

The Art of Lattice and Gravity Waves from Preheating

The 3rd IPhT/LPT COSMO meeting

Zhiqi Huang
IPhT, CEA/Saclay

Dec 13, 2010

Outline

Introduction: GW from preheating

The Art of Lattice

The Wisdom of Discretization

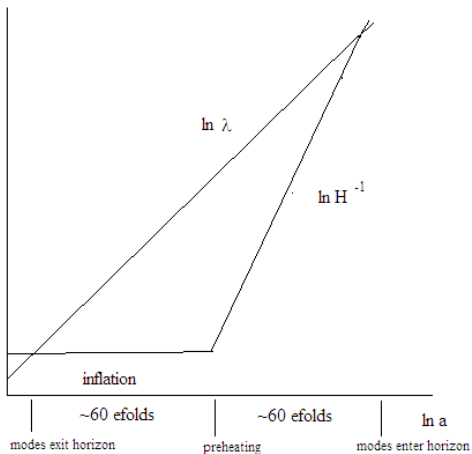
Symplectic Integrator

The Lattice Code

The GW Spectrum: Results and Comparison

Conclusions

Inflation and Preheating



GW from Preheating

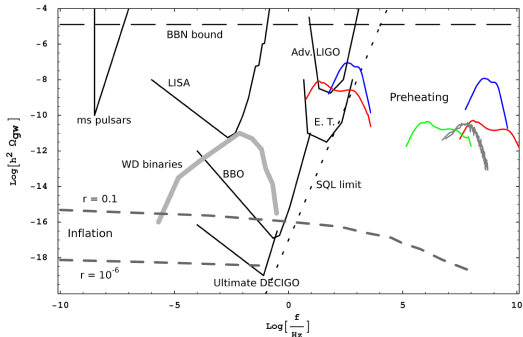


figure source: arXiv:0812.2917 (Dufaux, Felder, Kofman & Navros)

Tachyonic Preheating after Hybrid Inflation

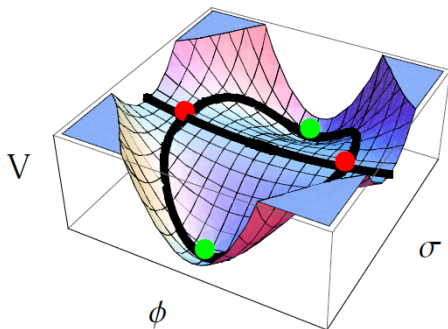


figure source: Dufaux, Felder, Kofman & Navros 2008

ϕ : Inflaton

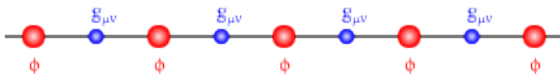
σ : a complex field ($= \sigma_1 + i\sigma_2$), it acquires a tachyonic mass ($m^2 < 0$) when ϕ rolls along the ridge.

$$V(\phi, \sigma) = \frac{\lambda}{4}(\sigma^2 - v^2)^2 + \frac{1}{2}g^2\phi^2\sigma^2 + V_{\text{inf}}(\phi)$$

The Wisdom of Discretization

The (1st order) discrete derivatives $\frac{\partial\phi}{\partial x_i}$ are defined on “displaced grids”, and so is $g_{\mu\nu}$.

1D example (actually done in 3D with more a sophisticated scheme):



- ▶ Discrete differential operators are commutable $\partial_i\partial_j = \partial_j\partial_i$. Operators such as ∇^{-2} are all well defined, so are the discrete scalar, vector, and tensor modes.
- ▶ (Discrete) scalar/vector/tensor terms in $\delta T^{\mu\nu}$ only excite (discrete) scalar/vector/tensor metric perturbations.
- ▶ ∇^{-2} is NOT equivalent to: discrete Fourier transformation & multiply by $(-k^{-2})$ & inverse discrete Fourier transformation. Need more sophisticated scheme to get traceless transverse (TT) component.

The Wisdom of Discretization

For $k_{\min} \ll k \ll k_{\max}$ the boundary effect and finite resolution effect should disappear. Why sophistications? Why not use projection in Fourier space to get GW (as done in all previous works)?

Two reasons:

- ▶ Realistic simulations have $k_{\max} \sim 10^2\text{-}10^3 k_{\min}$. So $k_{\min} \ll k \ll k_{\max}$ is never that well satisfied.
- ▶ Scalar metric perturbations \gg tensor metric perturbations. A tiny “leaking” from $g_{\mu\nu,\text{scalar}}$ to $g_{\mu\nu,\text{tensor}}$ can significantly change the GW spectrum.

Evolving A Classical System on the Computer

A trivial system the computer can evolve “exactly”: a free particle with mass m .

$$\text{Hamiltonian } H = \frac{p^2}{2m}.$$

Evolving from t to $t + \Delta t$:

$$p(t + \Delta t) = p(t) , \quad (1)$$

$$x(t + \Delta t) = x(t) + \frac{p(t)}{m} \Delta t . \quad (2)$$

Only round-off errors ($\lesssim 10^{-16}$ with double precision).

Evolving A Classical System on the Computer



A classical harmonic oscillator is already challenging:

$$H = \frac{p^2}{2} + \frac{q^2}{2}.$$

Evolving from t to $t + \Delta t$:

$$p(t + \Delta t) = p(t) - q(t)\Delta t + O(\Delta t^2), \quad (3)$$

$$q(t + \Delta t) = q(t) + p(t)\Delta t + O(\Delta t^2) \quad (4)$$

- ▶ Shrinking Δt usually does not help much.
- ▶ A better solution is to change the scheme to make the error term $\propto \Delta t^n$. Large $n \Rightarrow$ only round-off errors.

Evolving A Classical System on the Computer

For a general Hamiltonian $H(\mathbf{p}, \mathbf{q})$ and a variable $f(\mathbf{p}, \mathbf{q})$:

$$\frac{df}{dt} = \hat{\mathcal{H}}f, \quad (5)$$

where the Poisson bracket operator $\hat{\mathcal{H}} \equiv \{\cdot, H\}$.

The exact solution can be formally written as

$$f(t + \Delta t) = e^{\hat{\mathcal{H}}\Delta t} f(t) \quad (6)$$

- ▶ The computer knows how to calculate $e^{\hat{\mathcal{H}}\Delta t} f$ EXACTLY if $H = T(\mathbf{p}) \equiv \mathbf{p}^2/2$ or $H = V(\mathbf{q})$, but NOT in the case $H = T(\mathbf{p}) + V(\mathbf{q})$.
- ▶ This is because $\hat{\mathcal{T}} \equiv \{\cdot, T\}$ and $\hat{\mathcal{V}} \equiv \{\cdot, V\}$ are not commutable:

$$e^{\hat{\mathcal{H}}\Delta t} = e^{(\hat{\mathcal{T}}+\hat{\mathcal{V}})\Delta t} \neq e^{\hat{\mathcal{T}}\Delta t} e^{\hat{\mathcal{V}}\Delta t}. \quad (7)$$

Evolving A Classical System on the Computer

The idea of symplectic integrator (SI) is to find constant c-numbers $c_1, c_2, \dots, d_1, d_2, \dots$, such that:

$$e^{(\hat{T}+\hat{V})\Delta t} = e^{c_1\hat{T}\Delta t} e^{d_1\hat{V}\Delta t} e^{c_2\hat{T}\Delta t} e^{d_2\hat{V}\Delta t} \dots + O(\Delta t^{n+1}). \quad (8)$$

- ▶ The computer knows how to use the right-hand-side of Eq. (8) to evolve the system EXACTLY, if the $O(\Delta t^{n+1})$ term is dropped.
- ▶ The 2nd order SI ($n = 2$) is the well known “leap frog” algorithm (used in LatticeEasy & DEFROST when expansion of universe is ignored). Its explicit form is

$$e^{(\hat{T}+\hat{V})\Delta t} = e^{\frac{1}{2}\hat{T}\Delta t} e^{\hat{V}\Delta t} e^{\frac{1}{2}\hat{T}\Delta t} + O(\Delta t^3). \quad (9)$$

- ▶ Up to 8th order SI's are all known.

