



planck



DTU Space
National Space Institute



Science & Technology
Facilities Council



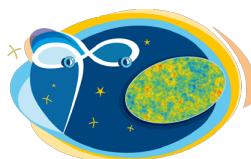
National Research Council of Italy



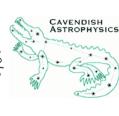
DLR Deutsches Zentrum
für Luft- und Raumfahrt e.V.



MAX-PLANCK-GESELLSCHAFT



HFi PLANCK
a look back to the birth of Universe



MilliLab



Bond since 1993, Canada since 2001, 1st CSA pre-launch contract 2002-09, post-launch 2010-11, 2011-15

Planck 2015 on Cosmic Photons, Phonons, Gravitons & Neutrinos



Dick Bond



CITA
ICAT

Canadian Institute for
Theoretical Astrophysics
L'institut Canadien
d'astrophysique théorique

**“To me every
hour of the light and
dark is a miracle.
Every cubic inch of
space is a miracle.”**

– Walt Whitman

IN EVERY teaspoon of air ~ 5 cubic cm

Ordinary Matter $\sim \text{amu}/\text{nm}^3$ 4.8% O₂ N₂; H, He

THE DARK

Dark Matter

$\sim \text{amu}/\text{m}^3$ $26.0 \pm 1\%$ compressed in MilkyWay $\sim 0.3 \text{ amu}/\text{cm}^3$;
for LHC@CERN-type relics ~ 1 every 10 cm

Dark Energy

\sim vacuum potential density $\sim 3 \text{ amu}/\text{m}^3$ $68.8 \pm 0.9\%$
inflaton-phonon condensate

THE LIGHT

cosmic radiation

the 1st light of the universe $412 \text{ photons}/\text{cm}^3$ 0.005%
cosmic neutrinos \sim cosmic photons Energy fraction $> 0.47\%$ \sim stars
cosmic gravity waves \ll cosmic photons

THE VACUUM

Higgs@CERN vacuum origin of mass

vacuum fluctuations in **phonons** origin of all cosmic structure we see
the vacuum is under **gravitational strain**, differentially accelerating
gravitons vacuum fluctuations a byproduct $r = \text{Tensor}/\text{Scalar} = ?$
isocons vacuum fluctuations in light scalars, ... axions

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Dark Energy

\sim vacuum potential density $\sim 3 \text{ amu}/\text{m}^3$ $68.8 \pm 0.9\%$
inflaton-phonon condensate

$S_{U,\gamma+\nu} \sim 10^{88.6}$

cf. $S_{\text{th},\text{cl}} \sim 10^{76}$ cf. $S_{\text{G},\text{DE}} \sim 10^{121.9}$

THE LIGHT

cosmic radiation

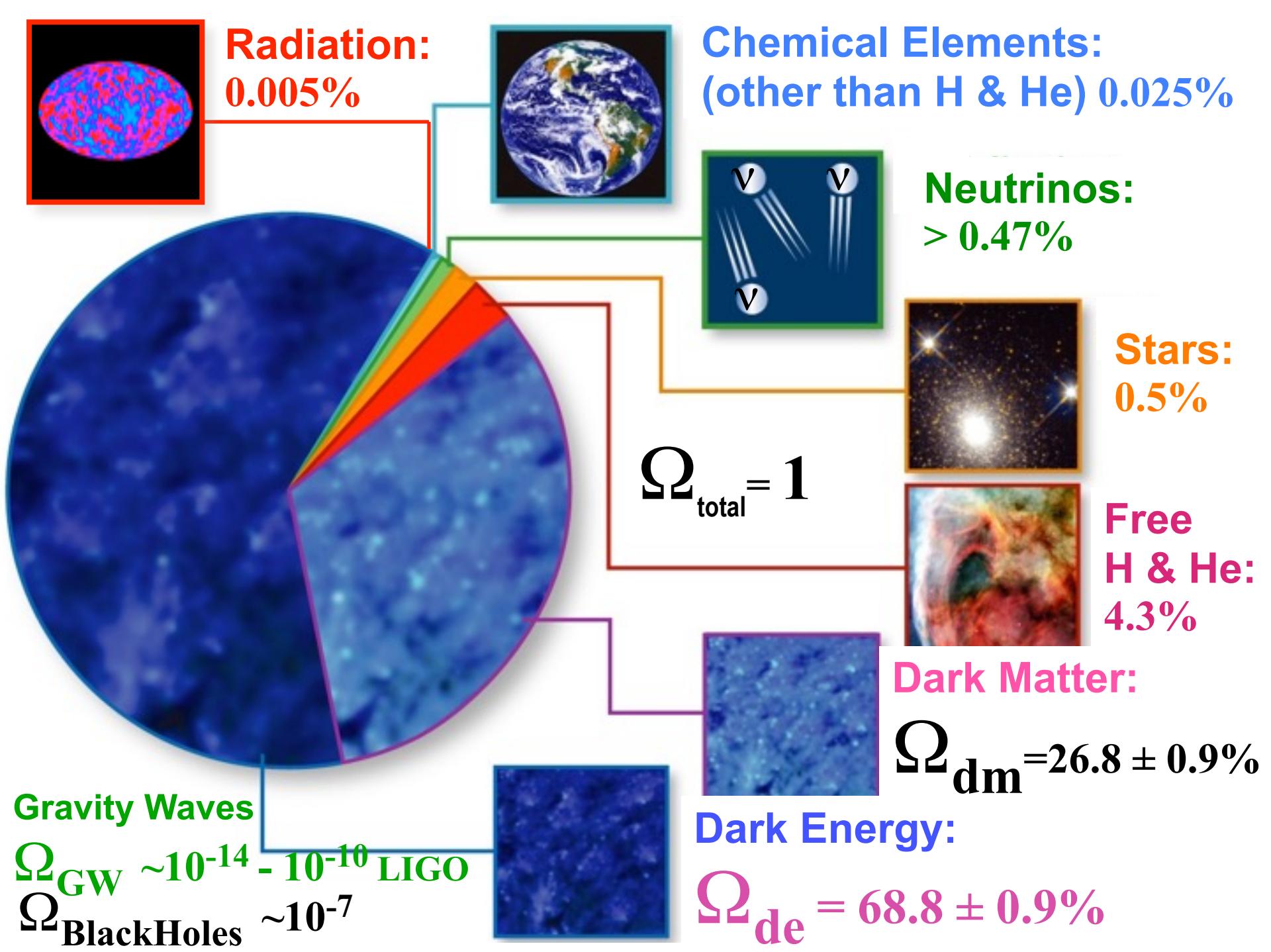
the 1st light of the universe 412 photons/cm³ 0.005% 5.2 bits/ γ
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cosmic gravity waves \ll cosmic photons

5.0 bits/v ($N_{\text{eff}}/3.046$)

THE VACUUM

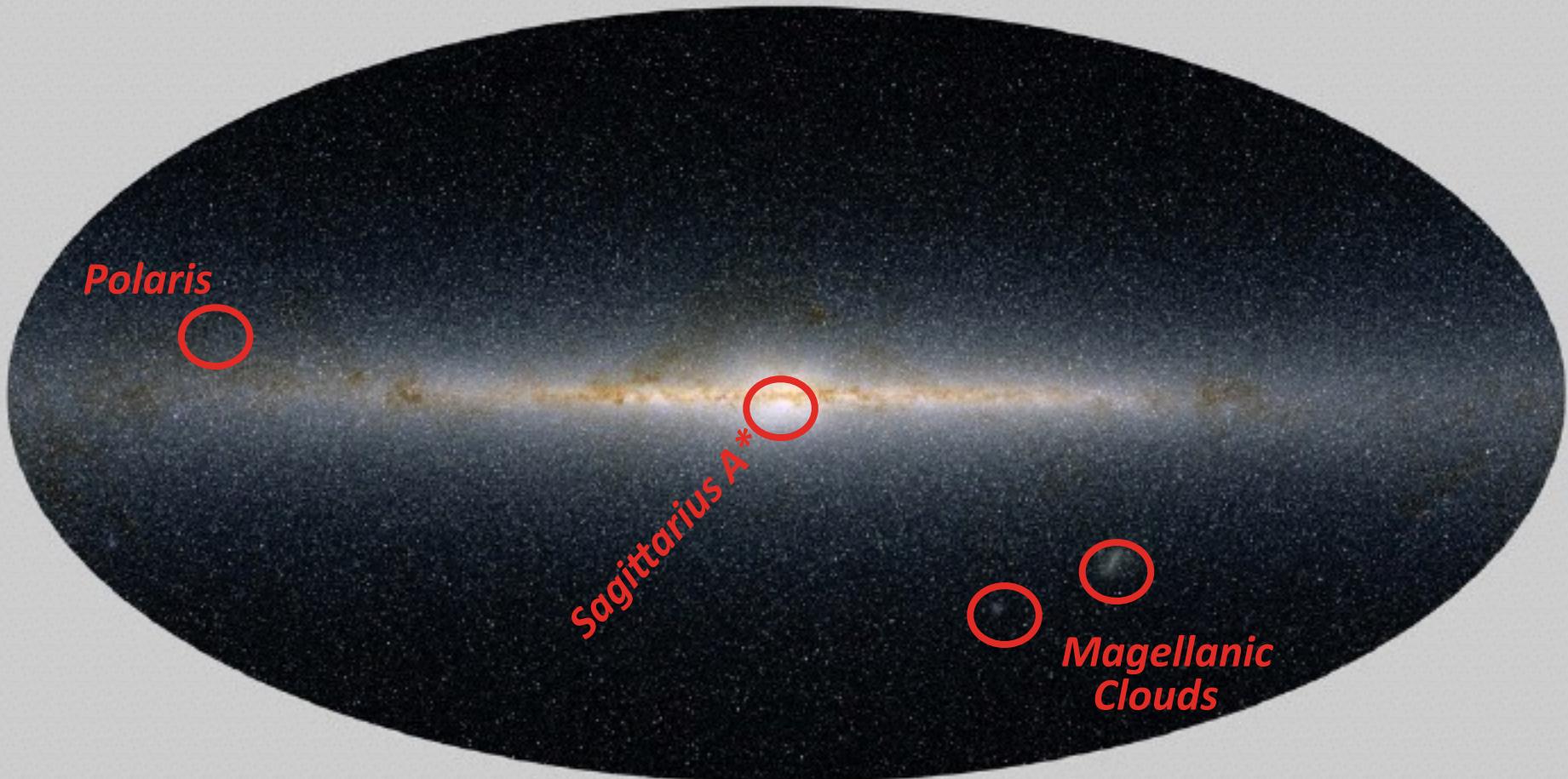
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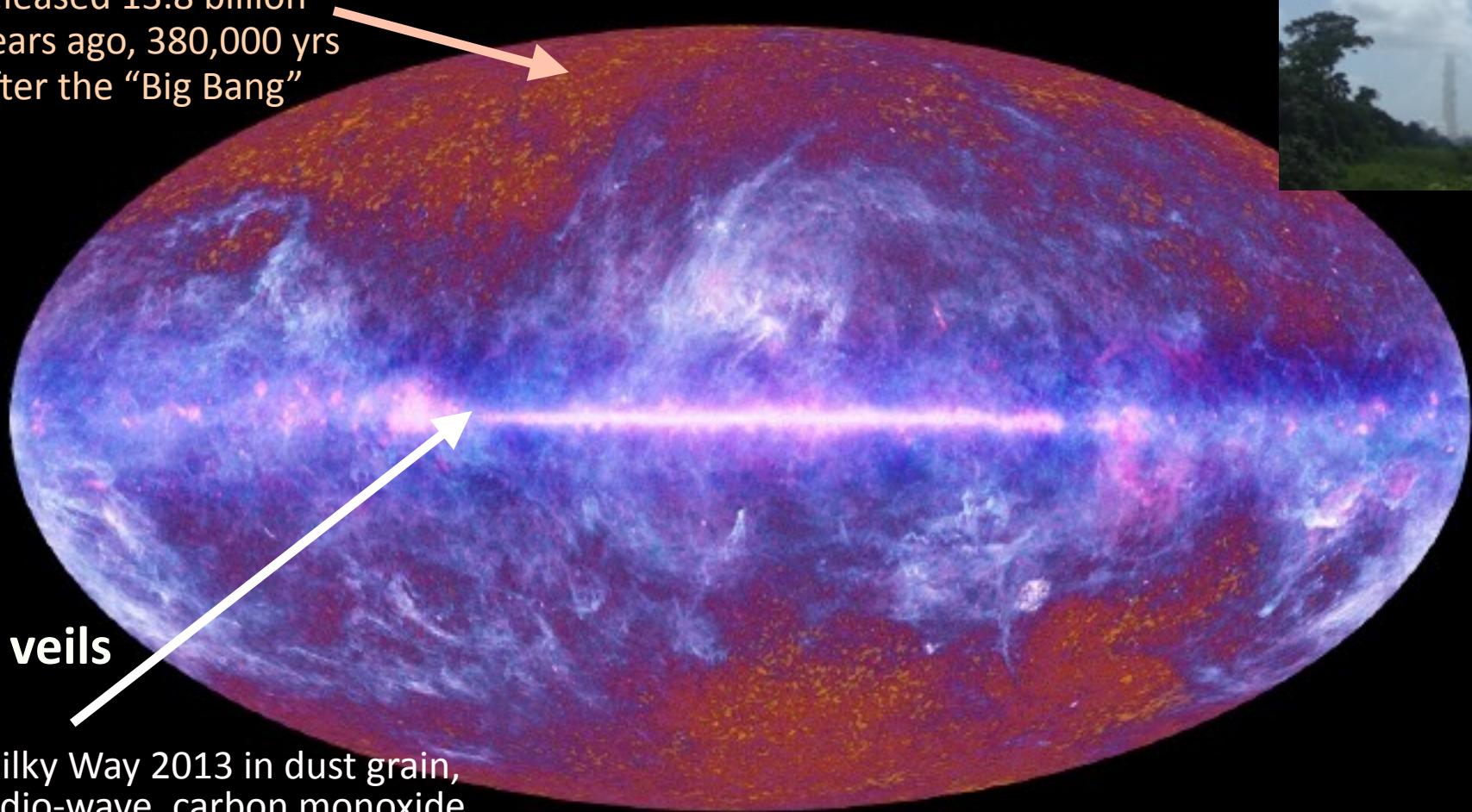


Milky Way in infra-red: half a billion stars, a disk galaxy



COMPLEXITY of here & now

the primordial light,
released 13.8 billion
years ago, 380,000 yrs
after the “Big Bang”

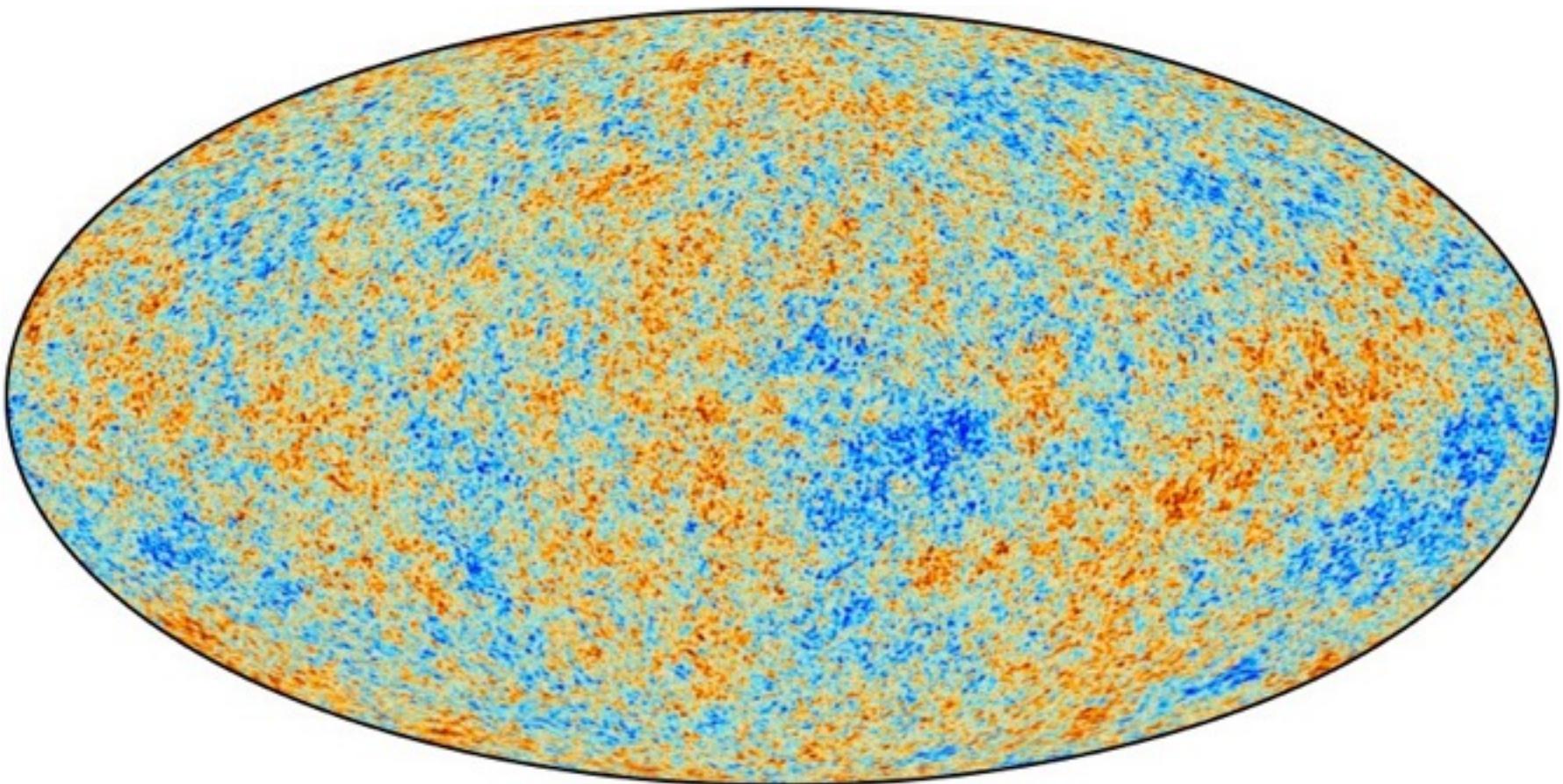


Milky Way 2013 in dust grain,
radio-wave, carbon monoxide
emissions; plus stellar, X-ray,
gamma ray, cosmic ray
emissions ...

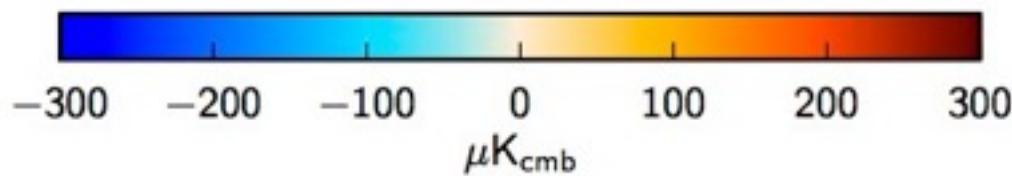
Planck's primordial light unveiled, Feb 5, 2015

reveals the **SIMPLICITY** of primordial cosmic structure

7⁺ numbers, 3 densities, 2+1 early-Universe inflation

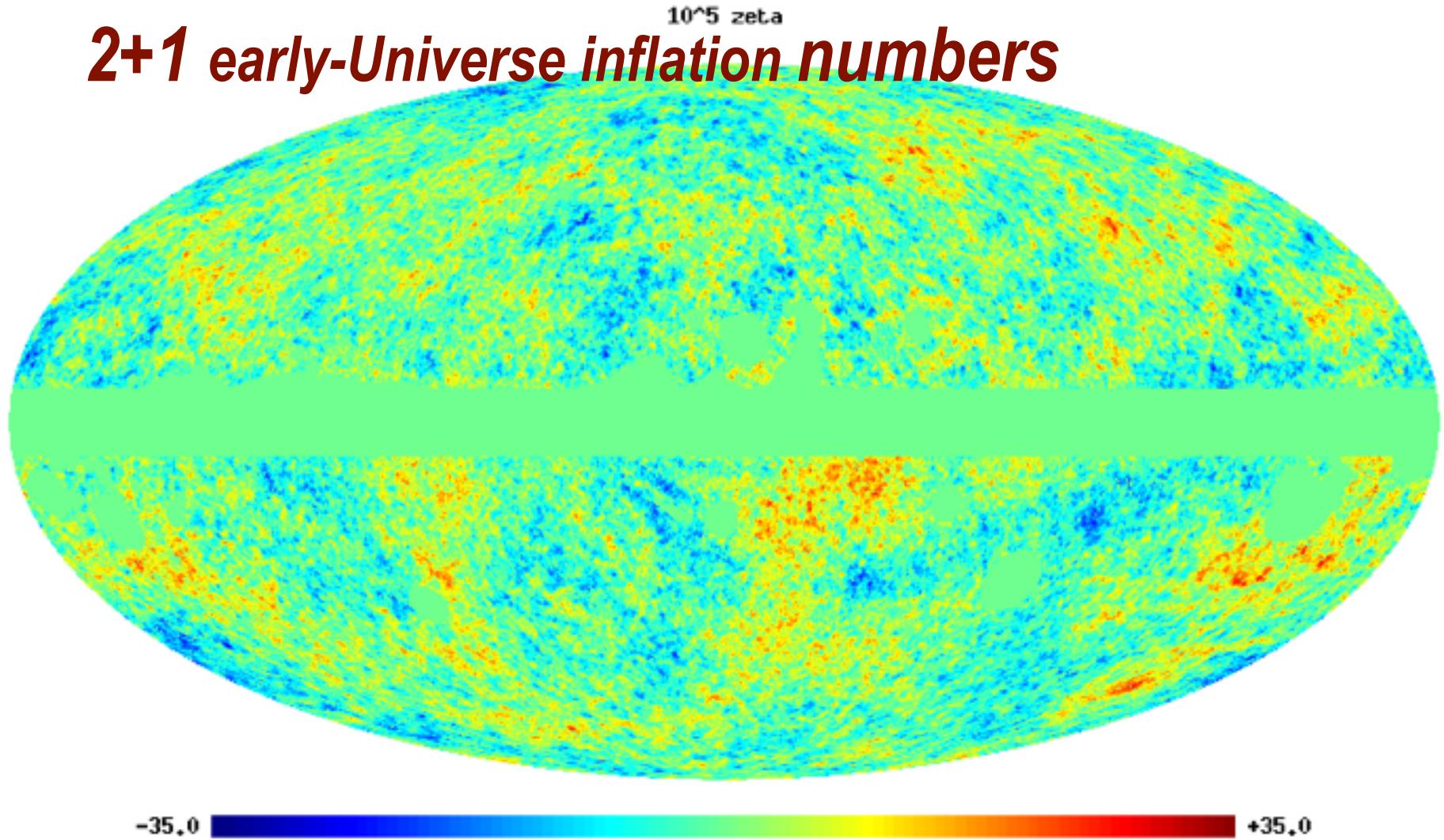


Temperature
changes in
micro-degrees



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reveals the **SIMPLICITY** of primordial cosmic structure

2+1 early-Universe *inflation numbers*



Planck 2015 Feb Papers

Planck_BICEP2_Keck.pdf BKP marginalize over dust polarization => primordial gravity wave constraint no detection

Planck_2015_Results_I_Overview_Products_Results.pdf

Planck_2015_Results_II_LFI_Data_Processing.pdf

Planck_2015_Results_IV_LFI(Beams).pdf

Planck_2015_Results_VI_LFI(Maps).pdf

Planck_2015_Results_VII_HFI_Data_Proc_TOI(Beams).pdf

Planck_2015_Results_VIII_HFI_Data_Proc_Calibration(Maps).pdf

Planck_2015_Results_X_Diffuse_Comp_Sep_Foreground_maps.pdf

*Planck_2015_Results_XIII_Cosmological_Parameters.pdf cf. PCP13 shifts <0.7 σ except $\tau = 0.066 \pm 0.016$,
 $z_{re} = 8.8 + 1.7 - 1.4$ is down;*

Planck_2015_Results_XIV_Dark_Energy_Mod_Gravity.pdf cosmological constant still works well

Planck_2015_Results_XV_Gravitational_Lensing.pdf 40sigma detection

Planck_2015_Results_XVII_Primordial_Non-Gaussianity.pdf limits similar to P13, polarization adds a bit

Planck_2015_Results_XVIII_Background_Geometry_Topology.pdf size/(2 distance to last scattering) > 1

Planck_2015_Results_XIX_Constraints_Primordial_Magnetic_Fields.pdf

Planck_2015_Results_XX_Inflation.pdf m2phi2 ruled out, conformally flattened potentials OK

Planck_2015_Results_XXI_Integrated_Sachs-Wolfe_Effect.pdf

Planck_2015_Results_XXII_Map_Thermal_SZ_Effect.pdf public now, agrees with cluster count cosmology

Planck_2015_Results_XXIV_Cosmology_SZ_Clusters_Counts.pdf tension remains with primary CMB; nu mass?

Planck_2015_Results_XXVII_Second_Planck_Catalogue_SZ_Sources.pdf PSZ2 1652,1203 confirmed

Planck_2015_Results_XXVIII_Planck_Catalogue_Galactic_Cold_Clumps.pdf

Planck+Herschel Launch
May14 09 French Guiana

1.5m telescope,
HFI bolometers @6freq

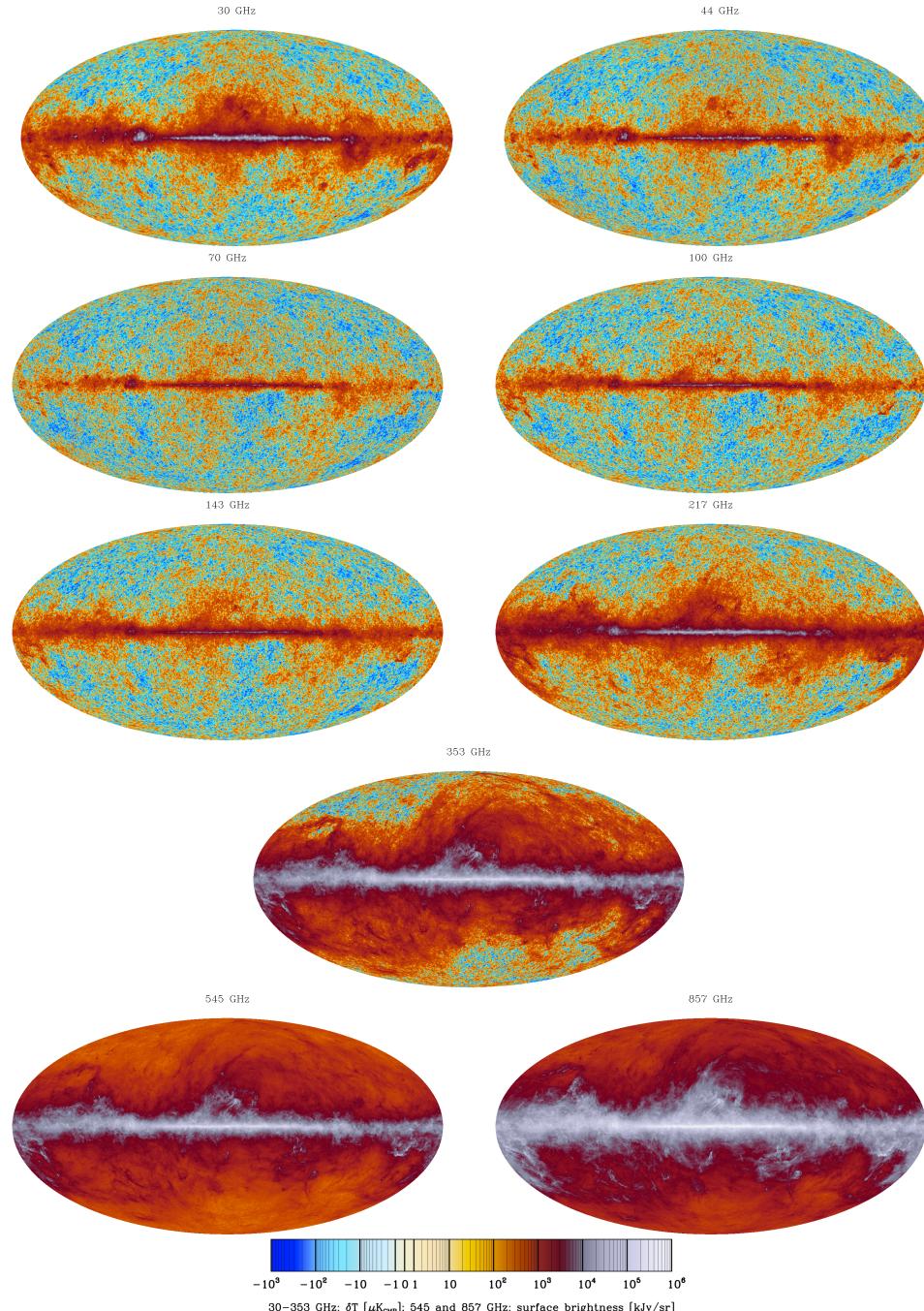
<100mK,

LFI HEMTs@3freq,

some bolometers & all
HEMTS are polarization
sensitive

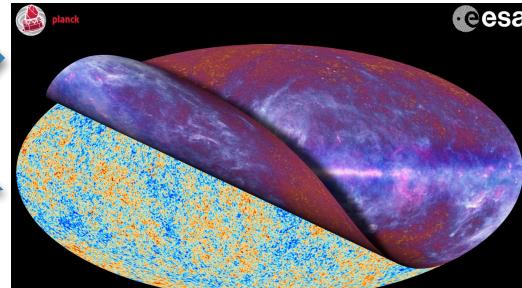
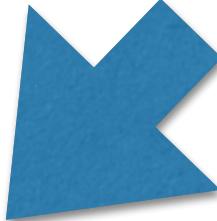
- Left earth at ~10 km/s, 1.5 million km in 45 days, cooling on the way (20K, 4K, 1.6K, 0.1K 4 stage). @L2 on July 2 09; Survey started on Aug 13 09
- spun@1 rpm, 40-50 minutes on the same circle, covered all-sky in ~6 month
- kicked out of L2 Oct13
- 5 HFI all-sky surveys (to Jan 2012) **29 months**
- 8 LFI surveys **48 months**
- Feb15 T some Q,U all-data, more Mar15, Q/U Jun15 refined final set Mar 2016

Planck 2013 Frequency Maps Mar13



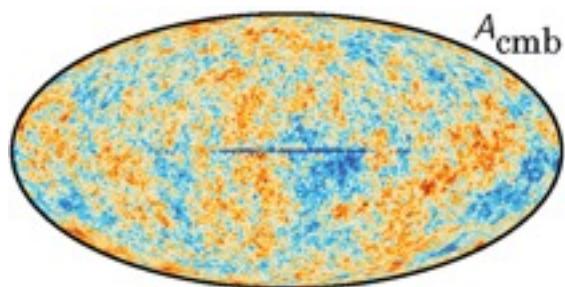
Feb 2015 Planck Component Separated Temperature Maps

esa

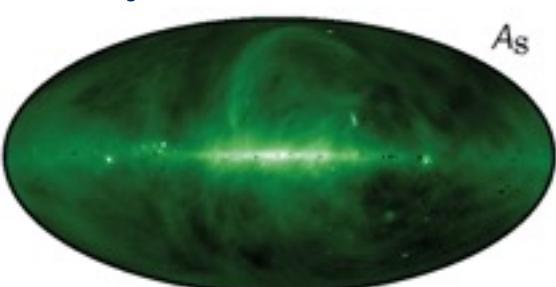


Planck unveils the Cosmic Microwave Background

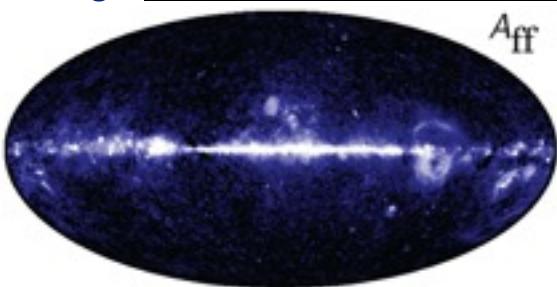
LF Synchrotron + bremsstrahlung



-250 μK 250

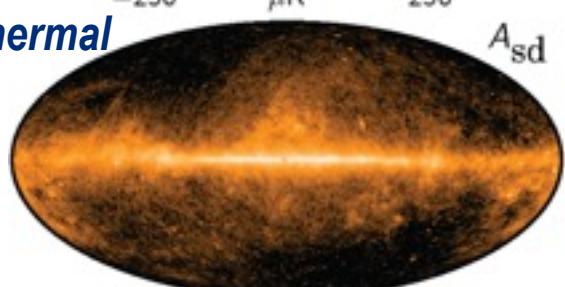


5 K @ 408 MHz 500



0 cm^{-6}pc 1000

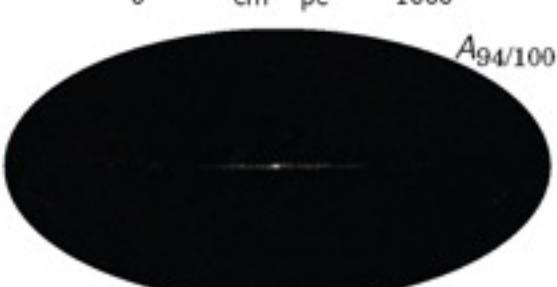
HF Thermal Dust



0.01 mK_{RJ} @ 30 GHz 10

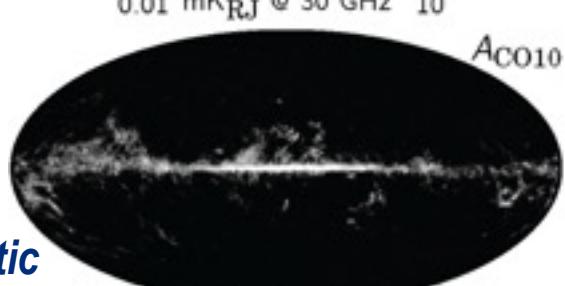


0.001 mK @ 545 GHz 10

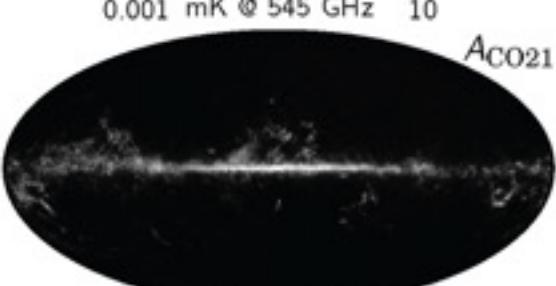


0 $\mu\text{K} @ 100\text{-ds1}$ 100

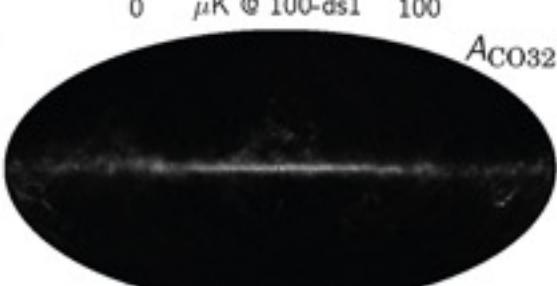
Galactic Carbon Monoxide



0 K km/s 100



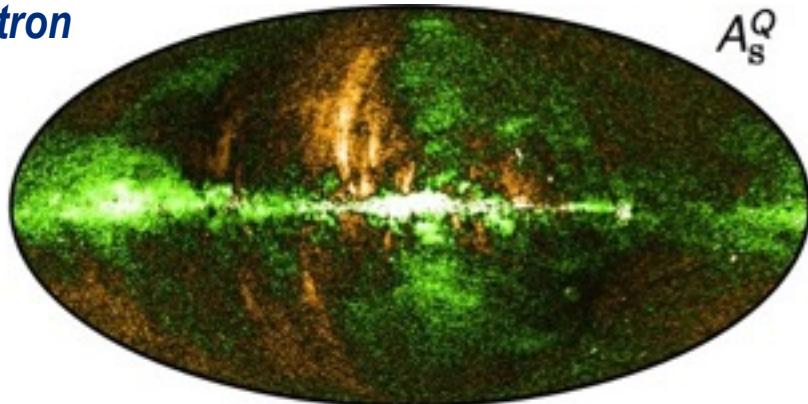
0 K km/s 100



0 K km/s 100

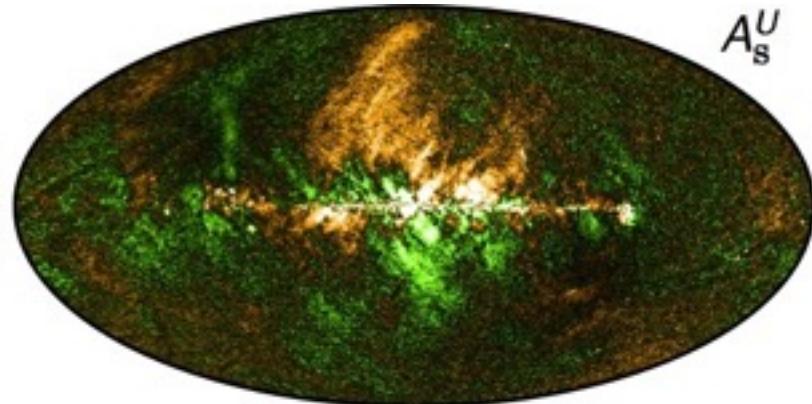
Feb 2015 Planck Component Separated Polarization Maps

LF Synchrotron

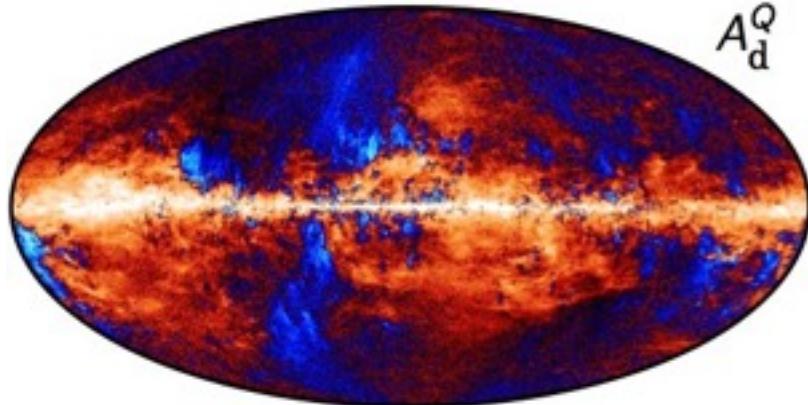


A_s^Q

A_s^U

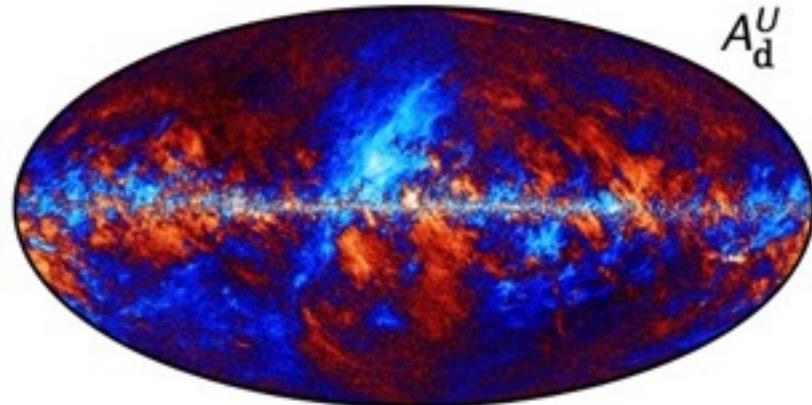


HF Thermal Dust



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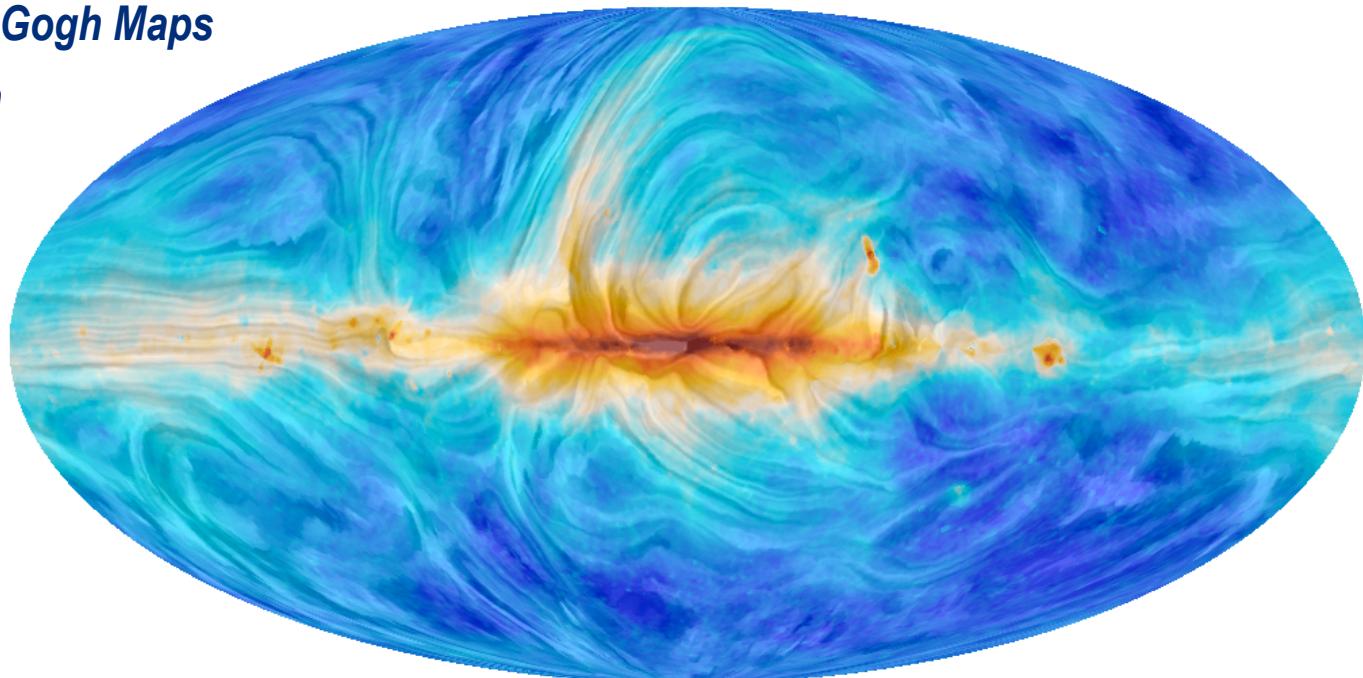


Galactic Dust Polarization Papers

4 in May 2014 on dust polarization, 1 in Sept 2014 power spectra, at high Galactic Latitude r_{dust} vs. BICEP2 claim of $r=0.2 \rightarrow .16 \text{ T/S}$ detection; Feb 15 BKP no r detection < 0.13 , P15 XX $r < .09$

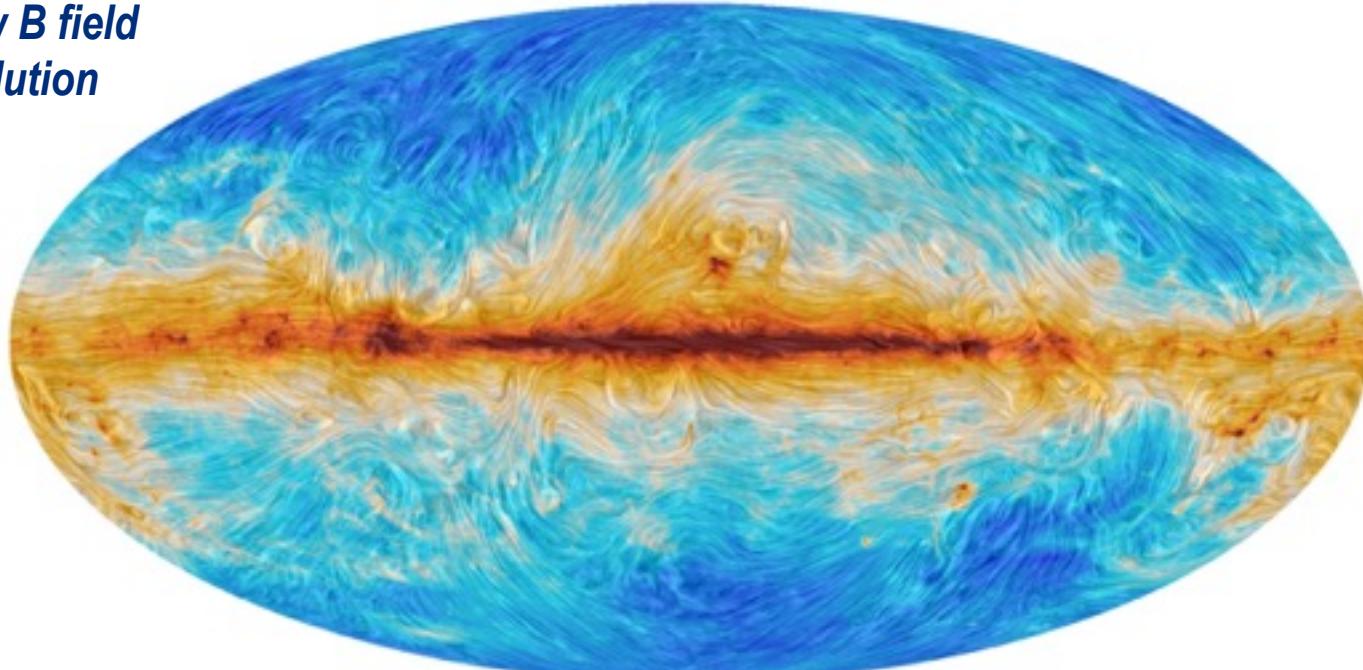
Planck T/P Combined van Gogh Maps

30 GHz LF Synchrotron



353 GHz HF Thermal Dust

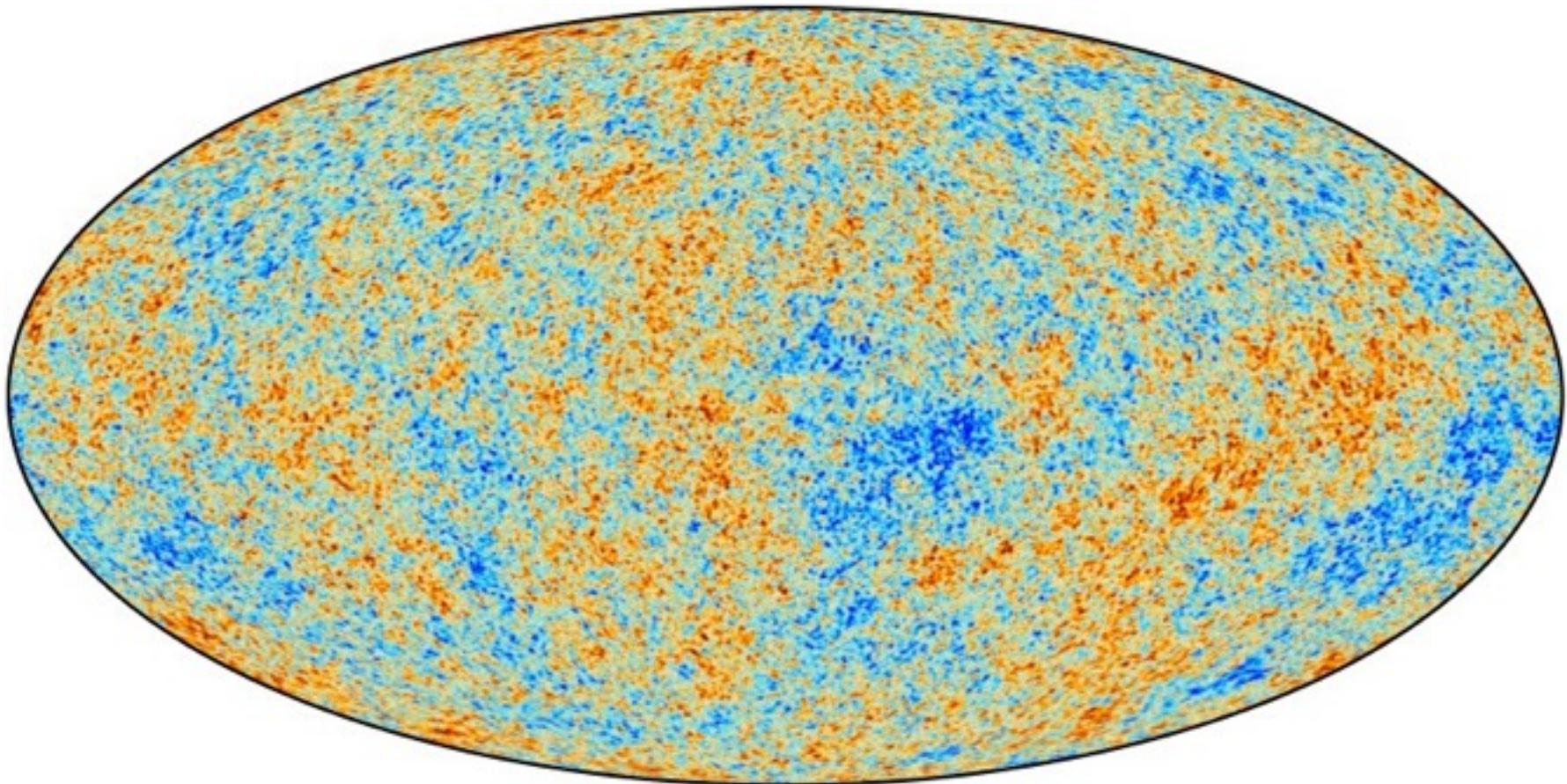
*Polarization used to follow B field
using Line Integral Convolution
a directional “flow” map*



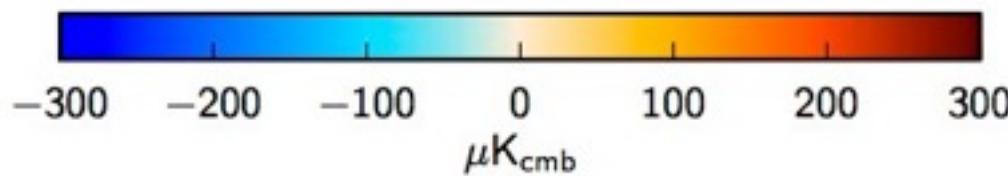
Planck's primordial light unveiled, Feb 5, 2015

reveals the **SIMPLICITY** of primordial cosmic structure

7⁺ numbers, 3 densities, 2+1 early-Universe inflation



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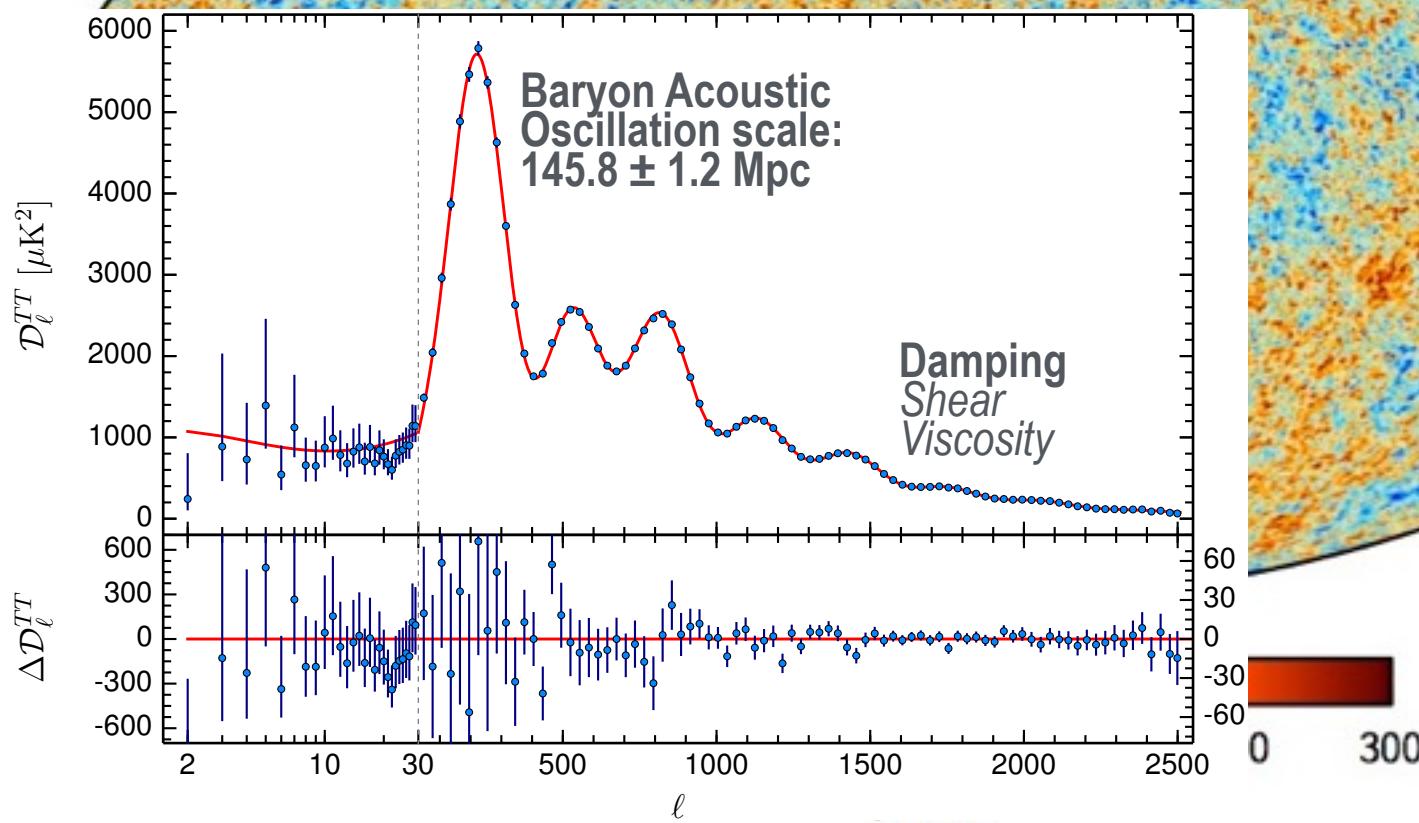


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harmonic analysis of the '*music of the spheres*'
=> *inharmonious, coloured noise in the CMB*

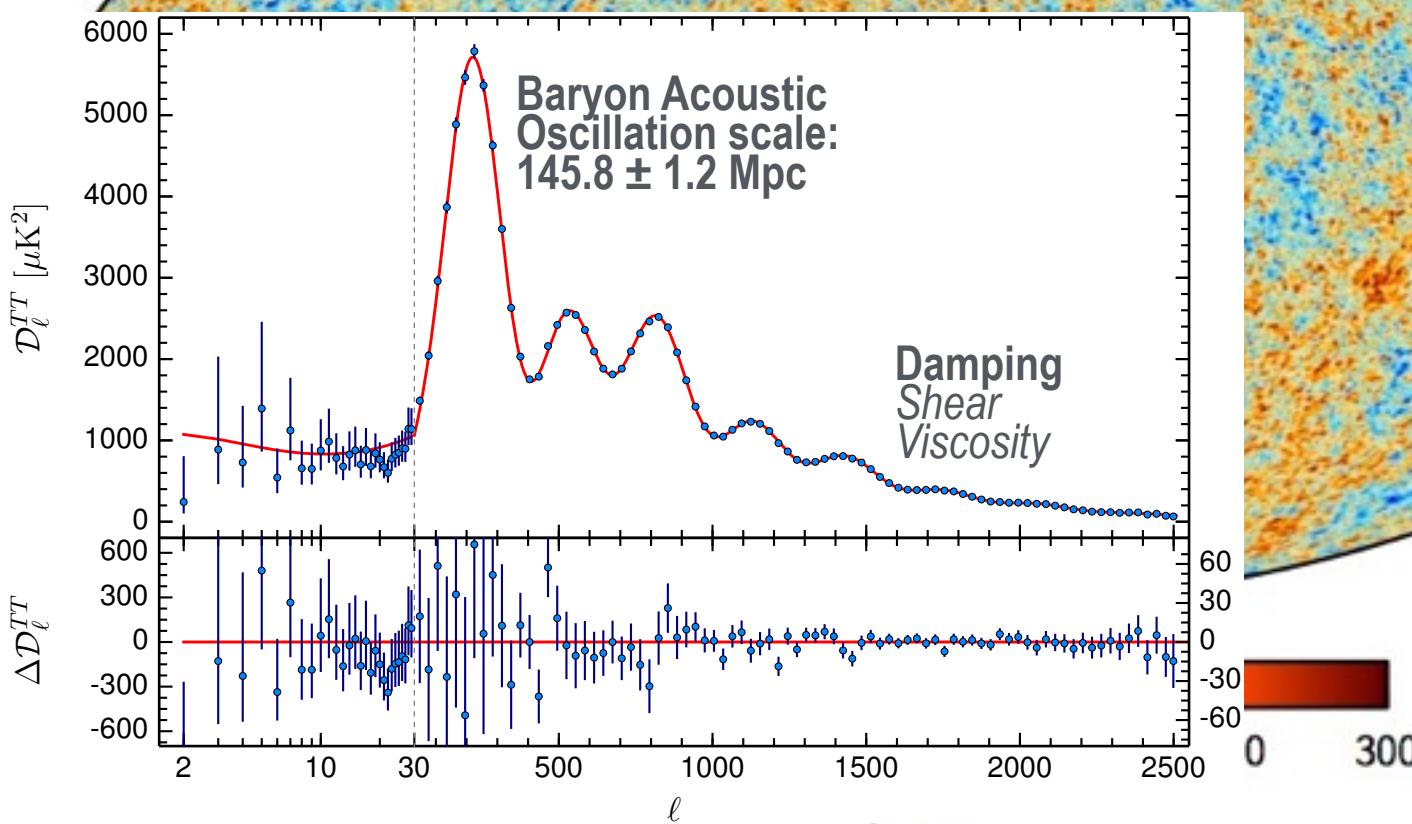


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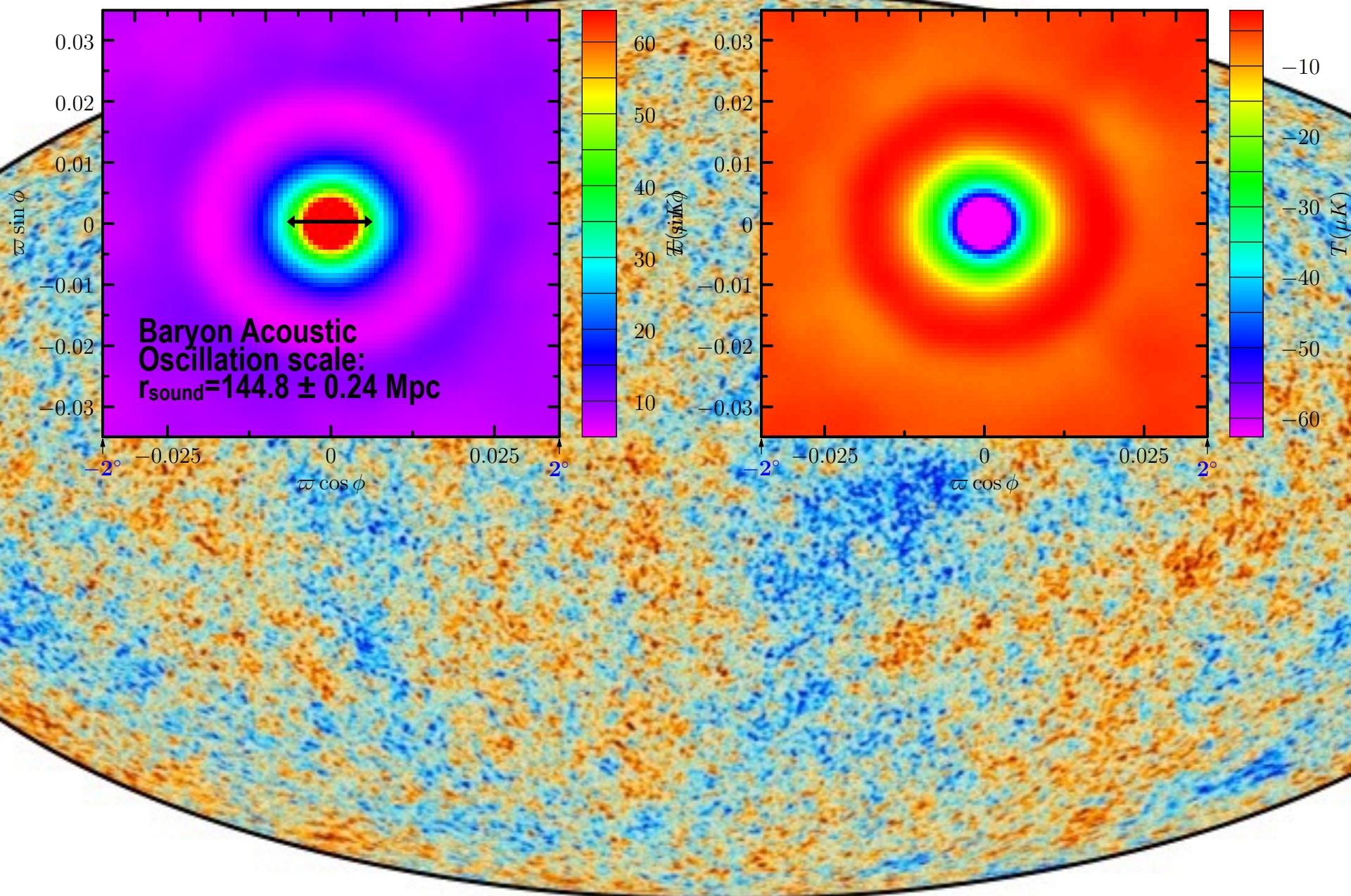
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cosmic sound
realization

Planck13+ reveals primordial sound waves BAO in matter at $a \sim e^{-7} \sim 1/1100$

24645 patches on T maxima, random orientation, threshold $\nu=0$ 24582 patches on T minima, random orientation, threshold $\nu=0$





ν decoupling ~ms

recombination

the nonlinear
COSMIC WEB

dSG/dt
I
N
F
L
A
T
I
O
N

primary anisotropies

- linear perturbations: scalar/density, tensor/gravity wave

$dS/dt > 0$



- tightly-coupled photon-baryon fluid:
Type to enter text oscillations $\delta\gamma$ $v\gamma$ $\pi\gamma$

- viscously damped

DarkM

- polarization $\pi\gamma$
- gravitational redshift Φ SW $d\Phi/dt$



dS/dt > 0



$z \sim 1100$

redshift z

$z \sim 10$

time t

10Gyrs

today

17 kpc
(19 Mpc)

secondary anisotropies

- nonlinear evolution
- weak lensing
- thermal SZ + kinetic SZ
- $d\Phi/dt$
- dusty/radio galaxies, dGs

DarkE



M
I
L
K
Y
W
A
Y



Bayesian flow
prior to posterior via likelihood

dS_{astro} < 0

dS/dt > 0





dS/dt

ν decoupling ~ms

recombination

the nonlinear
COSMIC WEB

I
N
F
L
A
T
T
I
o
N

67
N
dS/dt > 0

$dS/dt > 0$

primary anisotropies

- linear perturbations: scalar/density, tensor/gravity wave $dS/dt > 0$
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Lsound/
ksound

BAO

7

time t

10 Gyr

today

CMB
Lensing
FNS

H0

BAO (z)

cls
ISW

DarkE

0

$z=0$

Bayesian
flow

prior to
posterior
via
likelihood

$dS_{\text{astro}} < 0$

< 0

SIMPLICITY

at $a \sim e^{-7} \sim 1/1100 \Rightarrow$

at $a \sim e^{-67-60} \sim 1/10^{30+25}$

Planck2015 CMB map

reveals primordial sound waves in matter

\Rightarrow learn **contents & structure** at 380000 yr, $a \sim e^{-7}$

\Rightarrow infer the structure far far earlier $a \sim e^{-67-60}$

7⁺ numbers

Early Universe **STRUCTURE**: phonons/strain @ $a \sim 1/10^{30+25}$

“red” noise in phonons/strain: 2 numbers at $a \sim e^{-67-55}$

$$\ln \text{Power}_s \sim \ln 30.6 \times 10^{-10} \pm 0.025$$

$$n_s = 0.968 \pm 0.006 \quad 5\sigma \text{ from 1}$$

Tensor-to-Scalar ratio (GW)
 $r < 0.09$ P15+BKP

SIMPLICITY

at $a \sim e^{-7} \sim 1/1100 \Rightarrow$

at $a \sim e^{-67-60} \sim 1/10^{30+25}$

Planck2015 early U structure map

reveals primordial sound waves in matter

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\Rightarrow infer the structure far far earlier $a \sim e^{-67-60}$

10^5 zeta

2⁺ numbers

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-35.0

+35.0

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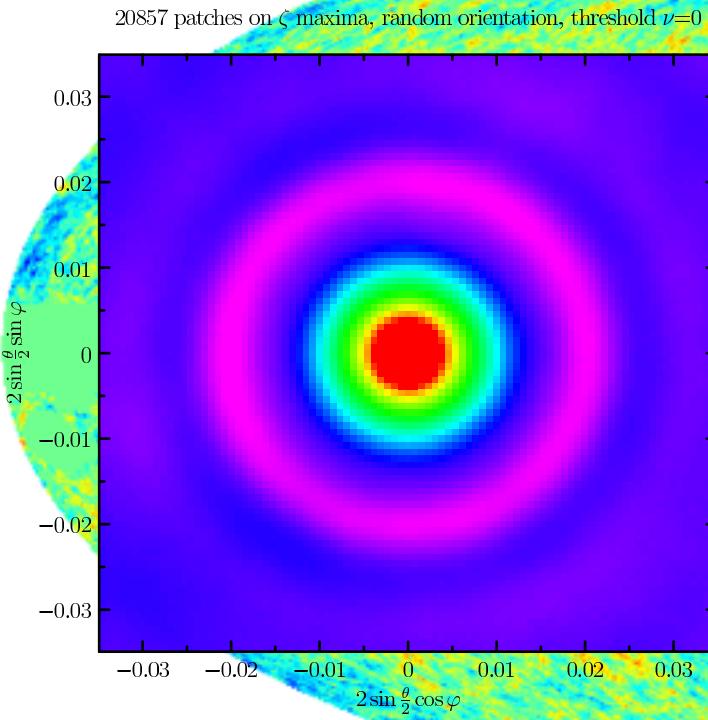
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stacked
 $\langle \zeta_{dv} | \zeta_{dv-pk} \rangle$



$10^5 \zeta$

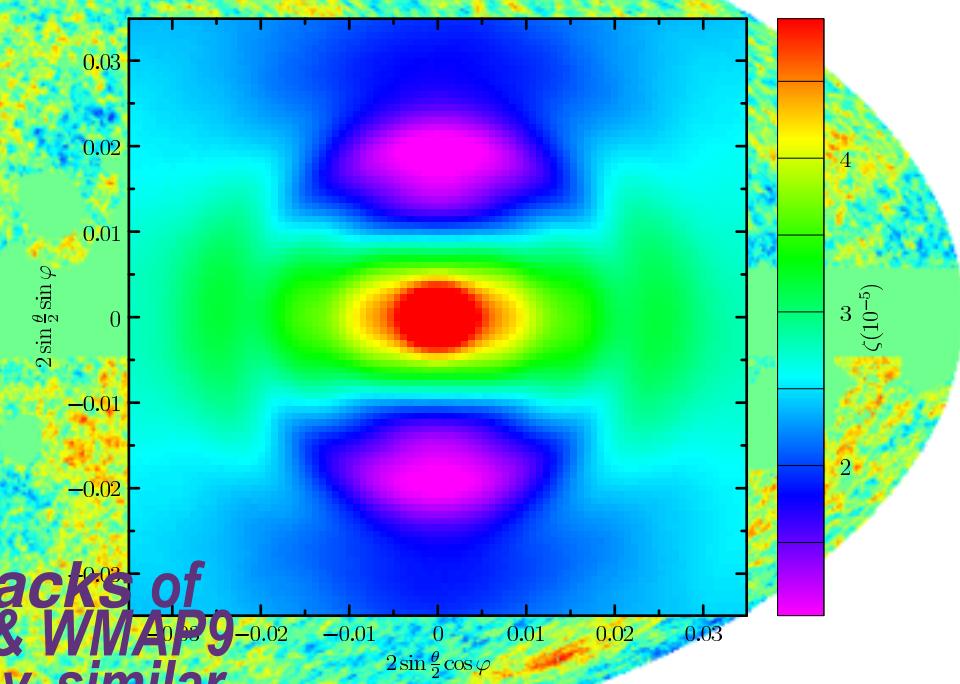
stacked

$\langle \zeta_{dv} | \text{oriented } \zeta_{dv-pk} \rangle$

20854 patches on ζ maxima, oriented, threshold $\nu=0$

2^+ numbers

ζ stacks of
P13 & WMAP9
look v. similar
simulations
look v. similar



-35.0

+35.0

Quadratic $\ln \mathcal{P}_\zeta(\ln k)$ Maps
aka Radical Compressions
=> ultra-early Universe sound/phonon spectrum

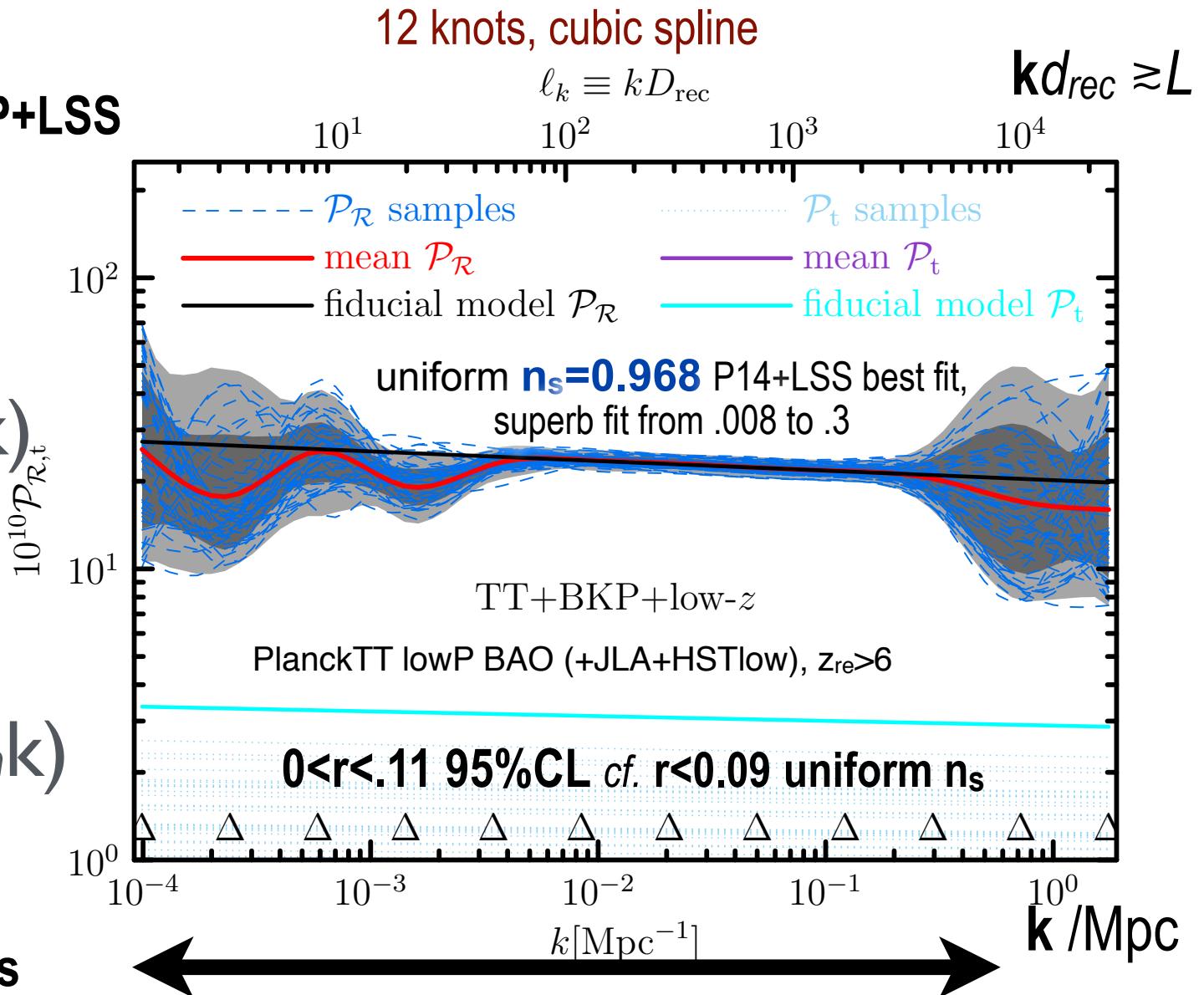
Planck15+BKP+LSS

$\ln \mathcal{P}_\zeta(\ln k)$

r - \mathcal{P}_ζ partial
degeneracy
if r floats

$\ln \mathcal{P}_{\text{GW}}(\ln k)$

9 e-folds



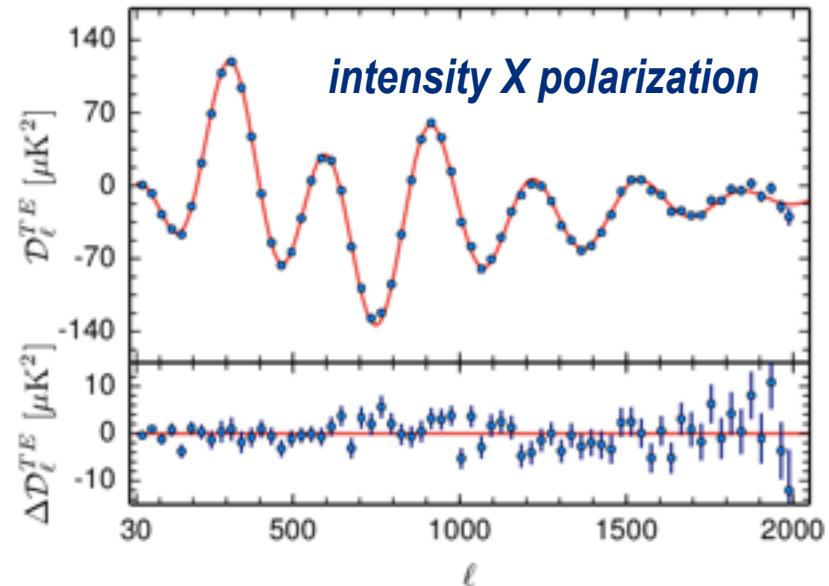
**Mar 2014: bicep2
GW detection $r \sim 0.2$
=>BKP Feb 2015**

**Planck: thou shalt
not ignore dust
polarization $r < 0.13$
P15+BKP $r < .09$ 95%CL**

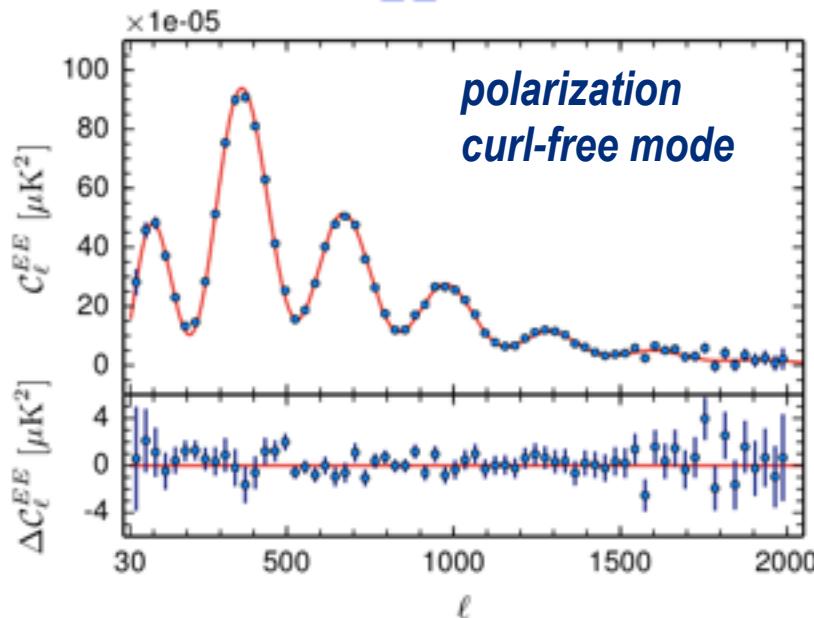
harmonic analysis of polarization

Planck 2015 TE/EE cf. TT => constrains subdominant primordial power contributions not phase-locked with the acoustic-peaks of the pure adiabatic case => constrain isocon spectra /parameters

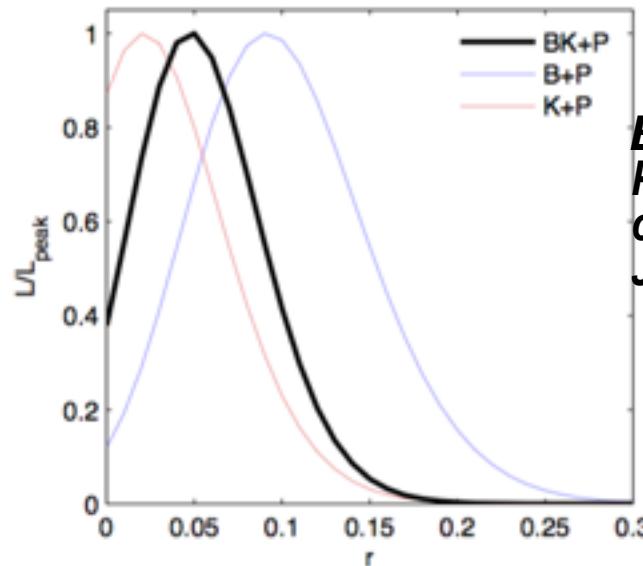
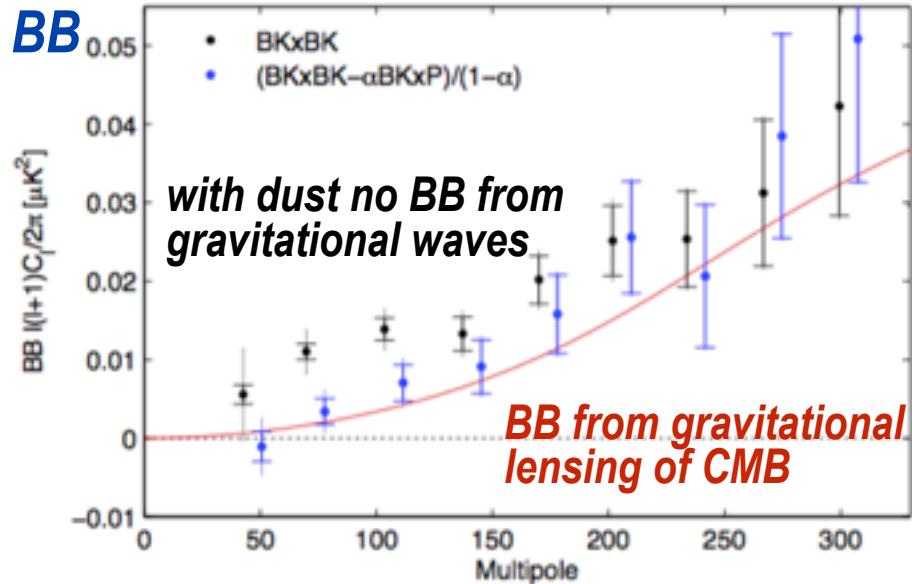
TE



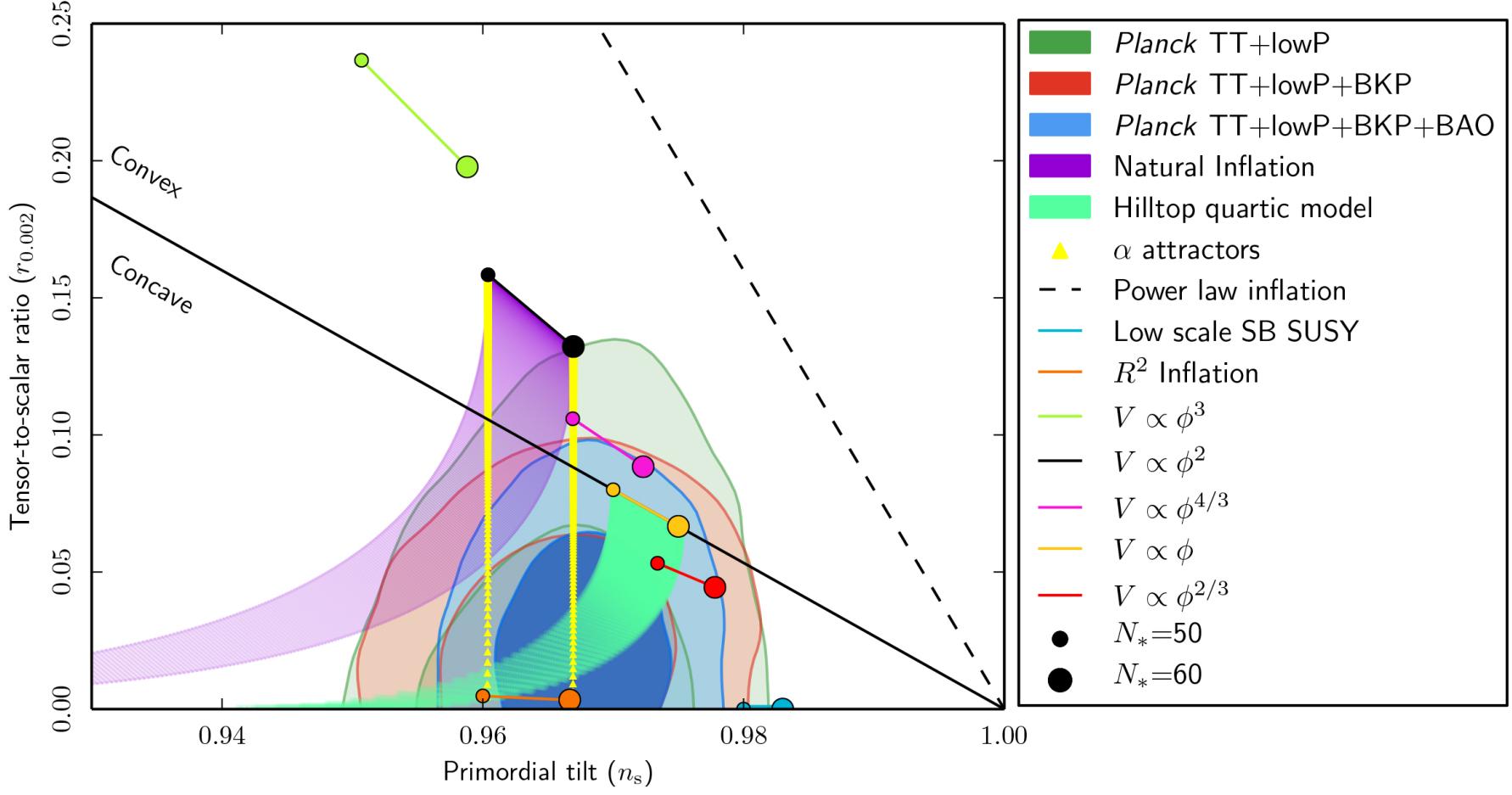
EE



BB

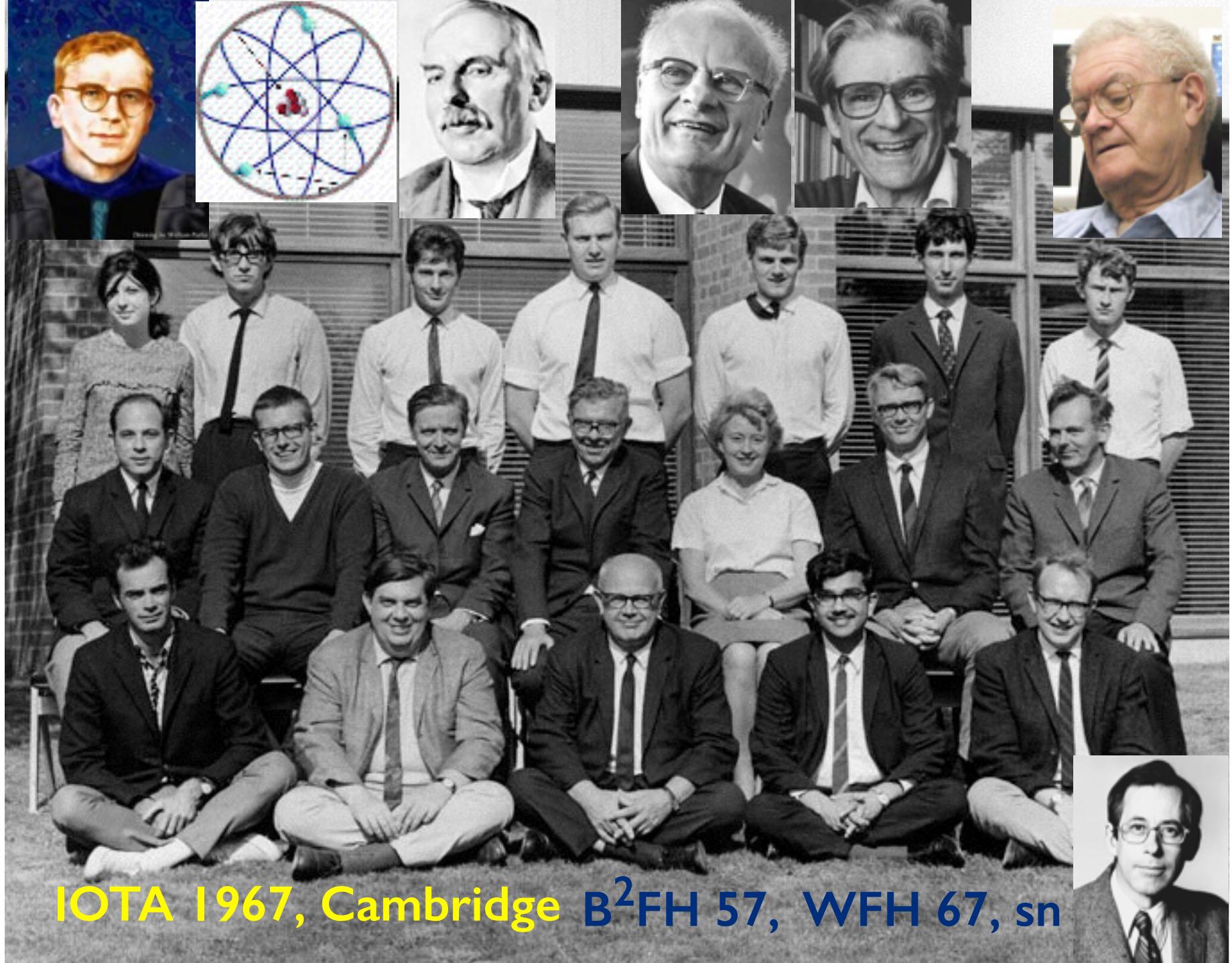


BKP Bicep2, Keck,
Planck cross
correlation analysis
Jan 30, 2015



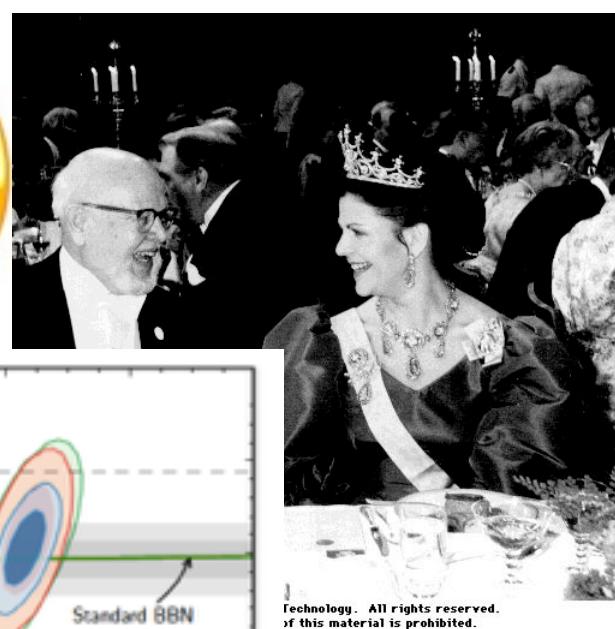
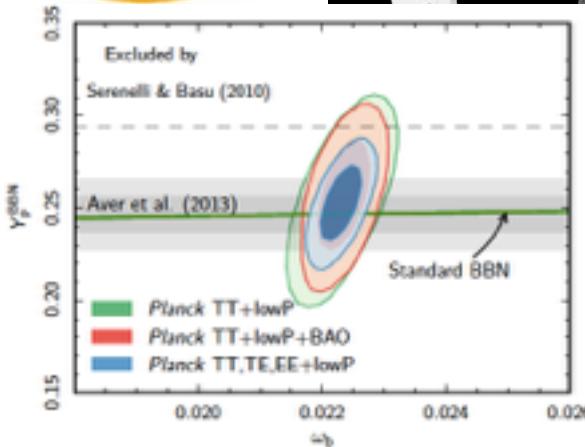
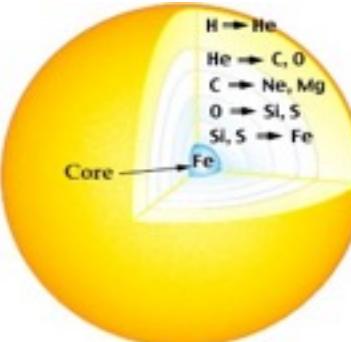
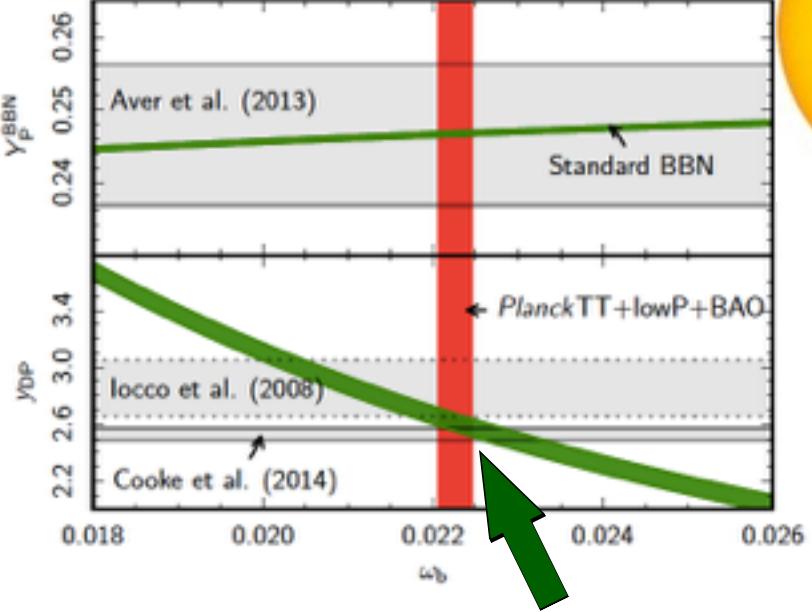
$0 < r < .11$ 95%CL P15+BKP 12 knots
cf. P15+BKP $r < 0.09$ uniform n_s

cf. P15+TE,EE loP $r < 0.10$ uniform n_s cf.
P15+loP+WMAP $r < 0.09$ uniform n_s



IOTA 1967, Cambridge B²FH 57, WFH 67, sn

Baryometers



	pre-boom	boom+	boom+cbi	boom+cbi+acbar	wmap1+
Ω _b h ²	January 2000 $0.0339^{+0.0443}_{-0.0246}$	January 2002 $0.0222^{+0.0025}_{-0.0021}$	June 2002 $0.0221^{+0.0024}_{-0.0020}$	January 2003 $0.0221^{+0.0023}_{-0.0018}$	March 2003 $0.0233^{+0.0013}_{-0.0013}$

0.0226 ± 0.0006 wmap3+acbar+cbi+... LSS

0.0233 ± 0.0005 wmap5+acbar+cbi+b03+.+WL+LSS+SNI+Lya

0.02214 ± 0.00024 Planck13+WP+hiL+BAO

0.02229 ± 0.00033 Planck15 TT,TE,EE +loP+BAO

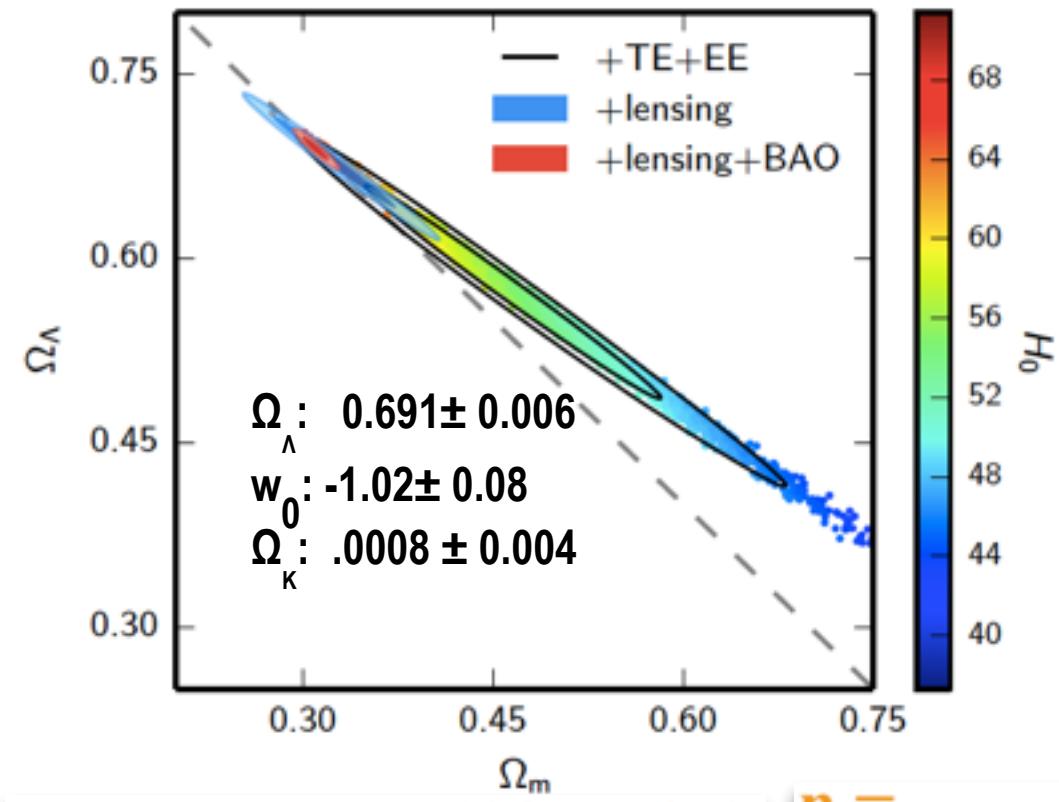
Dark Energy \Rightarrow inflation now

CMB lensing breaks “geometrical degeneracy”:

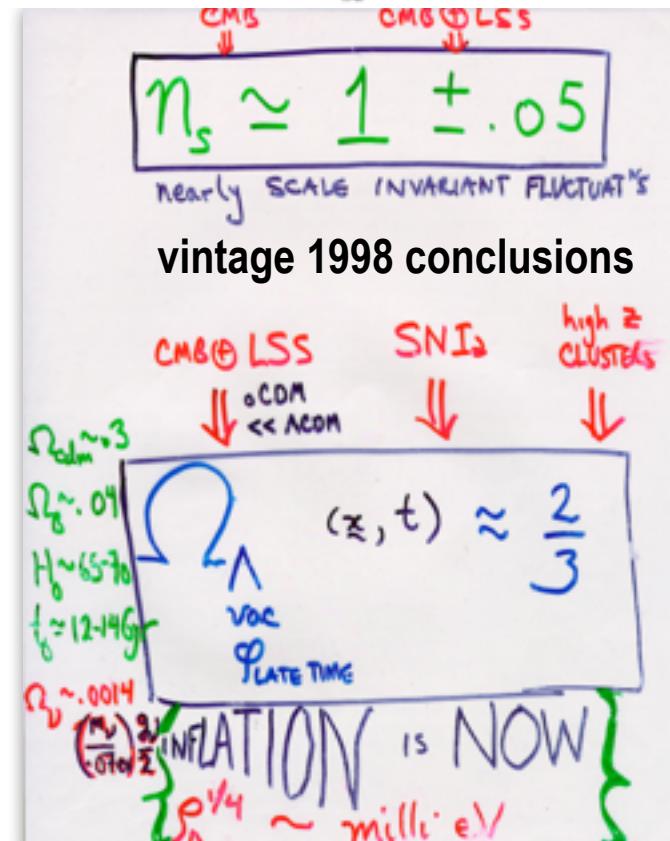
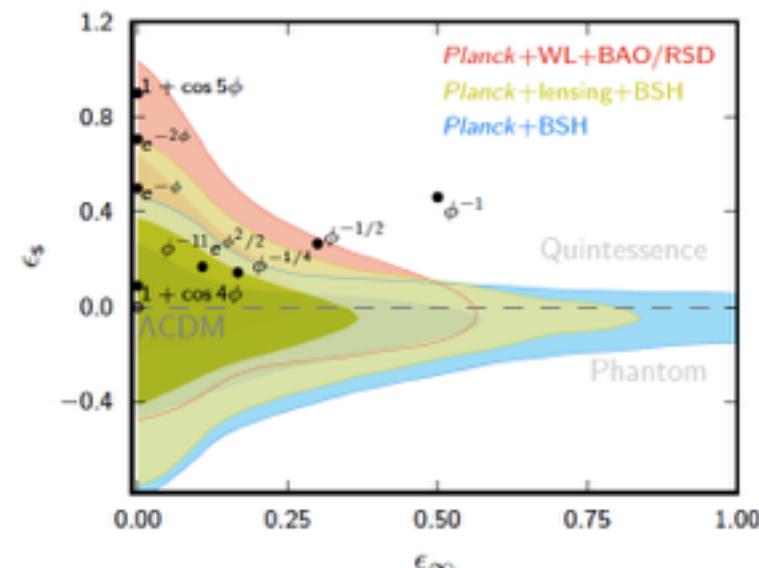
Planck alone cf. Planck+BAO

**Planck15 cf. CMB+LSS history
of $\Omega_\Lambda = PE_{de}/E_{crit}$**

also Sherwin+11: ACT



B+Jaffe'96, '98
 $\Omega_\Lambda \approx 2/3 \pm .07$ +LSS
 $n_s = .98 \pm .07$
 $.96 \pm .06$



NEUTRINOS: number of species, sum of masses

thermal SZ effect Compton cooling of high pressure / entropy electrons by the CMB

Planck2015 PSZ2: 1652 clusters, 1203 confirmed, SPT 224 =>747cls, ACT 91 cls

cf. X-ray sample from ROSAT+ All-sky distribution of MCXC clusters ~1600 (Piffaretti et 10)

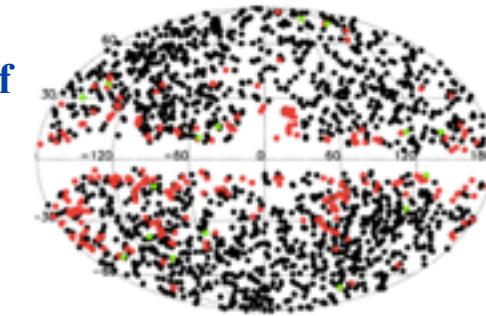
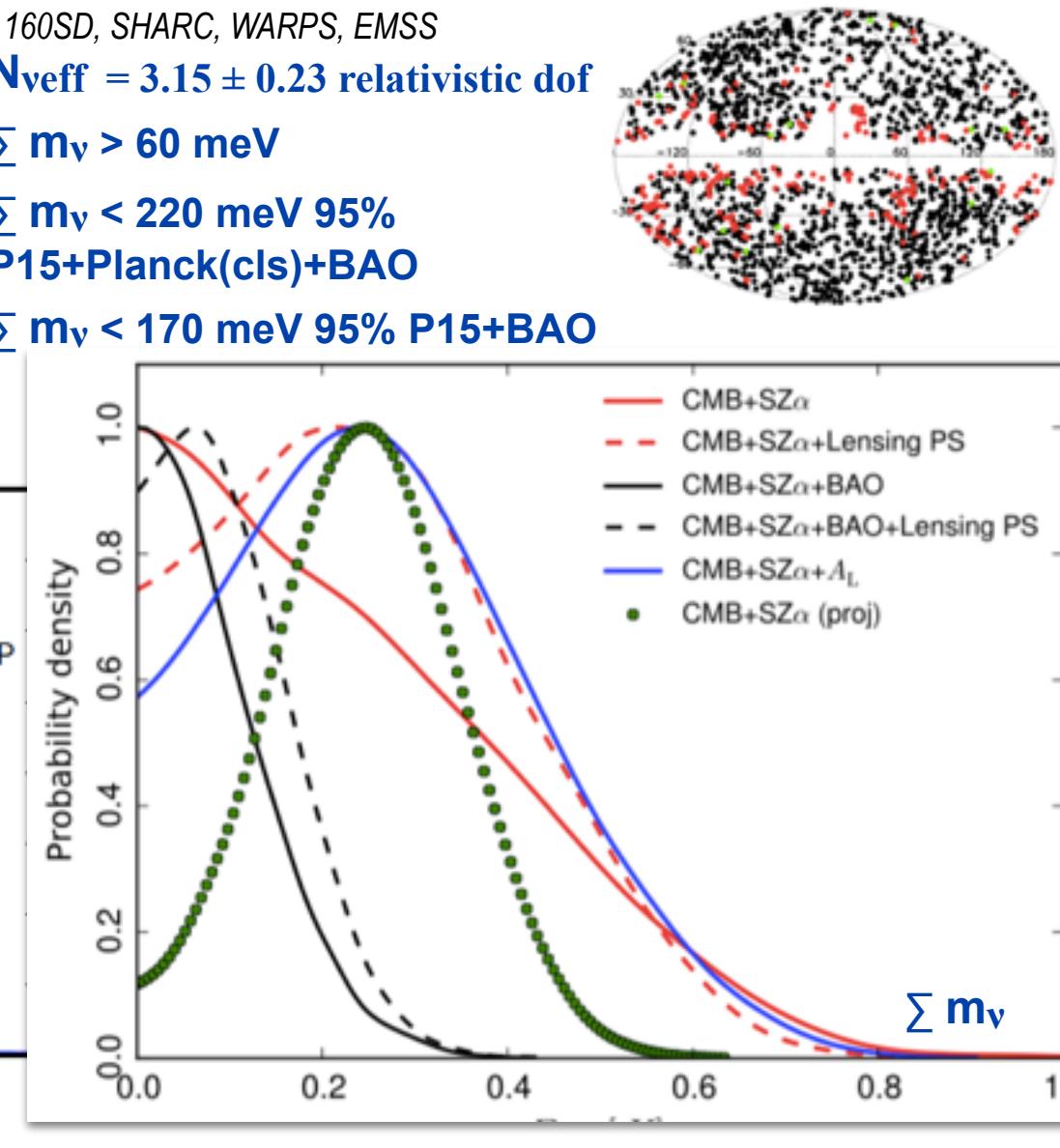
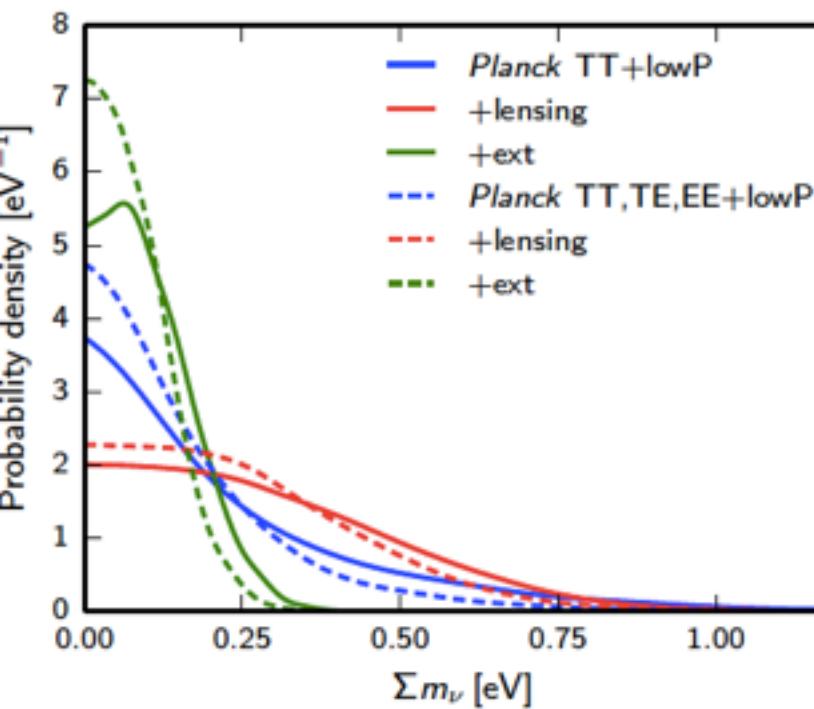
REFLEX, BCS, SGP, NEP, MACS, CIZA, 400SD, 160SD, SHARC, WARPS, EMSS

mild Tension: primary vs. clusters $N_{\text{eff}} = 3.15 \pm 0.23$ relativistic dof

$$\sigma_{8\text{SZ}} \sim 0.78$$

cf.

primary $\sigma_8 = 0.816 \pm 0.009$
gastrophysical problem with cls?
or higher ν mass relief of tension?

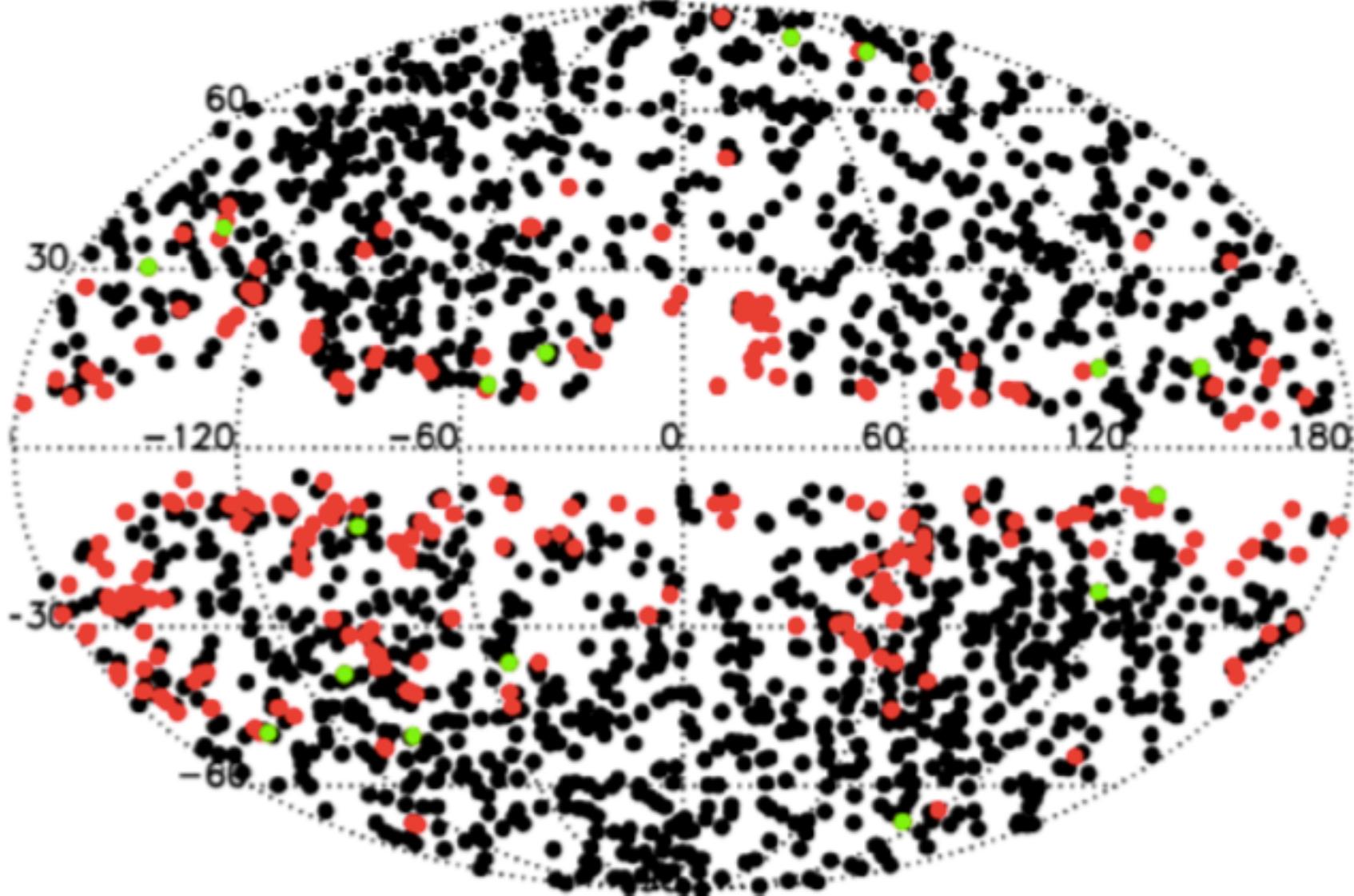


Compton cooling of high pressure / entropy electrons by the CMB

thermal SZ effect Planck2015 1652 clusters, SPT 224 =>747cls, ACT 91 cls

PSZ2: 1652 clusters, 1203 confirmed

cf. X-ray sample from ROSAT+ All-sky distribution of MCXC clusters ~1600 (Piffaretti et 10)
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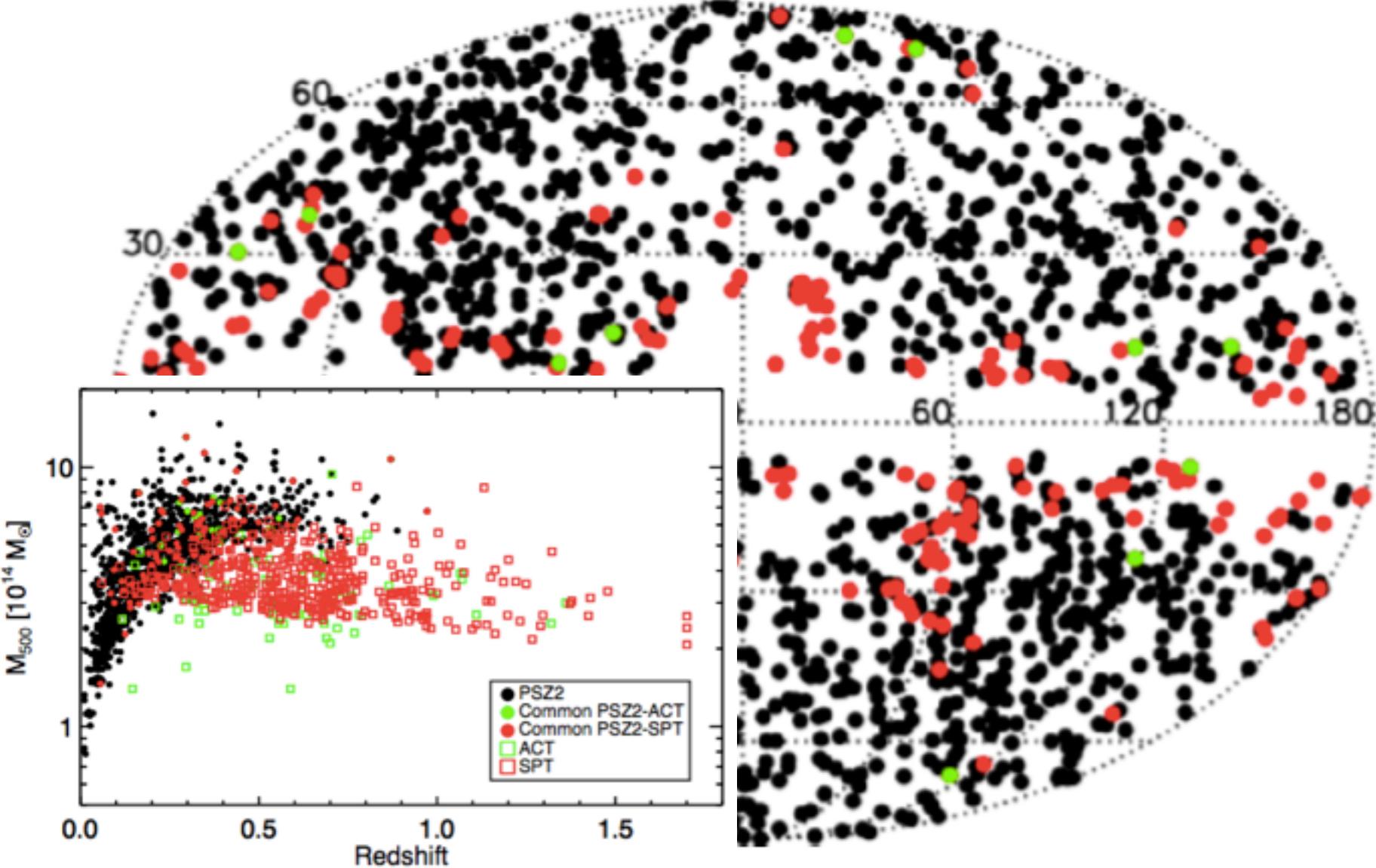


Compton cooling of high pressure / entropy electrons by the CMB

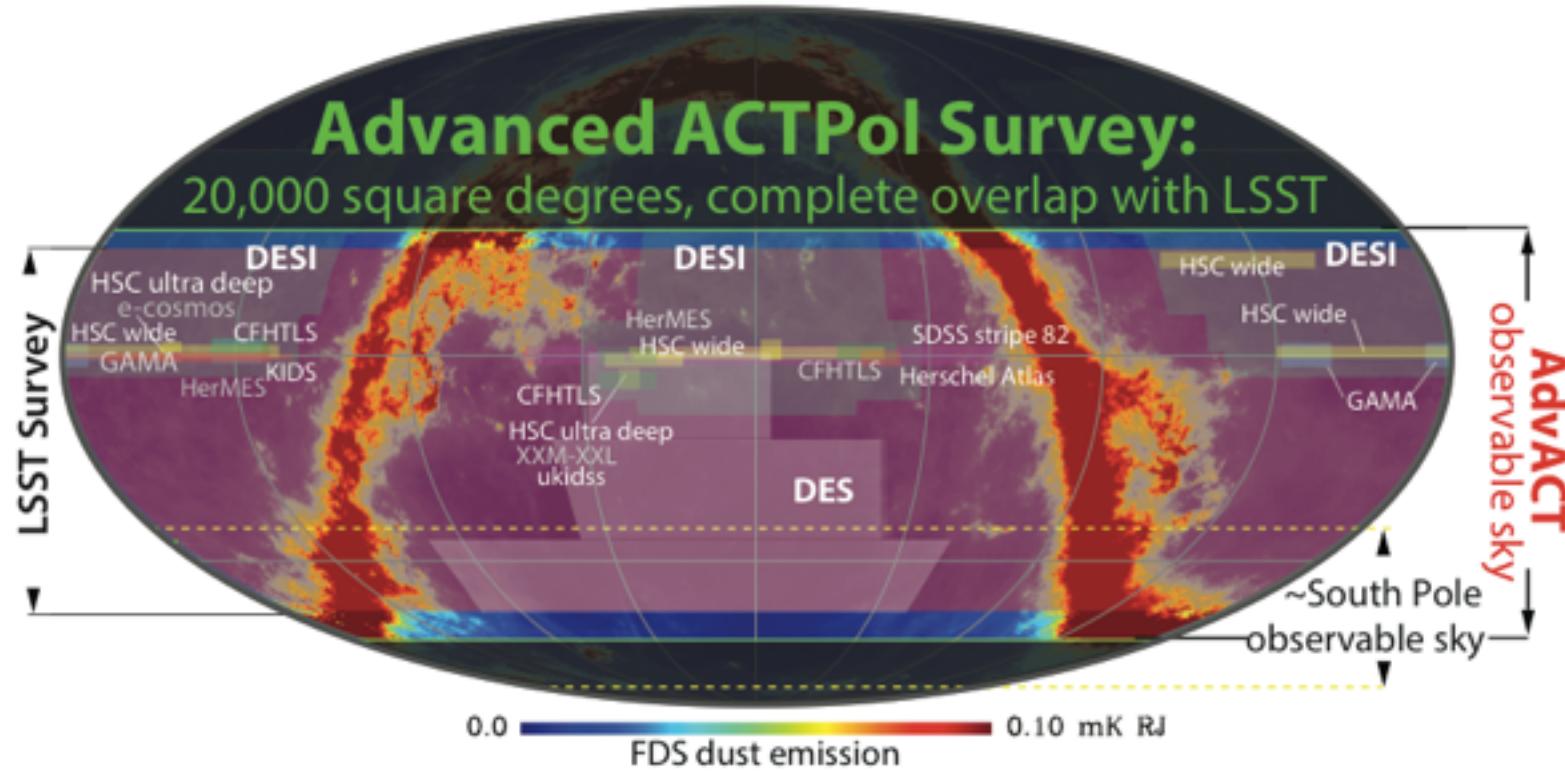
thermal SZ effect Planck2015 1652 clusters, SPT 224 =>747cls, ACT 91 cls

PSZ2: 1652 clusters, 1203 confirmed

cf. X-ray sample from ROSAT+ All-sky distribution of MCXC clusters ~1600 (Piffaretti et 10)
REFLEX, BCS, SGP, NEP, MACS, CIZA, 400SD, 160SD, SHARC, WARPS, EMSS



Advanced ACTPol (AdvACT) Observations



- $\sim 20,000 \text{ deg}^2$ survey ($f_{\text{sky}} \sim 0.5$) with complete LSST overlap as well as DES, ALMA, and other observatories located in Chile
- Substantial overlap with spectroscopic surveys (SDSS, PFS, DESI)

Planck 2015 Feb Papers

Planck_BICEP2_Keck.pdf BKP marginalize over dust polarization => primordial gravity wave constraint no detection

Planck_2015_Results_I_Overview_Products_Results.pdf

Planck_2015_Results_II_LFI_Data_Processing.pdf

Planck_2015_Results_IV_LFI(Beams).pdf

Planck_2015_Results_VI_LFI(Maps).pdf

Planck_2015_Results_VII_HFI_Data_Proc_TOI(Beams).pdf

Planck_2015_Results_VIII_HFI_Data_Proc_Calibration(Maps).pdf

Planck_2015_Results_X_Diffuse_Comp_Sep_Foreground_maps.pdf

*Planck_2015_Results_XIII_Cosmological_Parameters.pdf cf. PCP13 shifts <0.7 σ except $\tau = 0.066 \pm 0.016$,
 $z_{re} = 8.8 + 1.7 - 1.4$ is down;*

Planck_2015_Results_XIV_Dark_Energy_Mod_Gravity.pdf cosmological constant still works well

Planck_2015_Results_XV_Gravitational_Lensing.pdf 40sigma detection

Planck_2015_Results_XVII_Primordial_Non-Gaussianity.pdf limits similar to P13, polarization adds a bit

Planck_2015_Results_XVIII_Background_Geometry_Topology.pdf size/(2 distance to last scattering) > 1

Planck_2015_Results_XIX_Constraints_Primordial_Magnetic_Fields.pdf

Planck_2015_Results_XX_Inflation.pdf m2phi2 ruled out, conformally flattened potentials OK

Planck_2015_Results_XXI_Integrated_Sachs-Wolfe_Effect.pdf

Planck_2015_Results_XXII_Map_Thermal_SZ_Effect.pdf public now, agrees with cluster count cosmology

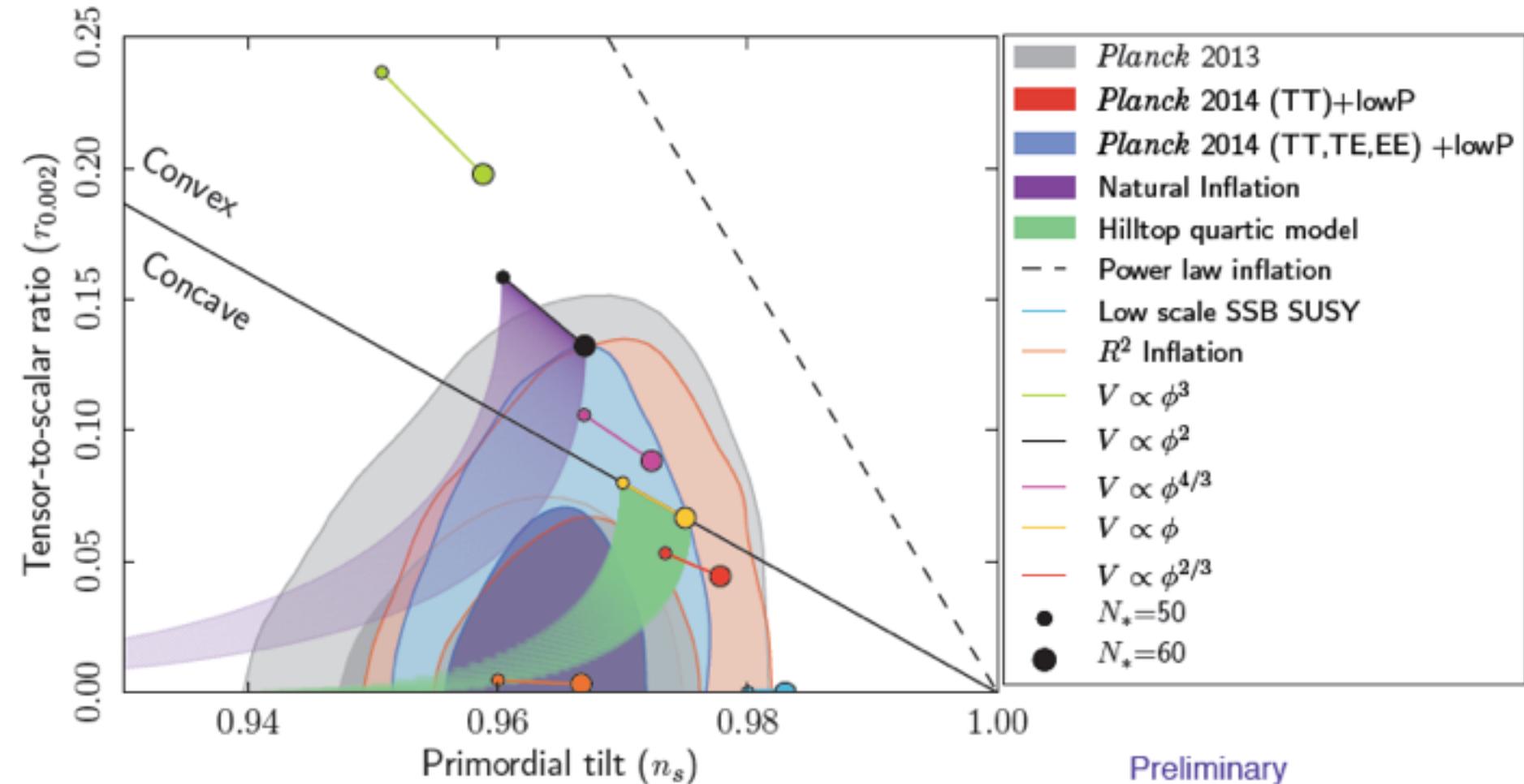
Planck_2015_Results_XXIV_Cosmology_SZ_Clusters_Counts.pdf tension remains with primary CMB; nu mass?

Planck_2015_Results_XXVII_Second_Planck_Catalogue_SZ_Sources.pdf PSZ2 1652,1203 confirmed

Planck_2015_Results_XXVIII_Planck_Catalogue_Galactic_Cold_Clumps.pdf

END

Inflationary models & Planck



$r_{0.002} < 0.10$ (95 %CL, Planck TT + lowP) Preliminary

$r_{0.002} < 0.11$ (95 %CL, Planck TT + lensing + lowP)

$r_{0.002} < 0.10$ (95 %CL, Planck TT, TE, EE + lowP)

$r_{0.002} < 0.09$ (95 %CL, Planck TT + lowP/wWMAP)

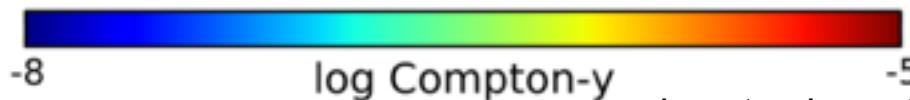
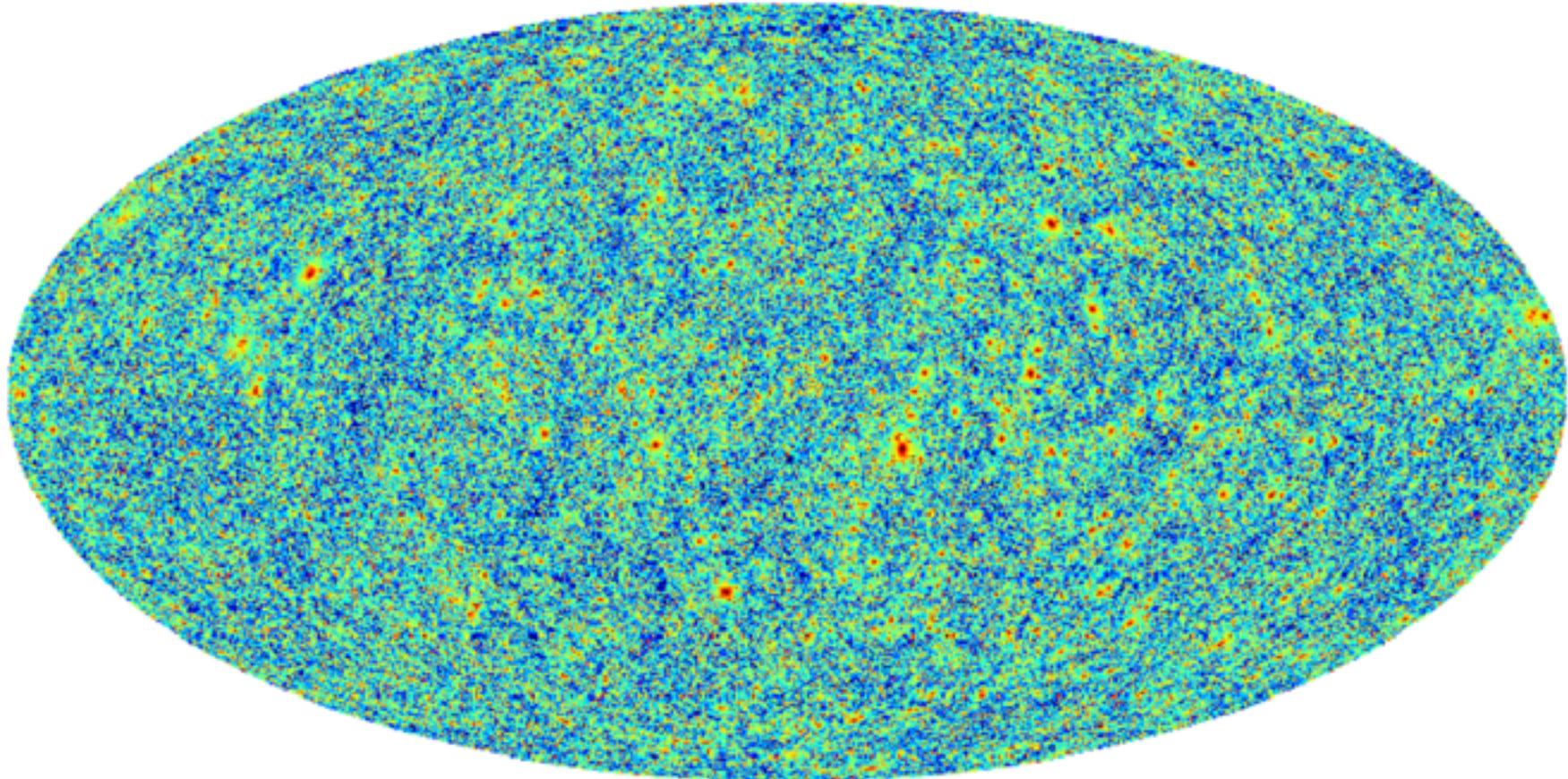
0 < r < .49 95%CL, $.2 \pm .15$ 1σ
cf. $r < 0.11$ uniform n_s

the Cosmic Web of Clusters, seen thru Compton cooling of high pressure electrons by the CMB

tsz
effect

Lightcone Simulation of Clusters & Groups $> 2 \times 10^{13} M_{\text{sun}}$ to $z=2$ in projected pressure

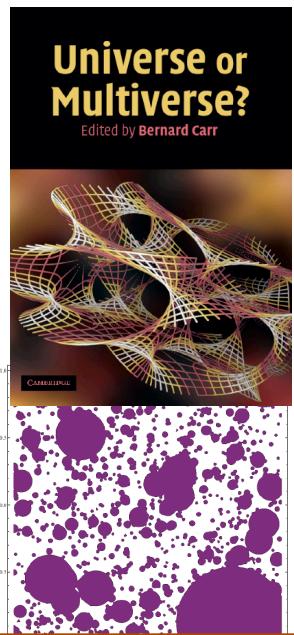
Alvarez, Bond, Hajian, Stein, Battaglia, Emberson,..2015



how to characterize map errors? by SIMs
inhomogeneous, CIB contamination, ..

ultra-Ultra Large Scale Structure of the Universe

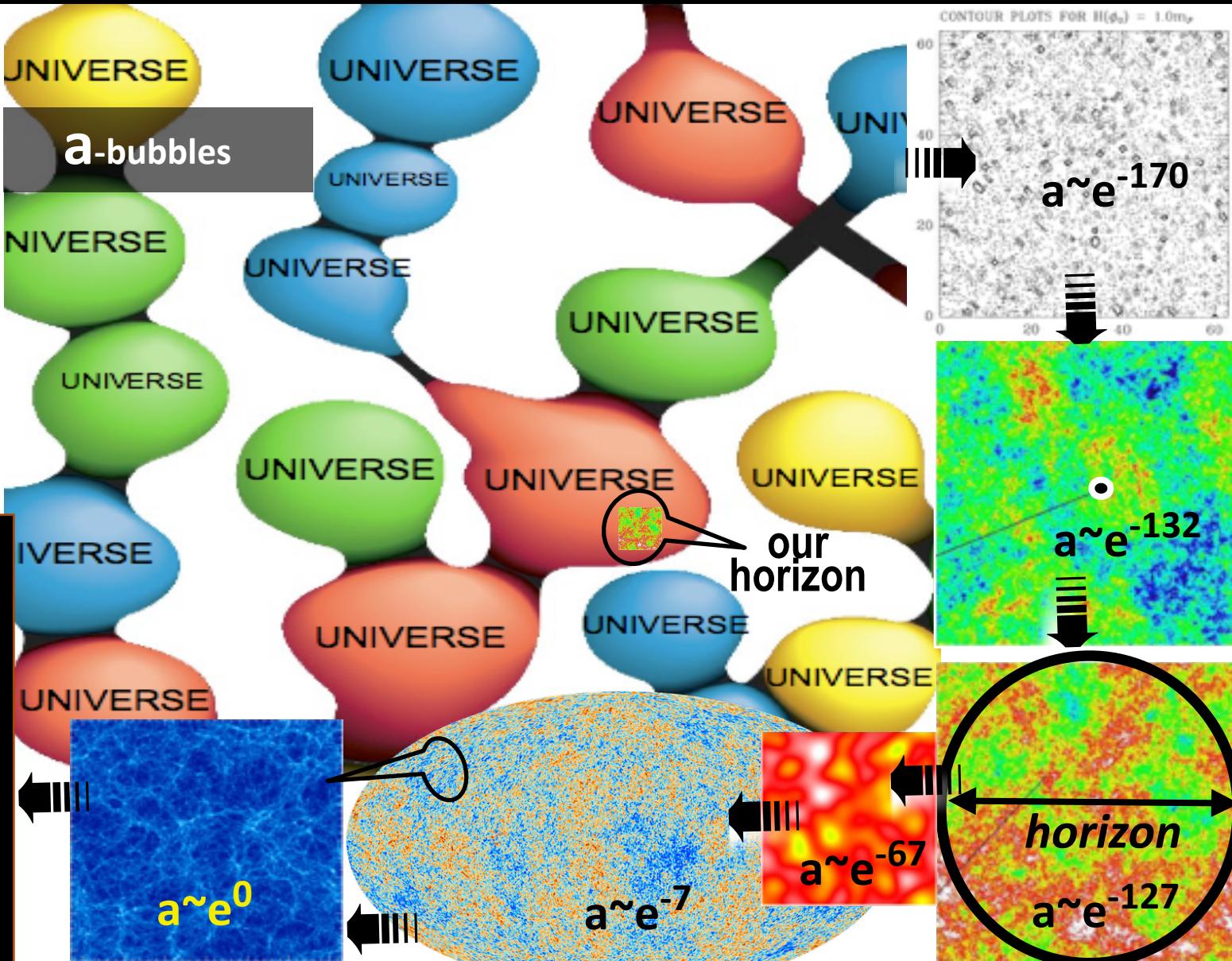
Horizons: the ultimate-speed constraint on light & information



END
a future DE-Void



$a \sim e^{+++}$



CBI pol to Apr'05 @Chile

53+35 cls (≥ 40)



CBI2

thermal SZ clusters

QUaD @SP

1652 cls

Planck09.4

52+ bolometers
+ HEMTs @L2
9 frequencies



WMAP @L2 to 2010

2004

2006

2008

LHC

2011

Bpol
@L2

2005

Acbar@SP

~1 blind

2007

AMIBA

6 cls

224 ($\Rightarrow 747$)

2009

SPT

1000 bolos
@SPole



ACT

23+68~91 cls

3000 bolos

3 freqs @Chile

SPTpol

ACTpol

ALMA

CCAT@Chile

>96

OVRO/BIMA array

38 cls

80s-90s
Ryle
OVRO

3 cls ($z > 1$), x?

AMI

7+1 cls $\geq 50+25$



APEX

~400 bolos @Chile

~25 cls

GBT Mustang

4 cls (~25 CLASH)



SCUBA2

12000 bolos

JCMT @Hawaii

LMT@Mexico

CBI pol to Apr'05 @Chile

53+35 cls (≥ 40)



CBI2 *thermal SZ clusters*

QUaD @SP

1652 cls

Planck09.4

Planck PSZ, cnts, ymap

1203 confirmed,

many $\sim 10^{15} M_{\odot}$ $0 < z < 0.8$



WMAP @L2 to 2010

2004

2006

Reichardt+12, Benson@ESLAB13
100 cl cosmology, 400 with S/N > 5
now, 747 summer 2013 2500 deg²

2005

Acbar@SP

~1 blind

2007

AMIBA

6 cls

224 ($\Rightarrow 747$)

SPT

1000 bolos
@SPole

Menanteau+12, Hasselfield+12

ACT Celestial Equator cls, 68 (49+19)

in SDSS, half $z > .5$, 1 $z \sim 1.1$ $10^{15} M_{\odot}$

502 sq deg $\Rightarrow 91$ in 952 deg², $0.1 < z < 1.3$

100% purity for S/N>5. 60% > 4.5

No significant evidence of SZ/BCG offset
Msz-N₂₀₀ weak correlation, large scatter

>96

OVRO/BIMA array

38 cls

80s-90s
Ryle
OVRO



APEX
~400 bolos@Chile

~25 cls

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