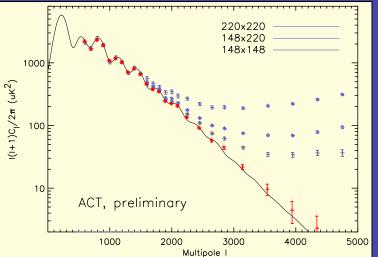
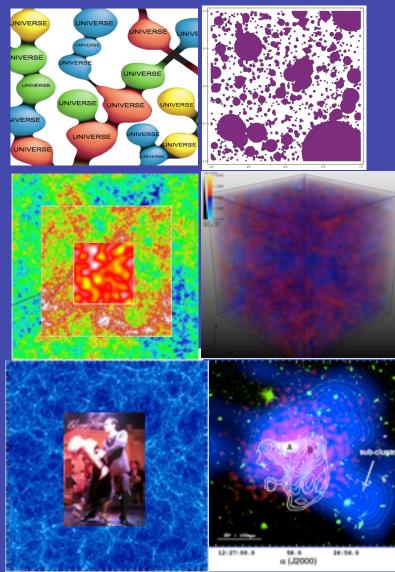
Dick Bond CIFAR@CITA with CITA aka Cosmic Information Theory & Analysis Probing the Cosmic Theory of Early & Late Universe Physics: from Simplicity to Complexity

IT from BIT from BITs in IT information quantity = entropy Shannon 1948 information quality = IQ essence info& primarily-earlyU =Bond@IAP 12.09.28 info& primarily-clusters/SZ =Bond@IAS 12.10.04 info& primarily-primaryCMB =Bond@APC 12.10.30 **Damping Tail & Recombination History** new ACT12+SPT12 + Planck13 to come



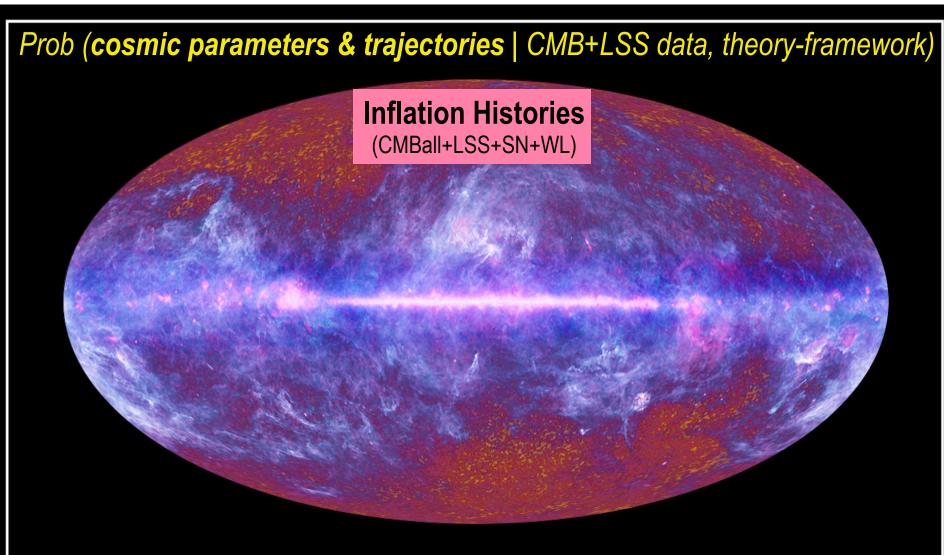


the coherent & the entropic, in all its forms, from the ultra-early-U to Now to the ultra-late-U

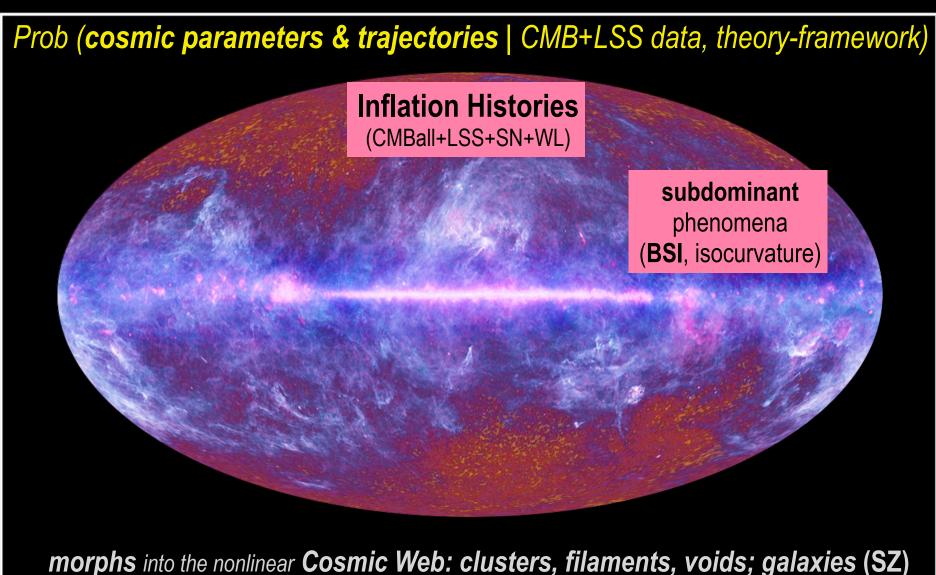
Thursday, 1 November, 12

Prob (cosmic parameters & trajectories | CMB+LSS data, theory-framework)

morphs into the nonlinear Cosmic Web: clusters, filaments, voids; galaxies (SZ) gastrophysical simulations with feedback from AGN / starbursts / SN .. confront CMB+LSS data The Planck one-year all-sky survey cesa (c) ESA, HFI and LFI consortia, July 2010 Beyond the standard model: tilted ACDM + x



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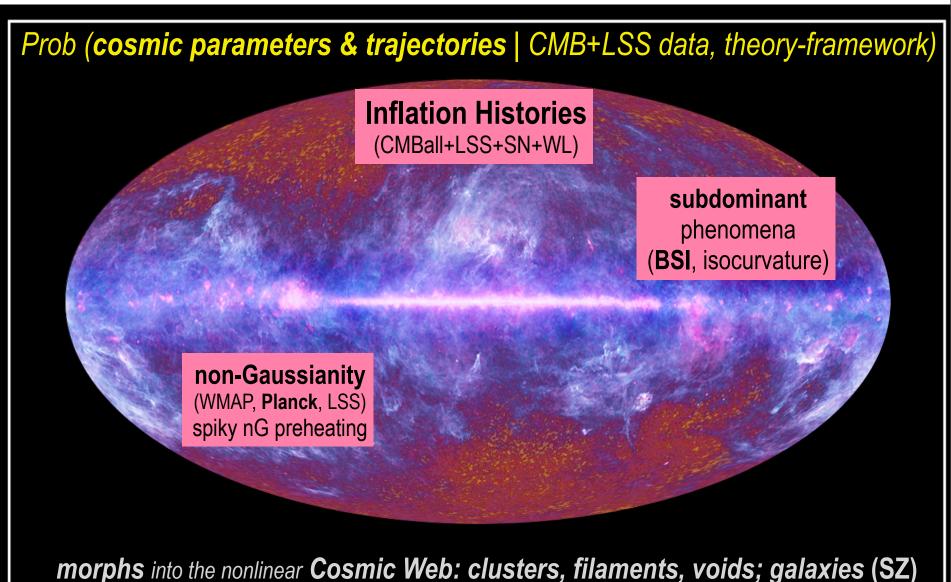


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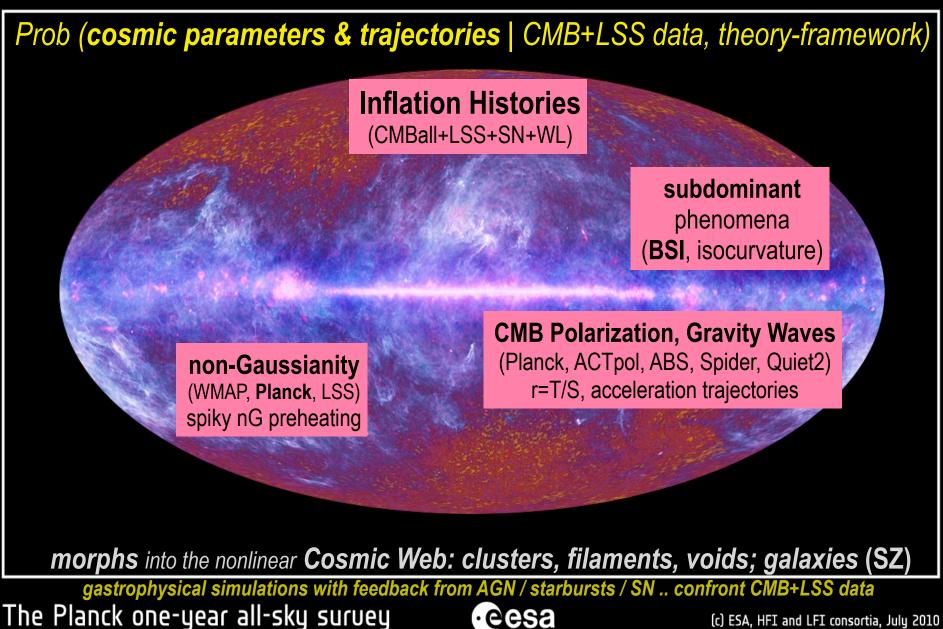
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 The Planck one-year all-sky survey

 Beyond the standard model: tilted ΛCDM + x



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> subdominant phenomena (**BSI**, isocurvature)

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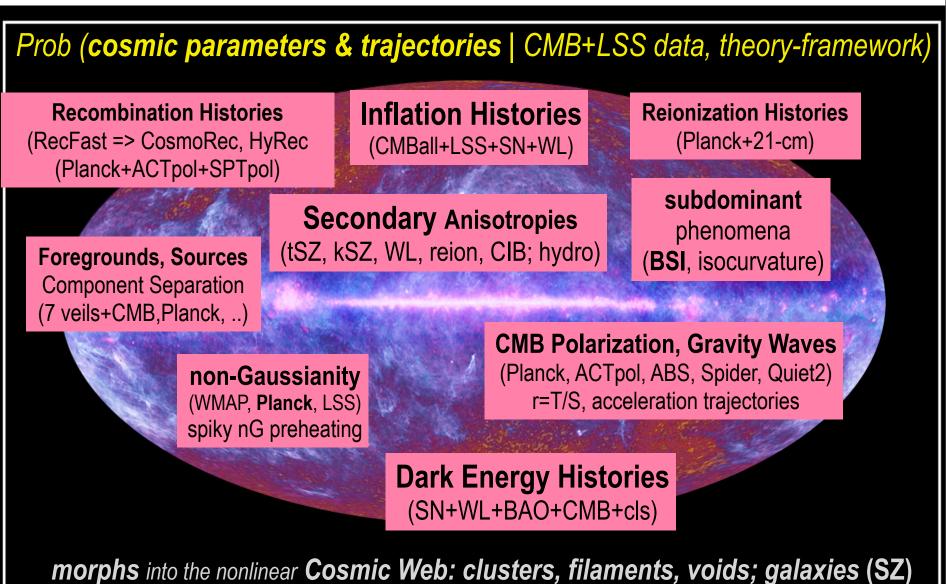
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Dark Energy Histories (SN+WL+BAO+CMB+cls)

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The Planck one-year all-sky survey cesa (c) ESA, HFI and LFI consortia, July 2010 Beyond the standard model: tilted ACDM + x

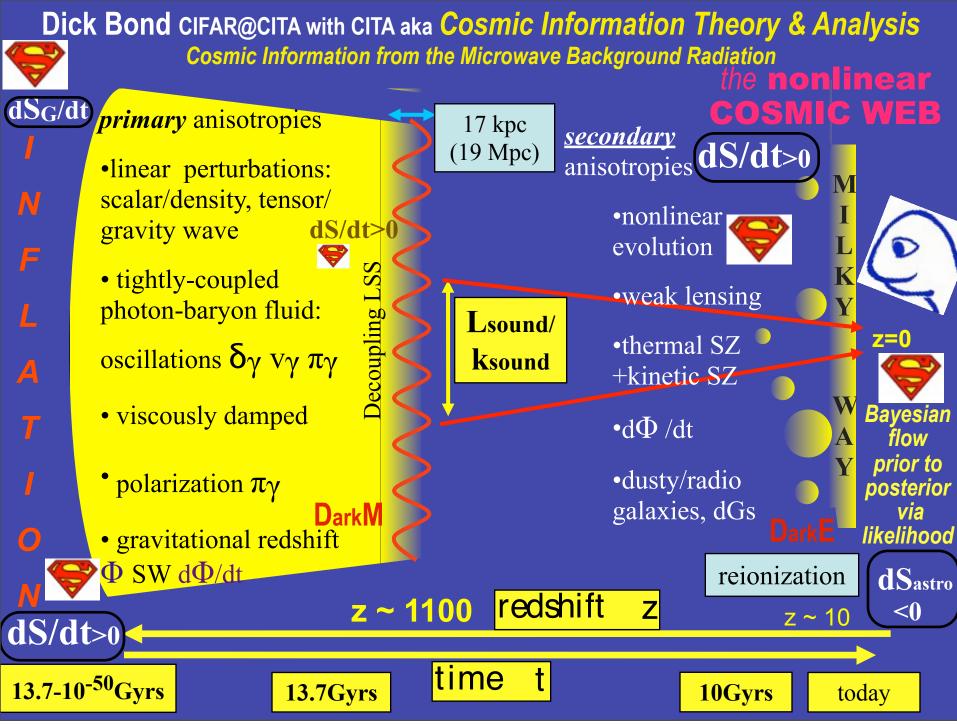


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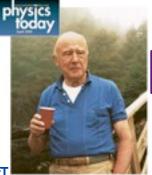
 (c) ESA, HFI and LFI consortia, July 2010

 Beyond the standard model: tilted ΛCDM + x



"Now I am in the grip of a new vision, that Everything Is Information. The more I have pondered the mystery of the quantum and our strange ability to comprehend this world in which we live, the more I see possible fundamental roles for logic and information as the bedrock of physical theory. ... I continue to search."

the medium is the message McLuhan 1964 UofT



the coherent and the entropic, in all its forms, from ultra-early-U to ultra-late-U

**SU,m+r ~10<sup>88.6</sup>** cf. **SG ~10<sup>121.9</sup>** asymptotic DE

Studying the Cosmic Tango en-TANGO-ment the dance of U=RUS

*Universe* =*System(s)*+*Reservoir* =Signal(s)+Residual *noise* =Effective Theory+*Hidden variables,* =*Data*+*Theory,* observer(s)+observed

