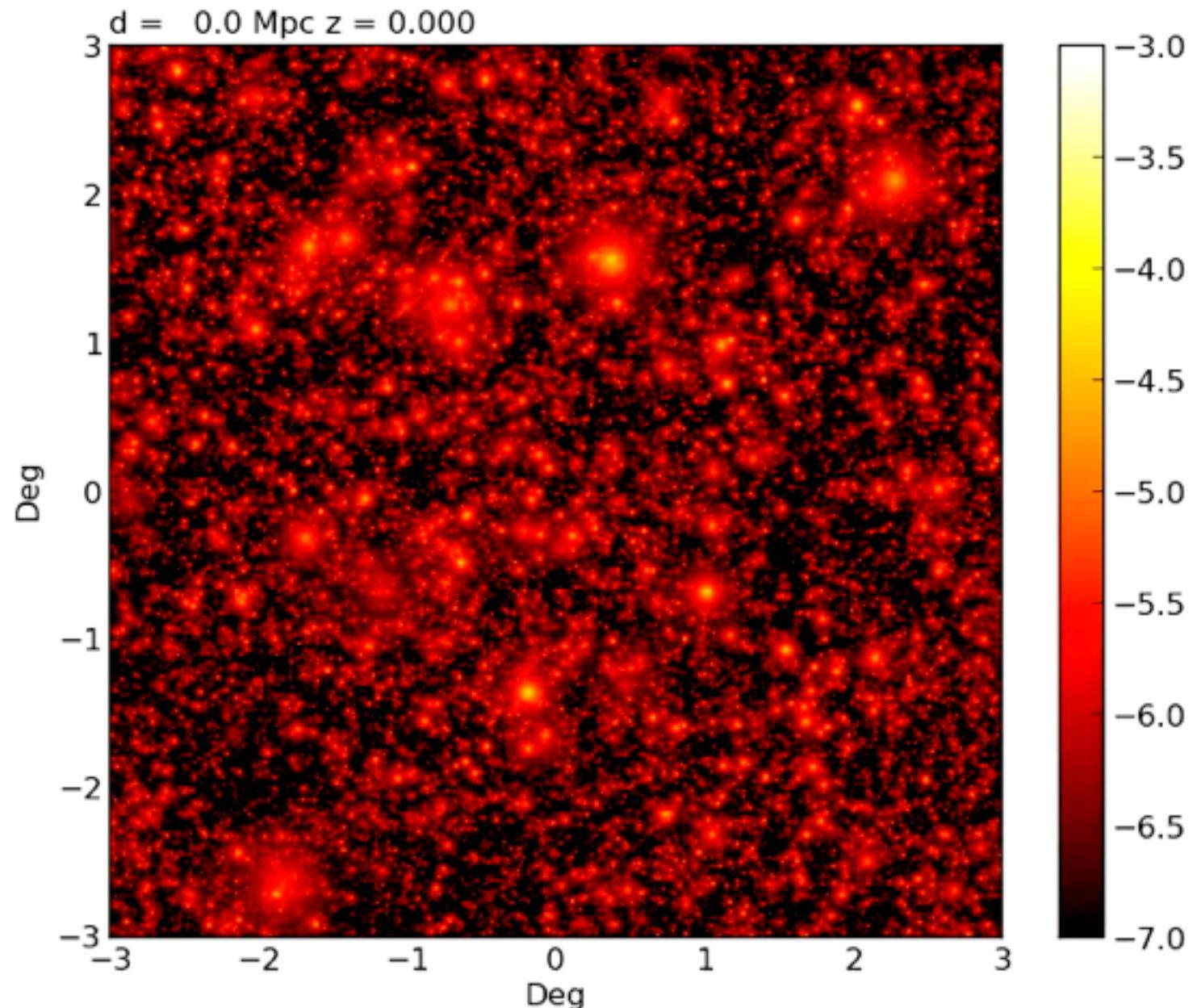
Alvarez, Bond, Hajian, Stein, Battaglia, Emberson,..2014

**1.5 hours on 256 cores on
SciNet, 30000 core IBM GPC**



Mocking Heaven: lightcone simulation for Λ CDM 36 sq deg to z=2

Planck all-sky tSZ mock 1.5 hours on 256 cores on SciNet, 30000 core IBM GPC



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SuperWeb of ultra-Ultra Large Scale Structure of the Universe

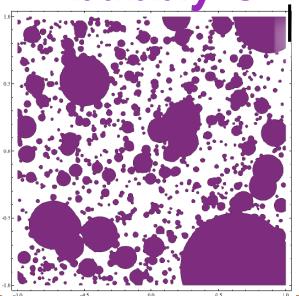
Horizons: the ultimate-speed constraint on light & information

a highly strained & stressed state in the universe-at-large (very, very), randomly simple in our Hubble patch, and highly entangled on small to medium scales

Universe or Multiverse?
Edited by Bernard Carr



quantum tunnels
= bubbly-U

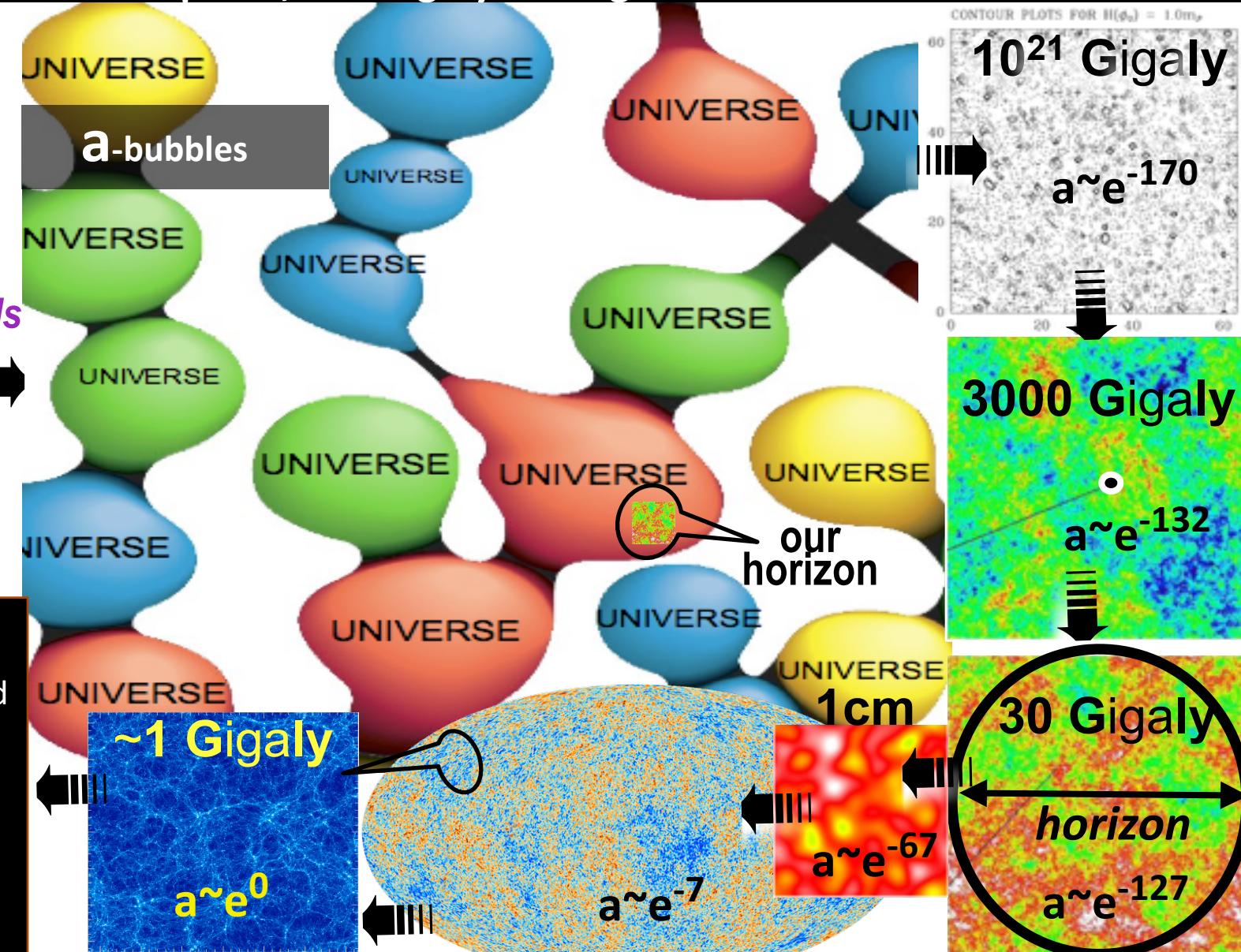


END

a future DE-Void



$a \sim e^{+++}$



END