The Lattice Code

Scalar fields on a lattice:

$$H(\phi, g_{\mu\nu}) = \hat{A} + \hat{B} + \hat{C}, \quad (10)$$

where the three noncommutative terms are

\hat{A} : the diagonal terms in the kinetic energy of gravity.

\hat{B} : the kinetic energy of the fields and the off-diagonal terms in the kinetic energy of gravity.

\hat{C} : the potential and gradient energy of the fields and gravity

- ▶ The high-order symplectic factorization of $e^{(\hat{A}+\hat{B}+\hat{C})\Delta t}$ is unknown (iterative expansion \Rightarrow hundreds of factors, practically useless.)
- ▶ Operator \hat{A} itself is non-canonical (i.e., it has some non-splittable terms containing both $p_{\mu\nu}$ and $g_{\mu\nu}$, the computer does not know how to calculate $e^{\hat{A}\Delta t}f$ exactly.

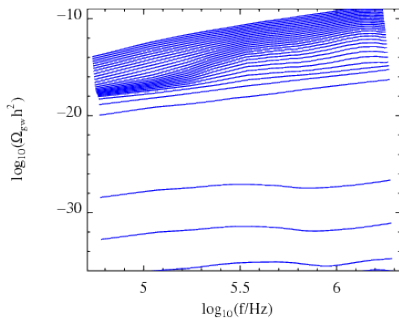
The Lattice Code

What I have done:

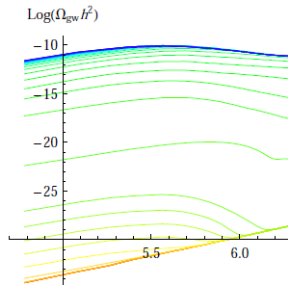
- ▶ Find the 4-th and 6-th order optimal symplectic factorizations of $e^{(\hat{A}+\hat{B}+\hat{C})\Delta t}$.
- ▶ Use 4-th order Runge-Kutta integrator (with much smaller time-steps) to calculate $e^{\hat{A}\Delta t}f$ (no cost of extra memory: because \hat{A} does not contain the gradient terms, I can calculate it grid-wise-ly.)

Features of the code:

- ▶ You can choose 2nd, 4th or 6th order symplectic integrators (CPU time cost – 1:3:7. Memory cost: all minimal), depending on how accurate you want your result to be.
- ▶ Three modes for the gravity: FRW background with metric perturbations; FRW background without metric perturbations; Minkowski spacetime.



HLattice output: metric perturbation feedbacks; evolution in real space; h^{TT} separated in real space

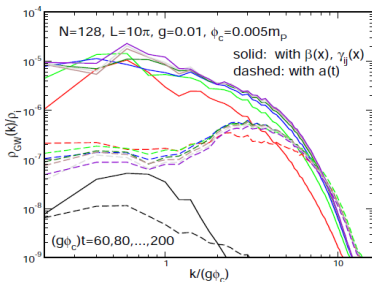


arXiv:0812.2917 (Dufaux, Felder, Kofman & Navros)

LatticeEasy output: no metric perturbation feedbacks; evolution in real space; TT component separated in Fourier space.

The opposite claim?

The opposite claim: non-linear enhancement of GW on large scales?



arXiv:1005.4054 (Bastero-Gil, Macias-Pérez and Santos).

TT component separated in Fourier space; solid/dashed: with/without metric perturbation feedbacks;

leaking from scalar to tensor? numerical noise?

Conclusions

- ▶ A discretization scheme allows the separation of scalar, vector and tensor modes.
- ▶ Symplectic integrator (with Runge-Kutta sub-integrator for non-canonical terms): evolve the discretized system almost exactly (numerical noise $\sim 10^{-14}$).
- ▶ Find lower IR tail of the GW spectrum. Is the extra IR power in the literature from non-perfect discretization?