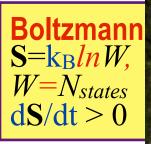
U=RUS ruled by (information) entropy in bits, entangled. *the fine grains in the coarse grains* 

entropy =<information-content> Quantity Shannon 1948

generalized parameter space {q} ~phase space

S<sub>f</sub>(D,T)=∫ dq P<sub>f</sub> In[P<sub>f</sub> <sup>-1</sup>]





equal a priori probability

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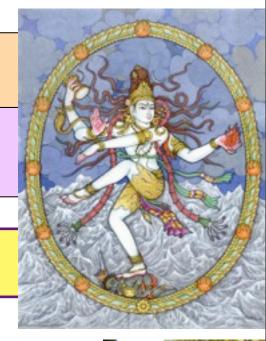
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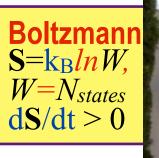
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A Long View of Particle Physics Frank Wilczek 2012, 25th Solvay:

Information as Foundation? There are, I think, significant hints that it should be. QITA Quantum Information Theory & Analysis

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Boltzmann

 $S = k_B ln W$ ,

 $W=N_{states}$ 

dS/dt > 0

**Bayes** measure

=>"d**S**f/dt<0"

Studying the Cosmic Tango en-TANGO-ment the dance of U=RUS

Glvi

our Cosmoticians' Agenda: Statistical Paths in Cosmic Theory & Data via the Bayesian chain drawing what we know of It from Its Bits

P(q|D,T) =P(D|q,T)P(q|T)P(T)/P(D|T) D=CMB,LSS,SN,..,Complexity, life T=baryon, dark matter, vacuum mass-energy densities,..., early & late inflation as low energy flows/trajectories on a (string) landscape

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 $S_f(D,T) = \int dq P_f \ln[P_f^{-1}]$ 

 $S_{fi}(D,T)=\int dq P_f \ln[P_f^{-1}P_i]$ 

relative Shannon entropy = - Kullback Leibler divergence  $P_f(q)$  probability density functional distribution function  $\Leftarrow$  quantum (von Neumann) S= -Tr  $\rho$  ln  $\rho$  density matrix

as System knowledge

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$$S_f(D,T)=\int dq P_f \ln[P_f^{-1}]$$

 $\sim - \ln \langle \rho^n \rangle_V / \langle \rho \rangle_V^n$ 

$$S_{fi}(D,T) = \int dq P_f \ln[P_f^{-1} P_i] = \langle \sigma_{fi} \rangle_f$$

*relative Shannon entropy* = - *Kullback Leibler divergence*   $P_f(q)$  probability density functional distribution function  $\leftarrow quantum$  (von Neumann) S= -Tr  $\rho \ln \rho$  density matrix  $-<\ln \rho >_{\rho}$ *relative RENYI entropy of order n* a concentration measure (1 is Shannon)

 $exp[-(n-1)S_{n,fi}(D,T)] = < exp[-(n-1)\sigma_{fi}] >_{f}$ 

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### **IQ**=information quality

**IQ~{minimal length messages/codes** | error tolerance} *Planck(E/T),* genetic code, recipes, axioms, algorithms, IC/BC/evolution eq<sup>n</sup>s

cat information\_overload.txt | grep fundamental | grep physics > exec\_summary.tex

filter, compress, reduce, marginalize



early U applications of "CITA" to cosmic-complexity

A the superhorizon measure problem & the Lambda-scape

☆ the emergence of the collective from the random! coherence from driven zero-point vacuum fluctuations ⇒ V inflaton, gravity waves; decohere

☆ let there be heat: entropy generation in preheating from the coherent inflaton (origin of all "matter")

P(q|D,T) =P(D|q,T)P(q|T)P(T)/P(D|T) D=CMB,LSS,SN,...,Complexity, life T=baryon, dark matter, vacuum mass-energy densities,..., early & late inflation as low energy flows/trajectories on a (string) landscape

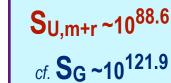
some non-early U applications of "CITA" to cosmic-complexity

information in nearly-Gaussian density/potential random fields of U,
 & in weakly and strongly non-linear fields. ergodic theorem & constrained fields

respatial coarse-grained CMB entropy & how we capture it

dark matter entropy, cluster & protocluster & cosmic web entropy MHD turbulence entropy with cooling & grain polarized emission - CMB fgnd

How Shannon info-entropy flows from CMB bolometer timestreams to marginalized cosmic parameters via Bayesian chains from prior to posterior. 1D & 2D & ... ΔS(q,DT) (cf. ACT10), q=r, w, n<sub>s</sub>, ...









S

info& primarily-earlyU=Bond@IAP 12.09.28 info& primarily-clusters/SZ=Bond@IAS 12.10.04 info& primarily-primaryCMB =Bond@APC 12.10.30 Damping Tail & Recombination History new ACT12+SPT12 + Planck13 to come

### some non-early U applications of "CITA" to cosmic-complexity

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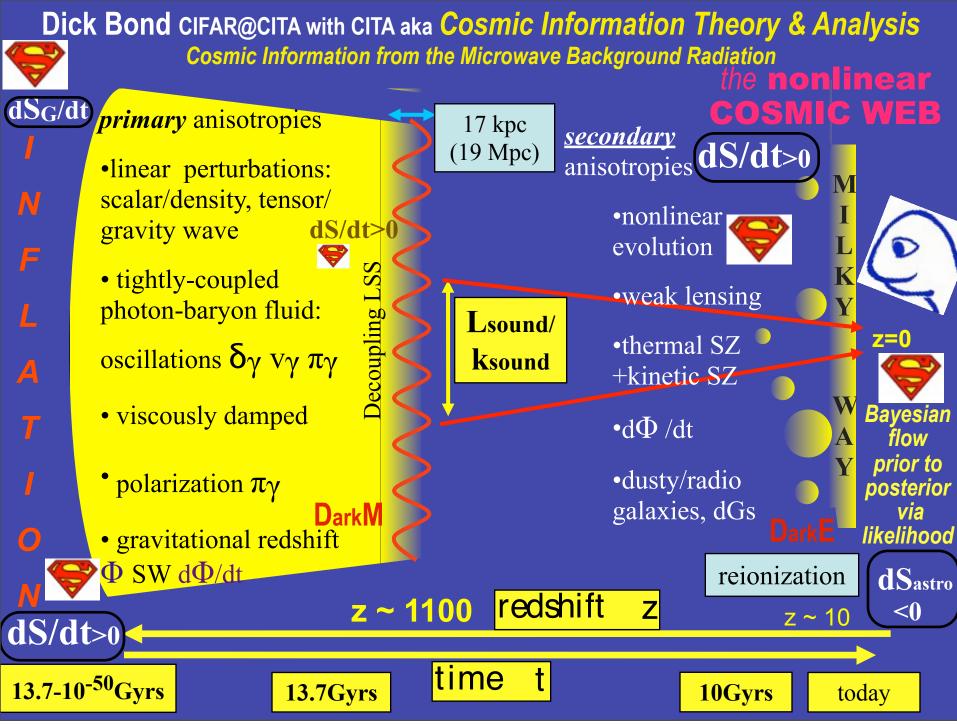
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Studying the

Cosmic



# the gatherers of cosmic information Cosmic Microwave Background + Large Scale Structure experimental probes

# then & now & then

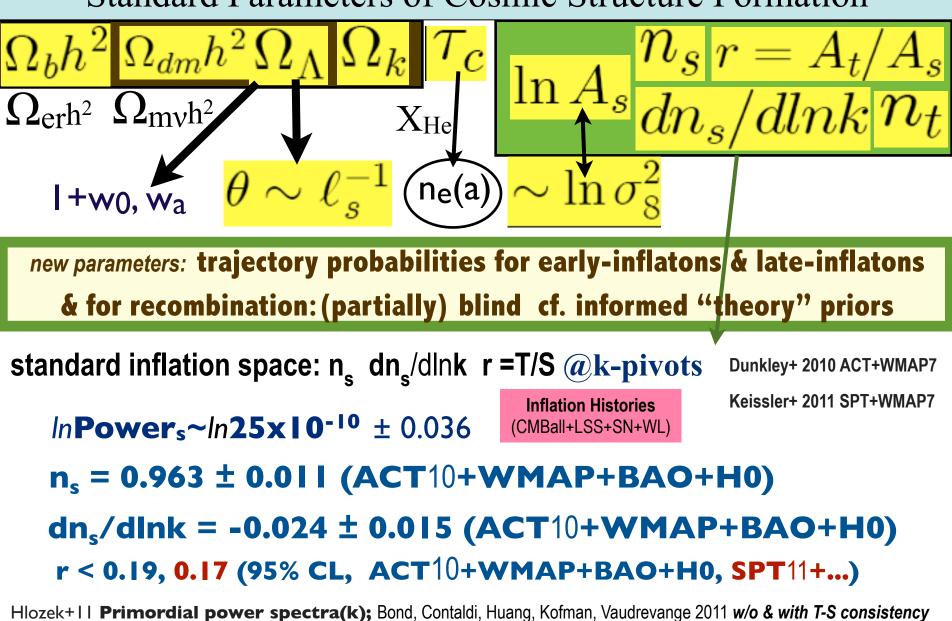
# 2012 cosmology => W/MAP9EXT 2013+ cosmology => PlanckEXT



EXT=many observatories & expts enabling the cosmology/astro ACT, SPT, Quiet, GBT, SSDS/BOSS, PanStarrs, ... ⊂ EXTi cosmology: n<sub>s</sub>(k), GW r(k), nonG f<sub>NL</sub>++, p<sub>de</sub>(t), m<sub>v</sub>, strings, isocurvature,... n<sub>e</sub>(t) ACTpol, SPTpol, ABS, Spider, Quiet-90, EBEX, Keck, GBT, PanStarrs,

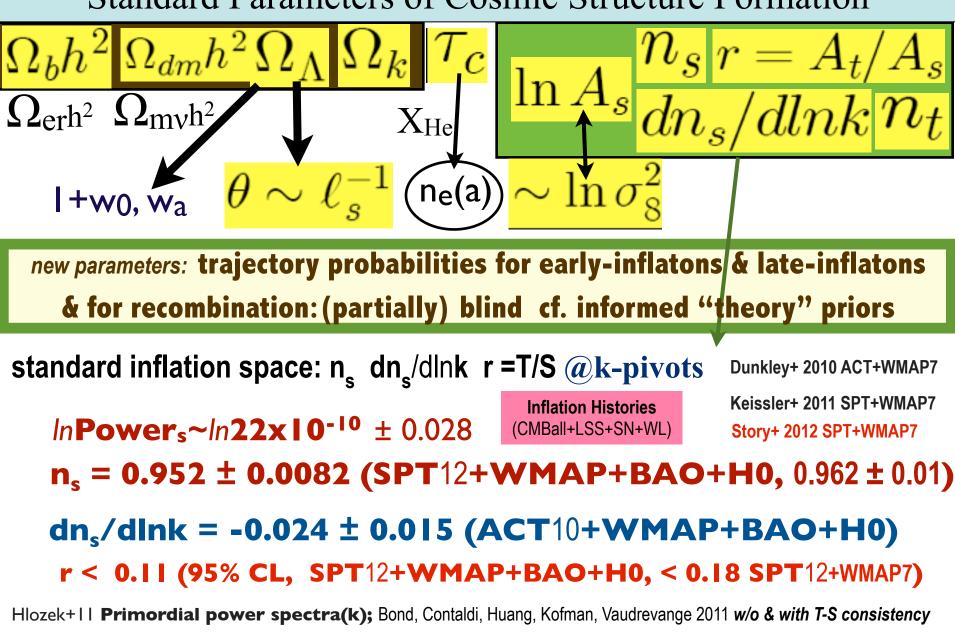
DES, HSC, *CHIME,* eRosita, **CCAT**, LSST, *EUCLID*, ... CEXTf

Standard Parameters of Cosmic Structure Formation



ACT12 final spectra now, final params in ~weeks, 1500 sq deg, ~1000 for params

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CITA = Cosmic Information Theory & Analysis: IT from BIT, from BITs in IT, Studying the Cosmic Tango en-TANGO-ment Universe=System+Res=Data+Theory =Signal(s)+noise=EFT+Hidden variables

we compress the Petabit++ observed cosmic info into a precious few bits encoding 6+ parameters of the Minimal Cosmic Standard model (tilted ACDM)

WMAP: 1.15 Tbits in 9yrs, cf. MyLifeBits, Gordon Bell, 1.28 Tbits in 9yrs, Planck 36 Tbits, ACT 304 Tbits. Radically Compress to high quality Bits. Terabit=10<sup>12</sup>bits=125 GigaBytes.

Shannon entropy difference  $\Delta S_{fi}(q, DT) = \int dq P_f \ln P_f^{-1} - \int dq P_i \ln P_i^{-1}$ 

a new figure of merit for experiments, <InVOLUME<sub>ps</sub>> ~ posterior Shannon entropy: how the (radically compressed) one-dimensional entropy of cosmic parameters, the high quality bits we quest, did/will change as the experiments became/become more & more precise:

CMB@CITA: Boomerang, Acbar, CBI1,2, WMAP, Planck, ACT, Spider, Blast, & ACTpol, ABS, QUIET2; GBT-Mustang2, CARMA/SZA, SCUBA2, ALMA, CCAT. CMB@CIFAR: these + APEX, SPT, SPTpol, EBEX

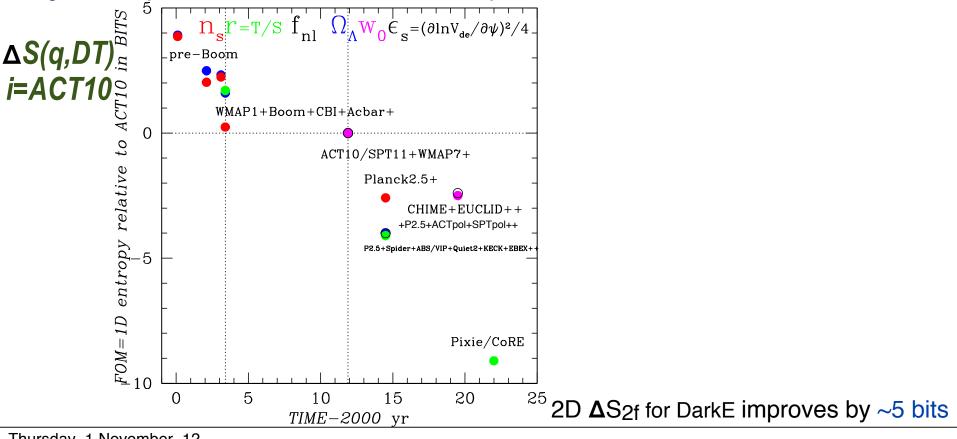
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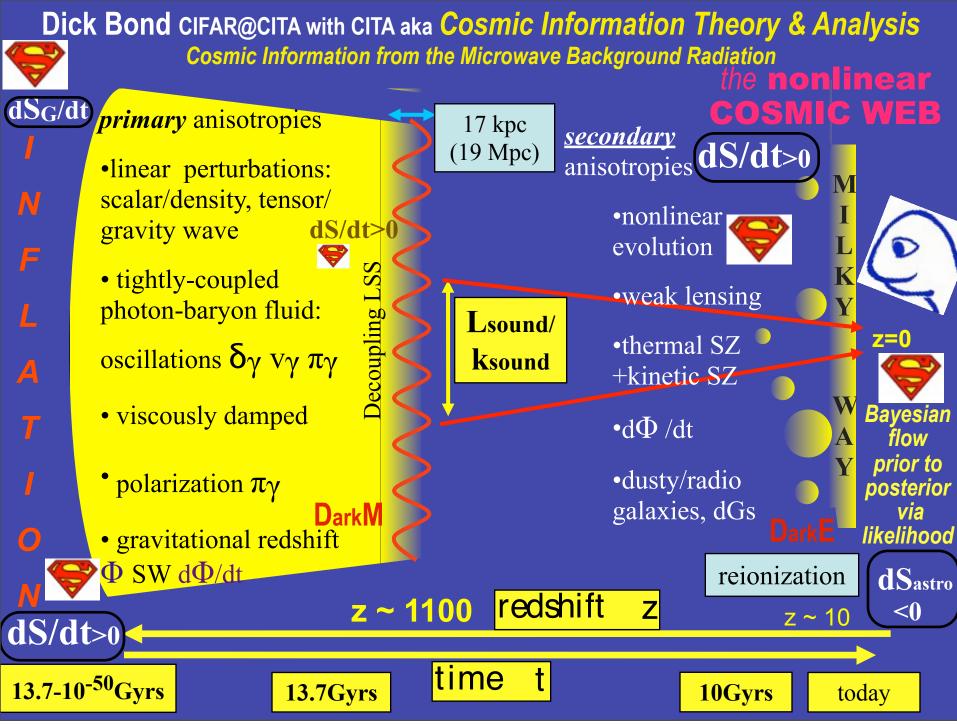
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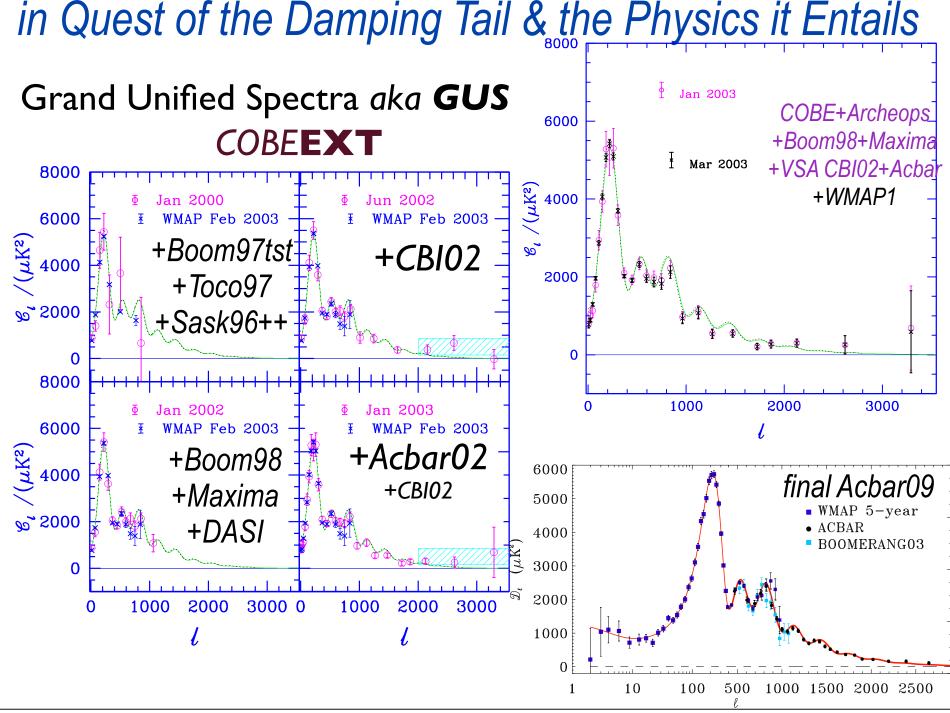
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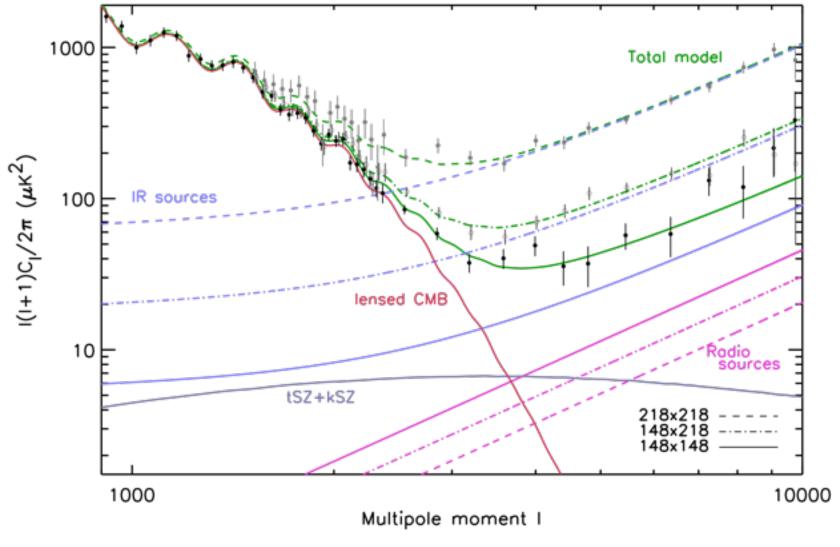
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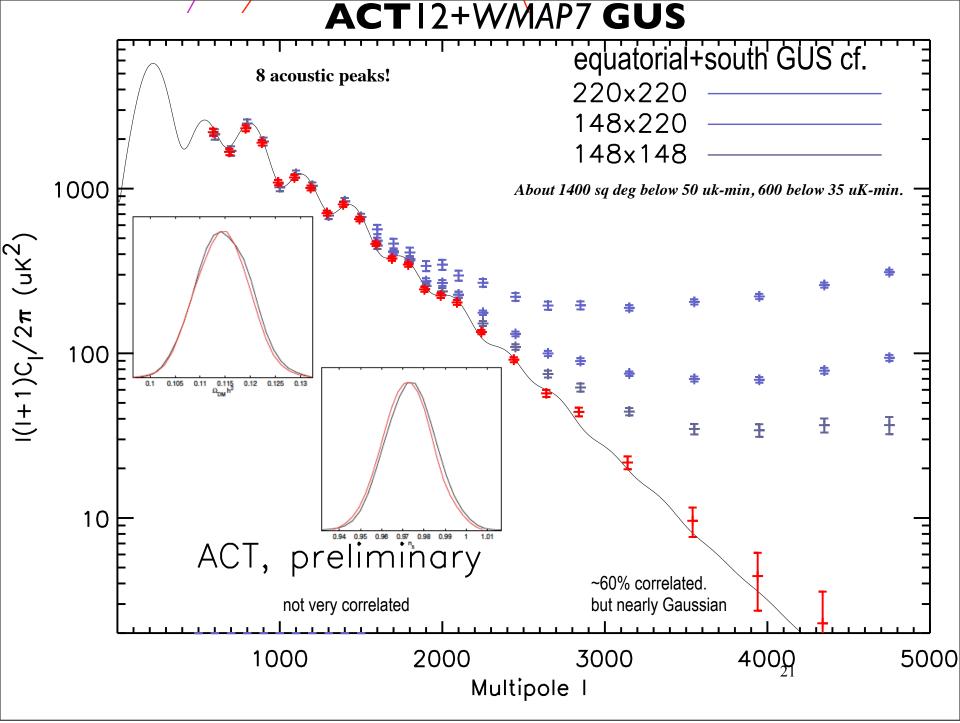


Thursday, 1 November, 12

primordial (lensed) CMB + veils, the veils = radio sources, the ClB, tSZ and kSZ (& Milky Way dust and synchrotron at lower multipoles)

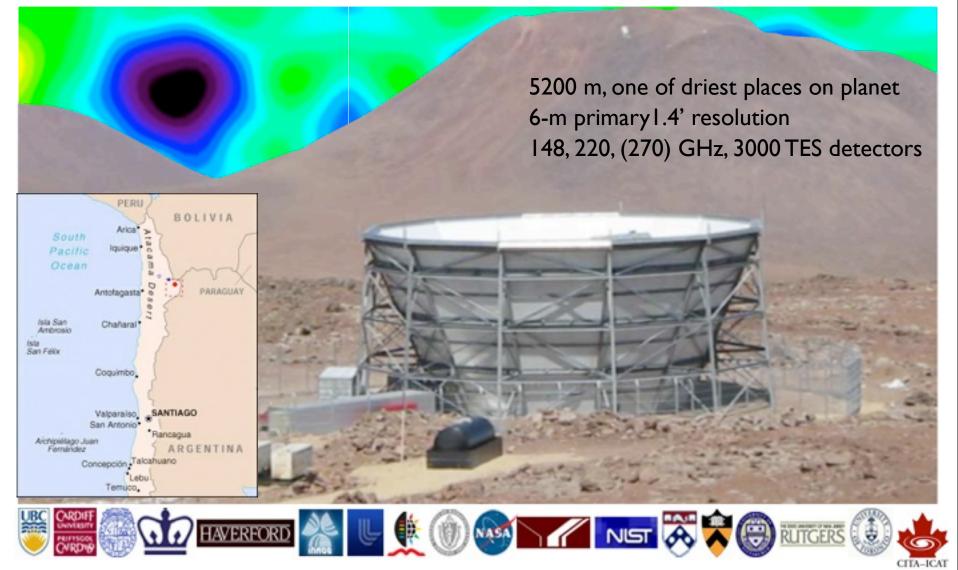


Dunkley+.2010



Thursday, 1 November, 12

# **Cosmology From 17,000 Feet:** Results From the Atacama Cosmology Telescope



CMB@CITA: Boomerang, Acbar, CBI1,2, WMAP, Planck, ACT, Spider, Blast, & ACTpol, ABS, QUIET2; GBT-Mustang2, CARMA/SZA, SCUBA2, ALMA, CCAT. CMB@CIFAR: these + APEX, SPT, SPTpol, EBEX

V.Acquaviva <sup>1,2</sup> R. Dunner<sup>4</sup> P.Ade<sup>3</sup> T. Essinger-Hileman<sup>6</sup> P.Aguirre<sup>4</sup> R.P. Fisher<sup>6</sup> M. Amiri<sup>5</sup> I.W. Fowler<sup>6</sup> J. Appel<sup>6</sup> A. Hajian <sup>6,8</sup> E. Battistelli 7,5 M. Halpern <sup>5</sup> N. Battaglia<sup>8</sup> M. Hasselfield <sup>5</sup> J. R. Bond<sup>8</sup> C. Hernandez-Monteagudo <sup>13,2</sup> B. Brown <sup>9</sup> G. Hilton 11 B. Burger <sup>5</sup> M. Hilton 14, 15 I. Chervenak<sup>10</sup> <sup>\*</sup> A. D. Hincks <sup>6,8</sup> S. Das <sup>29,6,1</sup> R. Hlozek <sup>12, 1</sup> M. Devlin<sup>2</sup> K. Huffenberger<sup>16,6</sup> S. Dicker<sup>2</sup> D. Hughes<sup>17</sup> W. B. Doriese <sup>11</sup> J. P. Hughes<sup>18</sup> I. Dunkley 12,6,1 <sup>1</sup> Princeton University Astrophysics (USA) <sup>2</sup> University of Pennsylvania (USA) <sup>3</sup> Cardiff University (UK) <sup>4</sup> Pontifica Universidad Catolica de Chile (Chile) <sup>5</sup> University of British Columbia (Canada) <sup>6</sup> Princeton University Physics (USA) <sup>7</sup> University of Rome "La Sapienza" (Italy) \* <sup>8</sup> CITA, University of Toronto (Canada) <sup>9</sup> University of Pittsburgh (USA) <sup>10</sup> NASA Goddard Space Flight Center (USA) <sup>11</sup> NIST Boulder (USA) <sup>12</sup> Oxford University (UK) <sup>13</sup> Max Planck Institut fur Astrophysik (Germany) <sup>14</sup> University of KwaZulu-Natal (South Africa)

L. Infante <sup>4</sup> K.D. Irwin <sup>11</sup> N. Jarosik <sup>6</sup> R. Jimenez <sup>19</sup> J.B. Juin <sup>4</sup> M. Kaul <sup>2</sup> J. Klein <sup>2</sup> A. Kosowsky <sup>9</sup> J.M. Lau <sup>20,6</sup> M. Limon <sup>21</sup> Y.T. Lin <sup>22,1,4</sup> R. Lupton <sup>1</sup> T.A. Marriage <sup>1,6</sup> D. Marsden <sup>2</sup>

P. Mauskopf <sup>3</sup> F. Menanteau <sup>18</sup> K. Moodley <sup>14</sup> H. Moseley <sup>10</sup> \* B. Netterfield <sup>24</sup> M.D. Niemack <sup>11,6</sup> \* M.R. Nolta <sup>8</sup> L.A. Page (PI) <sup>6</sup> L. Parker <sup>6</sup> B. Partridge <sup>25</sup> H. Quintana <sup>4</sup> B. Reid <sup>19,1</sup> N. Sehgal <sup>20,18</sup>

K. Martocci<sup>23,6</sup>

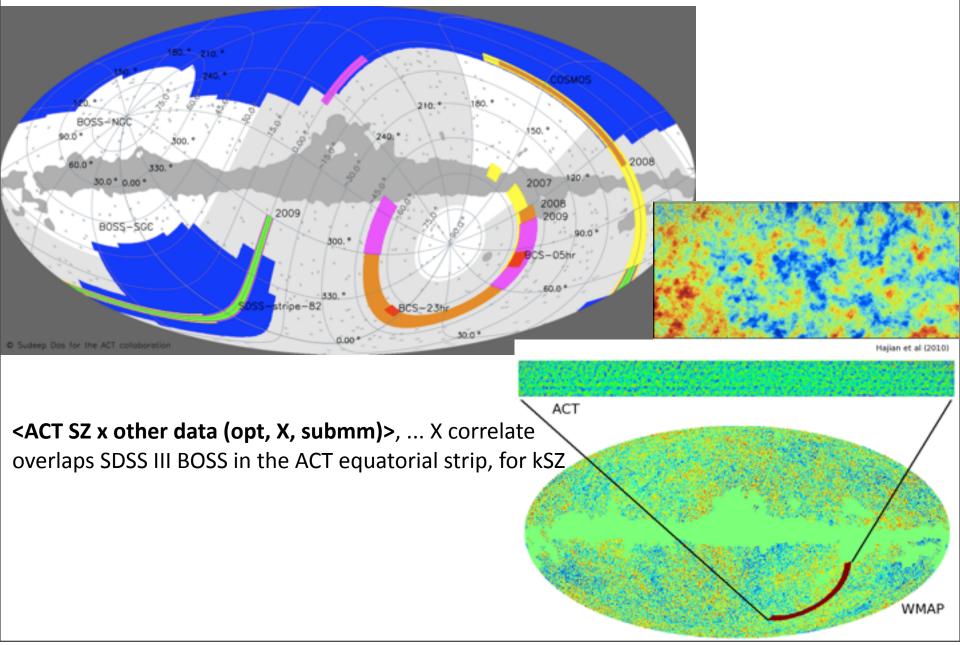
J. Sievers <sup>8,6</sup>
D. Spergel <sup>1</sup>
S.T. Staggs <sup>6</sup>
O. Stryzak <sup>6</sup>
D. Swetz <sup>2</sup>
E. Switzer <sup>23,6,8</sup>
R. Thornton <sup>26,2</sup>
H. Trac <sup>27,1</sup>
C. Tucker <sup>3</sup>
L. Verde <sup>19</sup>
R. Warne <sup>14</sup>
G. Wilson <sup>28</sup>
E. Wollack <sup>10</sup>
Y. Zhao <sup>6</sup>

- <sup>15</sup> South African Astronomical Observatory
  <sup>16</sup> University of Miami (USA)
  <sup>17</sup> INAOE (Mexico)
  <sup>18</sup> Rutgers (USA)
  <sup>19</sup> Institute de Ciencies de L'Espai (Spain)
  <sup>20</sup> KIPAC, Stanford (USA)
  <sup>21</sup> Columbia University (USA)
  <sup>22</sup> IPMU (Japan)
  <sup>23</sup> KICP, Chicago (USA)
  \* <sup>24</sup> University of Toronto (Canada)
  <sup>25</sup> Haverford College (USA)
  <sup>26</sup> West Chester University of Pennsylvania (USA)
  <sup>27</sup> Harvard-Smithsonian CfA (USA)
  <sup>28</sup> University of Massachusetts, Amherst (USA)
- <sup>29</sup> BCCP UC Berkeley and LBL (USA)

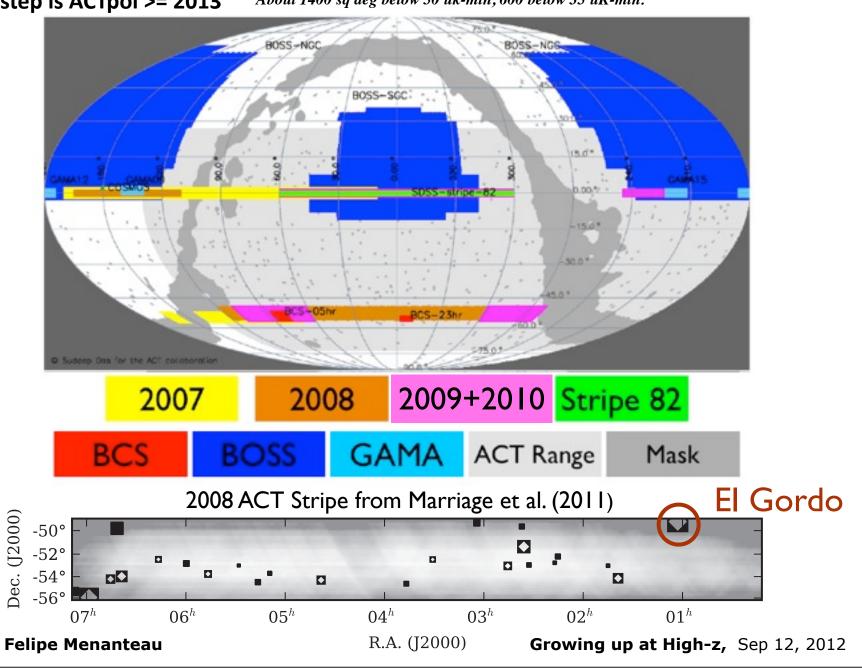


#### end observing 2011: ACT completed 3 full seasons, over ~1300 deg<sup>2</sup>, maps@CITA. next step is ACTpol

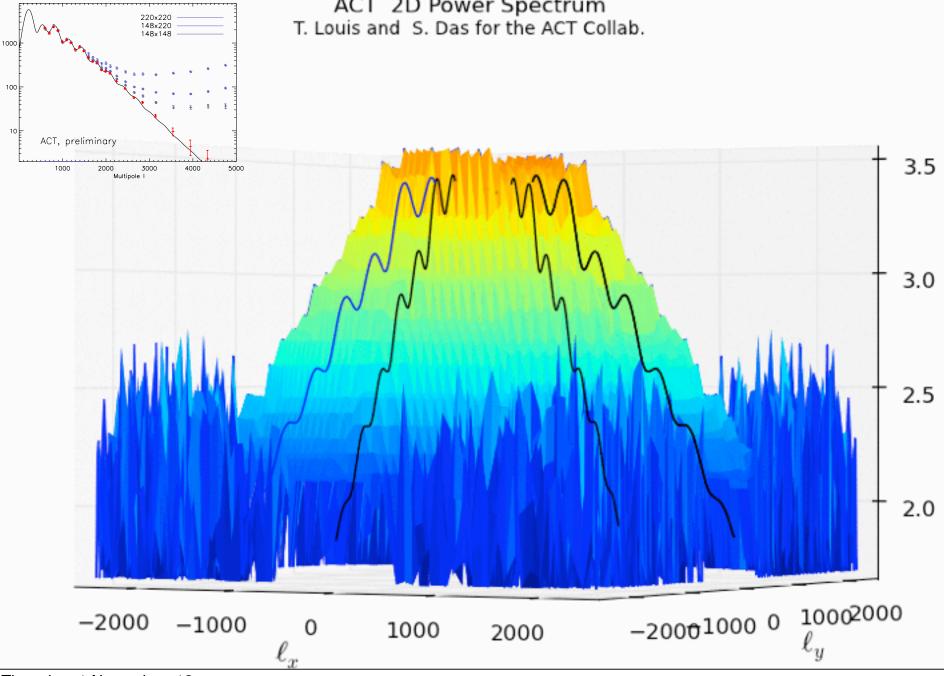
About 1400 sq deg below 50 uk-min, 600 below 35 uK-min.

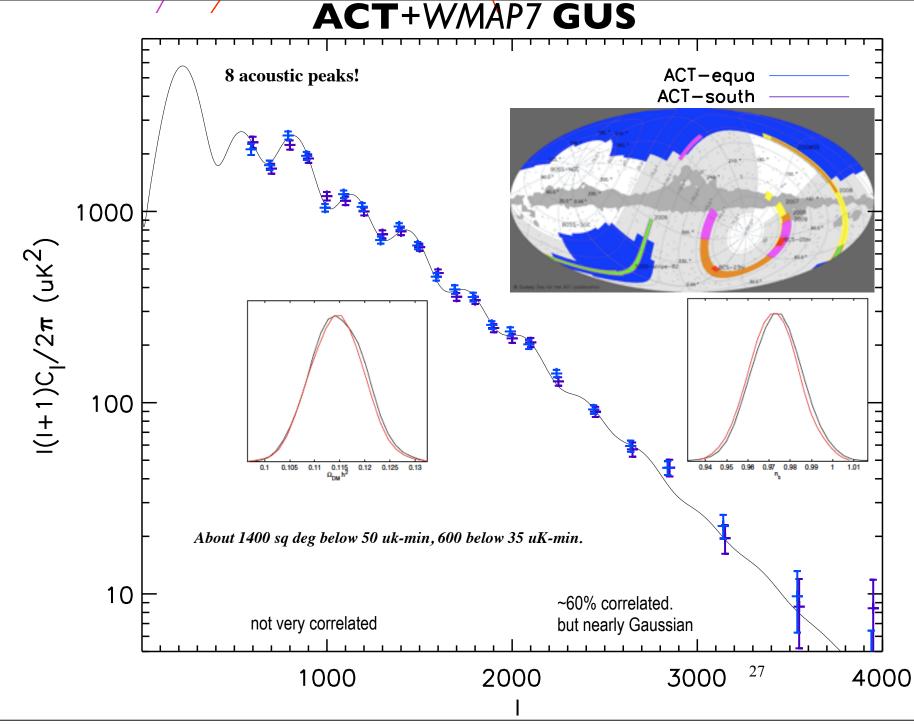


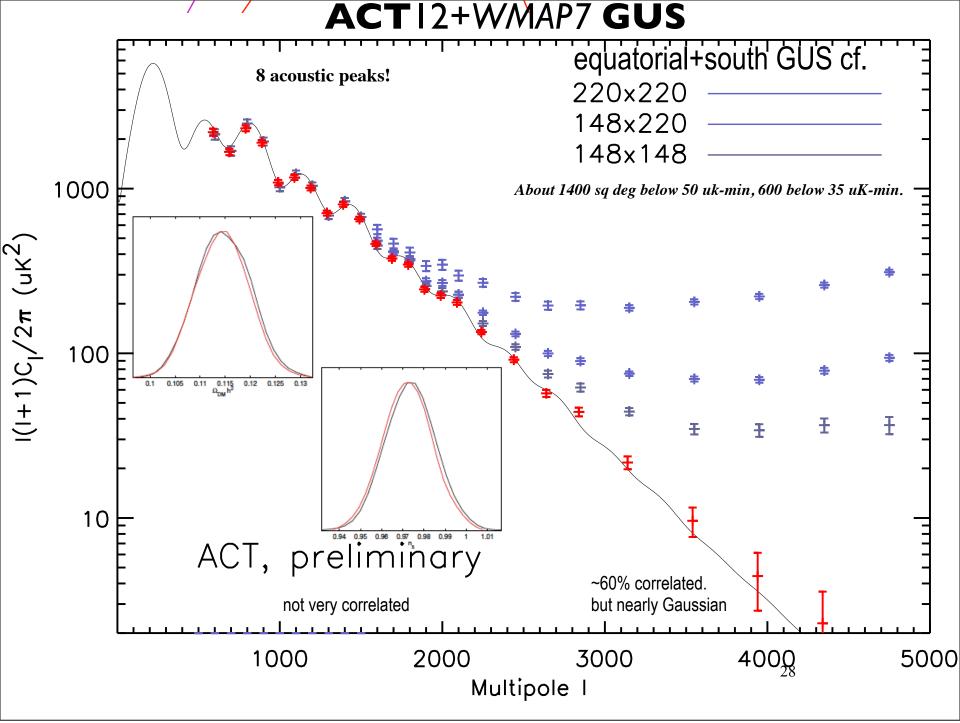
#### end observing 2011: ACT completed 3 full seasons, over ~1400 deg<sup>2</sup>, maps@CITA. next step is ACTpol >= 2013 About 1400 sq deg below 50 uk-min, 600 below 35 uK-min.



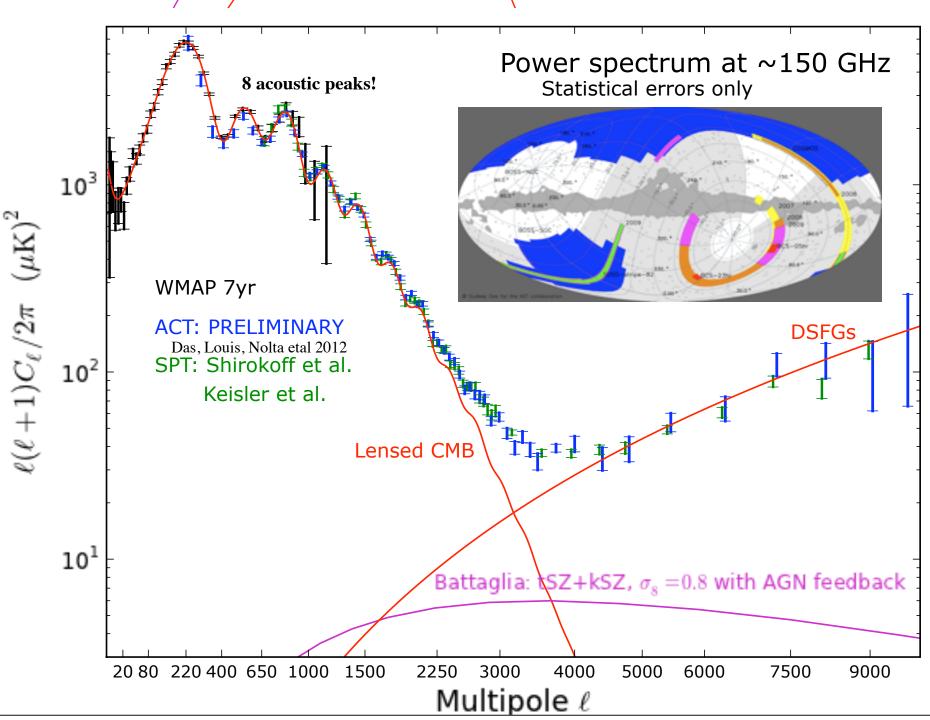
# in Quest of the Damping Tail & the Physics it Entails

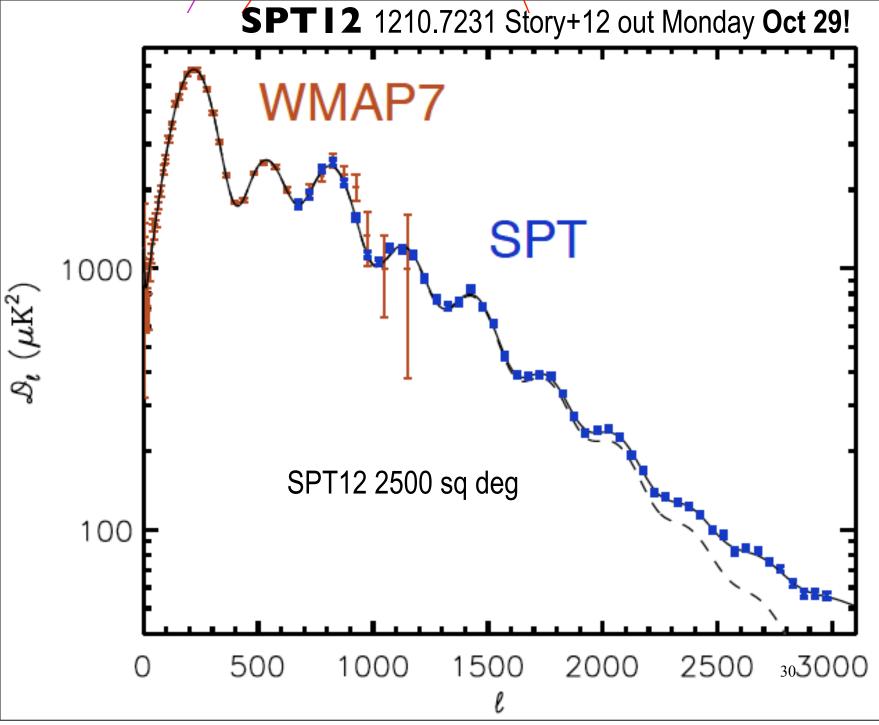


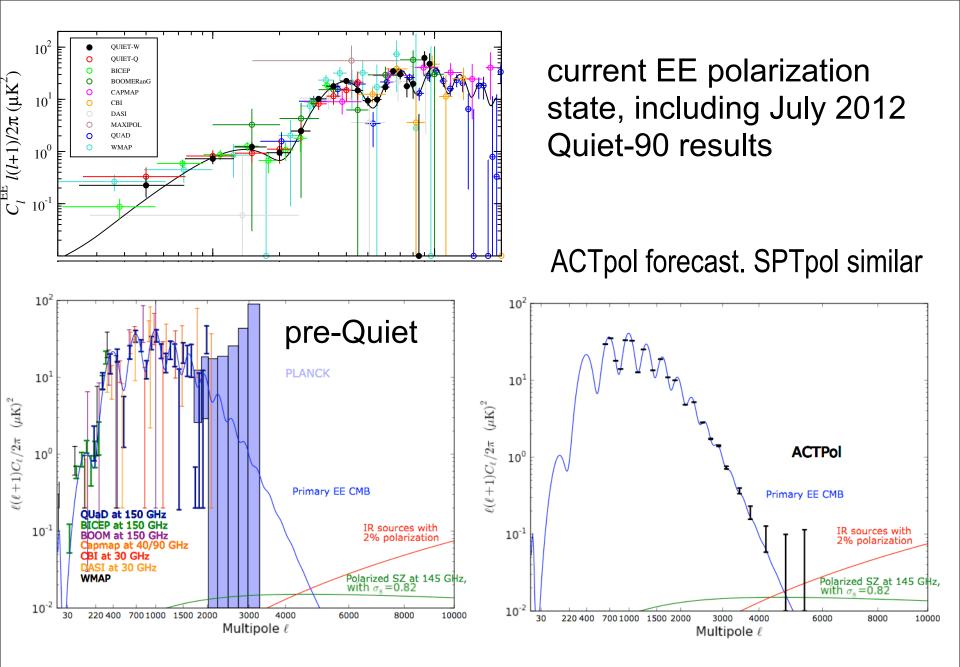




Thursday, 1 November, 12

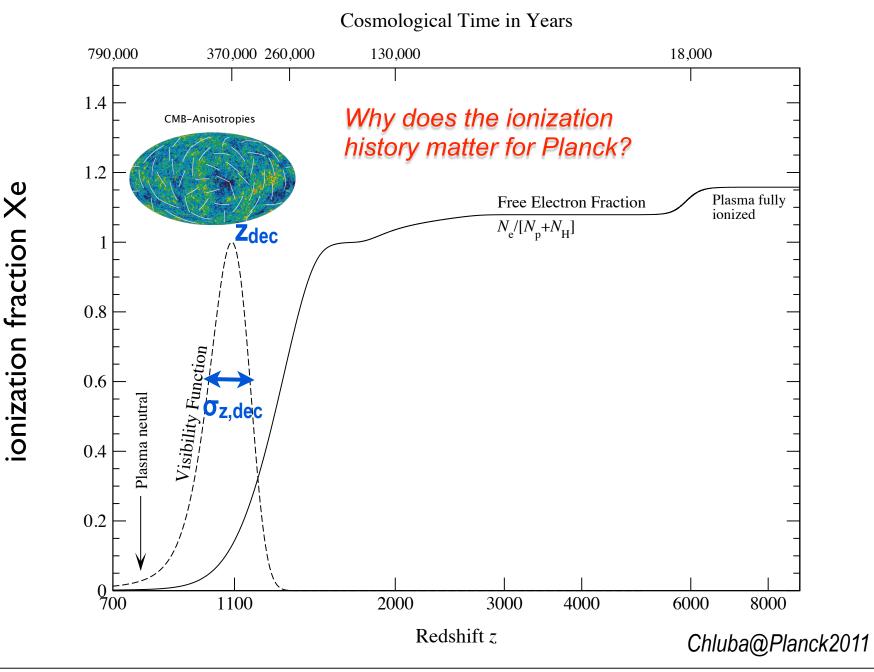




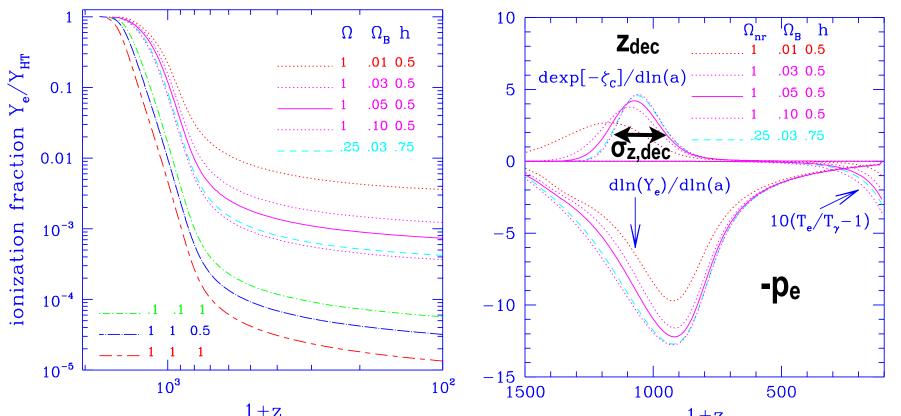


CMB is clean in EE polarization to much higher L than TT => ACTpol + SPTpol nicely complement Planck

# Standard Recombination History



# Standard Recombination History KSZ68,P68 => BE84,B96 => SSS99,00



running of the free electrons-per-baryon  $Y_e = n_e/n_b$ :  $p_e = 3dh^2/n_e/n_b$  /d/nhb  $p_e$  from 0 to 9@dec to max 12 to 0

differential visibility = running of the visibility  $n_e \sigma_T /H exp[-\int n_e \sigma_T /H d/na]$ with kinematic shear viscosity  $4/15C_s^2 / n_e \sigma_T$  thermal diffusion  $n_b S_{\gamma} / n_e \sigma_T$   $C_L \sim exp[-(L/L_D)^m_D]$  damping envelope  $m_D \sim 1.26$ ,  $L_D \sim 1350$  (6' fwhm) WKB baryon-photon tight coupling  $L_D \sim (p_e + 2)(1 + z_{dec})^{1/2} \sim (1 + z_{dec})^{1/2} / \sigma_{z,dec}$ 

## Jens Chluba@Planck2011

## **Getting Ready for Planck**

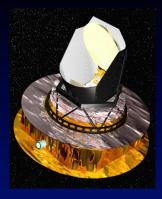
44 GHz

### Hydrogen recombination

- Two-photon decays from higher levels (Dubrovich & Grachev, 2005, Astr. Lett., 31, 359; Wong & Scott, 2007; JC & Sunyaev, 2007; Hirata, 2008; JC & Sunyaev 2009)
- Induced 2s two-photon decay for hydrogen (JC & Sunyaev, 2006, A&A, 446, 39; Hirata 2008)
- Feedback of the Lyman- $\alpha$  distortion on the 1s-2s two-photon absorption rate (Kholupenko & Ivanchik, 2006, Astr. Lett.; Fendt et al. 2008; Hirata 2008)
- Non-equilibrium effects in the angular momentum sub-states (Rubiño-Martin, JC & Sunyaev, 2006, MNRAS; JC, Rubiño-Martín & Sunyaev, 2007, MNRAS; Grin & Hirata, 2009; JC, Vasil & Dursi, 2010)
- Feedback of Lyman-series photons (Ly[n] → Ly[n-1]) (JC & Sunyaev, 2007, A&A; Kholupenko et al. 2010; Haimoud, Grin & Hirata, 2010)
- Lyman-α escape problem (atomic recoil, time-dependence, partial redistribution) (Dubrovich & Grachev, 2008; JC & Sunyaev, 2008; Forbes & Hirata, 2009; JC & Sunyaev, 2009)
- Collisions and Quadrupole lines (JC, Rubiño-Martín & Sunyaev, 2007; Grin & Hirata, 2009; JC, Vasil & Dursi, 2010; JC, Fung & Switzer, in prep.)
- Raman scattering (Hirata 2008; JC & Thomas , 2010; Haimoud & Hirata, 2010)

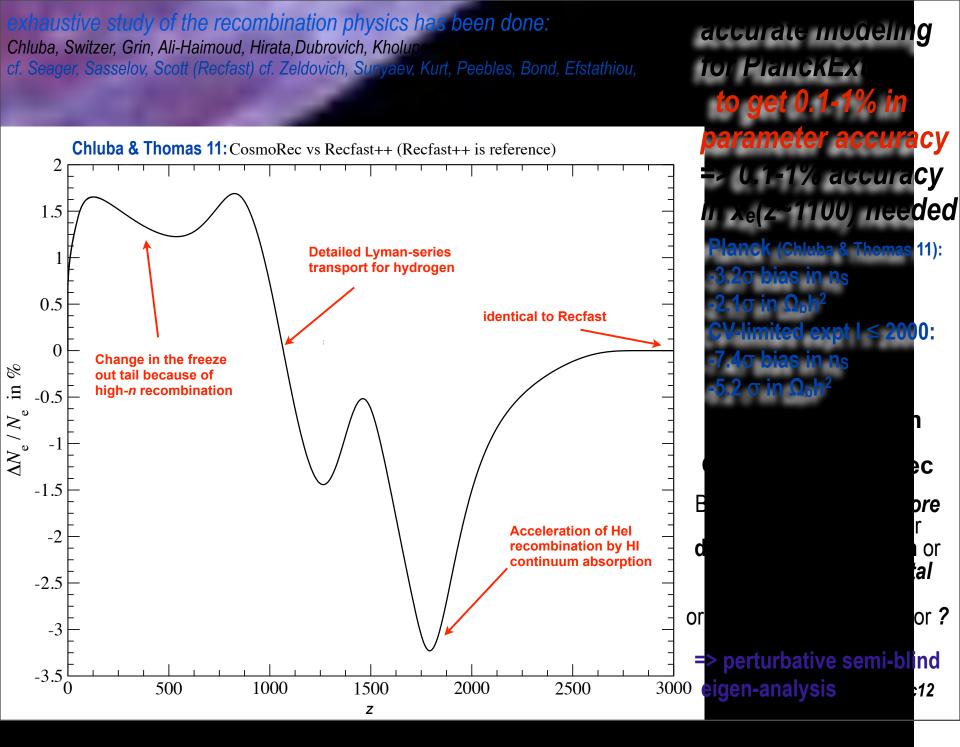
#### Helium recombination

- Similar list of processes as for hydrogen (Switzer & Hirata, 2007a&b; Hirata & Switzer, 2007)
- Spin forbidden 2p-1s triplet-singlet transitions (Dubrovich & Grachev, 2005, Astr. Lett.; Wong & Scott, 2007; Switzer & Hirata, 2007; Kholupenko, Ivanchik&Varshalovich, 2007)
- Hydrogen continuum opacity during He I recombination (Switzer & Hirata, 2007; Kholupenko, Ivanchik & Varshalovich, 2007; Rubiño-Martín, JC & Sunyaev, 2007)
- Detailed feedback of helium photons (Switzer & Hirata, 2007a; JC & Sunyaev, 2009, MNRAS)





HFI 100 GHz



Fisher information matrix, a weight matrix, the 'PRECISION':  $F_{ij} = \langle \partial \mathbf{s}_{\mathbf{f}} / \partial \mathbf{q}^{\mathbf{i}} \partial \mathbf{s}_{\mathbf{f}} / \partial \mathbf{q}^{\mathbf{j}} \rangle_{\mathbf{f}} = \langle \partial ln \mathbf{p}_{\mathbf{f}} / \partial \mathbf{q}^{\mathbf{i}} \partial ln \mathbf{p}_{\mathbf{f}} / \partial \mathbf{q}^{\mathbf{j}} \rangle_{\mathbf{f}}$ = average entropy-content fluctuations  $\mathbf{s} = ln \mathbf{p}^{-1}$  entropy= $\langle \mathbf{s} \rangle_{\mathbf{f}}$  $Fisher^{-1}$  = correlation matrix if Gaussian **Principal Component Analysis (PCA) of x<sub>e</sub>-perturbations**  $F_{ij} = \sum q^a XeM_{ai} XeM_{aj}$  ordered by decreasing weight, increasing error. q<sup>a</sup> now=amplitude of eigenmode XeM<sub>a</sub> only low order high IQ ones are measurable decide which ones by relative entropy criteria

saturate redshift space thru recombination with modes (100s) (**M4 B-splines**, Chebyshev, triangles, Fourier, Gaussians - doesn't matter which). modes of In **x**<sub>e</sub> uniform in z.

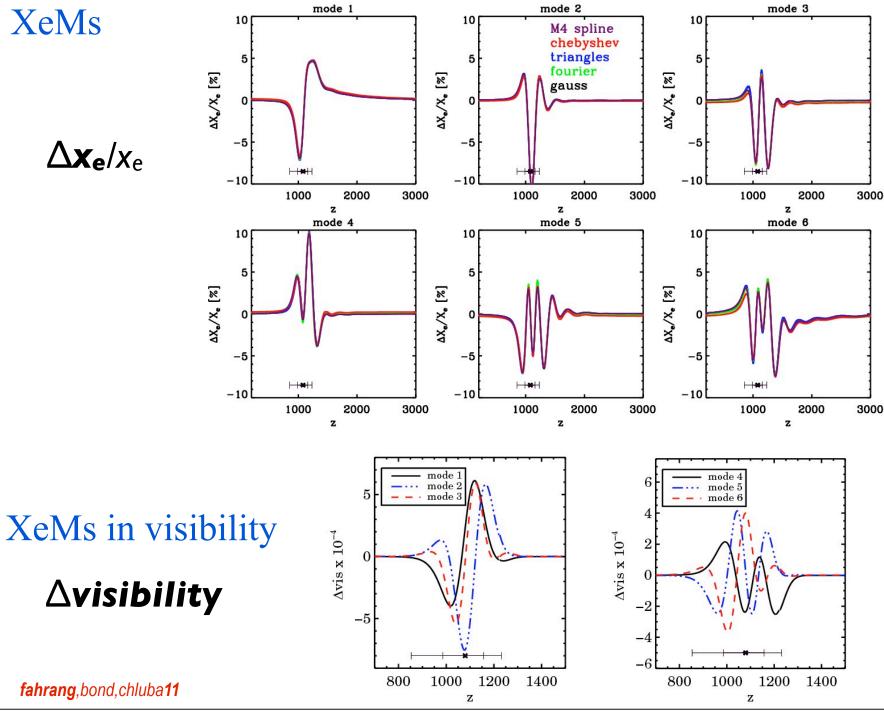
eXeM x<sub>e</sub>-perturbations marginalized over other cosmological parameters modify modes to focus on hi-z (Helium) or Io-z (freeze-out tail) recombination region,

e.g.,  $ln (\mathbf{x}_e + \boldsymbol{\sigma}_e)$ 

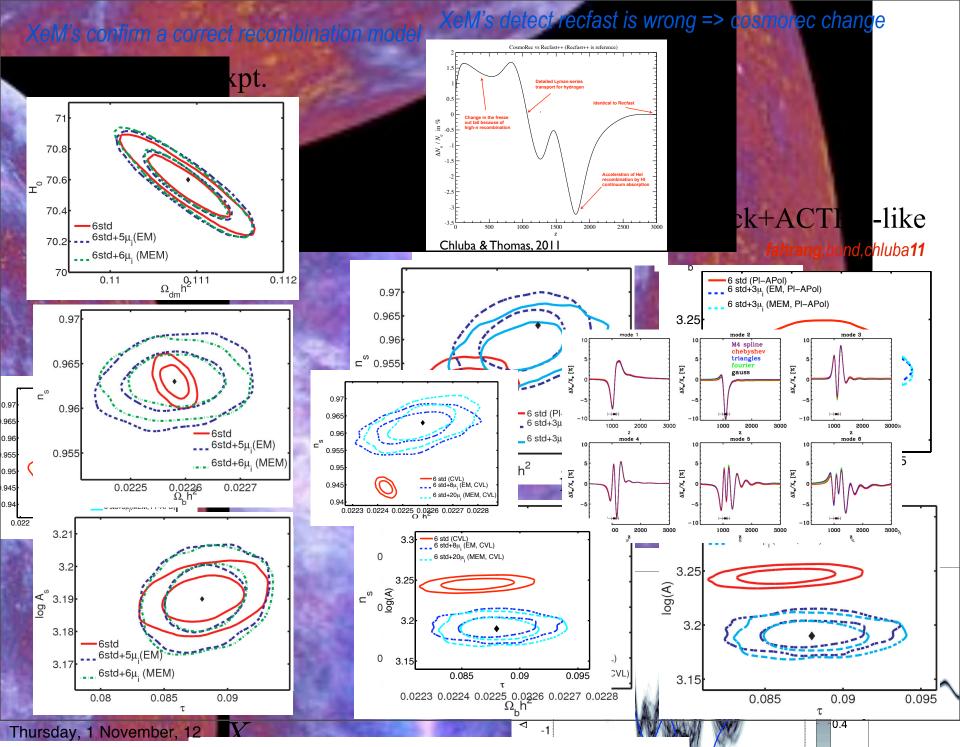
fahrang+bond+chluba11,f+b+switzer+c12







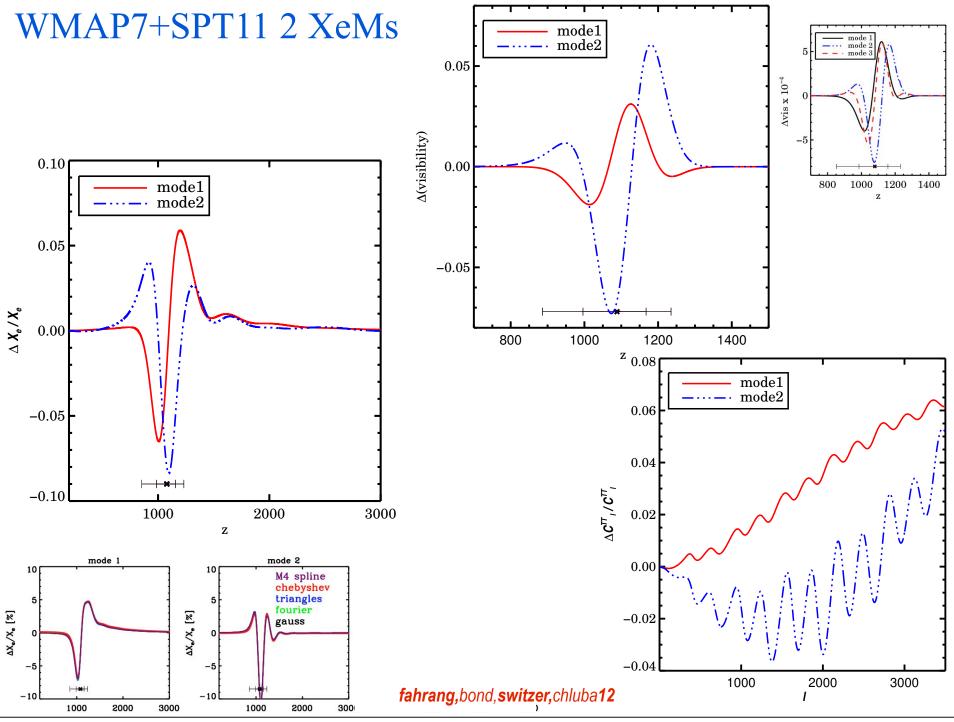
fahrang, bond, chluba11

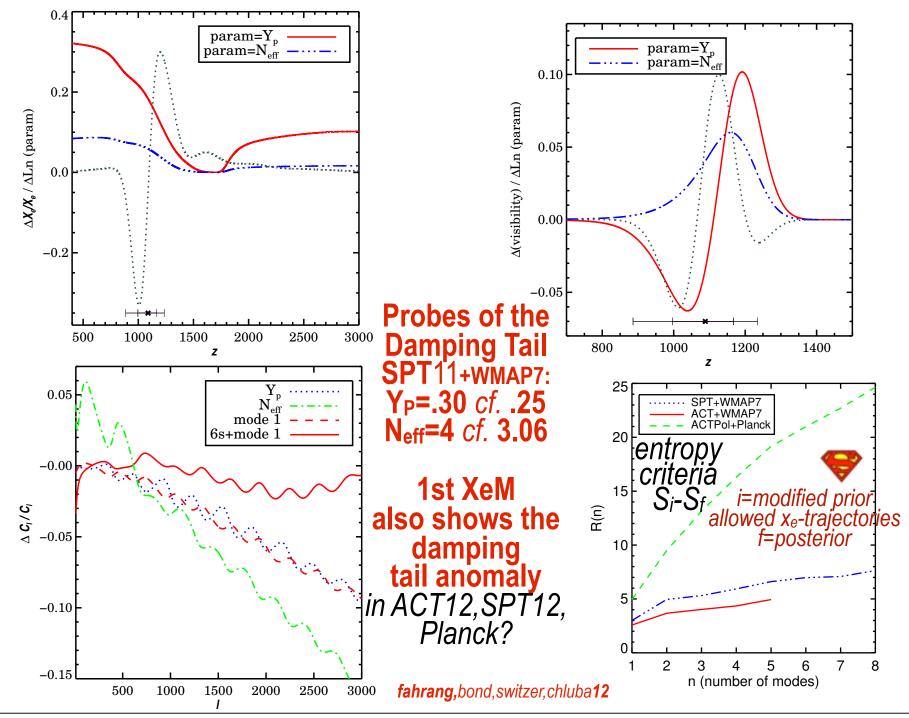


#### Reconstructed x<sub>e</sub>-perturbations Recombination correction is Helium recombination 'recovered correction is not visible EE c, [%] ideal expt х /Х ө 6 XeMs 0.2 500 1000 1500 2000 2500 3000 3500 500 1000 1500 2000 2500 3000 3500 500 1000 1500 2000 2500 3000 ideal expt % X /X [%] 5 10 eXeMs 02 1500 500 1000 2000 2500 3000 500 1000 1500 2000 2500 3000 3500 500 1000 1500 2000 2500 3000 3500 Planck+ACT/SPTPol-like % 3 eXeMs 1000 1500 2500 3000 1000 1500 2000 2500 3000 3500 500 2000 500 1000 1500 2000 2500 3000 3500 500

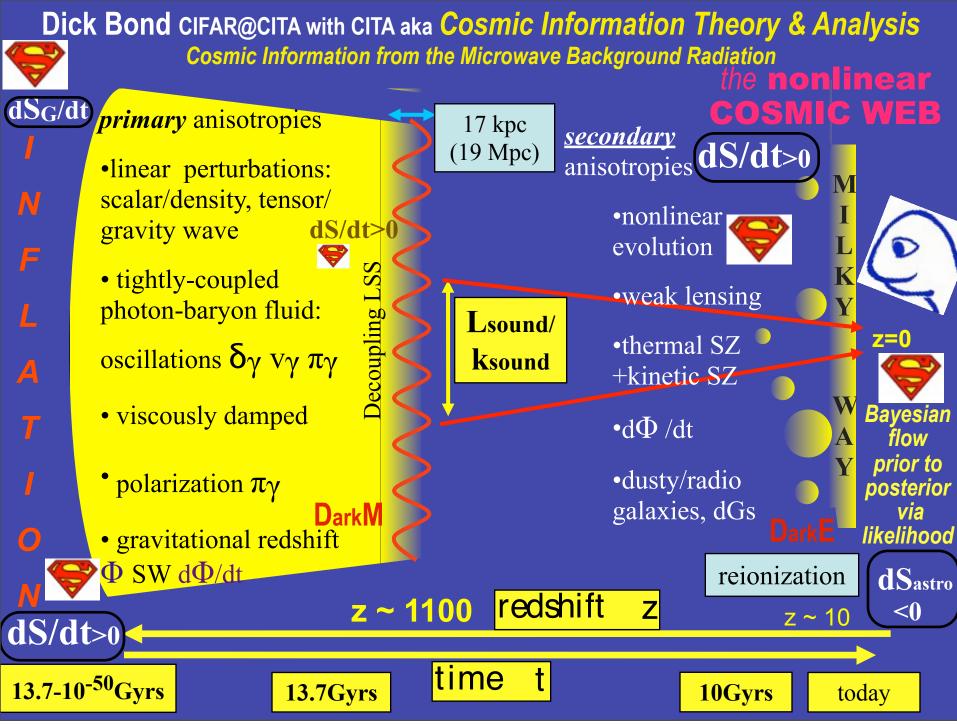
characteristic  $\Delta C_L/C_L$  shape = perturbed damping tail

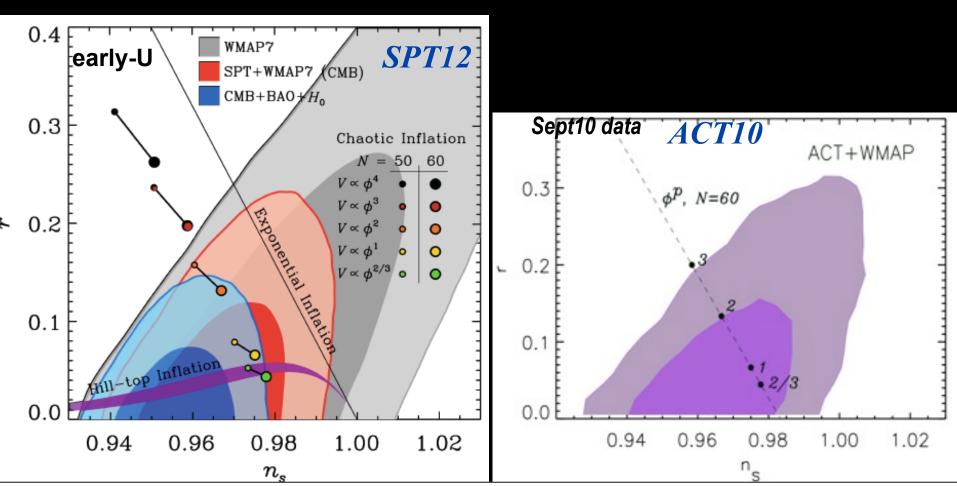
fahrang, bond, chluba11

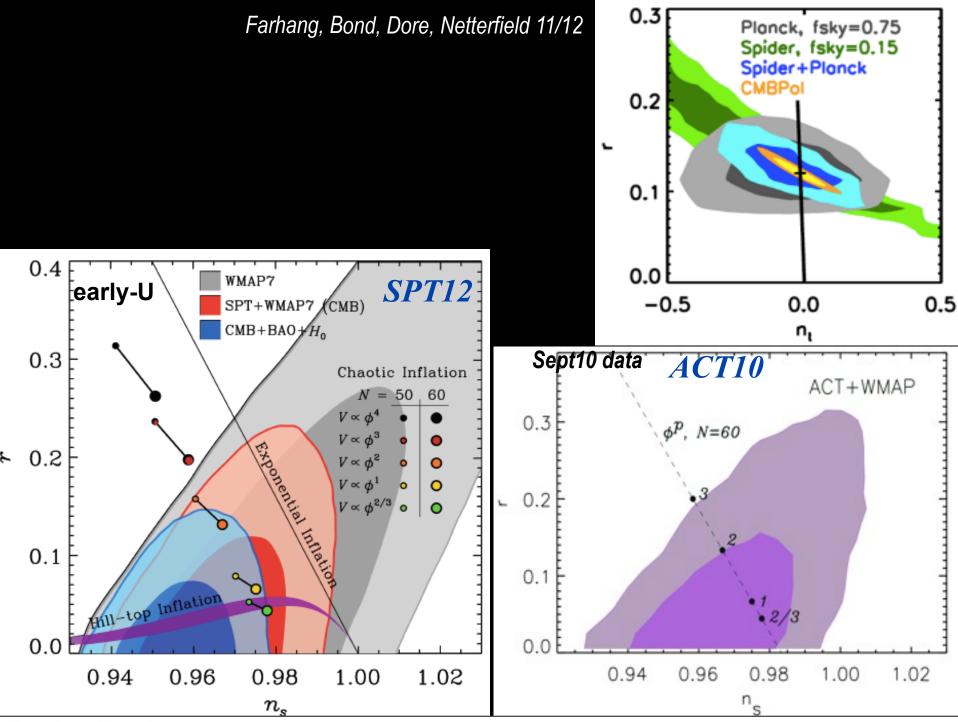




Thursday, 1 November, 12





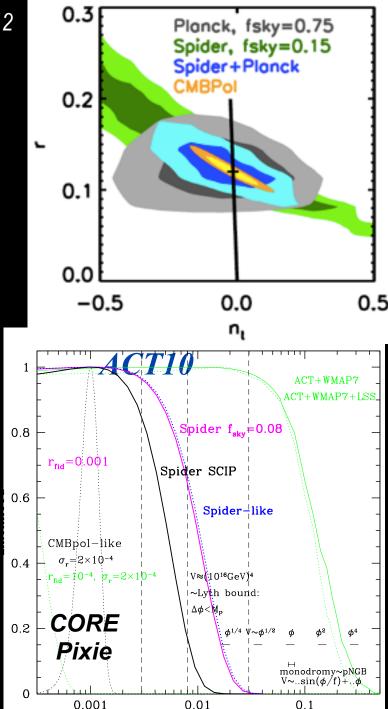


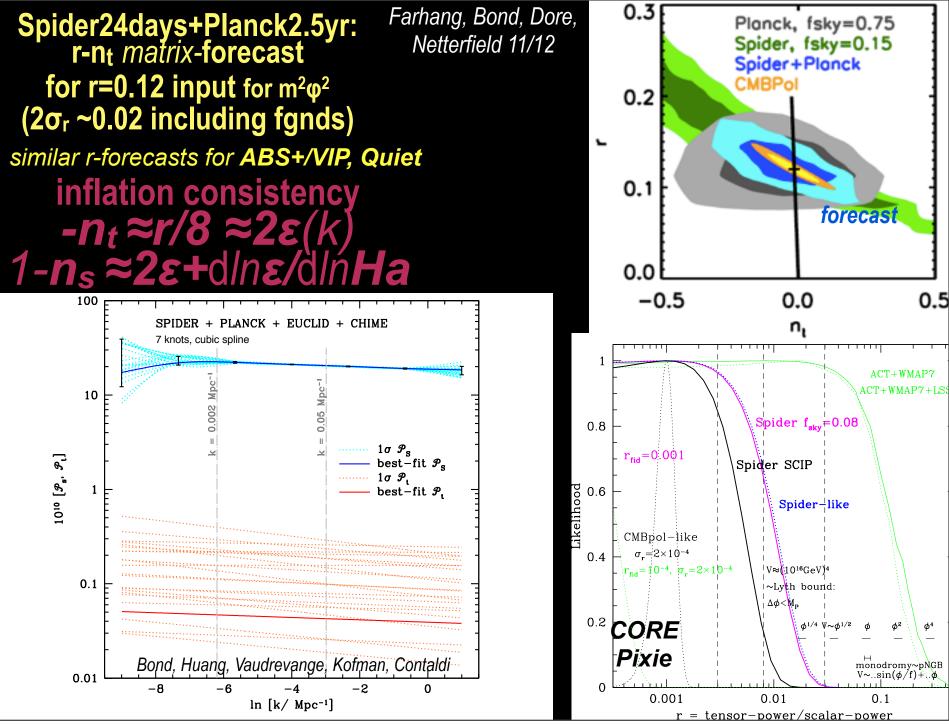
Farhang, Bond, Dore, Netterfield 11/12

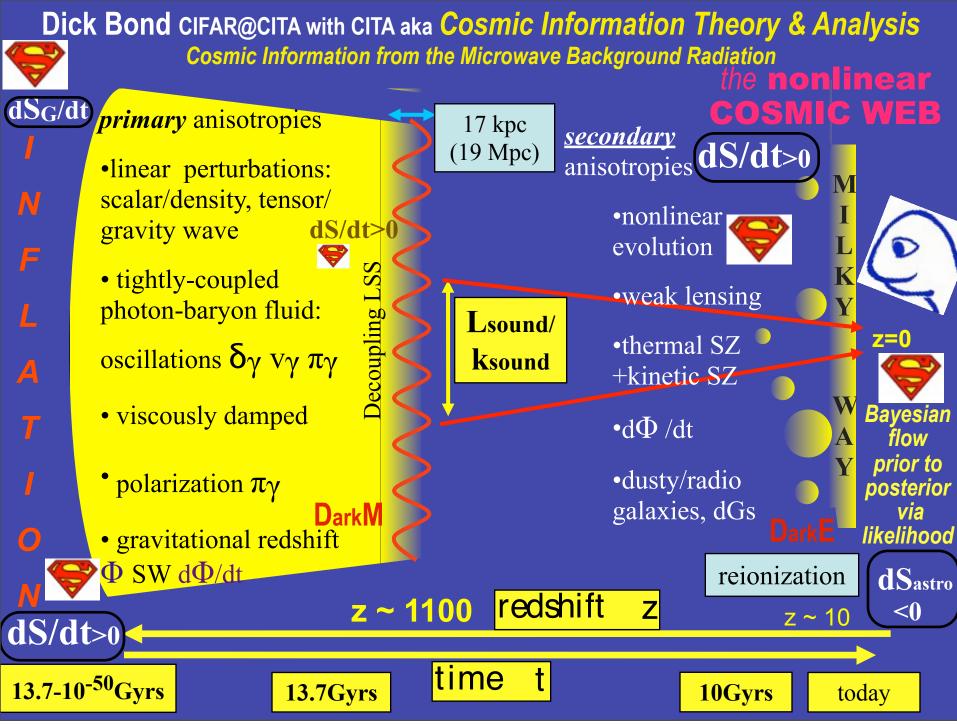
## Spider24days+Planck2.5yr: r-nt matrix-forecast for r=0.12 input for m<sup>2</sup>φ<sup>2</sup> (2σ<sub>r</sub> ~0.02 including fgnds)

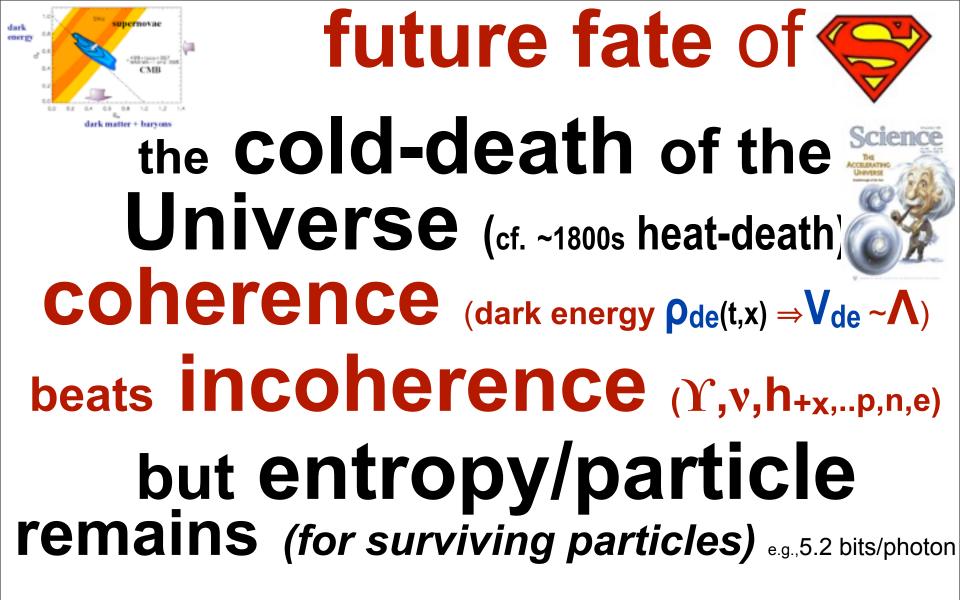
similar r-forecasts for ABS+/VIP, Quiet

## inflation consistency $-n_t \approx r/8 \approx 2\epsilon(k)$ $1-n_s \approx 2\epsilon + d/n\epsilon/d/nHa$



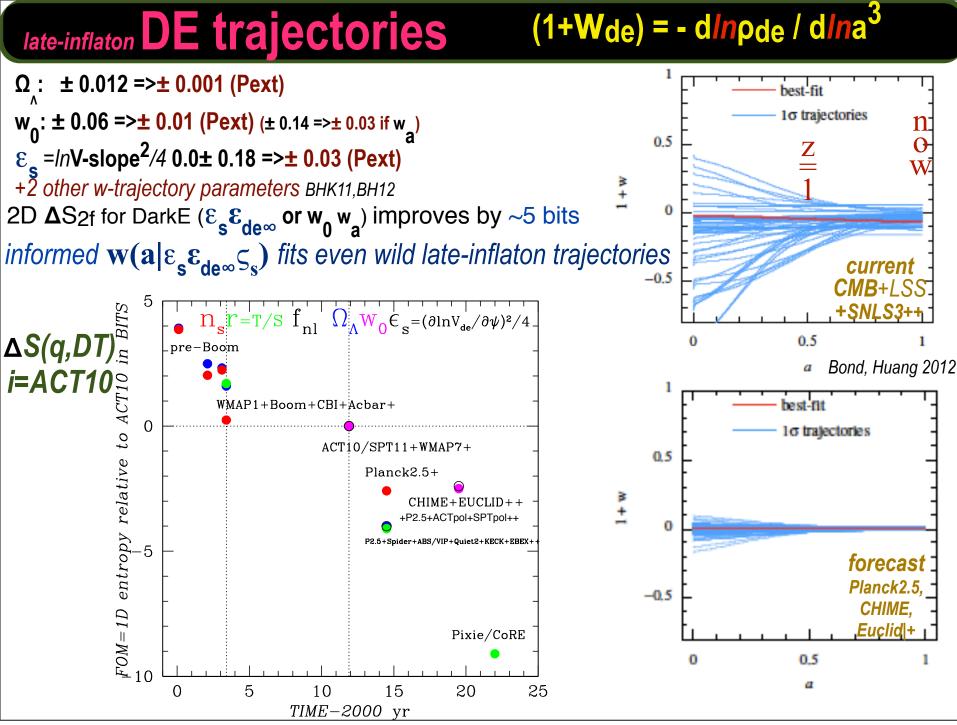




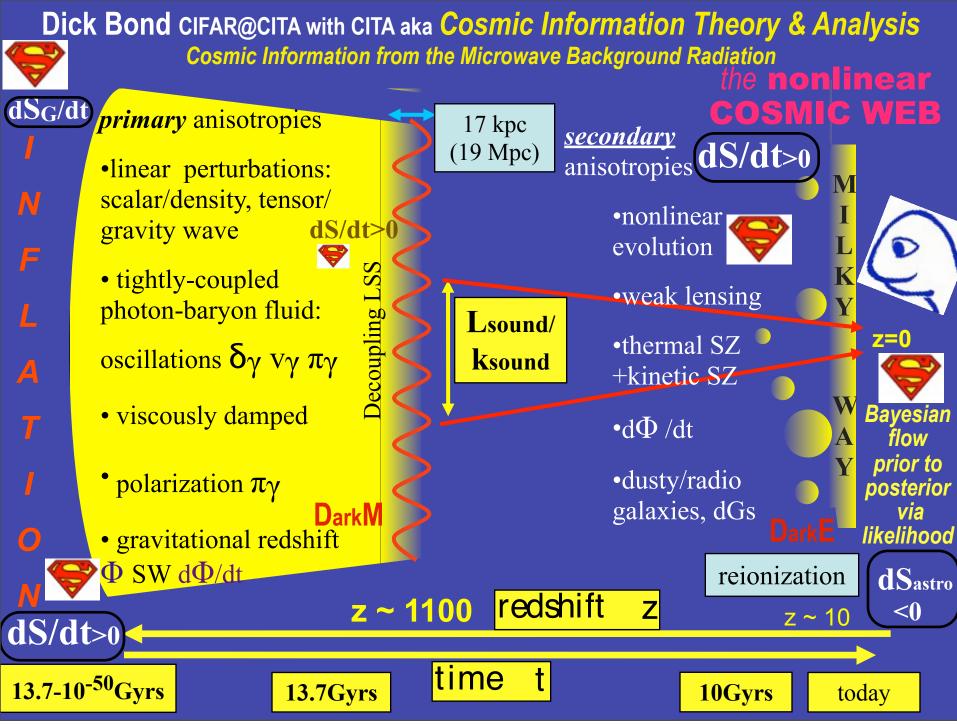


the **gravo-thermal catastrophe** = negative specific heat - goal to localize all mass into black holes & make accelerating voids **to straighten U** out, radiating entropy along the way

although  $S_G = M_{bh}^2 / 2M_P^2$  decays into radiation,  $S_G = M_P^2 / 2(H/2\pi)^2 \sim 10^{121.9}$  remains (until tunnel)



Thursday, 1 November, 12



early U applications of "CITA" to cosmic-complexity

the superhorizon measure problem & the Lambda-scape

☆ the emergence of the collective from the random! coherence from driven zero-point vacuum fluctuations ⇒ V inflaton, gravity waves; decohere

A let there be heat: entropy generation in preheating from the coherent inflaton (origin of all "matter")

P(q|D,T) = P(D|q,T)P(q|T)P(T)/P(D|T) D = CMB, LSS, SN, ..., Complexity, life*T*=baryon, dark matter, vacuum mass-energy densities,..., early & late inflation as low energy flows/trajectories on a (string) landscape **Old:** Theory prior = delta function of THE correct one&only **New: Theory prior = probability distribution** of

late-ish-flows on a LANDSCAPE



**S**<sub>U,m+r</sub>~10<sup>88.6</sup>

cf. SG~10<sup>121.9</sup>

Studying the





Ş

## modulating post-inflation entropy generation shocks via long range fields

isocon

**χ(x)** or **g(**σ(**x**)) or..

φ inflaton

preheating patch (~1cm)

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Parametric Resonance  $g^2/\lambda \sim 1$ 

1000 Gpc

**10 Gpc** 

### modulating post-inflation entropy generation shocks via long range fields



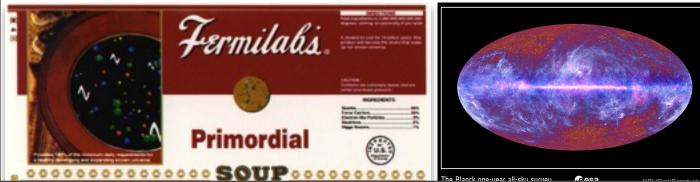
Coherent Inflation with Quantum Jitter to Hot Big Bang, an Incoherent Particle Soup

# how (most of) the entropy in matter

=> GUT plasma/quark soup =>  $S(\gamma, \nu)$  was

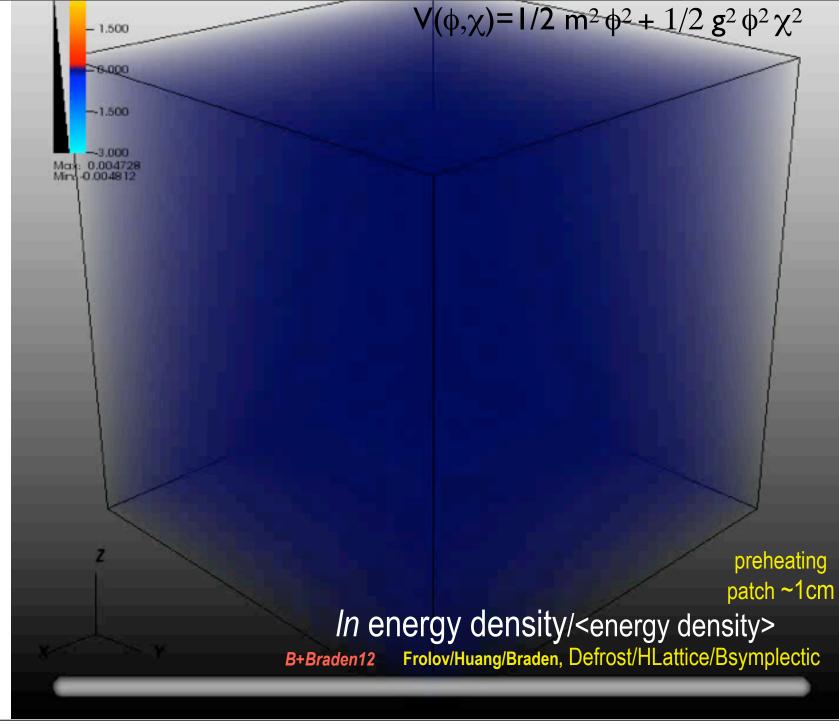
generated (through a shock-in-time) via nonlinear coupling of the inflaton to new interaction channels g, Xa ultimately to standard model degrees of freedom ∃ a role for decaying particles, 1st order phase transitions?

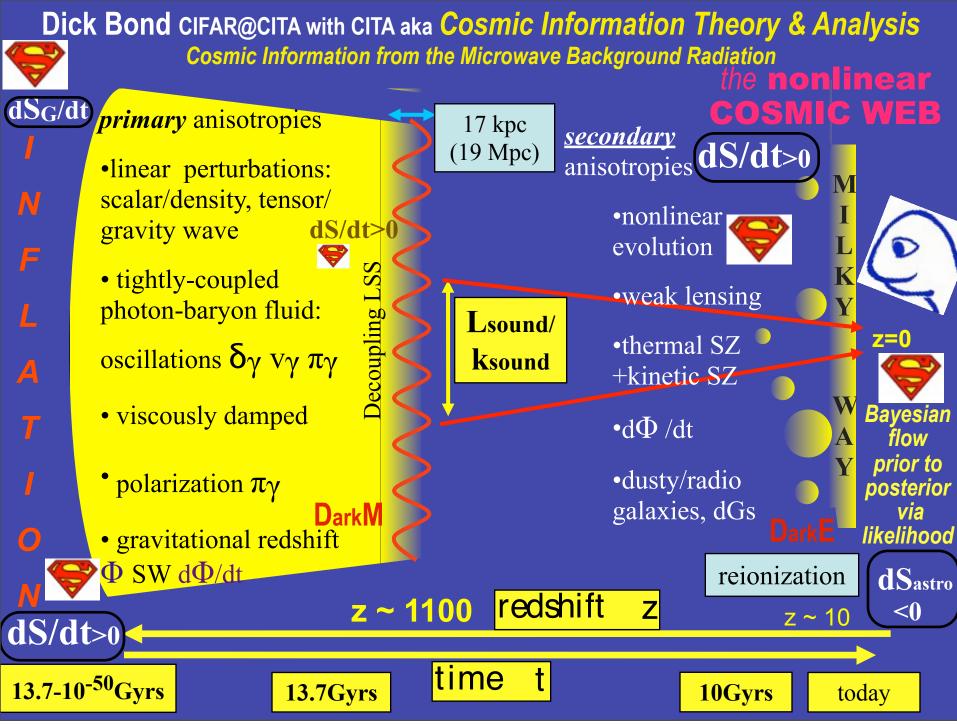
**exactly who, what, where, when, why?** we search for fossil "non-Gaussian" structures from this period with Planck +WMAP9



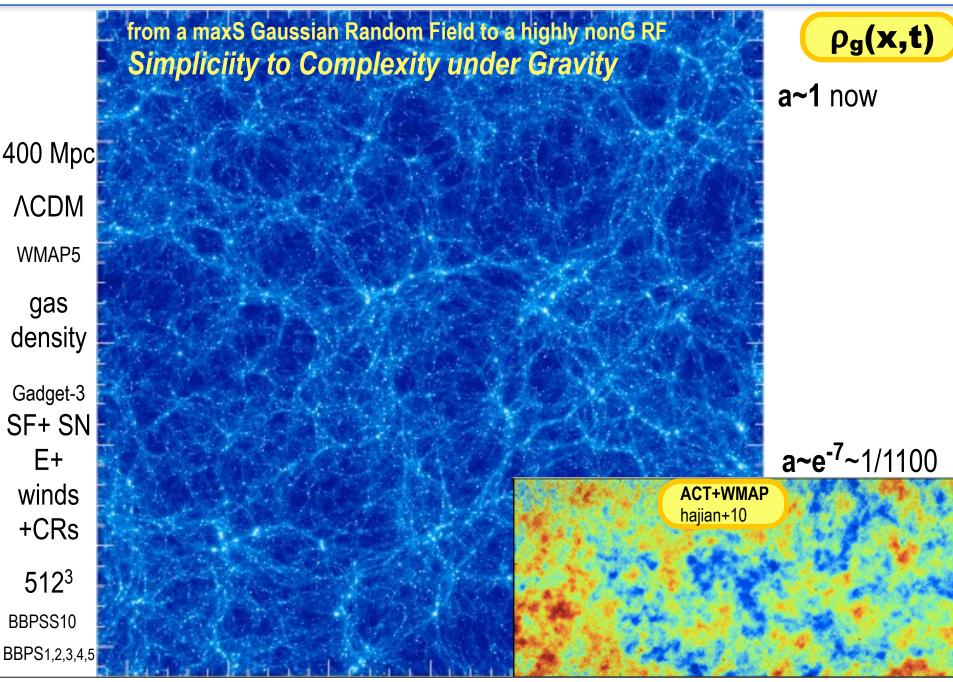
ashock(g)

non-Gaussianity (WMAP, Planck, LSS) spiky nG preheating



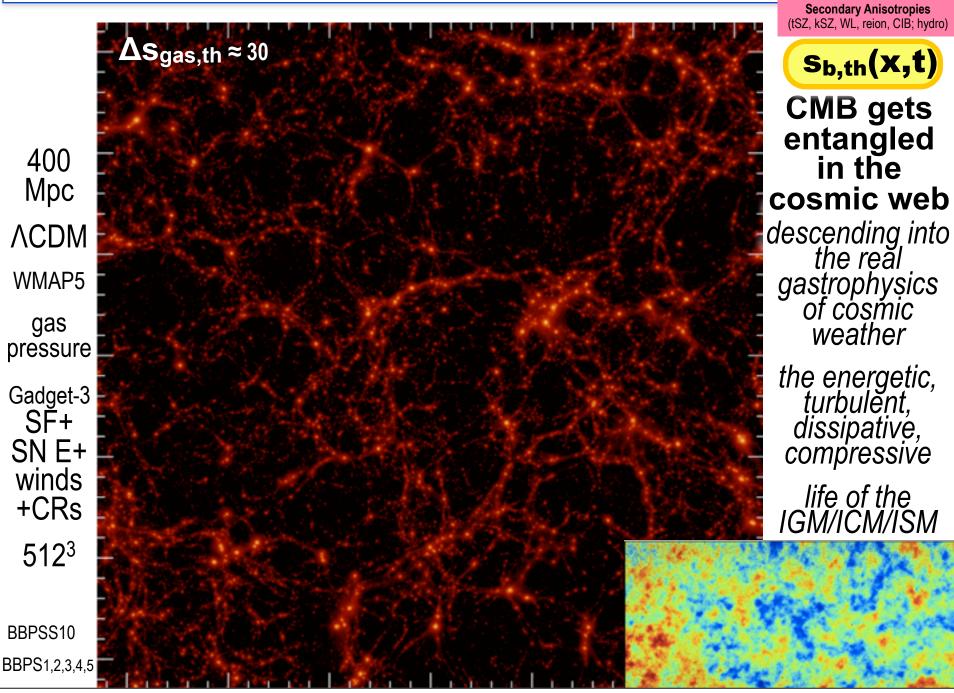


fluctuations in the early universe "vacuum" grow to all cosmic web structure

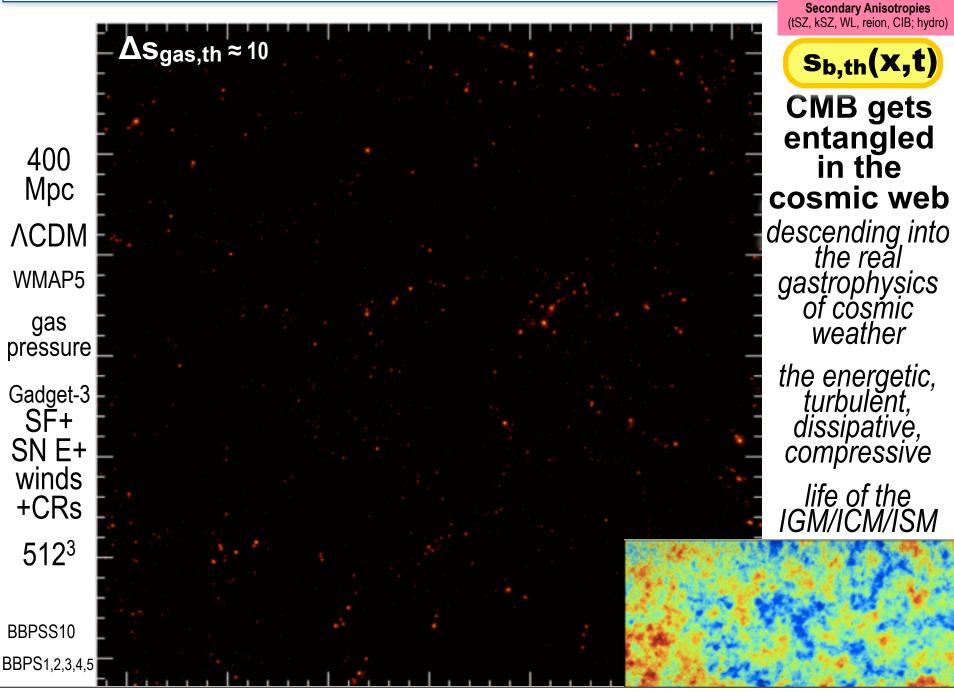


Thursday, 1 November, 12

entropy intermittency in the cosmic web, via gravitation-induced shocks (then E/S-feedback)

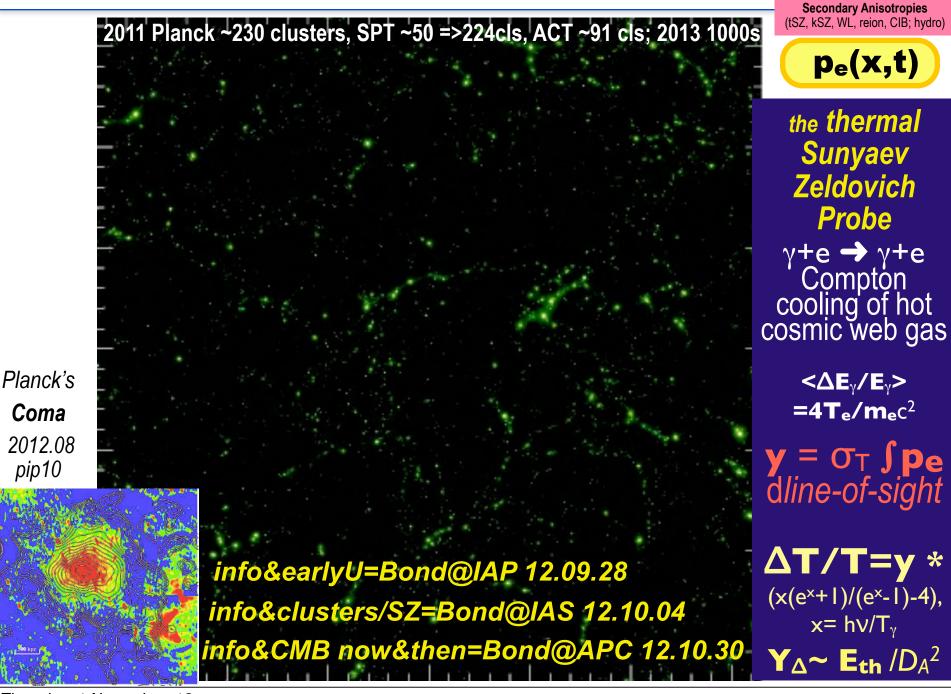


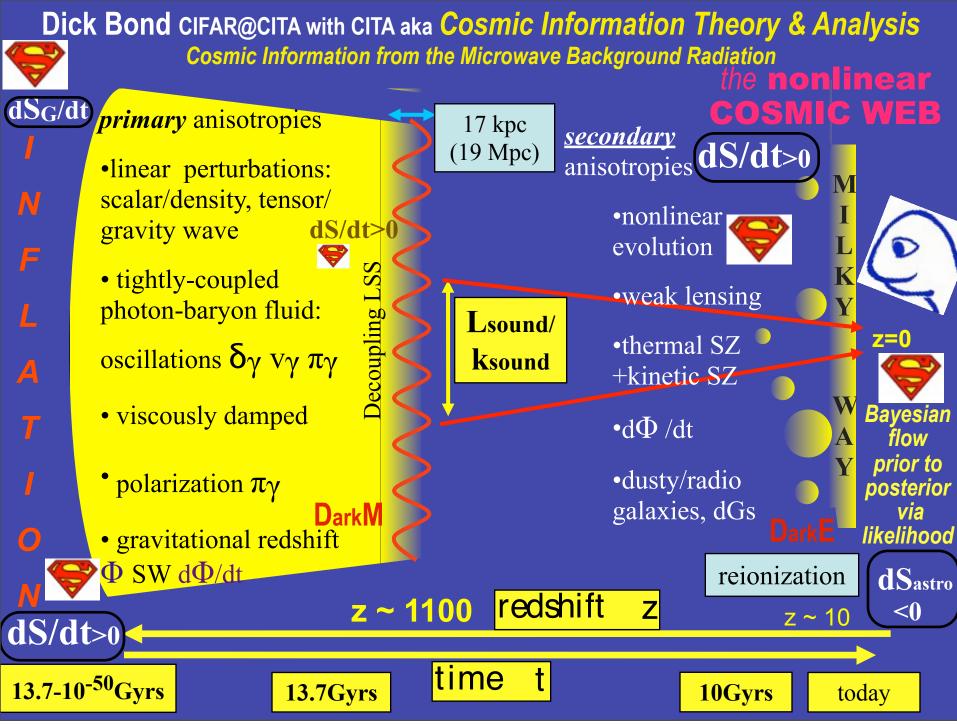
#### entropy intermittency in the cosmic web, via gravitation-induced shocks (then E/S-feedback)



Thursday, 1 November, 12

pressure intermittency in the cosmic web, in cluster-group concentrations probed by tSZ





# END