

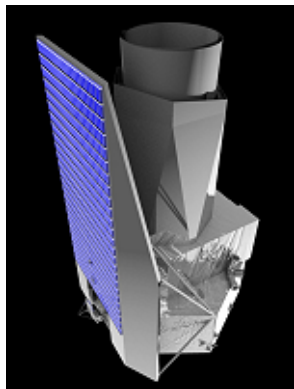
Future LSS Survey and Inflation Models

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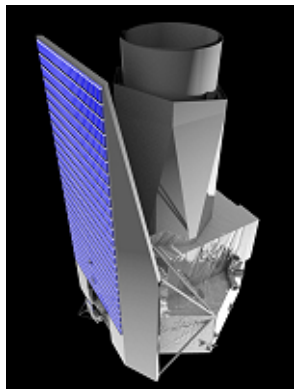
Why this is a right decision



EUCLID mission:

- ▶ Dark Energy
- ▶ Galaxies
- ▶

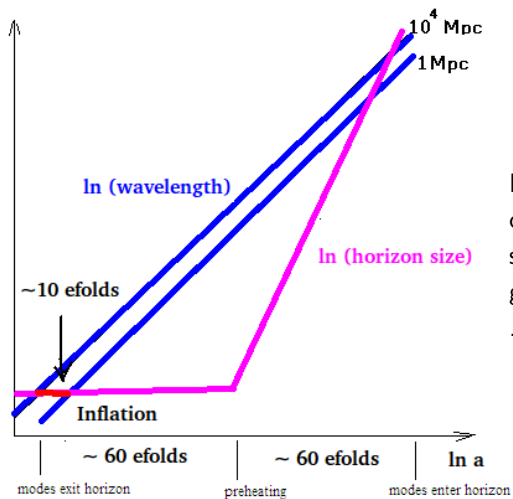
Why this is a right decision



EUCLID mission:

- ▶ Dark Energy
- ▶ Galaxies
- ▶ **Inflation**

Only about 10 efolds of inflations are observable



Inflation seeds the cosmological-scale structures (CMB, galaxy clustering, ...).

Can we learn something?

Two classes of models (natureness v.s. falsifiability).

- ▶ Simple models: scale-invariant power spectrum, adiabatic perturbations, ...
natural; featureless; difficult to distinguish; general paradigm without concrete physics
- ▶ Complicated models: Starobinsky potential (Starobinsky 1992), double inflation (Polarski & Starobinsky 1992), particle production during inflation (Barnaby, Huang, Kofman, Pogosyan 2009, Barnaby & Huang 2009), modulated preheating (Bond, Frolov, Huang, Kofman 2009), extra dimensions (Bean et al 2008, McAllister et al 2010, Flauger et al 2010).
crazy; falsifiable; most are related to concrete physics above TeV or in extra dimensions

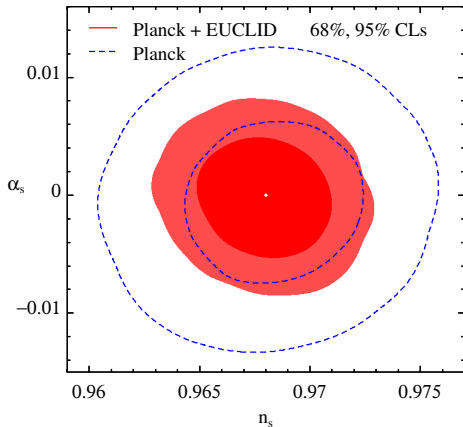
The bench-mark experiments

- ▶ CMB
Planck (3 channels 70GHz, 100GHz, 143GHz, 2.5yr integration);
- ▶ LSS
A EUCLID-like LSS forecast model.
The galaxy power spectrum:

$$P_g(k, z, \mu) = \left(b + \frac{d \ln D}{d \ln a} \mu^2 \right)^2 D^2(z) P_m(k, z=0) e^{-k^2 \mu^2 \sigma^2} .$$

8 redshift bins \times 30 k bins \times 20 μ bins
marginalize over 16 nuisance parameters: $b_1, b_2, \dots, b_8; \sigma_1, \sigma_2, \dots, \sigma_8$.
cut-off at quasi nonlinear scales ($k \sim 0.2 \text{ Mpc}^{-1}$).

EUCLID and n_s , α_s

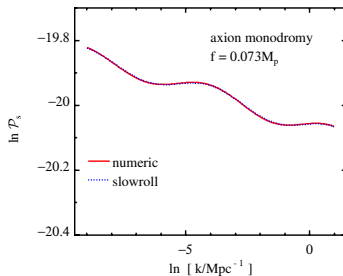
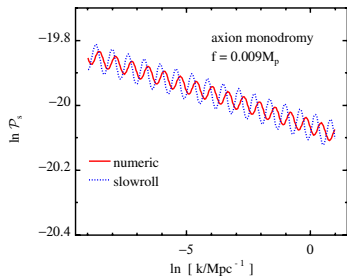


Planck will measure n_s and α_s very accurately. EUCLID can only mildly improve it.

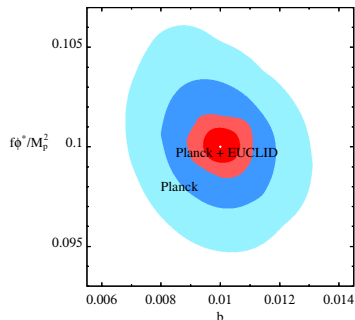
Models that have features in the potential

Axion monodromy inflation model:

$$V(\phi) = \mu^3 \left[\phi + bf \cos\left(\frac{\phi}{f}\right) \right].$$



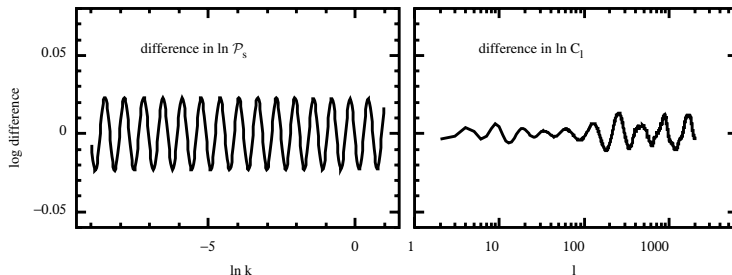
EUCLID and axion monodromy



For $\delta \ln k \sim 0.1$, EUCLID improves the FOM by a factor of about 10.

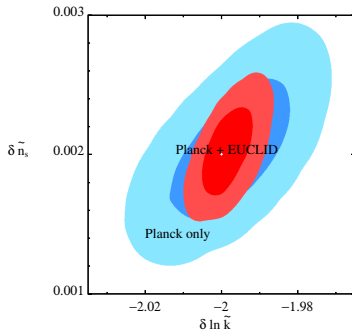
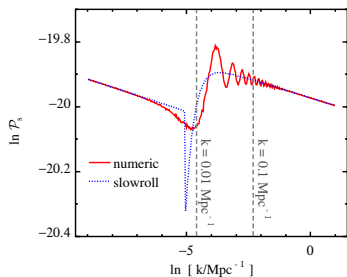
Why EUCLID can measure sharp features better

The difference between monodromy and a linear potential with the same μ^3



CMB measures integrated angular correlations. Sharp features are suppressed.

Another example, ringing feature from Starobinsky potential.



Take-home Message

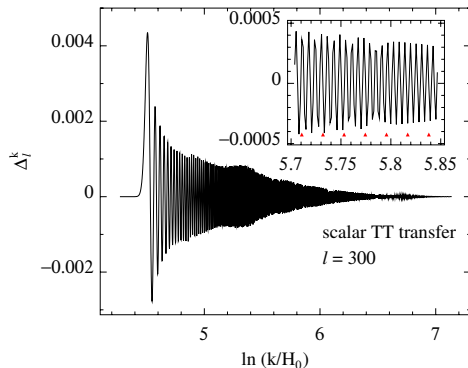
- ▶ Interesting physics can manifest itself in the form of glitches in the primordial scalar power spectrum.
- ▶ EUCLID is a crucial experiment to measure the glitches.

Some Technical Details (for Julien Lesgourgues)...

Towards the quick-oscillation $\delta \ln k \sim 10^{-3}$

$$C_\ell = \int (\Delta_\ell^k)^2 \mathcal{P}_s(k) d \ln k$$

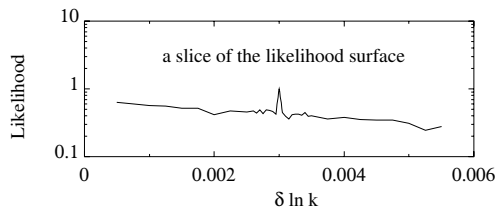
Integration scheme in CAMB, CLASS etc.



- ▶ Brute-force increment of ℓ and k resolutions $\Rightarrow \sim 10^3$ times slower.
- ▶ Better integration scheme using the recurrence relation of $j_\ell(x)$ (implemented in my code).

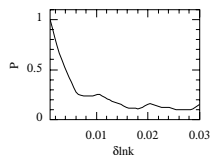
Monodromy likelihood surface: searching the radio band

CMB only
fiducial $\delta \ln k =$
0.003.



- ▶ Quadratic approximation fails (Fisher-matrix does not work).
- ▶ MCMC converges much slower than usual.
- ▶ "Tuned-in channel" can be easily destroyed by varying other parameters. Marginalization is important.

CMB Only:



LSS + CMB:

