

**Unveiling Fundamental Physics
from the Cosmic First Light:
from COMPLEXITY
to SIMPLICITY
to COMPLEXITY
to SIMPLICITY,
the Universe at Large**

the BOUNDED flow of information
the BOUNDless thought of man

Dick Bond

“To me every
hour of the light and
dark is a miracle.
Every cubic inch of
space is a miracle.”

– Walt Whitman

IN EVERY teaspoon of air ~ 5 cubic cm
Ordinary Matter $\sim \text{amu} / \text{nm}^3$ 4.8% O₂ N ; H,He

THE DARK

Dark Matter

$\sim \text{amu} / \text{m}^3$ $26.0 \pm 1\%$ compressed in MilkyWay $\sim 0.1 \text{ amu} / \text{cm}^3$;
for LHC@CERN-type relics ~ 1 every 10 cm

Dark Energy

\sim vacuum potential $\sim 3 \text{ amu} / \text{m}^3$ $69.2 \pm 1.0\%$

THE LIGHT

cosmic radiation

the 1st light of the universe $412 / \text{cm}^3$ 0.005%

cosmic neutrinos \sim cosmic photons $> 0.47\%$

cosmic gravity waves \ll cosmic photons

THE VACUUM

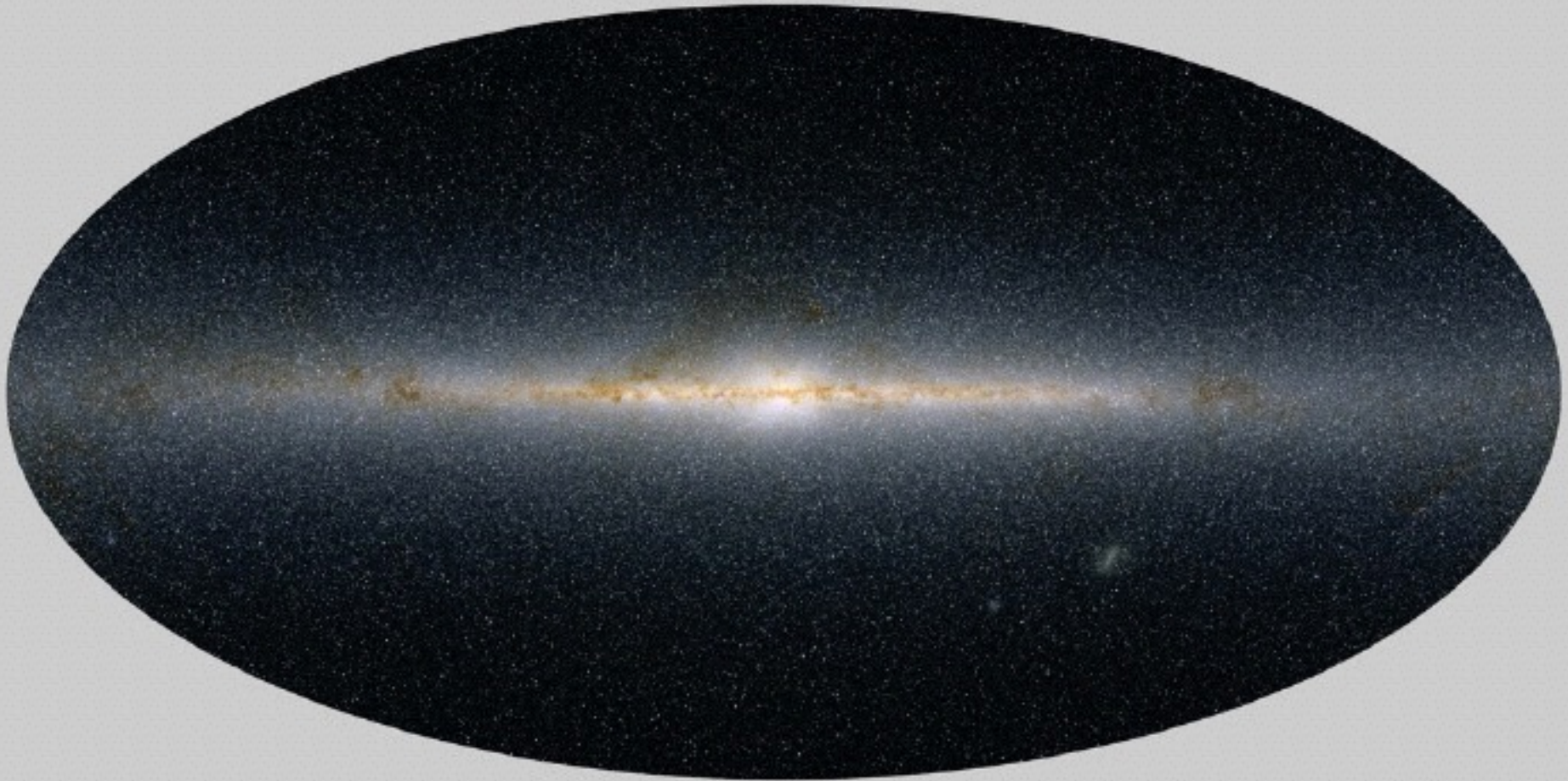
Higgs@CERN vacuum origin of mass

vacuum fluctuations origin of all the cosmic structure we see

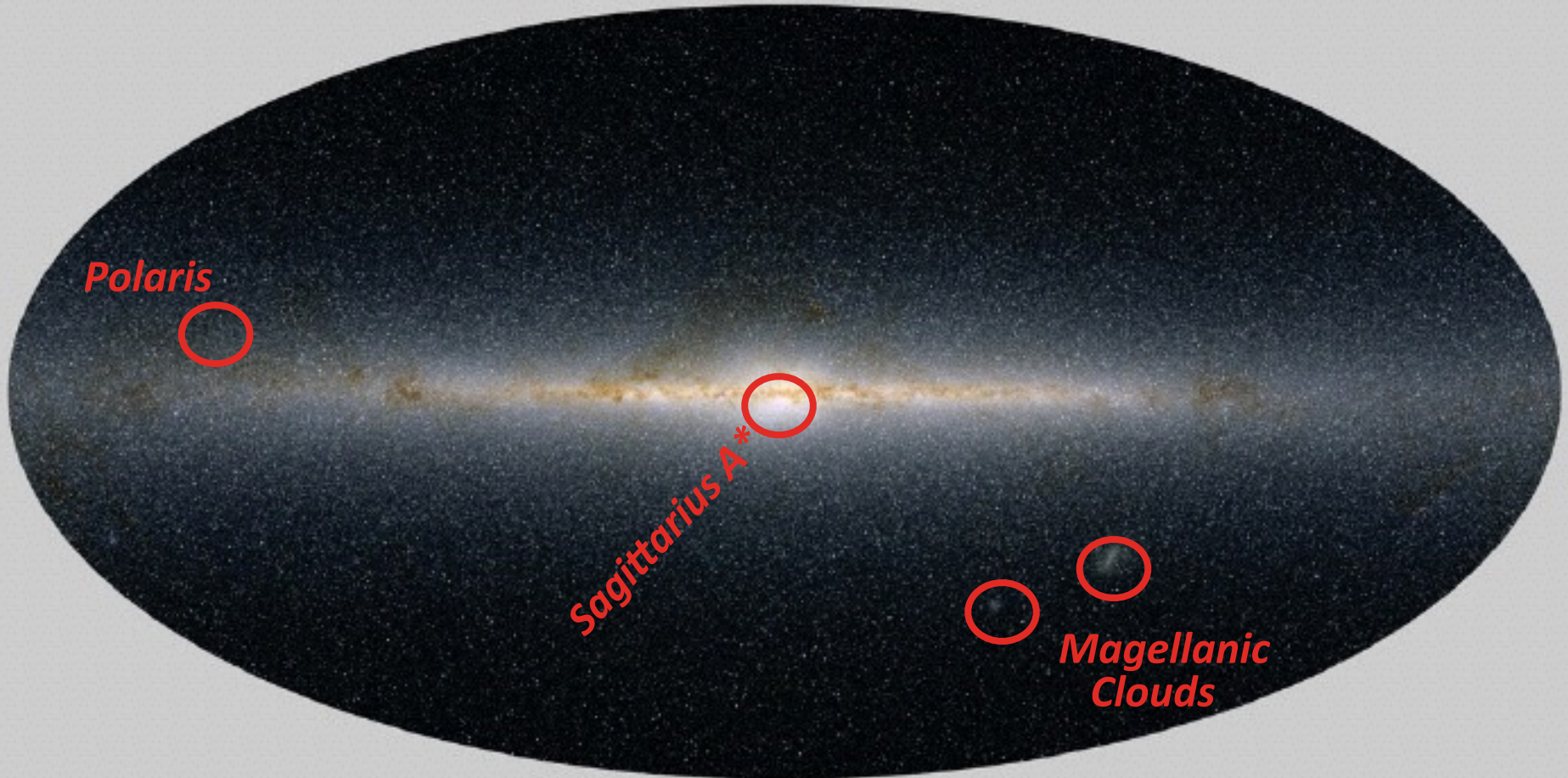
the vacuum is under gravitational strain, differentially accelerating



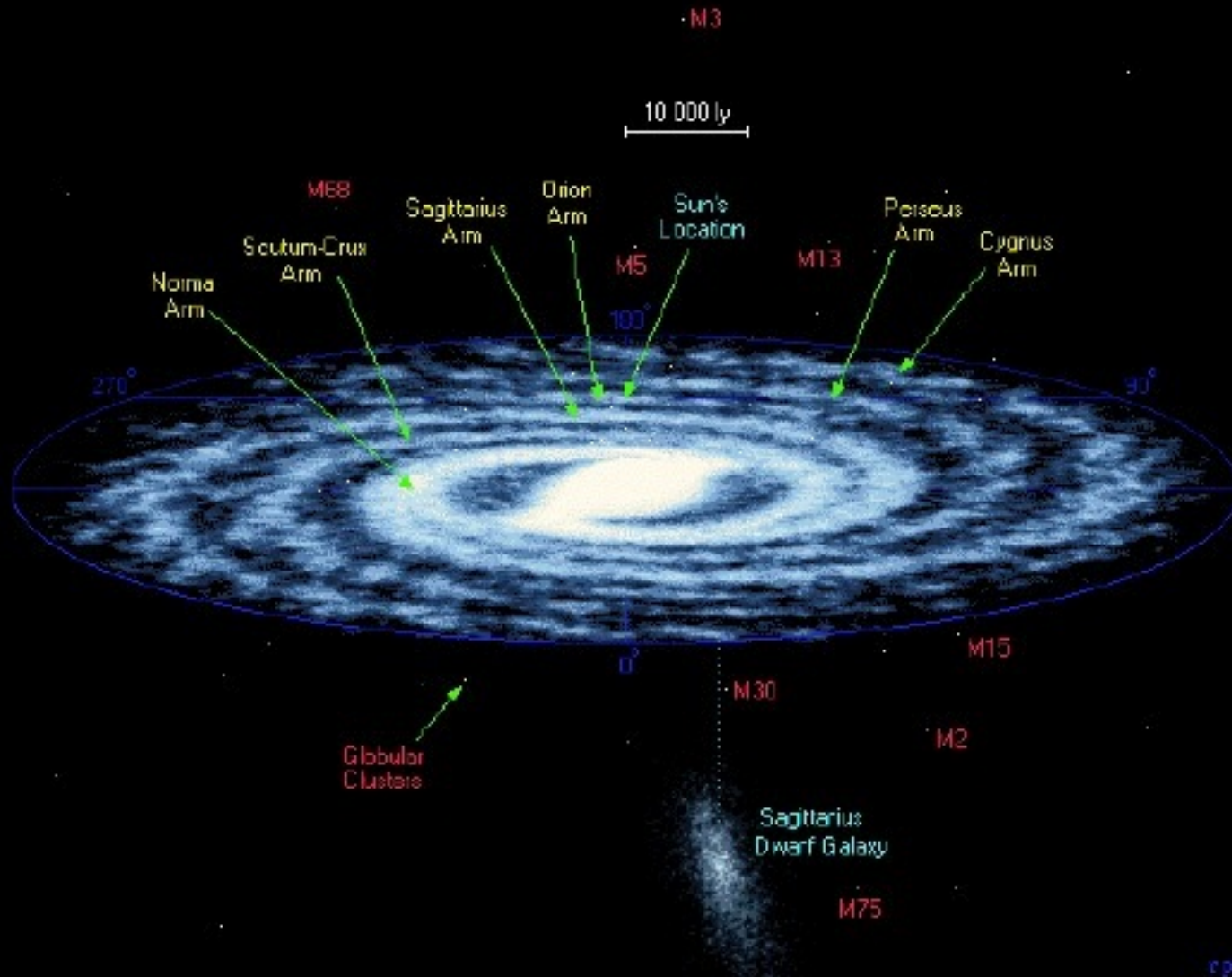
Milky Way in infra-red: half a billion stars, a disk galaxy



Milky Way in infra-red: half a billion stars, a disk galaxy



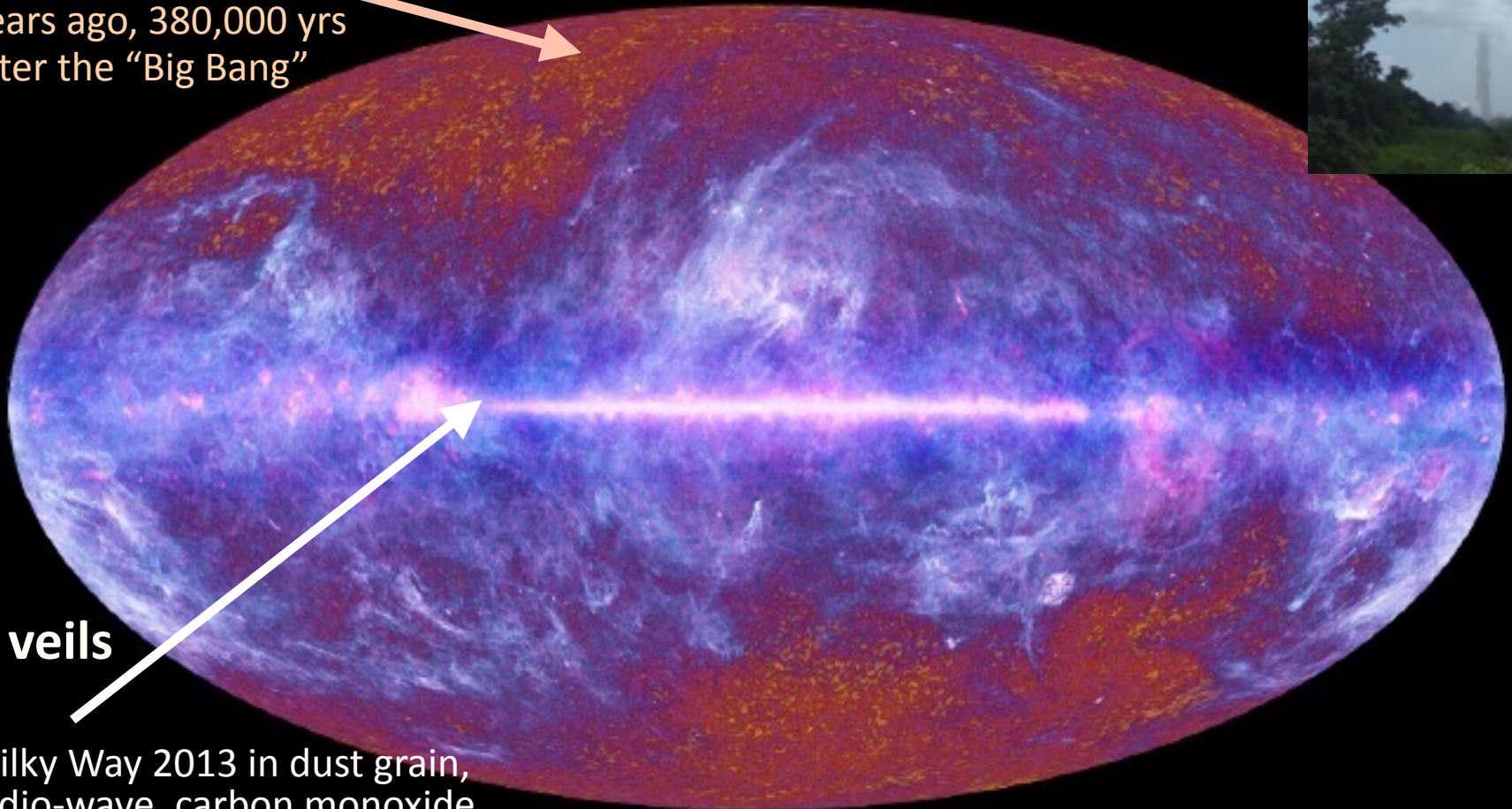
Milky Way in 3D: a disk galaxy with a large dark matter halo



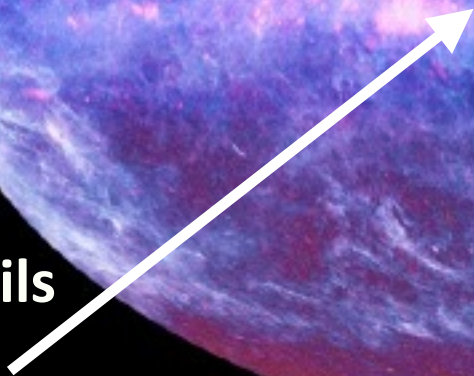


COMPLEXITY of here & now

the primordial light,
released 13.8 billion
years ago, 380,000 yrs
after the "Big Bang"



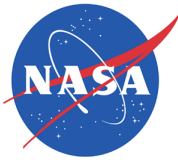
7 veils



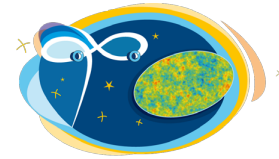
Milky Way 2013 in dust grain,
radio-wave, carbon monoxide
emissions; plus stellar, X-ray,
gamma ray, cosmic ray
emissions ...



planck



DTU Space
National Space Institute



HFI PLANCK
a look back to the birth of Universe



Science & Technology
Facilities Council



National Research Council of Italy



Deutsches Zentrum
für Luft- und Raumfahrt e.V.



UK SPACE
AGENCY



INSU
Observer & comprendre



IN2P3
Les deux infinis



MilliLab



Bond since 1993, Canada since 2001, 1st CSA pre-launch contract 2002-09, post-launch 2010-11, 2011-15

Planck 1.3yr Frequency Maps

Planck+Herschel Launch
May14 09 French Guiana

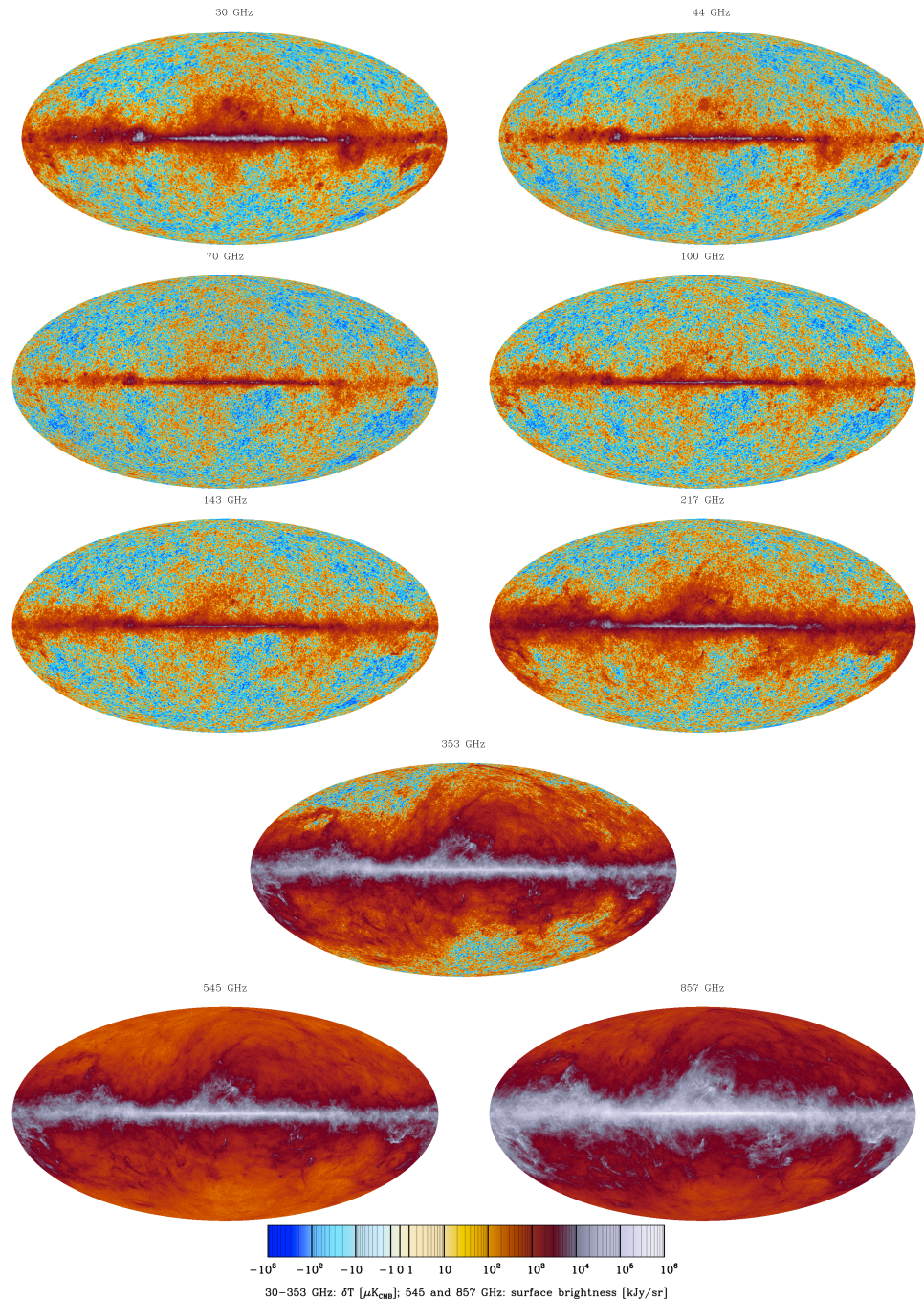
1.5m telescope,

HFI bolometers @6freq
<100mK,

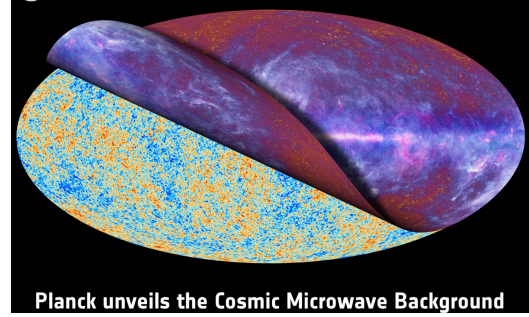
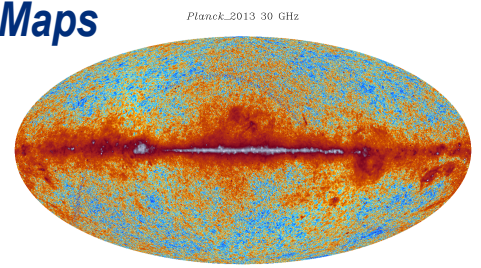
LFI HEMTs@3freq,

some bolometers & all
HEMTs are polarization
sensitive

- Left earth at ~10 km/s,
1.5 million km in 45
days, cooling on the
way (20K, 4K, 1.6K, 0.1K
4 stage). @L2 on July 2
09; Survey started on
Aug 13 09
- spin@1 rpm, 40-50
minutes on the same
circle, covered all-sky in
~6 month
- ~5 HFI all-sky surveys
(to Jan 2012)
- ~8 LFI surveys
- kicked out of L2 Oct
2013



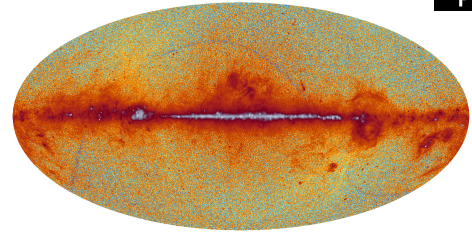
Some Planck Component Separated Maps



Commander: Low-Frequency Emission Amplitude @ 30 GHz

C/R: Low-Frequency Emission Amplitude @ 30 GHz

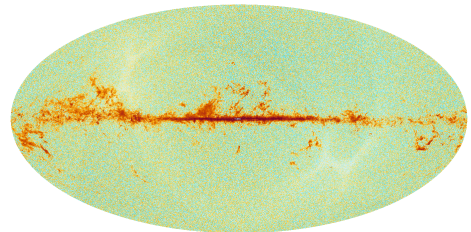
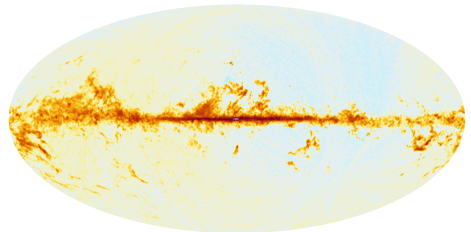
**LF Synchrotron +
bremsstrahlung**



Commander: "discovery" CO map @ 100 GHz

C/R: "discovery" CO map @ 100 GHz

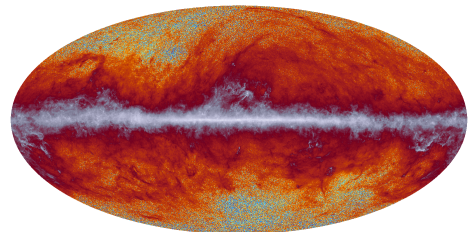
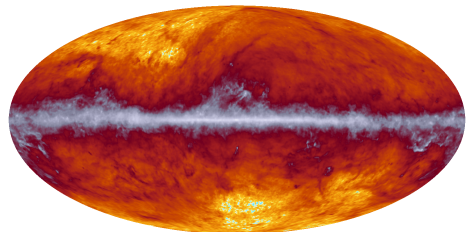
**Galactic Carbon
Monoxide**



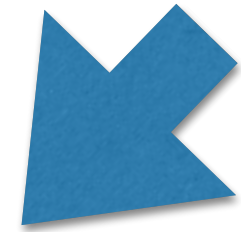
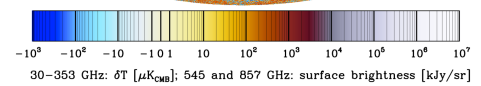
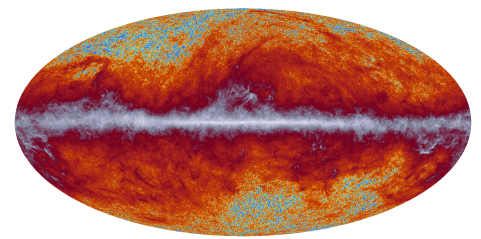
Commander: Dust Amplitude @ 353 GHz

C/R: Dust Amplitude @ 353 GHz

**HF Thermal Dust
Emission**



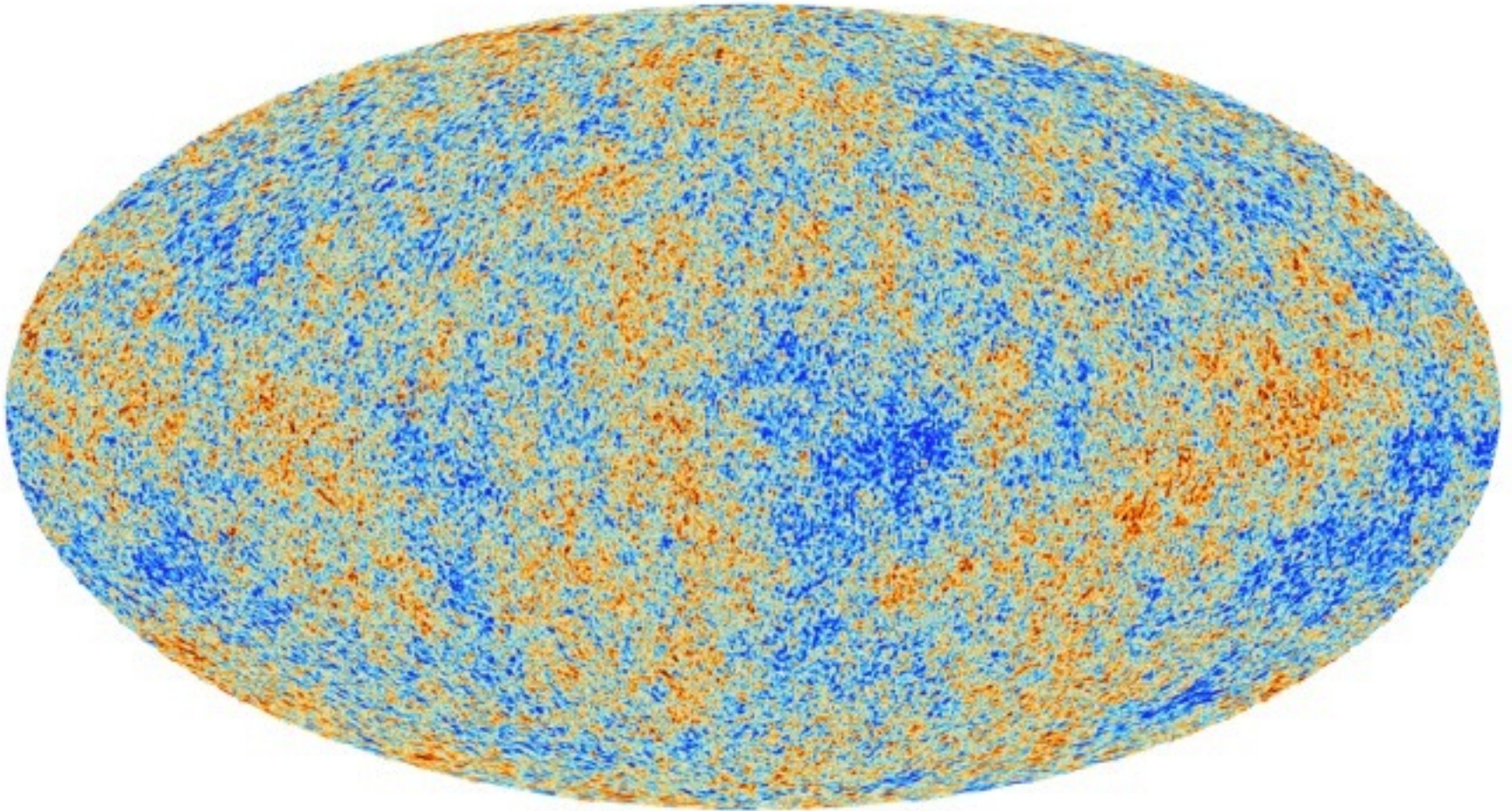
Planck_2013 353 GHz



Planck's primordial light unveiled, March 21, 2013

reveals the **SIMPLICITY** of primordial cosmic structure

7⁺ numbers, 3 densities, 2+1 early-Universe inflation



**Temperature changes
in micro-degrees**

Google “Planck Satellite 2013 results” yields ~ **1 million links**

Google “gravity waves from inflation 2014” yields ~ **0.3 million links**”

THE GLOBE AND MAIL 

SPACE

New glimpses of ancient light fuel cosmic debate

 Government of Canada / Gouvernement du Canada

Canada 

Canadian Space Agency

[Home](#) > [Audiences](#) > [Media](#) > [News releases](#) > 2013

> Canadian astronomers reveal surprising new portrait of the Early Universe

Canadian astronomers reveal surprising new portrait of the Early Universe

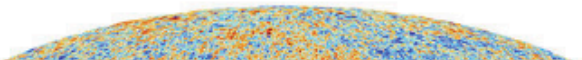
The New York Times

Space & Cosmos

WORLD U.S. N.Y. / REGION BUSINESS TECHNOLOGY SCIENCE HEALTH SPORTS OPINION

ENVIRONMENT SPACE & COSMOS

Universe as an Infant: Fatter Than Expected and Kind of Lumpy



L'enfance de l'Univers dévoilée

LE MONDE | 21.03.2013 à 11h27 • Mis à jour le 21.03.2013 à 13h44

gravity waves from inflation

 UNIVERSITY OF TORONTO

U of T News

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Plancking at U of T: space |

CIFAR
CANADIAN
INSTITUTE
FOR
ADVANCED
RESEARCH

**CIFAR
cosmologists
contribute to
new portrait
of the Early
Universe**



NEWS ARCHIVE

PLANCK
Light

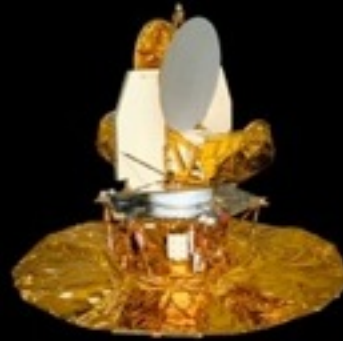
<http://www.nytimes.com> *Space Ripples Reveal Big Bang's Smoking Gun* By DENNIS OVERBYE MARCH 17, 2014

Comparison of CMB Space Experiments: Resolution, 420', 12.5', ~5-7'

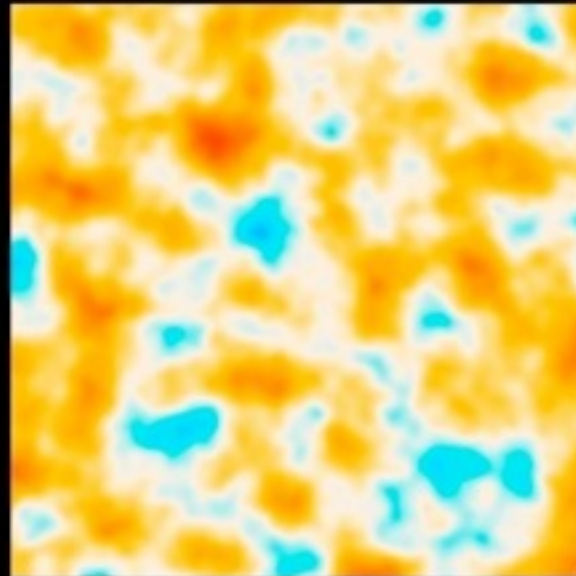
COBE 89 launch

WMAP 01 launch

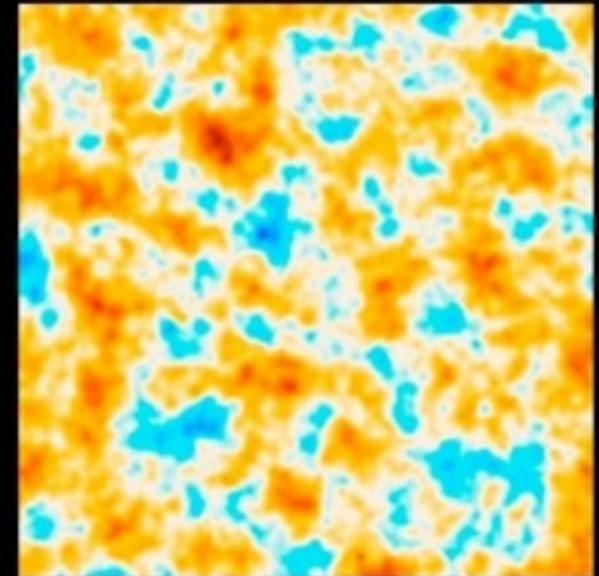
Planck 09 launch



COBE



WMAP



Planck

goal: high enough resolution to plumb all cosmic parameter information. but high L foregrounds, extragalactic sources => higher L expts ACT (1.4'), SPT (1') = PlanckEXT to nail the "nuisance"

SIMPLICITY

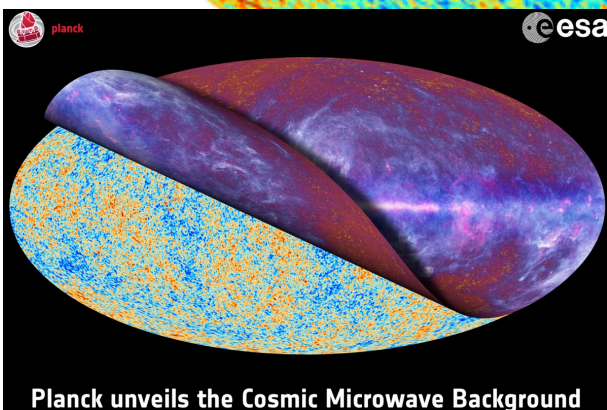
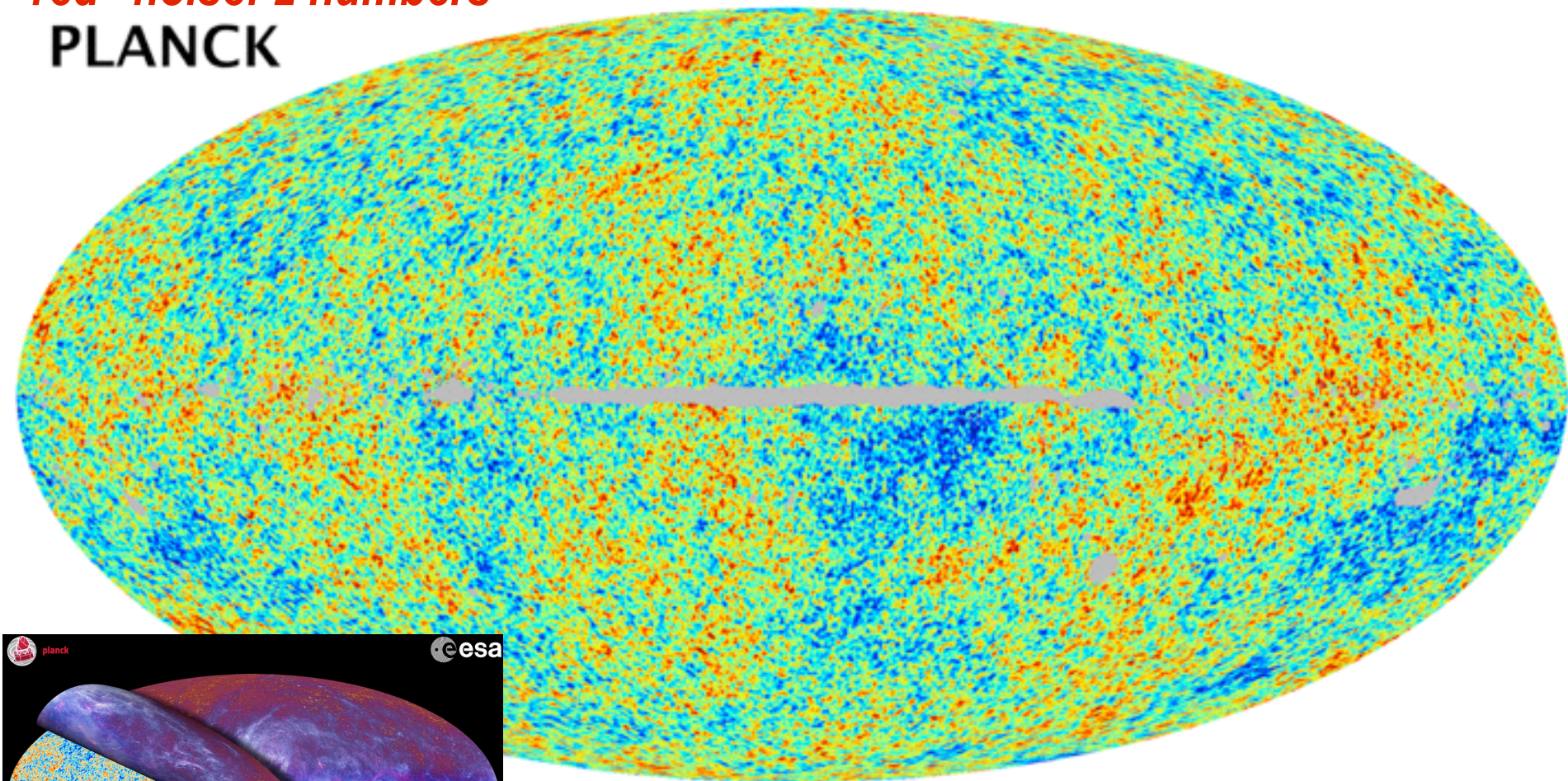
Planck 09 launch

at $a \sim e^{-7} \sim 1/1100 \Rightarrow$

at $a \sim e^{-67+60} \sim 1/10^{30+25}$
"red" noise: 2 numbers

PLANCK

Planck SMICA Map CMB-data Concordance



Planck CMB/SMICA map, $\sim 5'$ resolution
+ NILC, SEVEM, C-R 3 independent component
separated CMB maps show the same features

SIMPLICITY

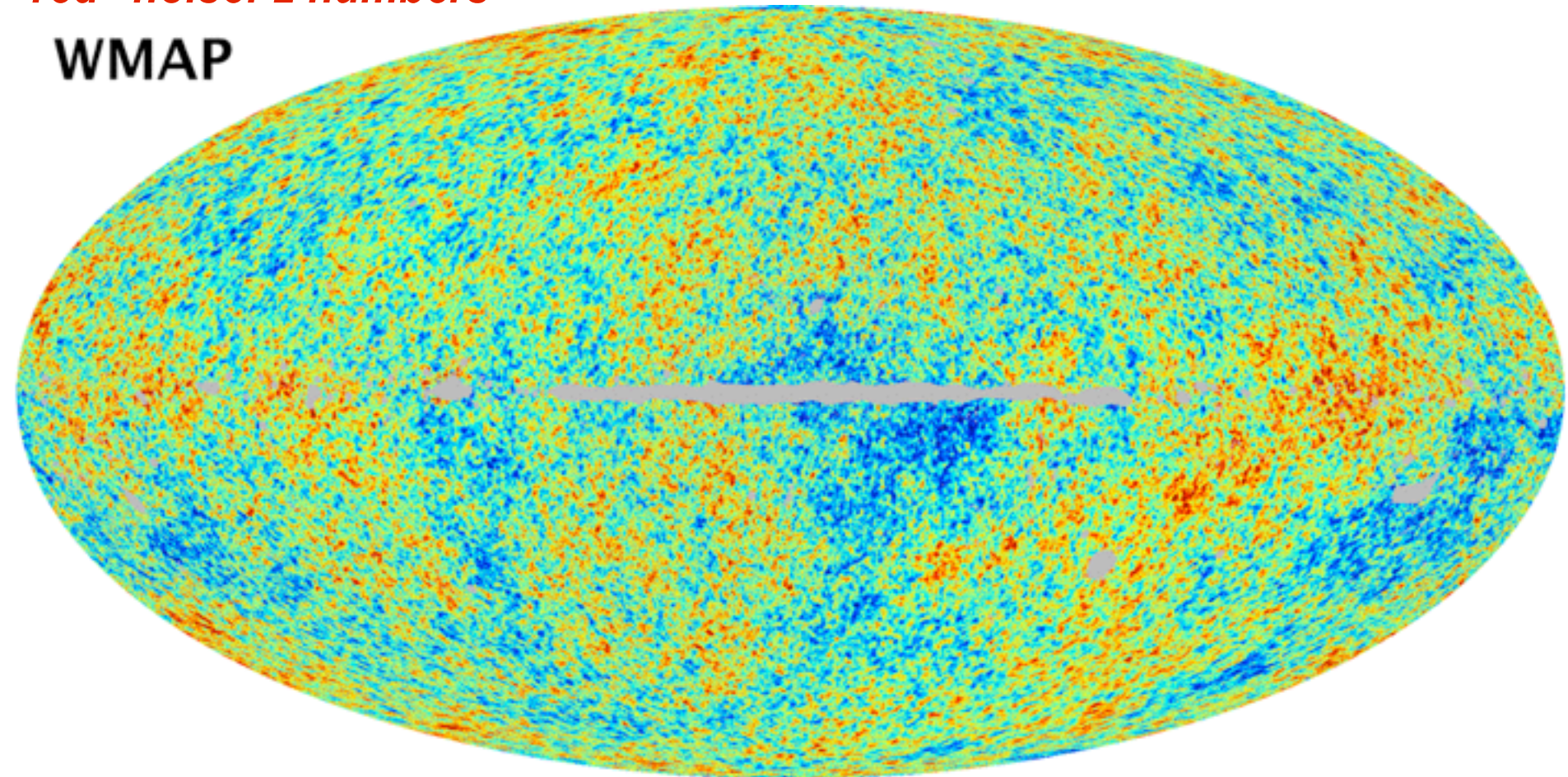
WMAP 01 launch

**WMAP W-band,
Template Cleaned
CMB-data Concordance**

at $a \sim e^{-7} \sim 1/1100 \Rightarrow$

at $a \sim e^{-67+60} \sim 1/10^{30+25}$
“red” noise: 2 numbers

WMAP



Cleaned with Planck 353 GHz dust map and low-frequency templates. 12' resolution.
similar tremendous agreement with the much higher (5X) resolution ACT & SPT maps
total focus on the 1.2% difference in “calibration” between P13 (HFI & LFI) & WMAP9
Planck’s information > 4X WMAP9 in multipoles

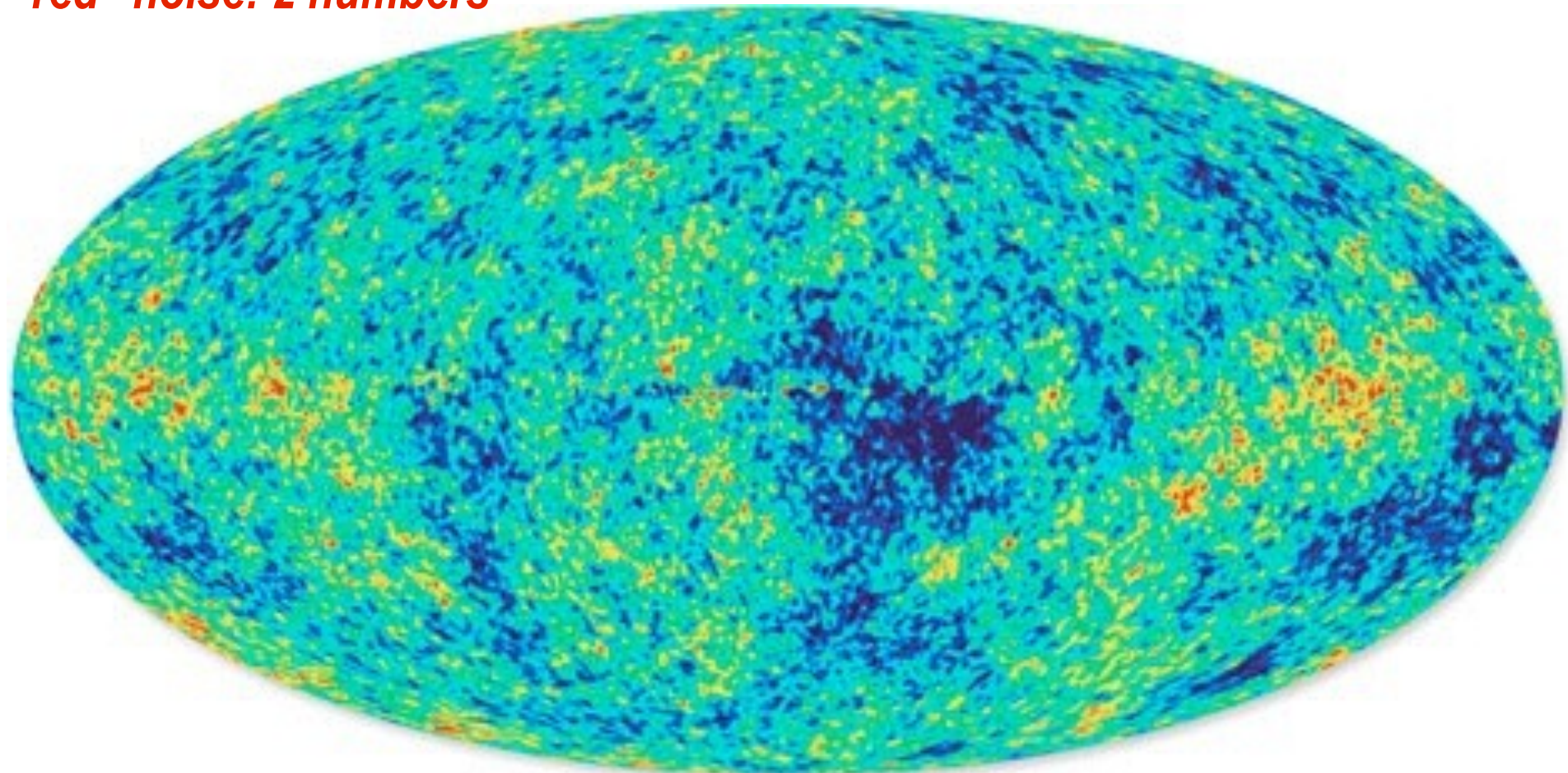
SIMPLICITY

WMAP 01 launch

WMAP W-band, Template Cleaned CMB-data Concordance

at $a \sim e^{-7} \sim 1/1100 \Rightarrow$

at $a \sim e^{-67+60} \sim 1/10^{30+25}$
“red” noise: 2 numbers



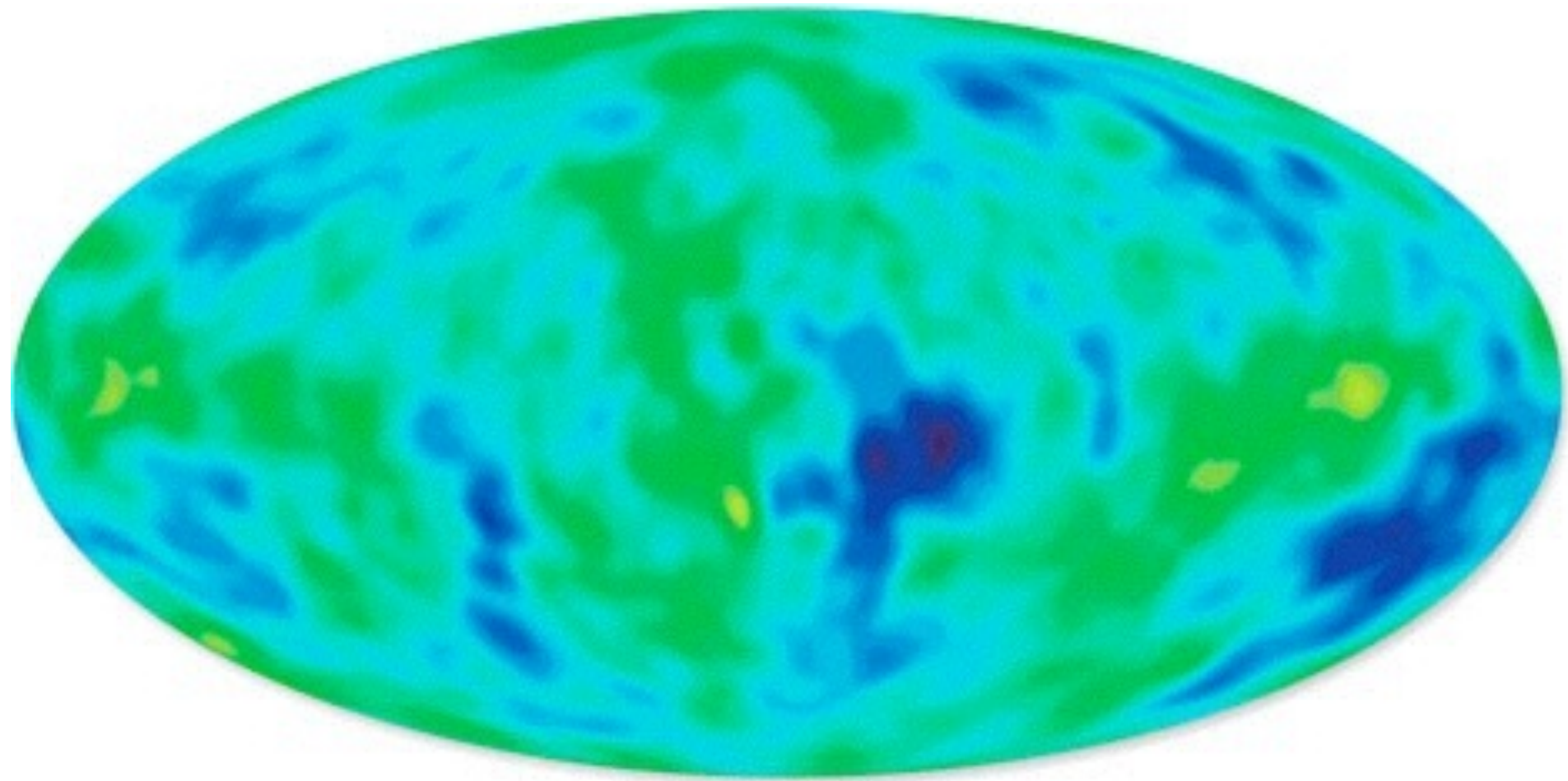
Cleaned with low-frequency templates only.

similar tremendous agreement with the much higher (5X) resolution ACT & SPT maps
total focus on the 1.2% difference in “calibration” between P13 (HFI & LFI) & WMAP9

COBE 89 launch

COBE

CMB-data Concordance



**ACT (1.4 arcmin res) vs Planck1.3 (~5.5 arcmin res @217) in limited sky region
=> excellent agreement; cross correlations also look great**

ACT collaboration: Louis+14

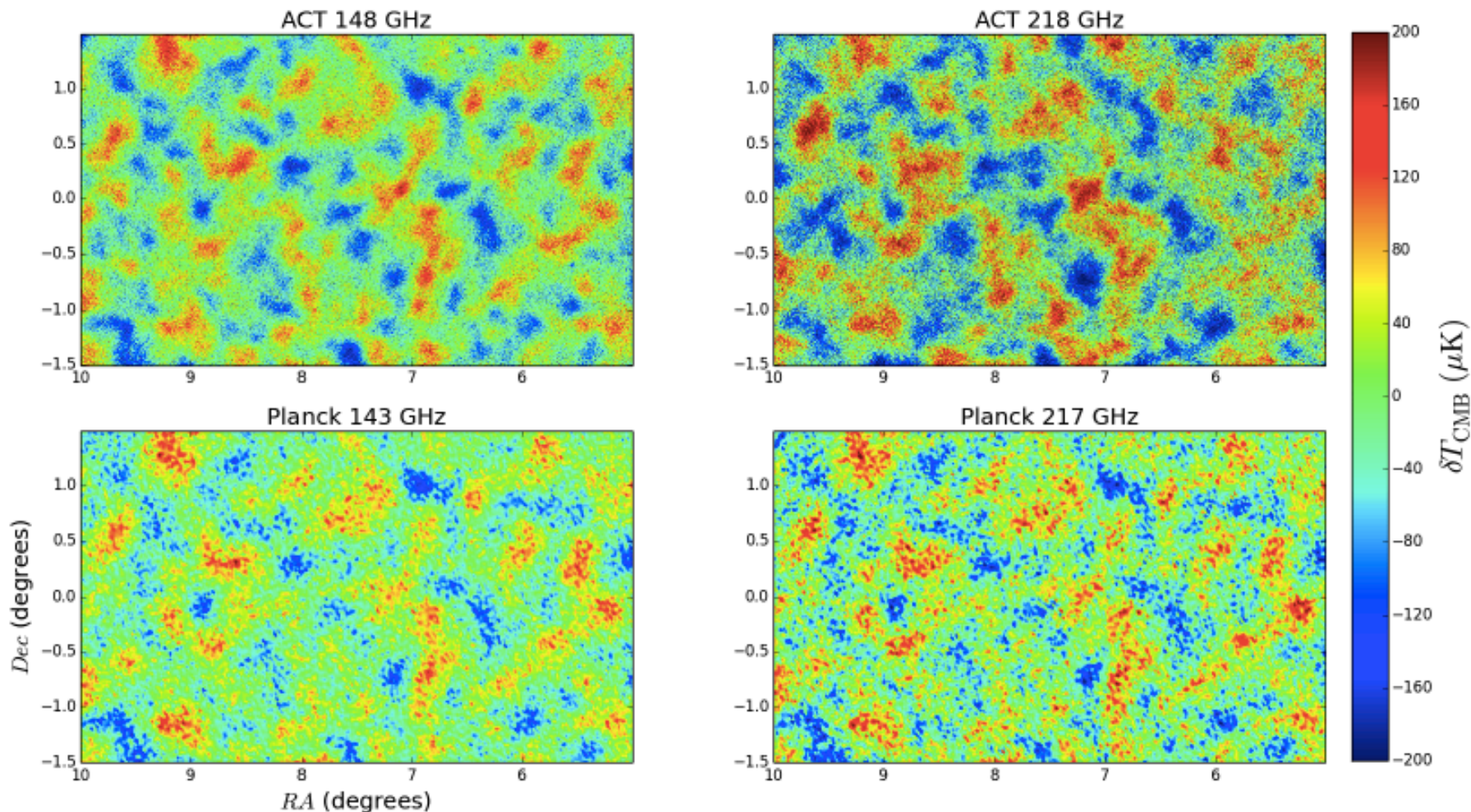
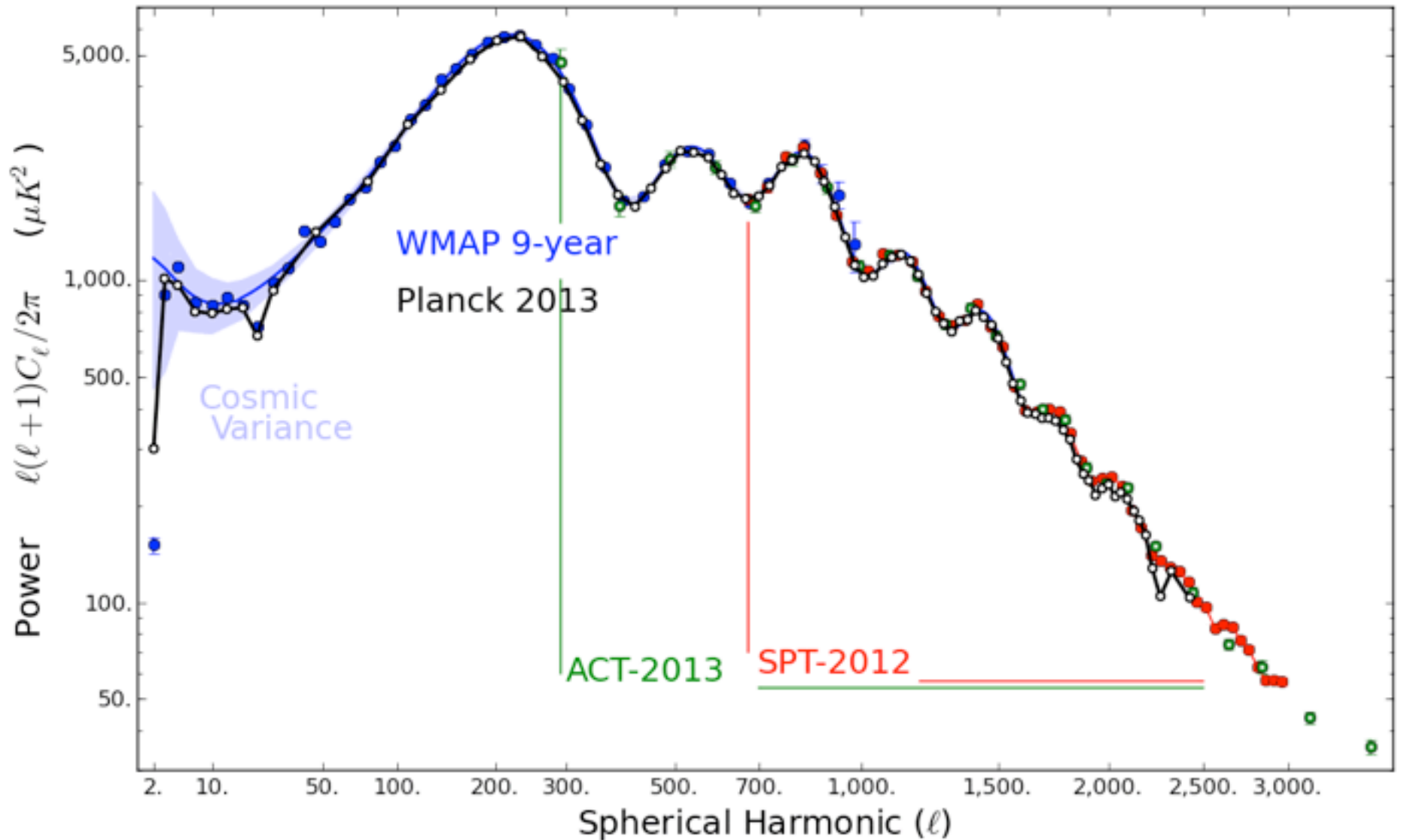


FIG. 1.— Comparison of ACT (*top*) and *Planck* (*bottom*) maps for a 15 deg² patch in the ACT Equatorial region. The maps are the inverse variance weighted combination of all ACT data at 148 GHz (*left*) and 218 GHz (*right*) and all *Planck* data at 143 GHz and 217 GHz. All maps have been filtered with a high pass filter (for modes on scales: $\ell < 500$). Artifacts of the HEALpix pixelization are seen in the *Planck* maps. The agreement is visually excellent.

Boomerang 2000, 2003 also agree, as does SPT in the overlap region

harmonic analysis of the 'music of the spheres' => inharmonious, coloured noise in the CMB



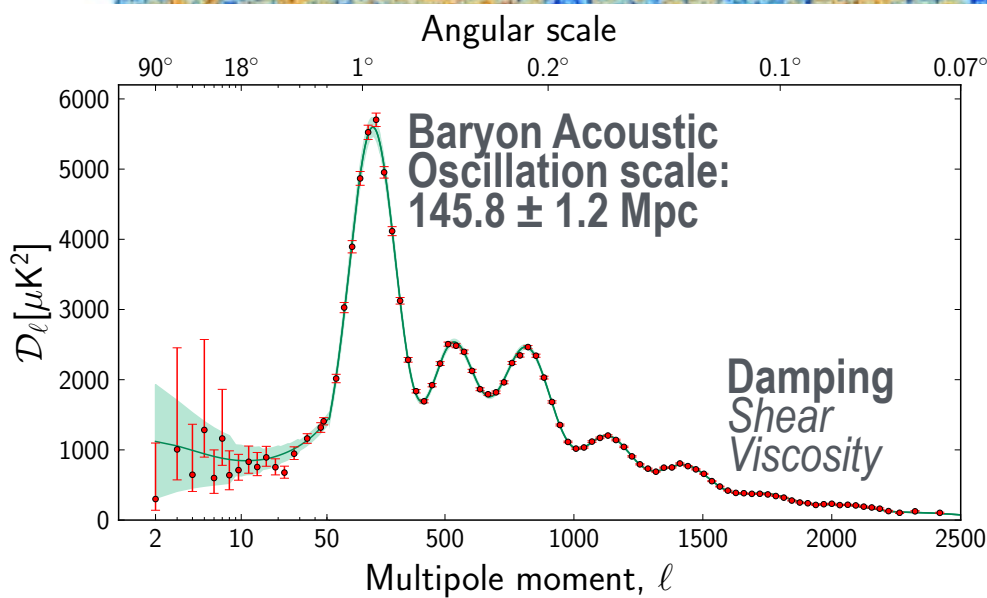
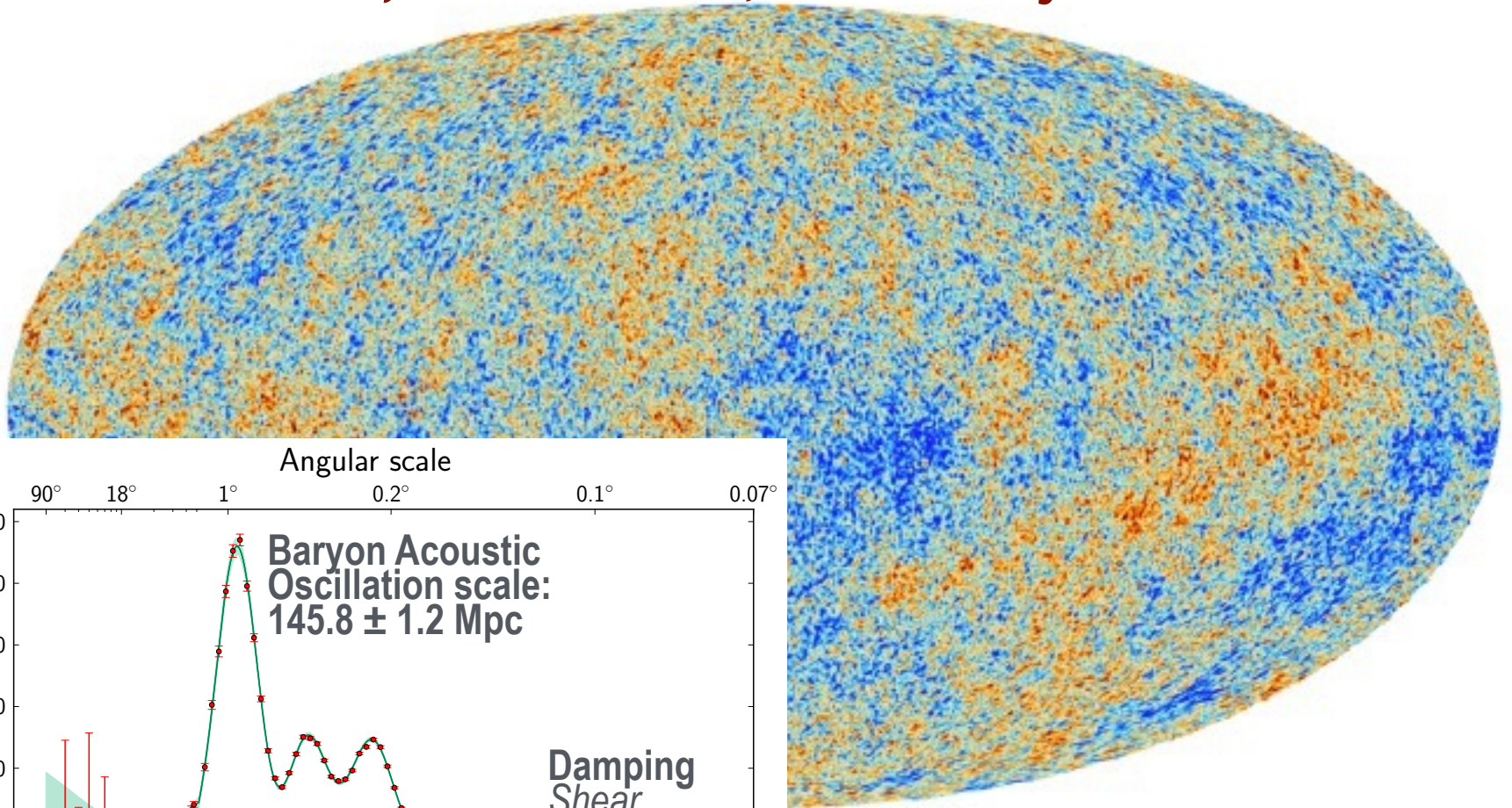
WMAP9 + ACT + SPT cf. Planck2013 = Planck1.3yr

Halpern13 gif

reveals **primordial sound waves**

=> the inharmonious *'music of the spheres'*

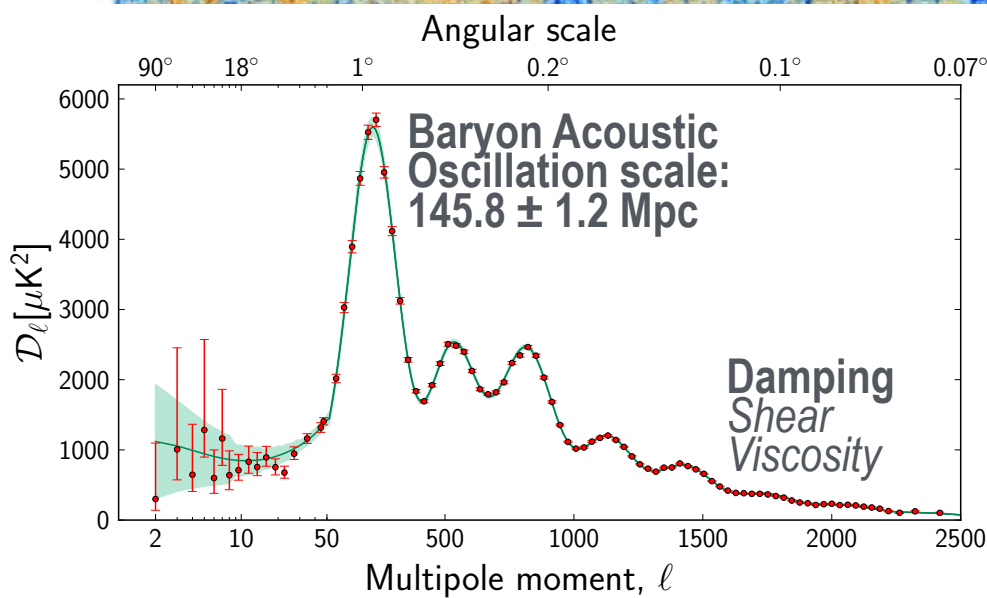
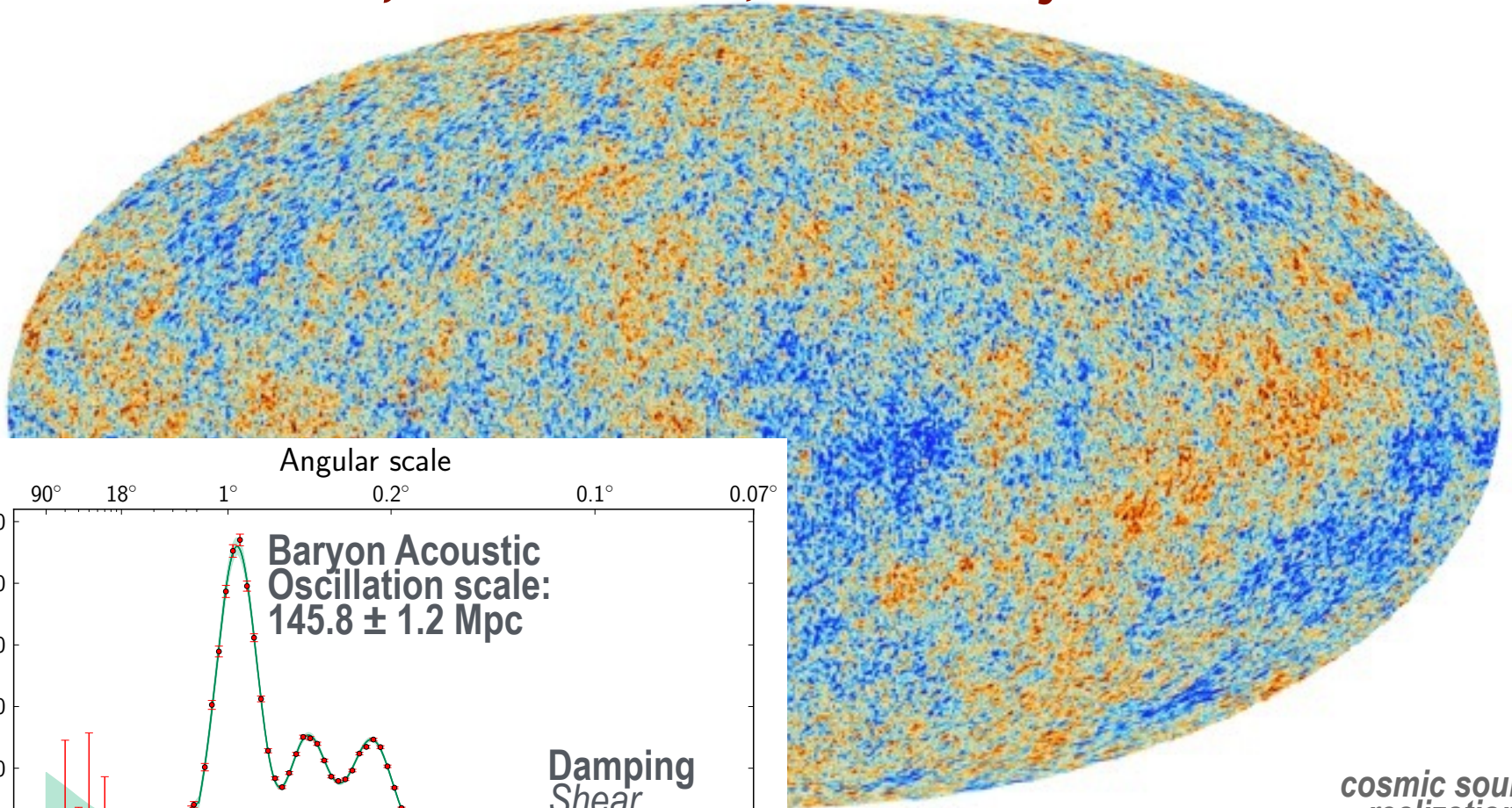
7⁺ numbers, 3 densities, 2+1 early-Universe inflation



reveals **primordial sound waves**

=> the inharmonious *'music of the spheres'*

7⁺ numbers, 3 densities, 2+1 early-Universe inflation



cosmic sound realization

SIMPLICITY

at $a \sim e^{-7} \sim 1/1100 \Rightarrow$

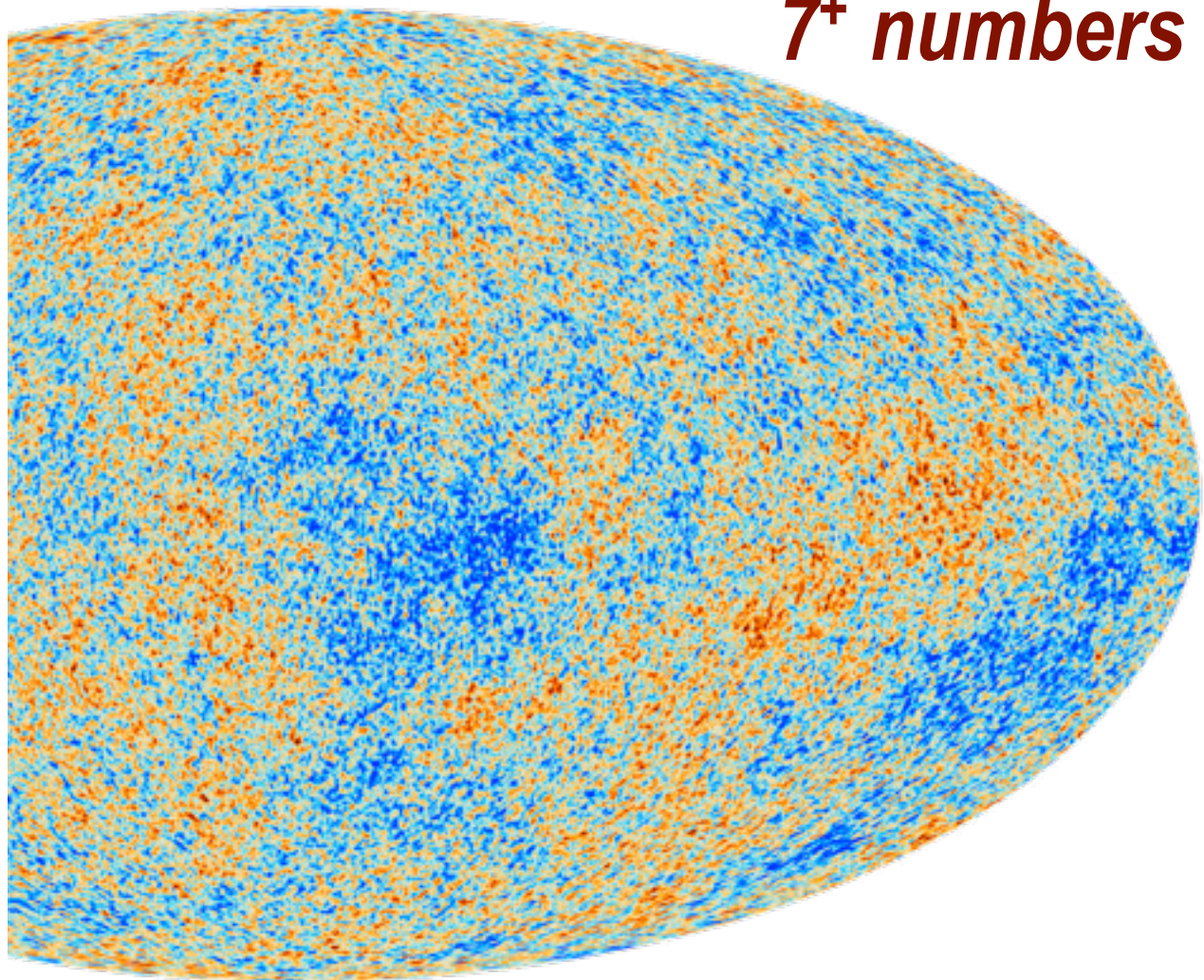
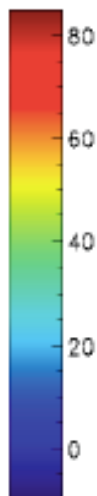
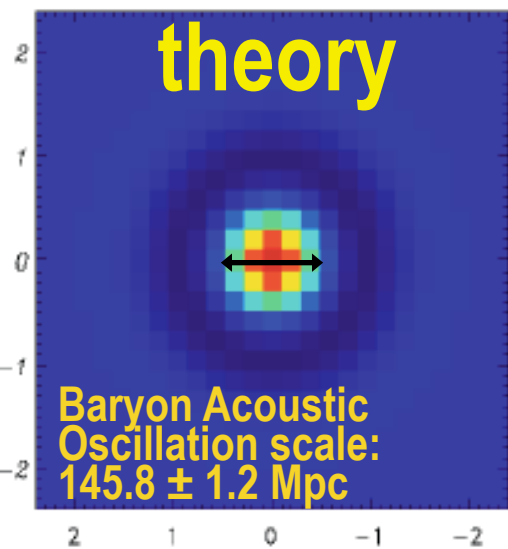
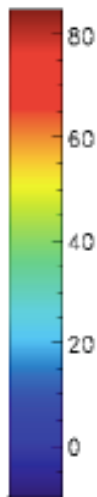
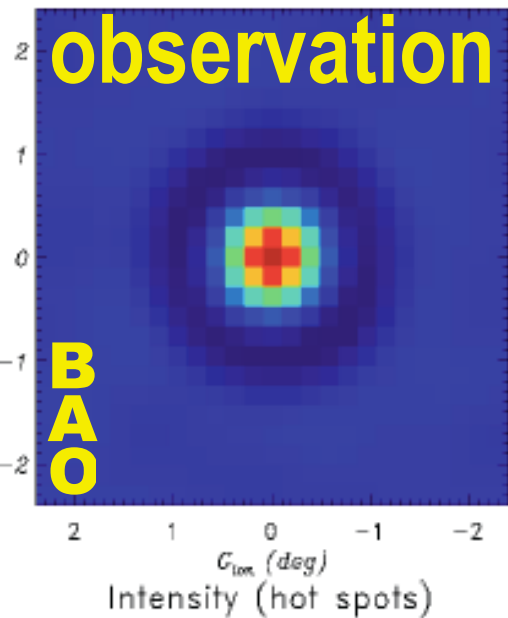
at $a \sim e^{-67+60} \sim 1/10^{30+25}$

reveals primordial sound waves in matter

\Rightarrow learn **contents & structure** at 380000 yr, $a \sim e^{-7}$

\Rightarrow infer the structure far far earlier $a \sim e^{-67+60}$

7⁺ numbers



SIMPLICITY

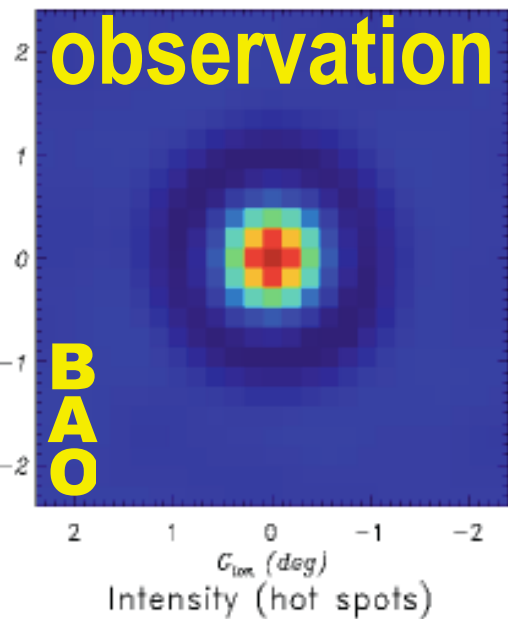
reveals primordial sound waves in matter

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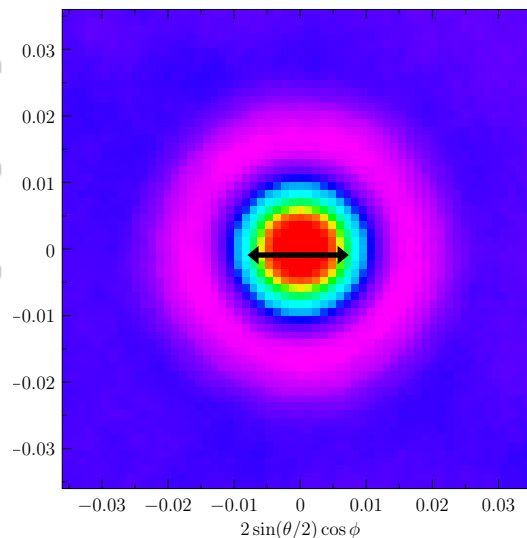
Planck2013

oriented peaks,
anisotropic CMB strain

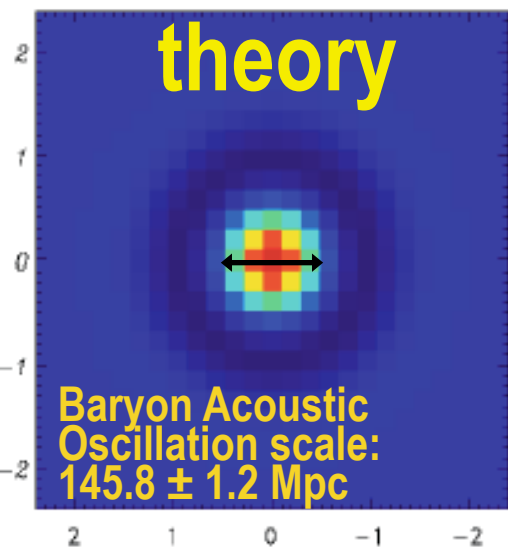
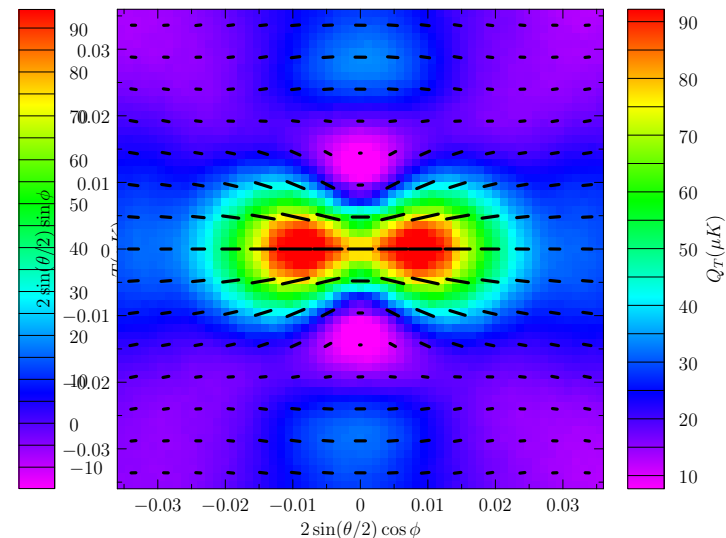


B2FHAP14, Planck2013

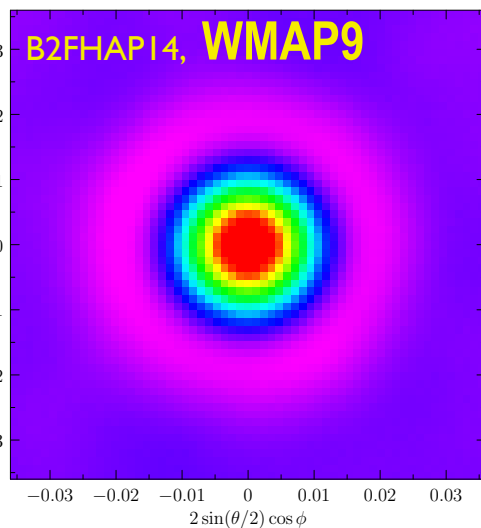
11113 patches on T maxima, random orientation



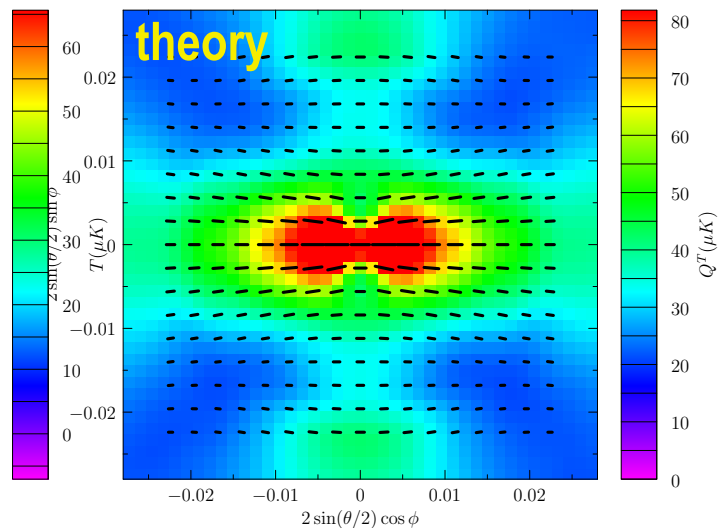
10825 Q_T patches on T maxima, oriented



9257 patches on T maxima, random orientation



63165 patches on T maxima, oriented





a scale of the Universe

strained photons redshift

$$= 1 / (1 + \text{redshift})$$

now = 1 when we **observe** the **1st light**

then = 1/1100 when the **1st light**
was **released from matter**,
billion X denser

galaxies forming ~ 1/4

there were **no galaxies** when $a < 1/20$

mean (isotropic) number of
e-foldings of scale $\equiv \langle \ln a \rangle$

a scale of the Universe

$\langle \alpha \rangle$

0

now = 1 when we observe the 1st light

then = 1/1100 when the 1st light
was released from matter,
billion X denser

7

galaxies forming $\sim 1/4$ 1 \Downarrow 2

there were no galaxies when $a < 1/20$ 3

light nuclei

Dark Matter

21 \Downarrow 35

Heat: matter & radiation

67

quantum noise

67 \Downarrow 127



$\mathbf{a}_J^i(r,t)$ scale-tensor of the Universe

$$d\mathbf{X}^i(r,t) = \mathbf{a}_J^i(r,t) dr_{eq}^J$$

$$\mathbf{a}_J^j \equiv \exp(\boldsymbol{\alpha})_J^j$$

$$\boldsymbol{\alpha}_J^j \equiv \langle \ln a \rangle \delta_J^j + \boldsymbol{\epsilon}_J^j$$

$\boldsymbol{\epsilon}$ =strain tensor

$$d\mathbf{V}^i(r,t) = \mathbf{H}_J^i(r,t) d\mathbf{X}^i(r,t)$$

\mathbf{H}_J^i =Hubble aka shear = $d\boldsymbol{\alpha}_J^j / dt$
general relativity

Earth under Strain:
earthquakes, seismic waves

$\boldsymbol{\varepsilon}$ =strain tensor

Universe under Strain:
space-quakes, gravity waves

elastic deformation $d\mathbf{x}^i = \mathbf{e}_J^i d\mathbf{r}_{eq}^J$ $e_{J^i} = a_{J^i} / \langle a \rangle$
anisotropic strain, shear waves $\boldsymbol{\varepsilon} - \text{Trace}(\boldsymbol{\varepsilon})/3$
isotropic strain, sound $\text{Trace}(\boldsymbol{\varepsilon})$

scale-deformation a_{J^i}
anisotropic strain, gravity waves
isotropic strain, sound

linear: strain \propto *tide*
cosmic web story

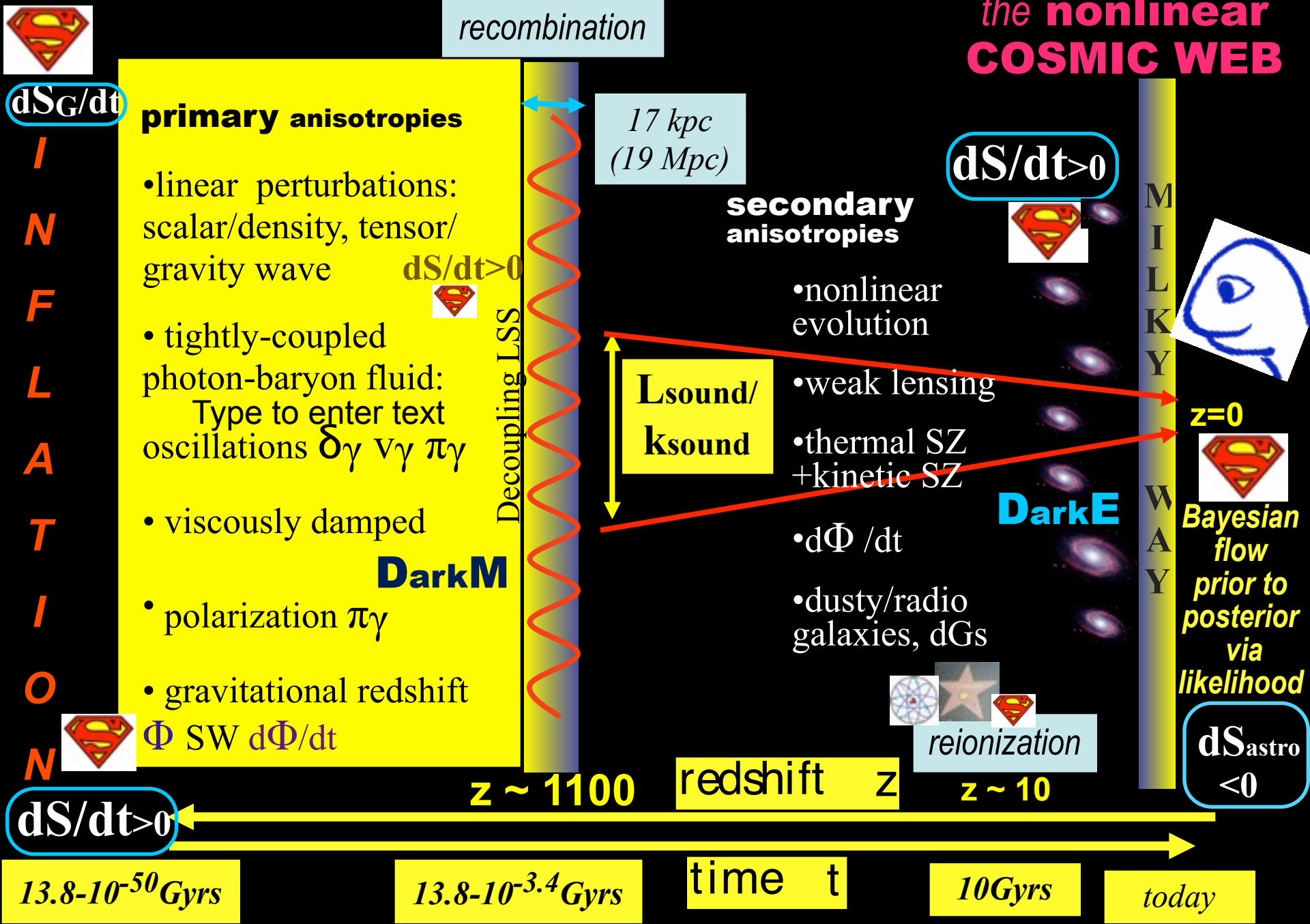


light and gravity are entangled: wavelength stretches under space-strain: redshift
the vacuum is modified under space-strain: inflation theory

general relativity => \mathbf{a} = dreibein, triad, Lagrangian-space metric $\mathbf{g} = \mathbf{a}\mathbf{a}^\dagger$

the flow of time => 4D vierbein spacetime-strain \mathbf{a}_b^β $b, \beta = 0, 1, 2, 3$

the **nonlinear** COSMIC WEB



the **nonlinear** COSMIC WEB



recombination

dS_G/dt

primary anisotropies

• linear perturbations: scalar/density, tensor/gravity wave $dS/dt > 0$

• tightly-coupled photon-baryon fluid: oscillations $\delta\gamma v\gamma \pi\gamma$

• viscously damped

• polarization $\pi\gamma$

• gravitational redshift

Φ SW $d\Phi/dt$

DarkM

Decoupling LSS

17 kpc (19 Mpc)

secondary anisotropies

• nonlinear evolution

• weak lensing

• thermal SZ + kinetic SZ

• $d\Phi/dt$

• dusty/radio galaxies, dGs

L_{sound}/k_{sound}

BAO
7

B
C
S
L
L
S

S
N
a
L
L
S

M
I
L
K
Y



z=0
Bayesian flow prior to posterior via likelihood

$dS_{astro} < 0$

H0
BAO
(z)
cls
ISW

DarkE

$dS/dt > 0$

$z \sim 1100$

redshift z

$z \sim 10$

13.8-10⁻⁵⁰ Gyrs

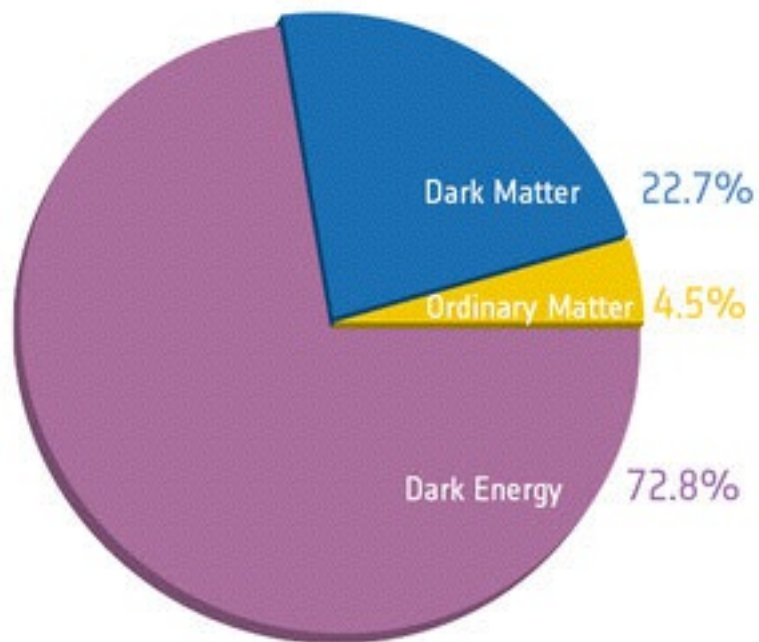
13.8-10^{-3.4} Gyrs

time t

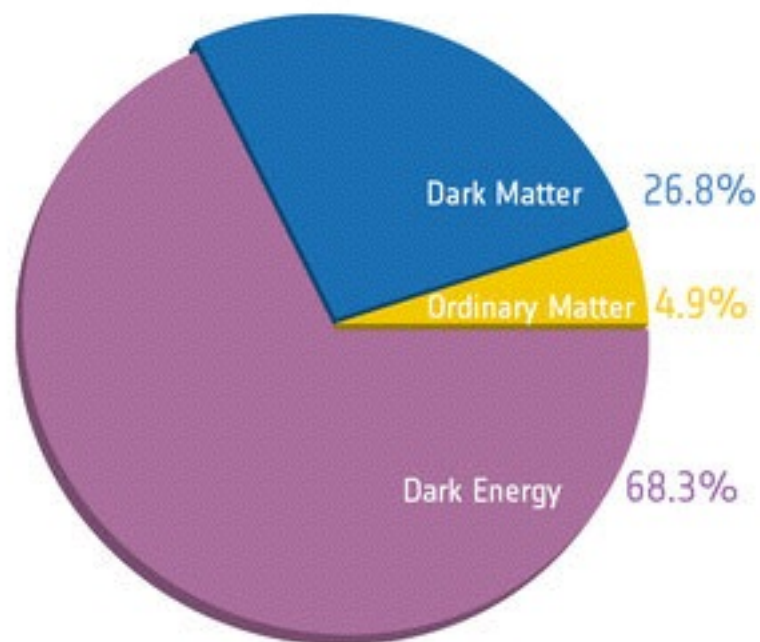
10 Gyrs

today

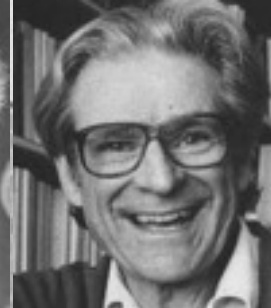
small shift in the pie chart make-up of the Universe



Before Planck

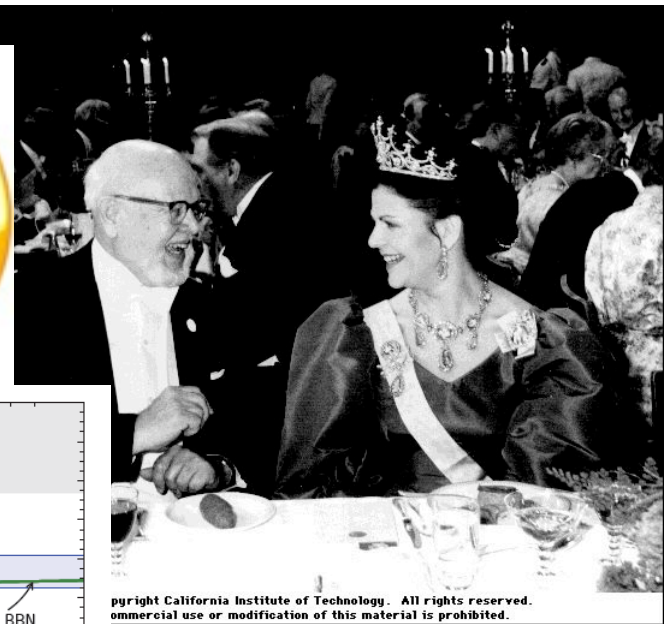
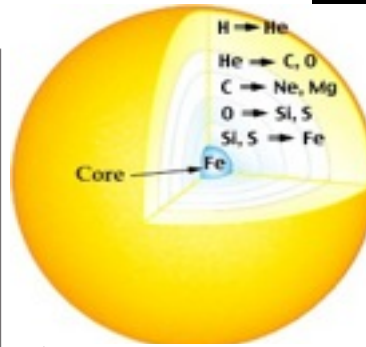
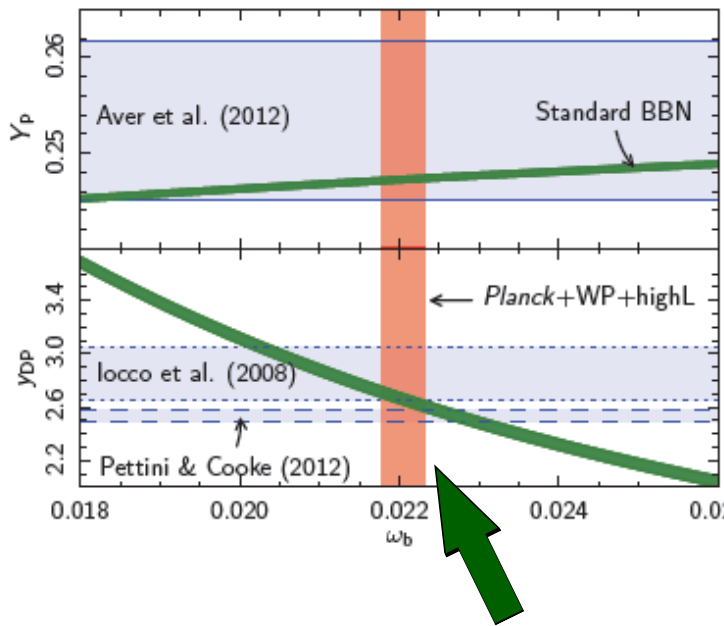


After Planck

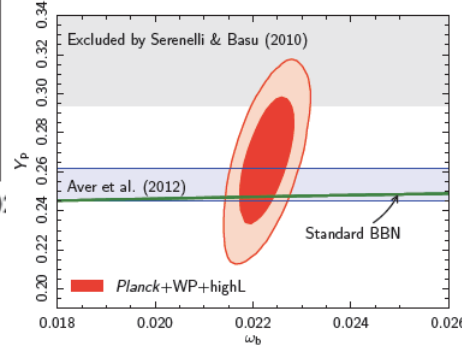


IOTA 1967, Cambridge B²FH 57, WFH 67, sn

Baryometers



Nobel Prize 84
Willy Fowler + Chandrasekhar

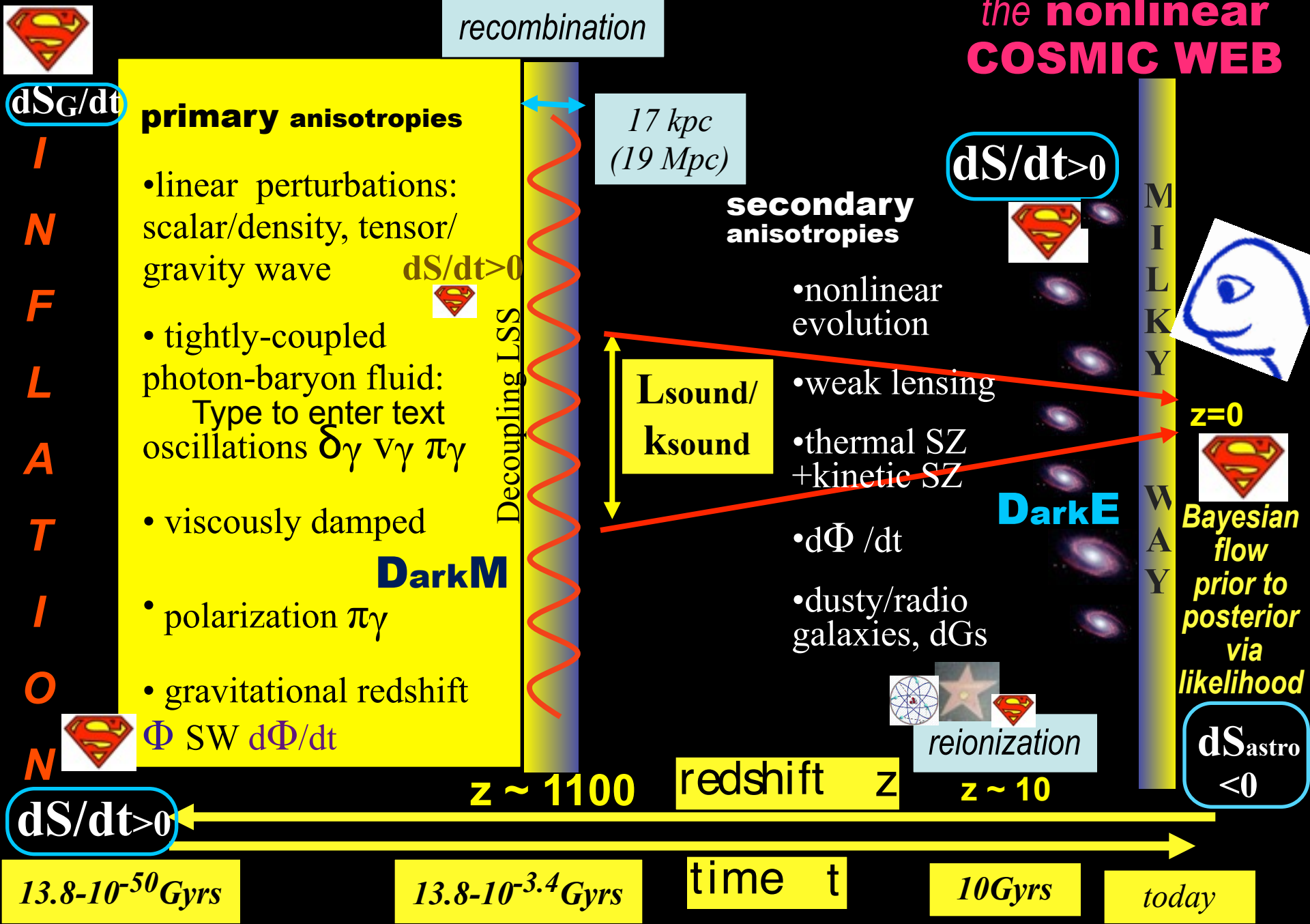


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	pre-boom	boom+	boom+cbi	boom+cbi+acbar	wmap1+
	January 2000	January 2002	June 2002	January 2003	March 2003
$\Omega_b h^2$	$0.0339^{+0.0443}_{-0.0246}$	$0.0222^{+0.0025}_{-0.0021}$	$0.0221^{+0.0024}_{-0.0020}$	$0.0221^{+0.0023}_{-0.0018}$	$0.0233^{+0.0013}_{-0.0013}$

- 0.0226 +/- 0.0006 wmap3+acbar+cbi+... LSS
- 0.0233 +/- 0.0005 wmap5+acbar+cbi+b03+...+WL+LSS+SNI+Lya
- 0.02217 +/- 0.00033 Planck I3+CMB Lensing
- 0.02214 +/- 0.00024 Planck I3+WP+hiL+BAO

the **nonlinear COSMIC WEB**



Simulation of the 7⁺ numbers

begets the **Cosmic Web** of clusters
now $a \sim 1$ & galaxies then $a \sim 1/4$

SIMPLICITY to COMPLEXITY under Gravity

void

filament

cluster

supercluster

\sim billion light years

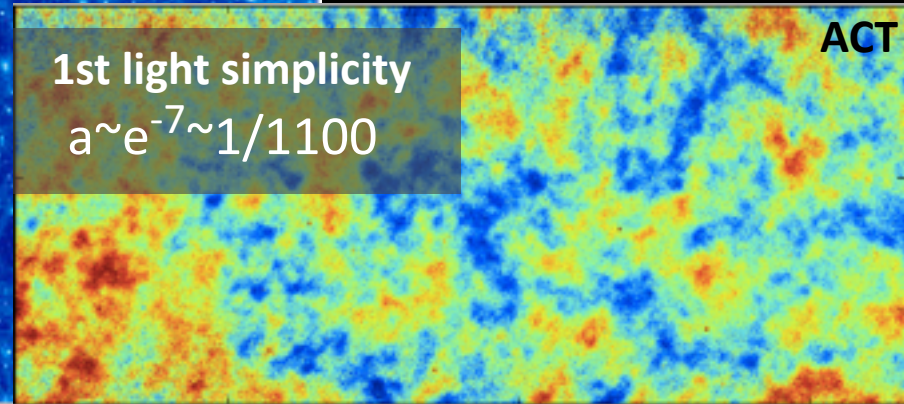
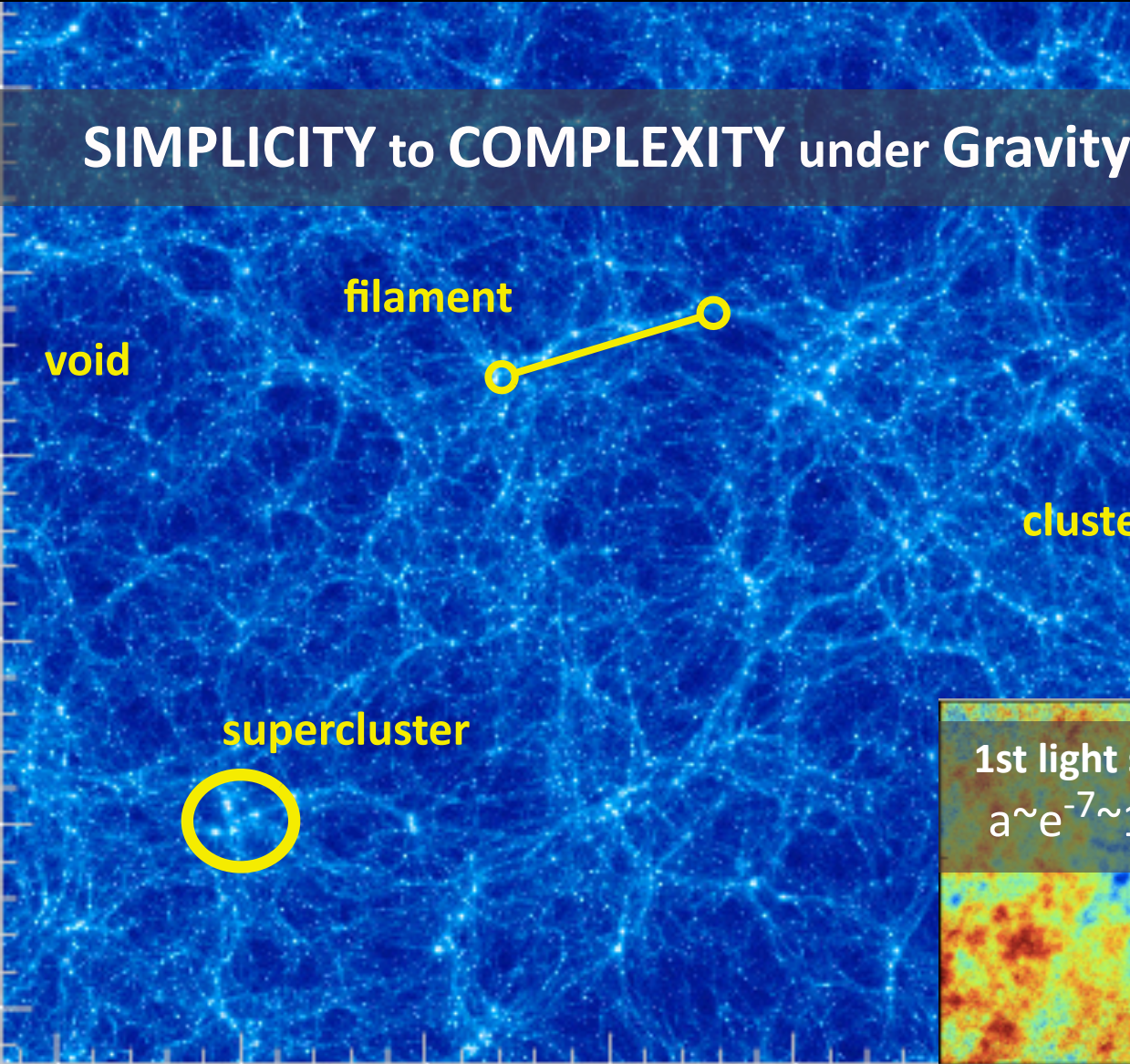
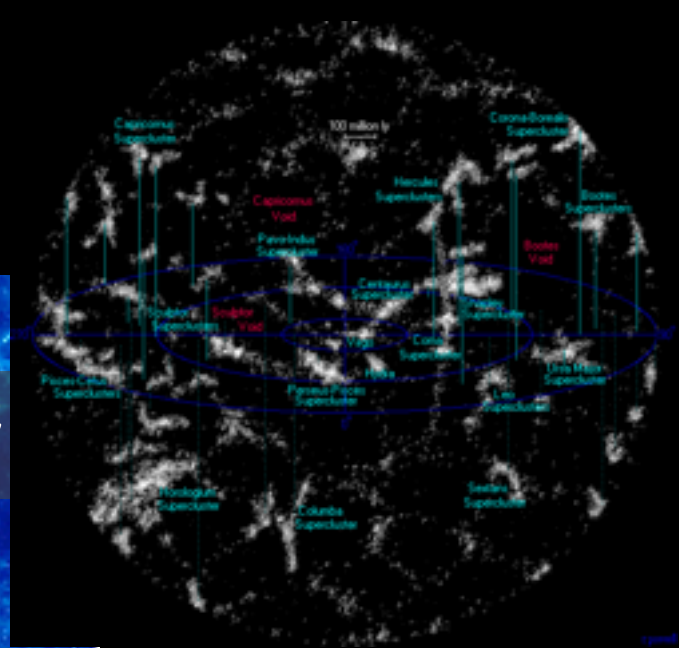
state of the art simulations
 $a \sim 1$ to $1/1.1$

ordinary matter
dark matter
dark energy

1st light simplicity

$a \sim e^{-7} \sim 1/1100$

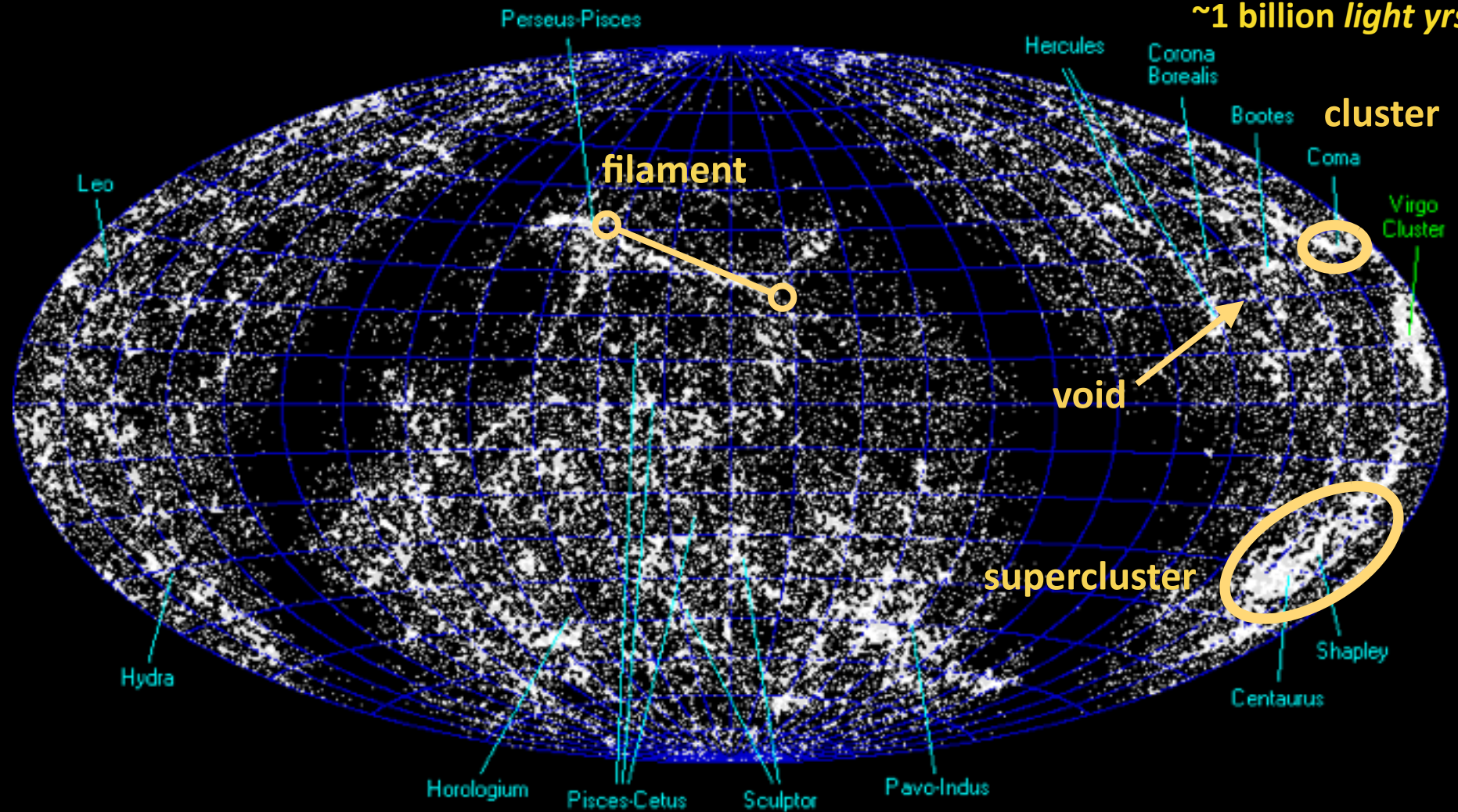
ACT



Cosmic Web of 60,000 nearby galaxies: exhibits “local” COMPLEXITY

$$a \sim e^{-0.1} = 1/1.1$$

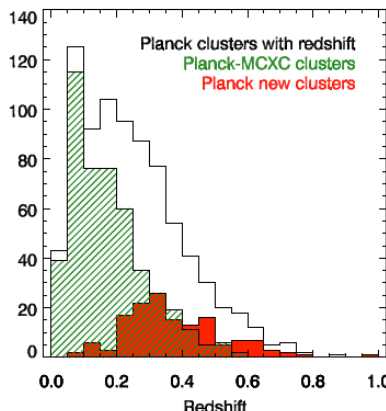
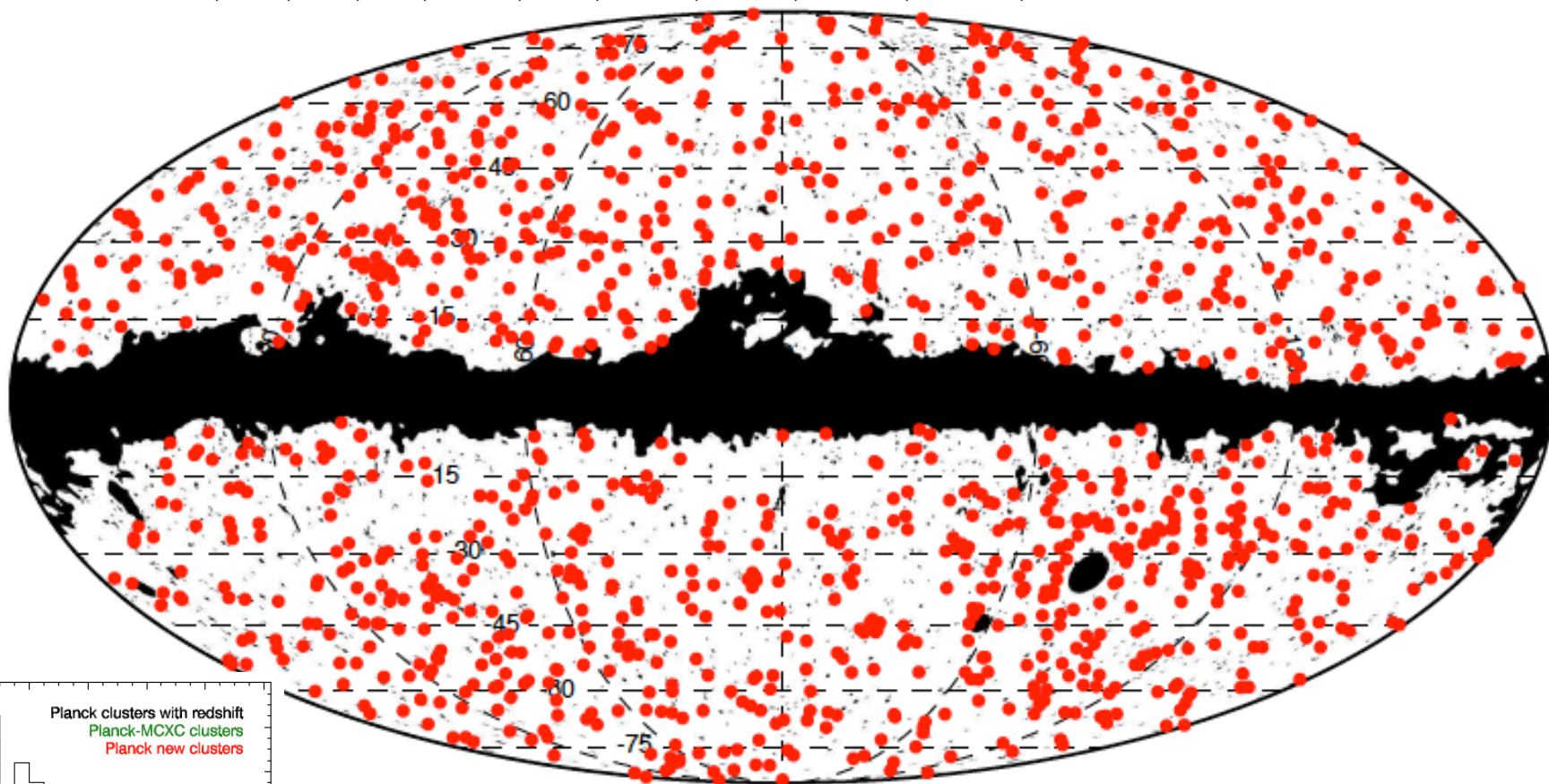
~1 billion light yrs



Compton cooling of high pressure / entropy electrons by the CMB

thermal SZ effect Planck2013 1227 clusters, SPT 224 =>747cls, ACT 91 cls
PSZ: 1227 clusters, 861 confirmed, 178 by Planck + 683 known, rest in class 1, 2, 3

cf. X-ray sample from ROSAT+ All-sky distribution of MCXC clusters ~1600 (Piffaretti et al 10)
REFLEX, BCS, SGP, NEP, MACS, CIZA, 400SD, 160SD, SHARC, WARPS, EMSS



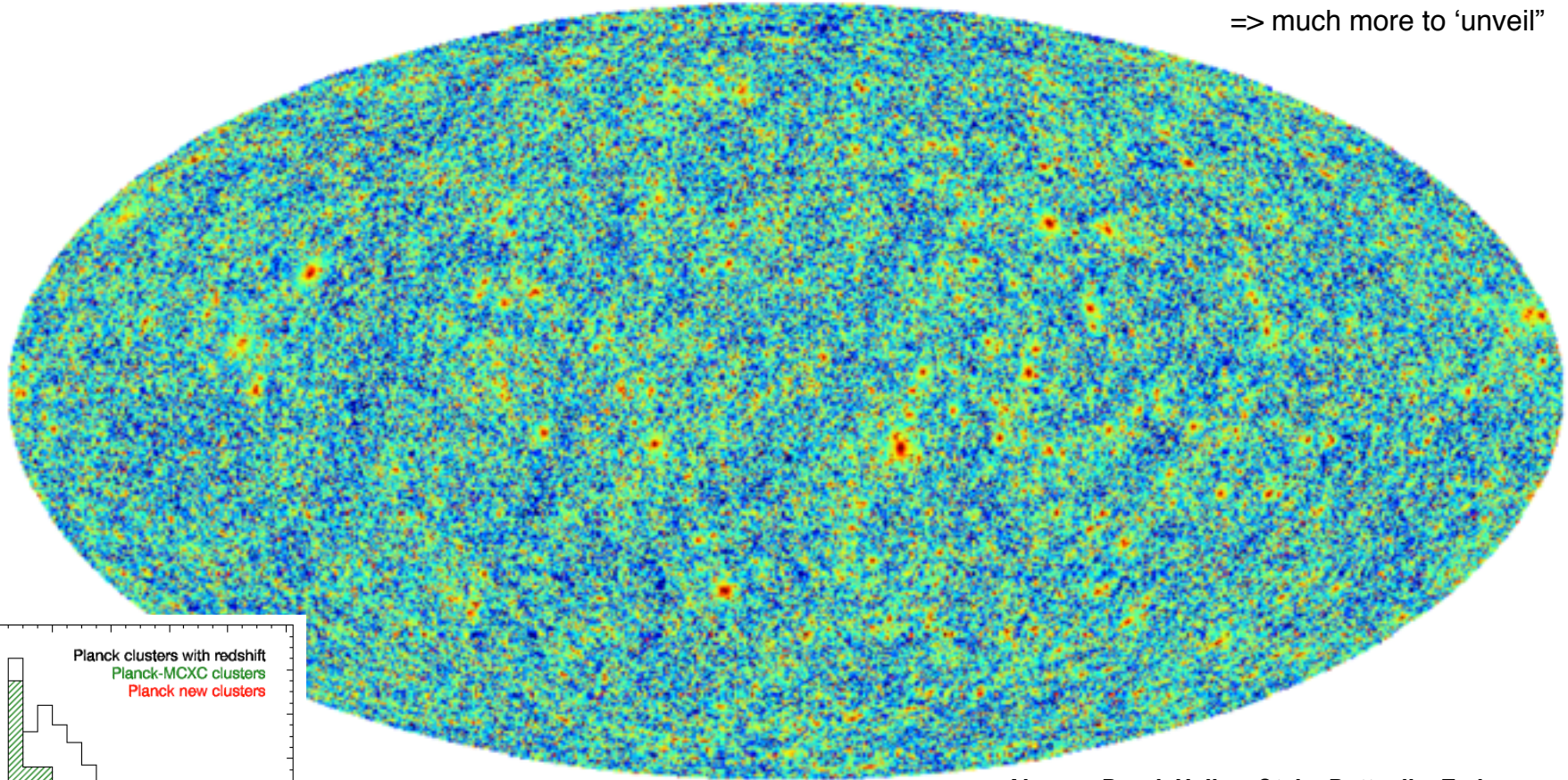
the Cosmic Web of Clusters, seen thru Compton cooling of high pressure electrons by the CMB

tSZ
effect

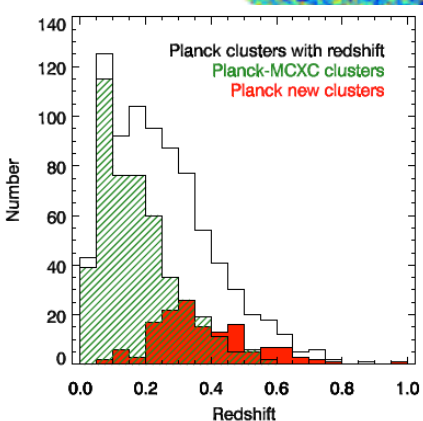
Lightcone Simulation of 35000 Clusters $> 2 \times 10^{13} M_{\text{sun}}$ to $z=0.5$ in projected pressure

Planck all-sky, $\sim 5'$, ACT ~ 1000 sq deg $1.4'$, SPT ~ 2500 sq deg $1'$

=> much more to 'unveil'



Alvarez, Bond, Hajian, Stein, Battaglia, Emberson,..2014

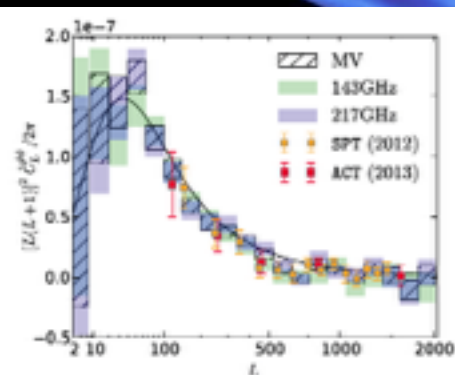
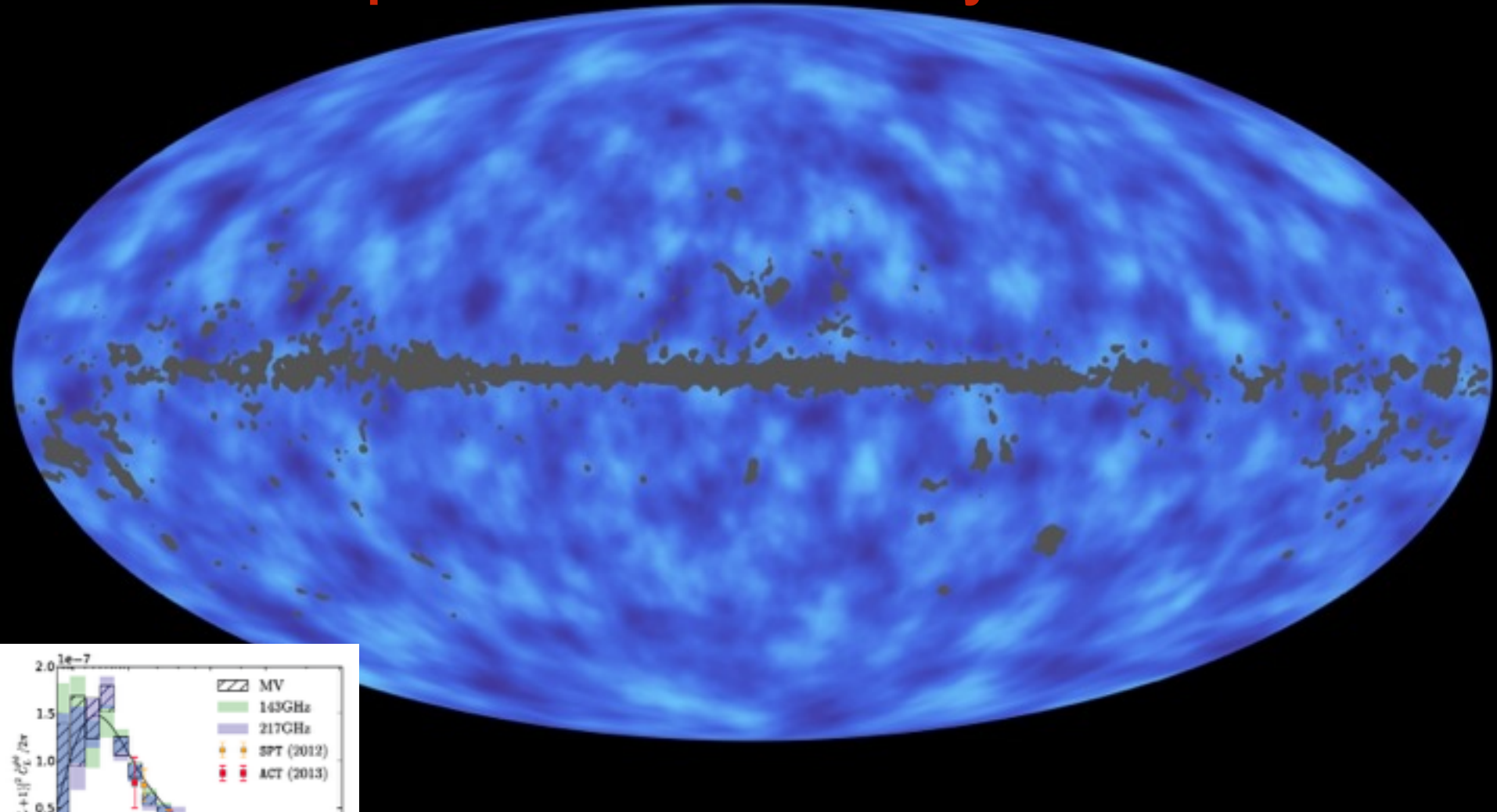


-8 -5
log Compton-y

wall clock ~ 1.5 hrs on 256 cores of SciNet@UofT,
painted-on BBPS mean-fields from hydro sims with AGN feedback

Planck1.3 CMB Lensing: reconstructed projected gravitational potential map (!)
~ **dark+baryonic matter map, mean-field map** = Wiener filter (beware: fluctuations about mean-field)

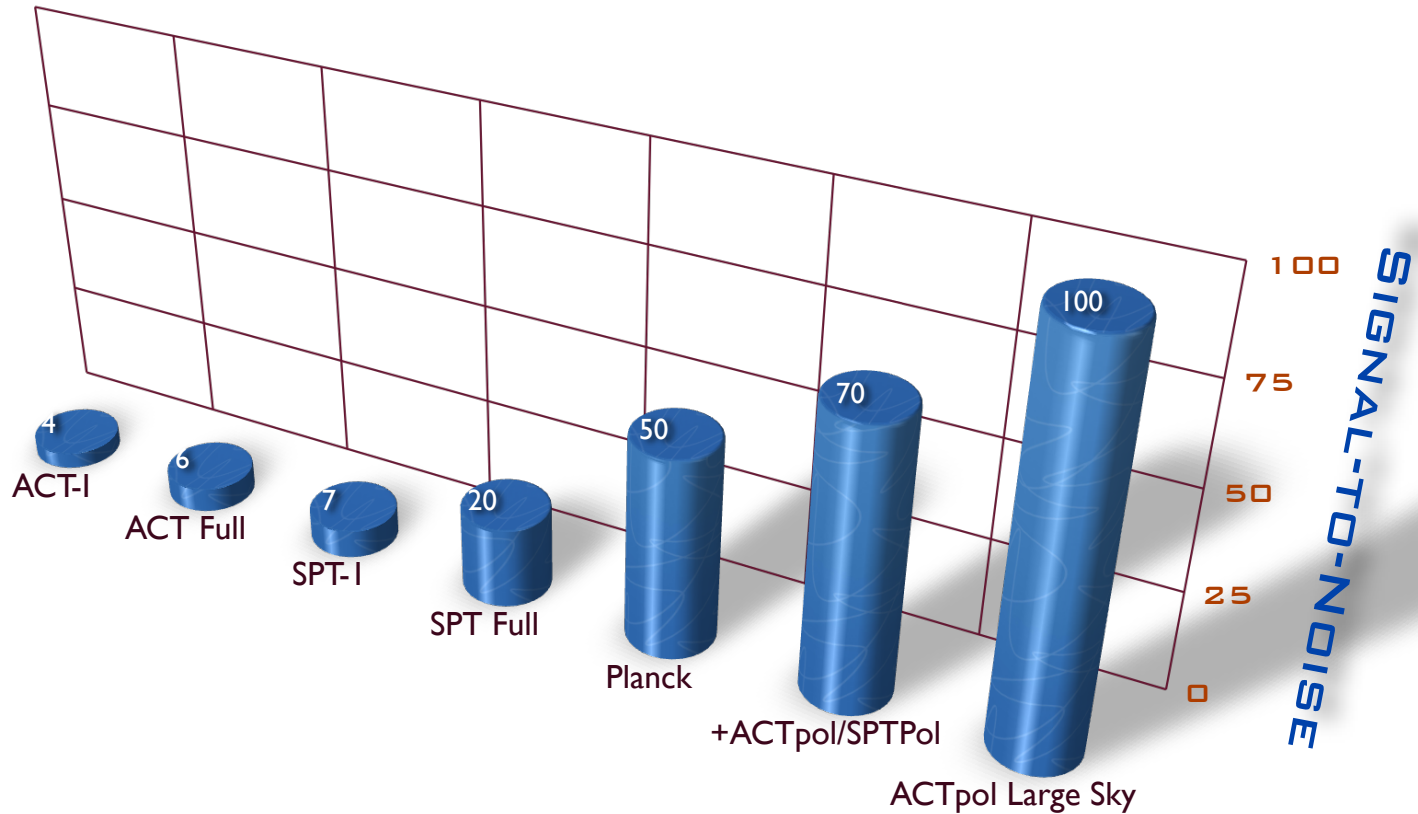
photons under strain by tides



related to primordial scalar curvature map

photons under strain by tides

**CMB LENSING IS GOING TO EXPLODE AS A FIELD
IN THE NEXT FEW YEARS**



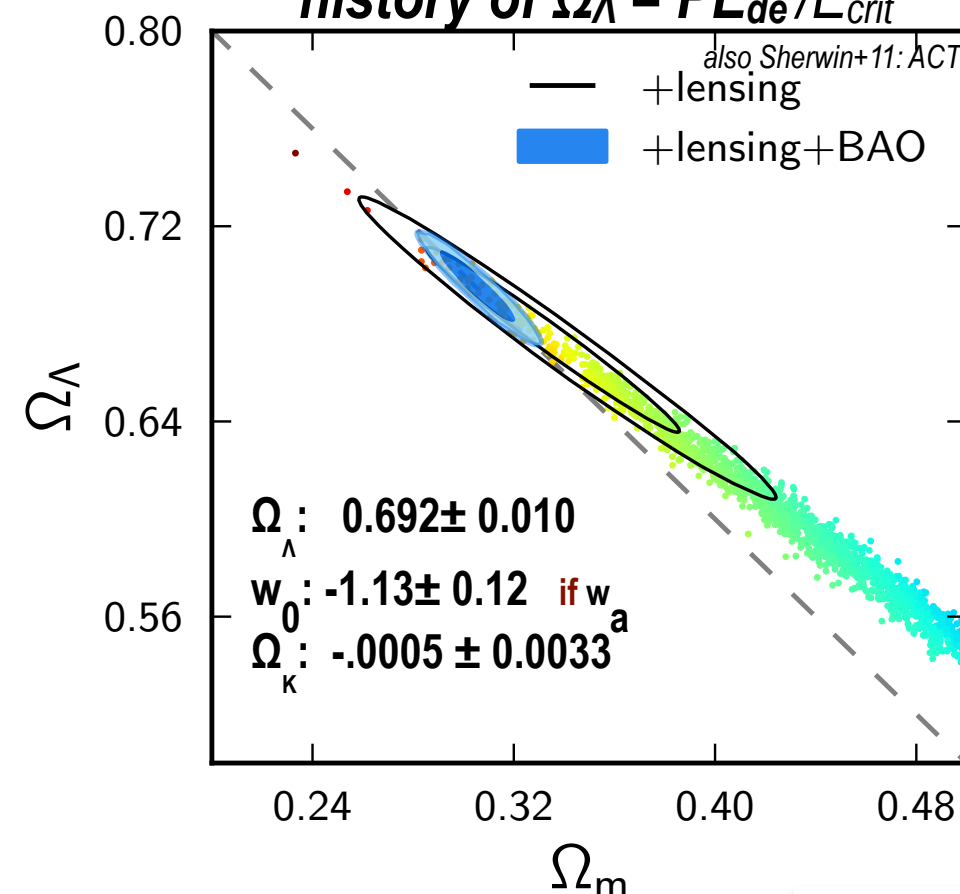
Dark Energy => inflation now

CMB lensing breaks "geometrical degeneracy":

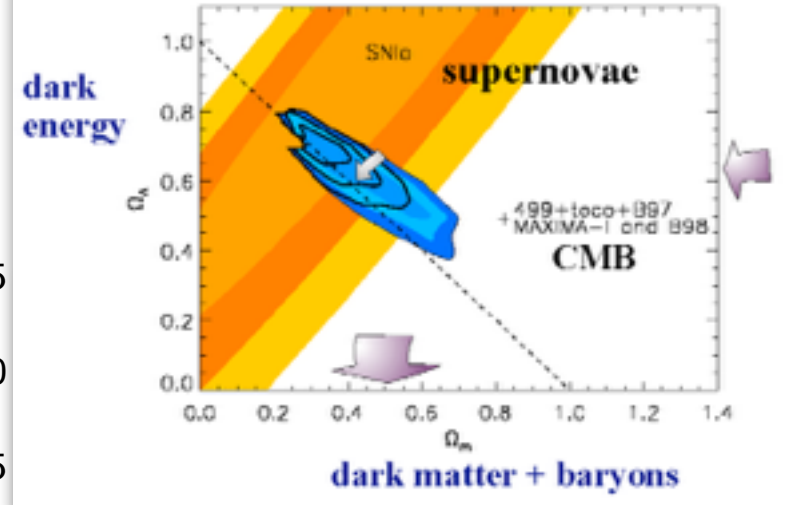
Planck alone cf. Planck+BAO

Planck1.3 cf. CMB+LSS

history of $\Omega_\Lambda = PE_{de}/E_{crit}$



BOOM 2000



vintage 1998 conclusions

H_0

CMB @ LSS \downarrow $\omega_{CDM} \ll \Lambda_{CDM}$

SN Ia \downarrow

high z clusters \downarrow

$\Omega_{cdm} \sim 0.3$

$\Omega_b \sim 0.04$

$H_0 \sim 65-70$

$t_0 \sim 12-14 Gyr$

$\Omega_\Lambda \sim 0.014$

$(\frac{M_p}{\sigma_{100}^2})^{2/3}$

$\Omega_\Lambda(z, t) \approx \frac{2}{3}$

vac

PLATE TIME

INFLATION IS NOW

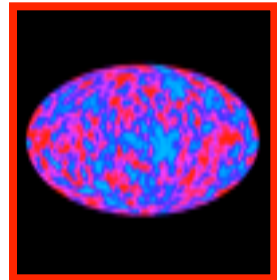
$\rho \sim 10^{14} \sim \text{milli-eV}$

B+Jaffe '96, '98

$\Omega_\Lambda \approx 2/3 \pm .07$ +LSS

$n_s = .98 \pm .07$

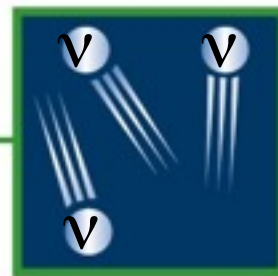
$.96 \pm .06$



Radiation:
0.005%



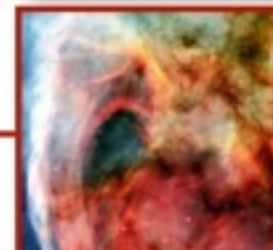
Chemical Elements:
(other than H & He) 0.025%



Neutrinos:
> 0.47%

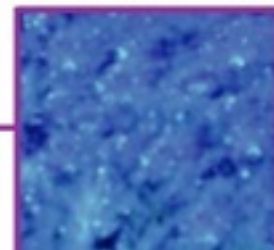


Stars:
0.5%



**Free
H & He:**
4.3%

$\Omega_{\text{total}} = 1$



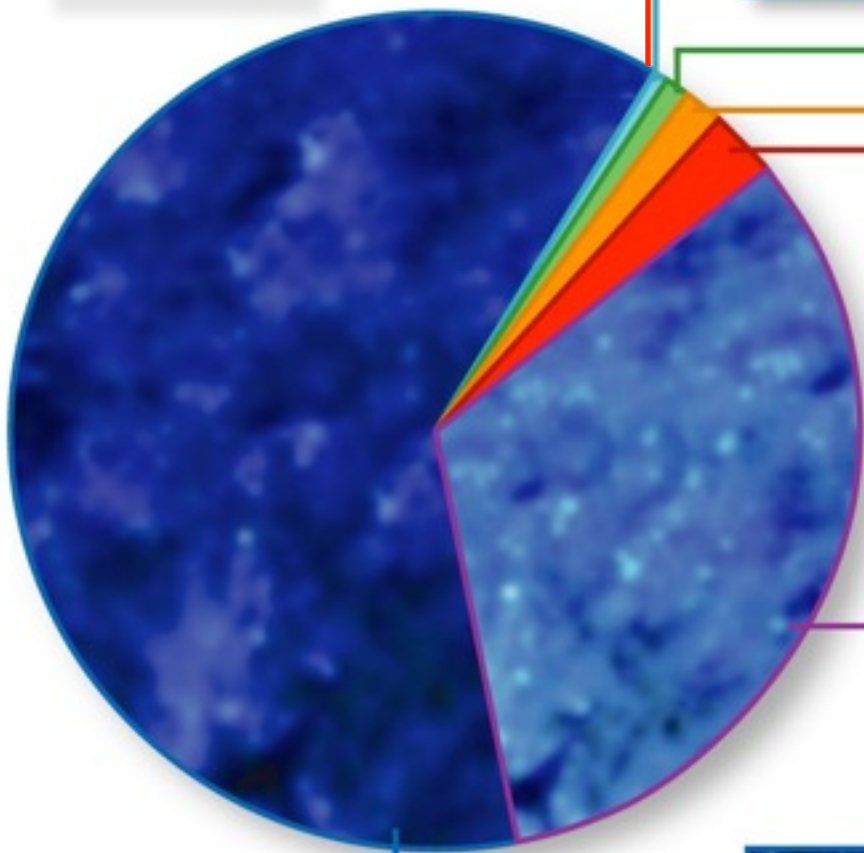
Dark Matter:

$\Omega_{\text{dm}} = 26.0 \pm 1\%$



Dark Energy:

$\Omega_{\text{de}} = 69.2 \pm 1.0\%$



Gravity Waves

$\Omega_{\text{GW}} \sim 10^{-14} - 10^{-10}$ LIGO

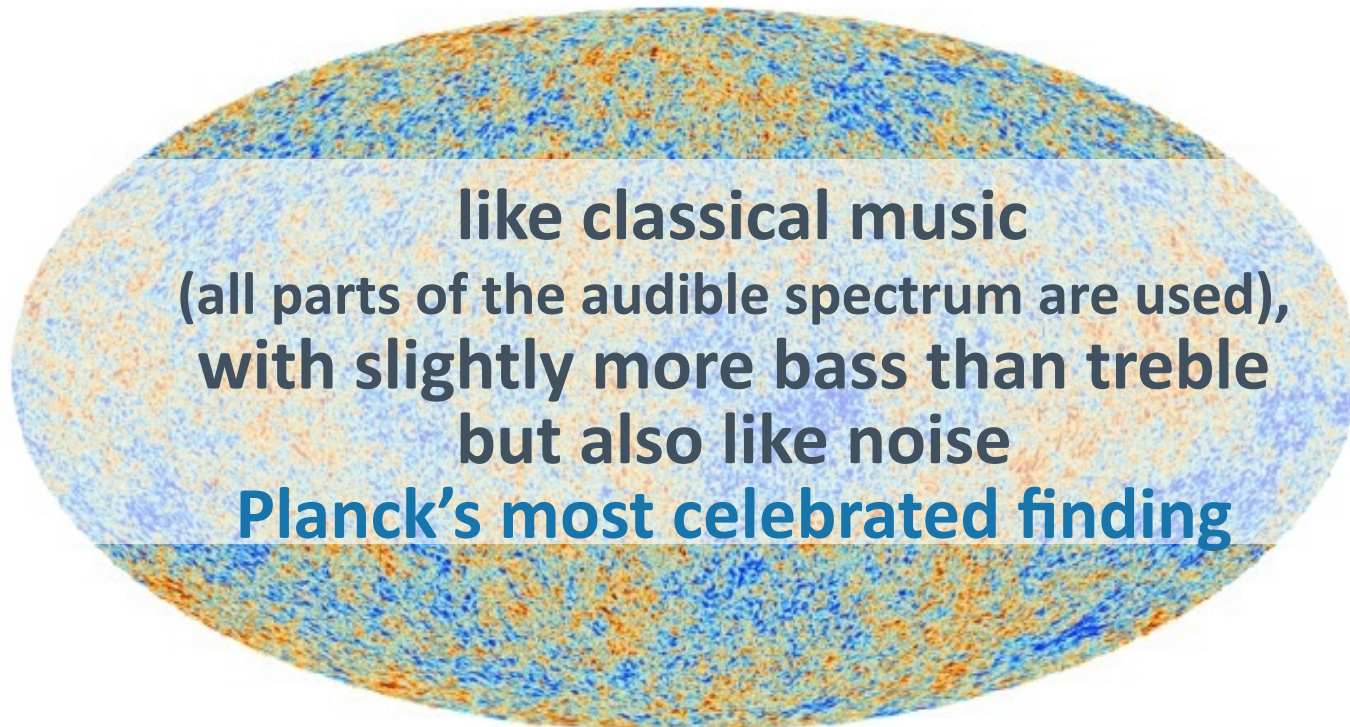
$\Omega_{\text{BlackHoles}} \sim 10^{-7}$

CMB reveals **ultra-early Universe sound waves**

=> the inharmonious '*music of the spheres*' **in 7+ numbers**

=> learn **matter & energy content & structure** at $a \sim e^{-7}$ 380000 yr

=> infer structure **far far earlier** $a \sim e^{-127} \sim 1/10^{55}$ **in 2 numbers**



like classical music
(all parts of the audible spectrum are used),
with slightly more bass than treble
but also like noise
Planck's most celebrated finding

standard inflation space: $P_s n_s \frac{dn_s}{d \ln k} r = T/S$ @k-pivots

5 σ from 1 $n_s = 0.9608 \pm 0.0054$ -0.014 ± 0.009 $r < 0.12$ cf. **BICEP2**

$\ln Power_s \sim \ln 22.0 \times 10^{-10} \pm 0.025$

$r = 0.20^{+0.07}_{-0.05}$

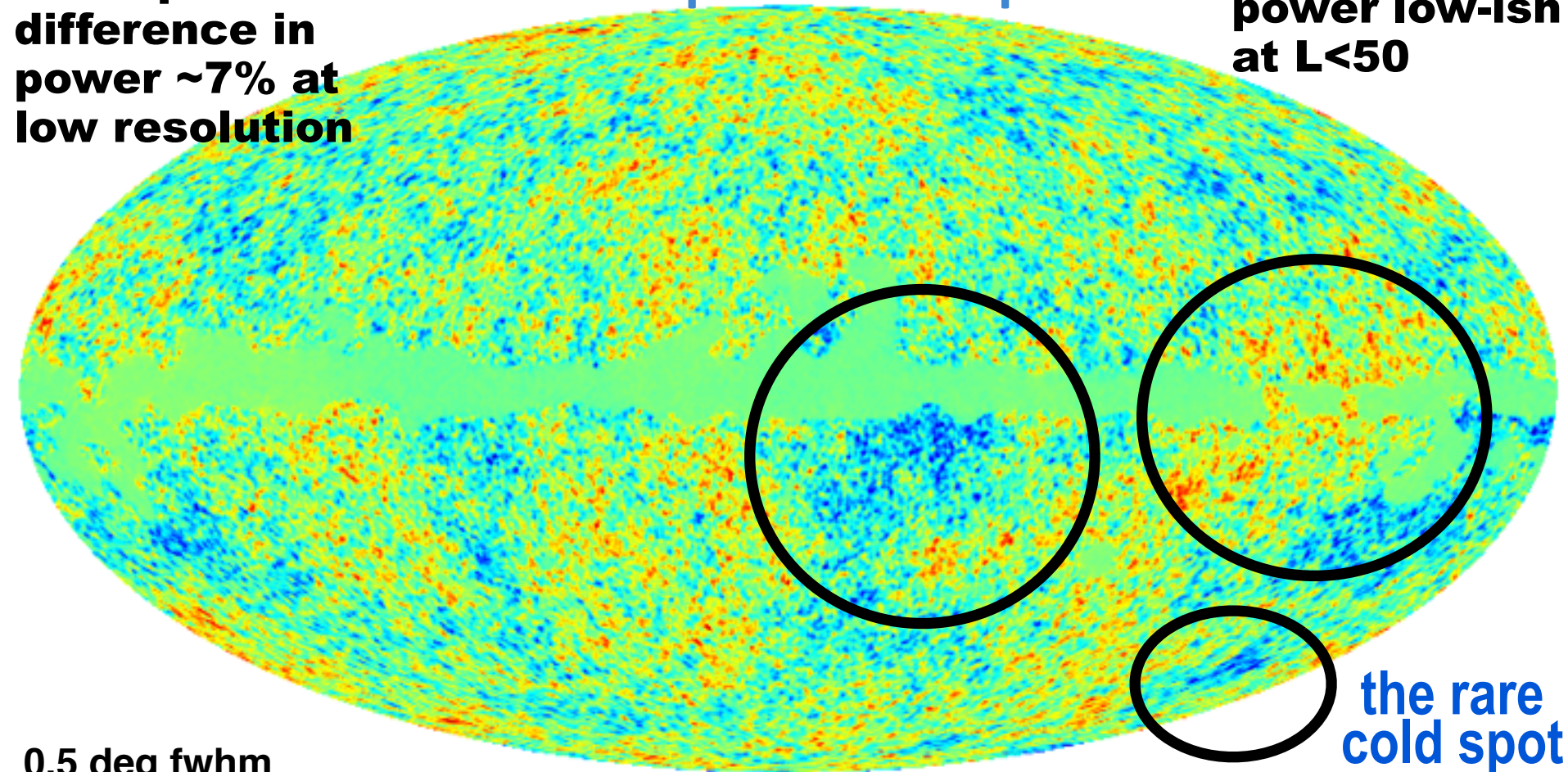
Gaussian to high precision for high multipole,
anomalies at low multipoles, non-Gaussian, anisotropic
anomalies => inflation COMPLEXITY at $t \sim 10^{-36}$ seconds?

mean temperature, 1000 realizations, smooth scale fwhm = 30 arcmin,

temperature map

power low-ish
at $L < 50$

hemisphere
difference in
power $\sim 7\%$ at
low resolution



-355.

+340.

Grand Unified Theory of Anomalies? TBD

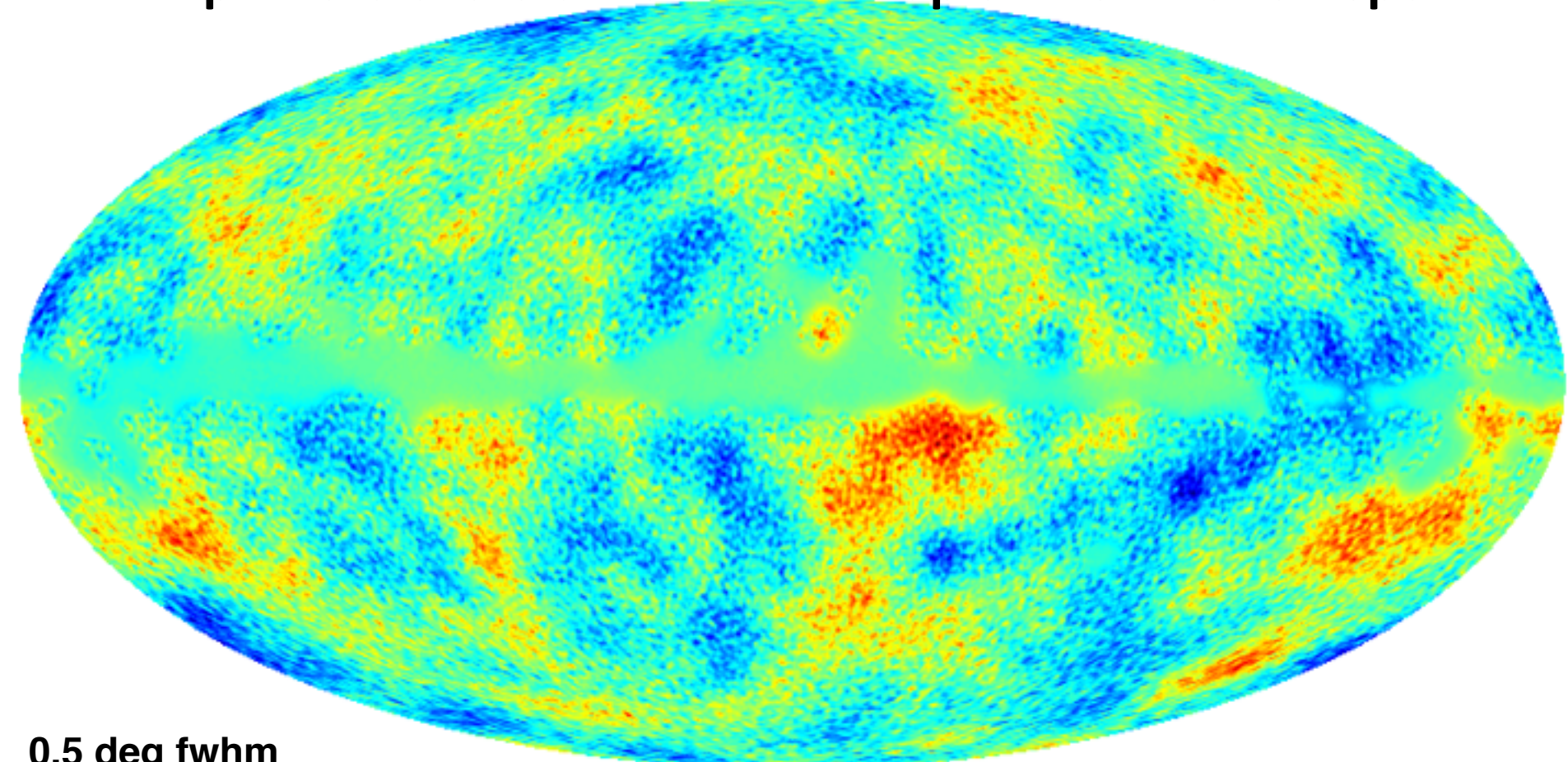
intermittent strain-power bursts (in curvature)?

reveals map of **primordial isotropic strain / phonons**

$\int d\text{visibility}(\text{distance}) < \text{Trace}(\alpha) | \text{Temp} > \text{ (angles, distance)}$

mean zeta, 1000 realizations, smooth scale fwhm = 30 arcmin,

=> **primordial scalar curvature map of the inflation epoch**



0.5 deg fwhm

Reconstructing the Early Universe

-4.70

+5.18

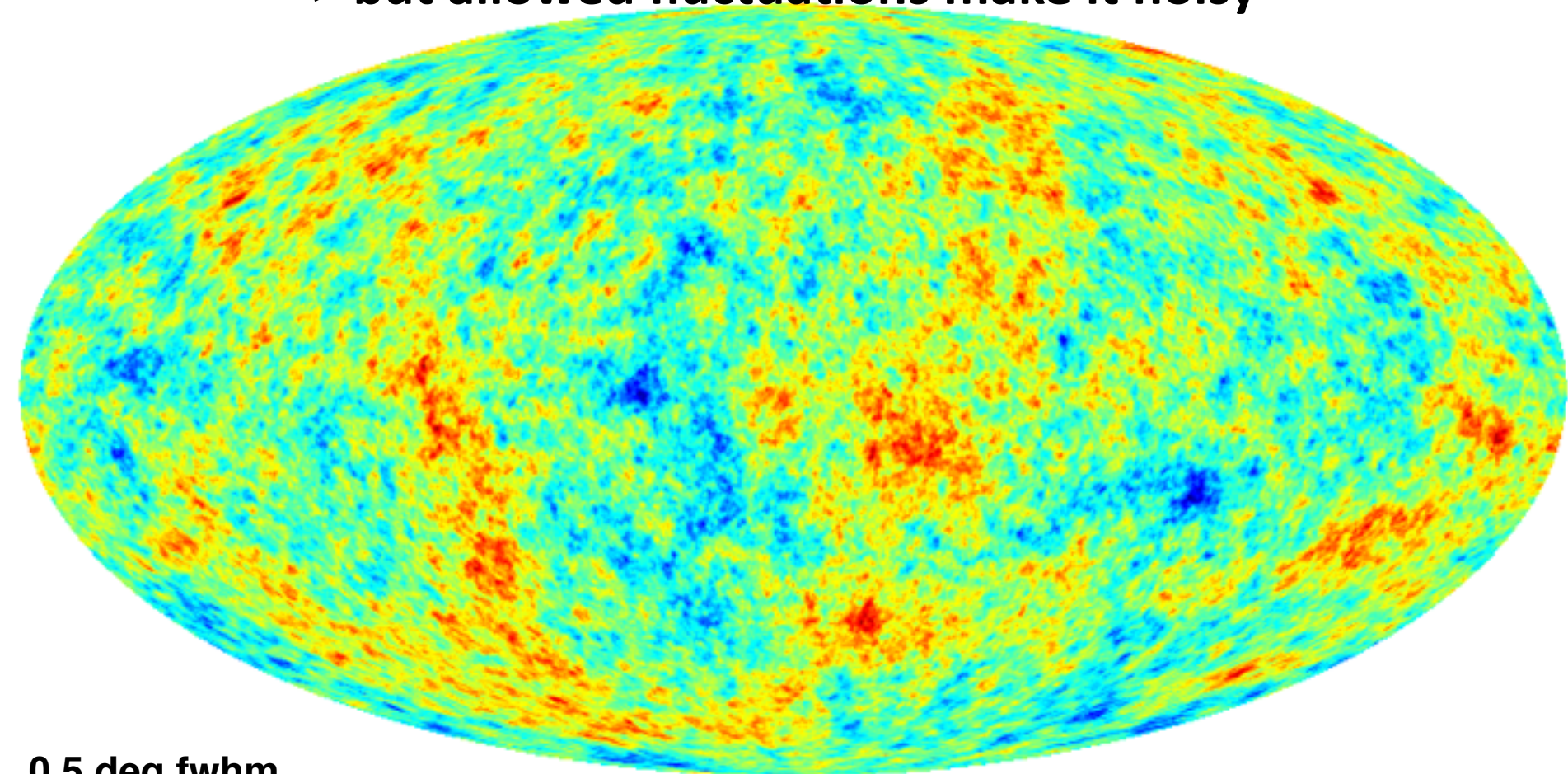
visibility mask

reveals map of **primordial isotropic strain / phonons**

$$\int d\text{visibility}(\text{distance}) \langle \text{Trace}(\boldsymbol{\alpha}) | \text{Temp} \rangle + \delta \text{Trace}(\boldsymbol{\alpha})$$

one realization of fullsky zeta, fwhm = 30 arcmin

=> but allowed fluctuations make it noisy



0.5 deg fwhm

-8.61

+7.54

Reconstructing the Early Universe

visibility mask

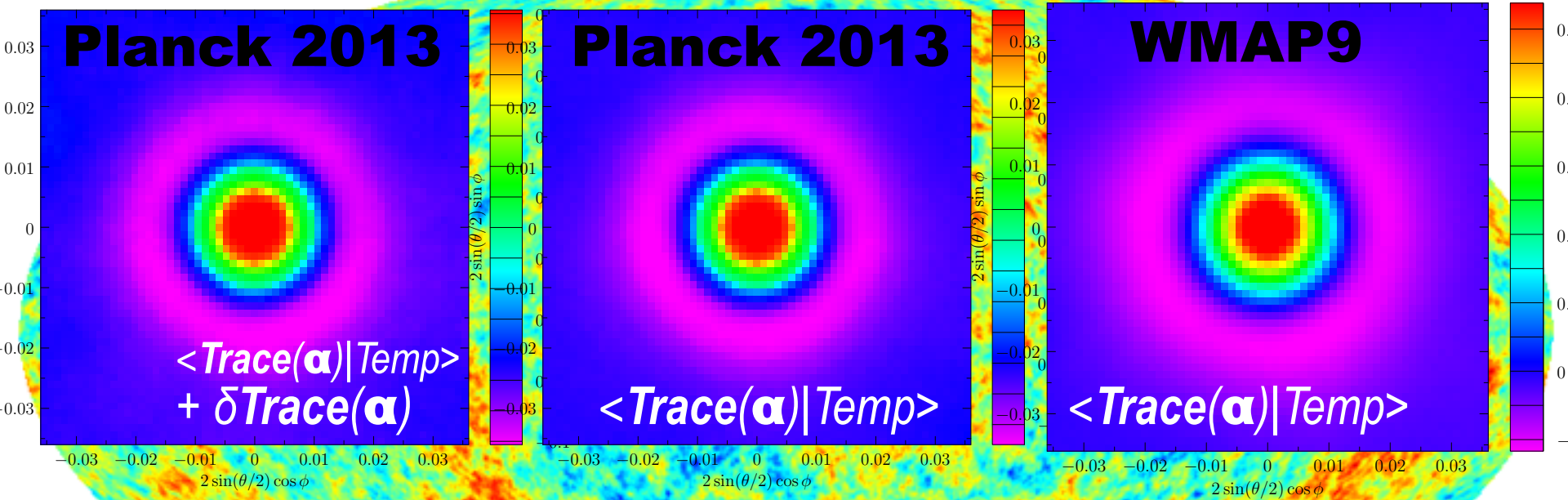
reveals map of **primordial isotropic strain / phonons**

$$\int d\text{visibility}(\text{distance}) < \text{Trace}(\boldsymbol{\alpha}) | \text{Temp} > + \delta \text{Trace}(\boldsymbol{\alpha})$$

one realization of fullsky zeta, fwhm = 30 arcmin

=> but allowed fluctuations make it noisy

one realization of ζ map, 11113 patches on T maxima, random orientation stacking mean ζ map, 11113 patches on T maxima, random orientation 9257 mean ζ patches on T maxima, random orientation



0.5 deg fwhm



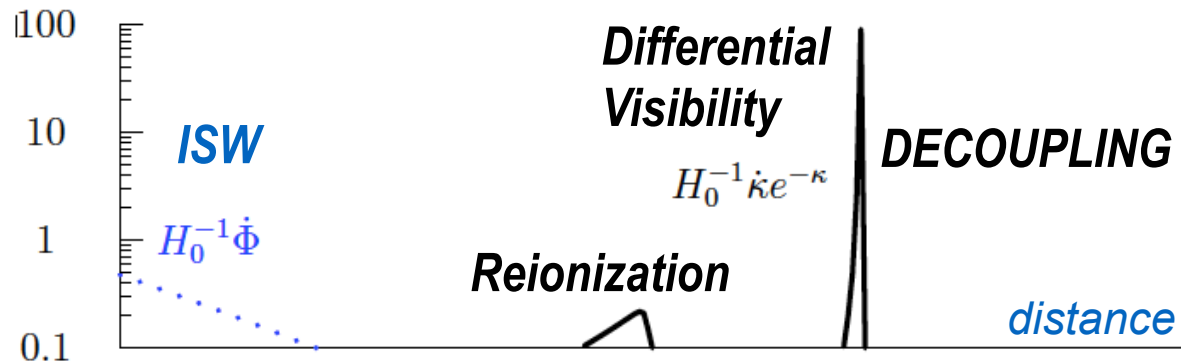
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=> but allowed fluctuations make it noisy



CMB-probe no tomography (radial distance (redshift)):

CMB-probe ~ differential visibility
at decoupling/recombination (all L)
reionization/reheating (low L)

CMB-probe ~ changing gravitational potential

Integrated Sachs Wolfe effect (low L), Rees-Sciama effect (hi L)

available modes: $f_{\text{sky}} L_{\text{max}}^2 - f_{\text{sky}} L_{\text{min}}^2$ $L_{\text{max}} \sim L_{\text{damp}}$

Large Scale Structure Galaxy Surveys

available modes $\sim f_{\text{sky}} L_{\text{max}}^2 k_{\text{max}} d_{\text{max}}$
 $\sim f_{\text{sky}} (k_{\text{max}}^3 d_{\text{max}}^3), \quad k_{\text{min}} \sim 2\pi/d_{\text{max}} \quad V_{\text{com}} \sim d_{\text{max}}^3$

ultra-early Universe sound spectrum $\ln P_s(\ln k)$

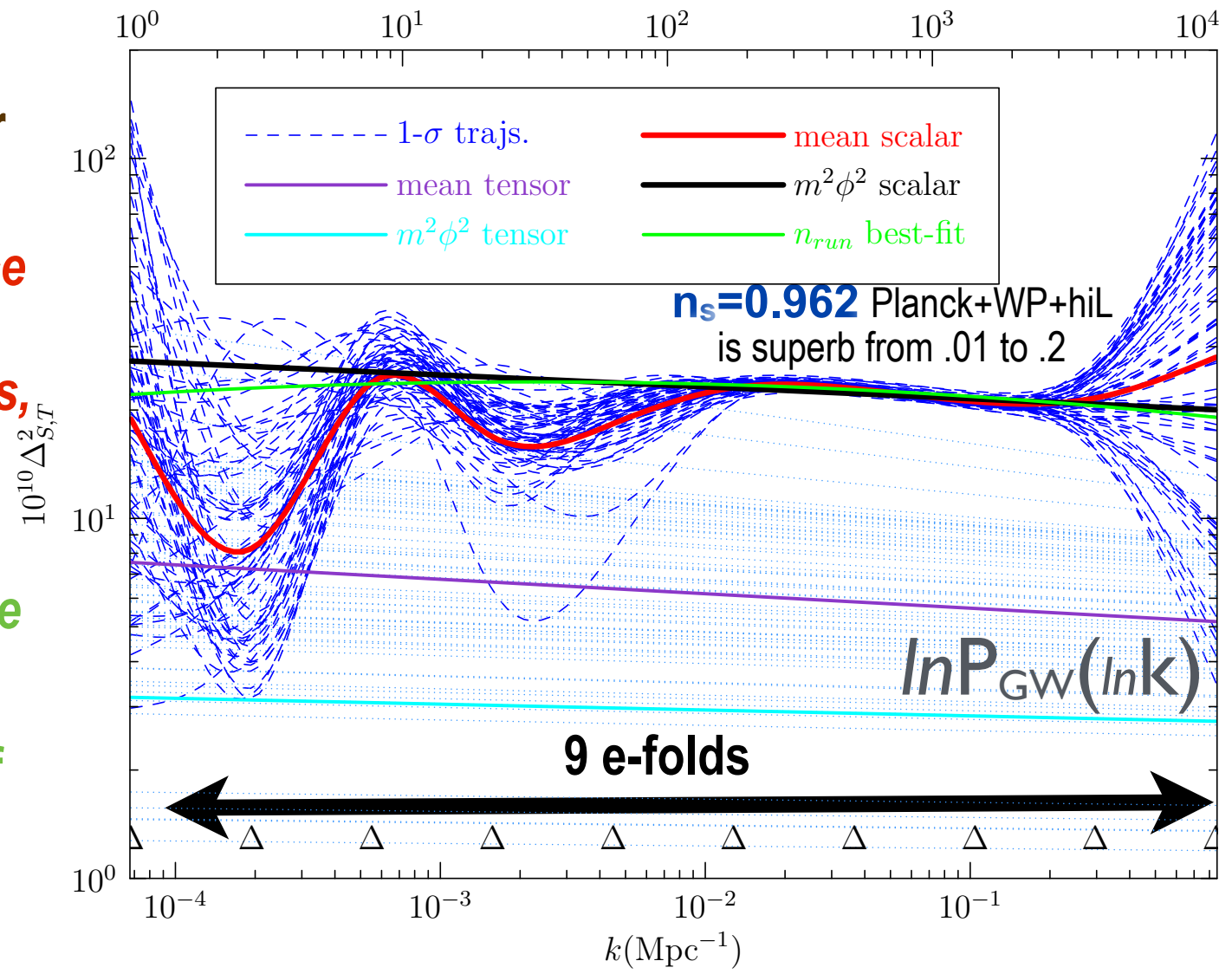
$$\ell \equiv kD_{\text{rec}}$$

**new parameters:
trajectory
probabilities for
early-inflatons**

**no strong evidence
for oscillation
patterns, cutoffs,
local features**

**but hints of
change on large
 $L < 100$ scales**

**PS: running of
 P_s is a bad fit**



Bond, Braden, Huang, Frolov, Vaudrevange 2014

scan $\ln P_s(\ln k)/A_s$, $\ln A_s = \ln P_s(k_{\text{pivot},s})$, $r(k_{\text{pivot},t})$; consistency \Rightarrow reconstruct $\epsilon(\ln H a)$, $V(\psi)$

ultra-early Universe sound spectrum $\ln P_s(\ln k)$

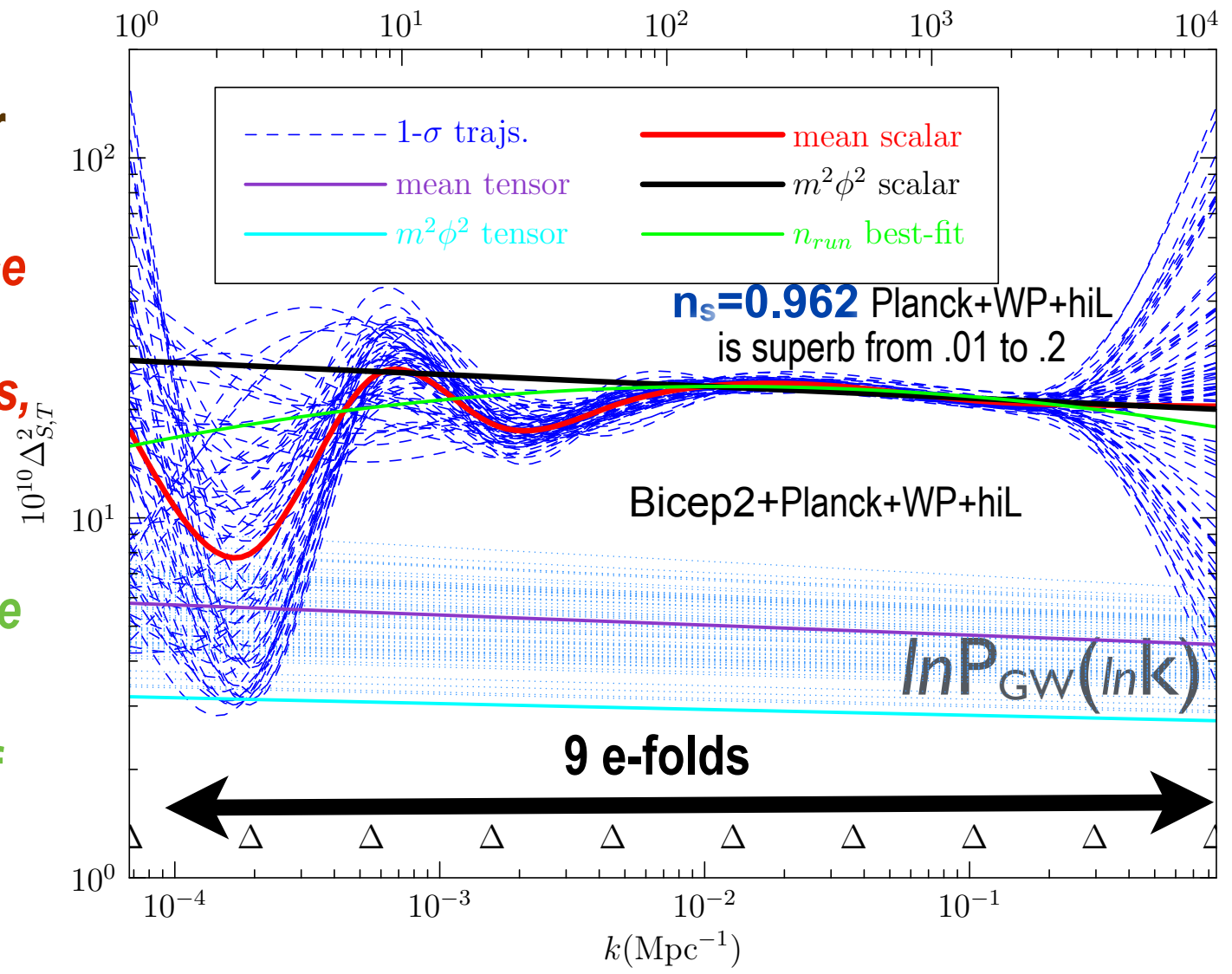
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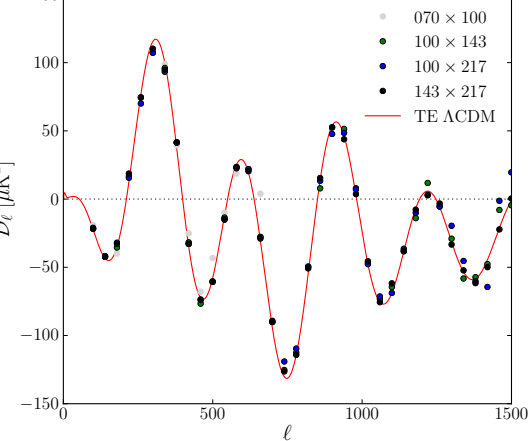
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CMB Peak Statistics

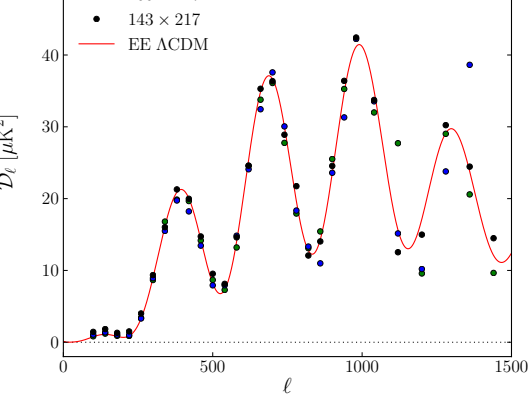
temperature stacked on temperature Peaks

polarization rotated & stacked on temperature Peaks

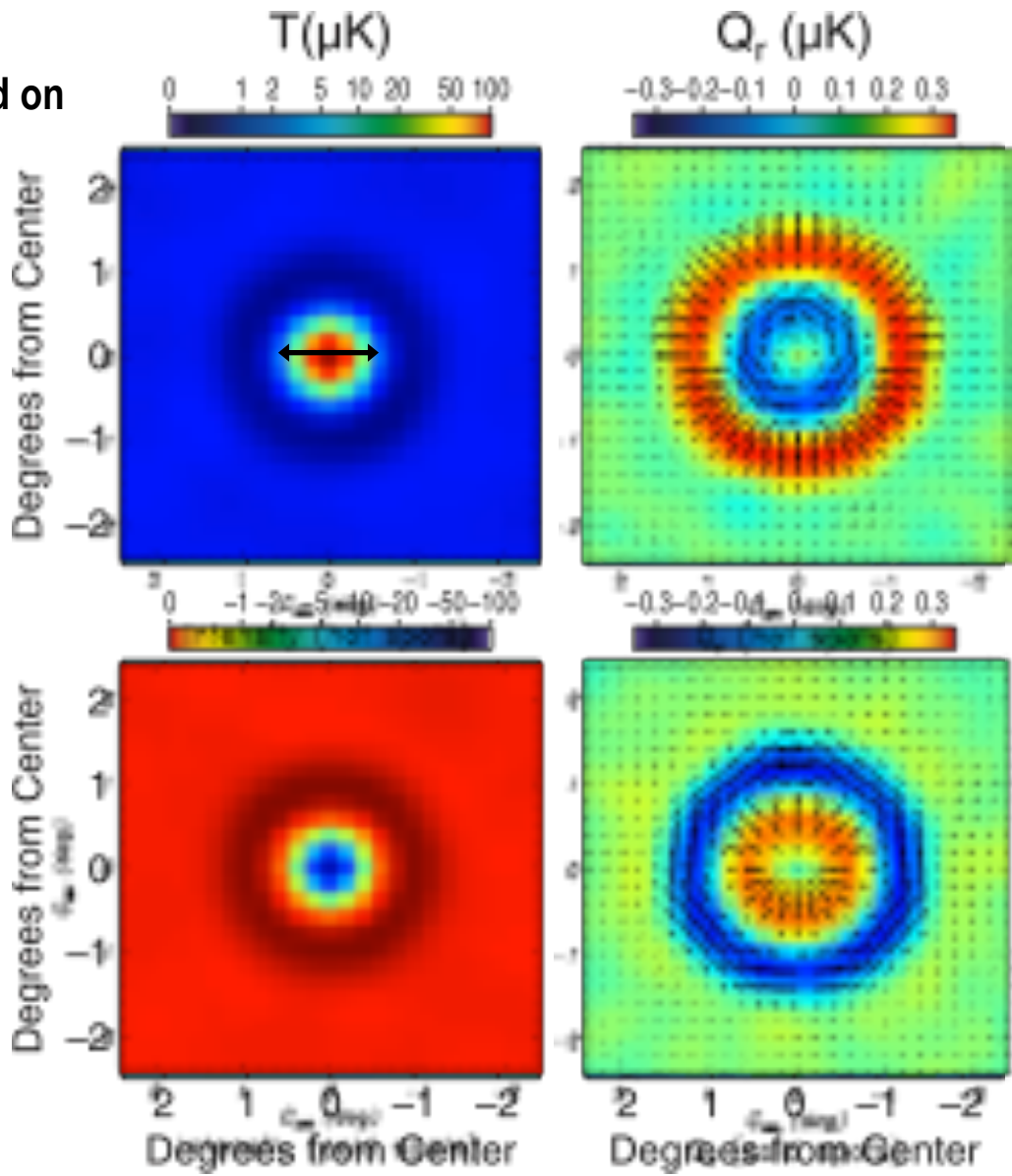
TE intensity X polarization



EE polarization



CMB Polarization BAO in the CMB - Planck2013



Planck2013 teaser for Planck2014 polarization release

E mode patterns



no B here

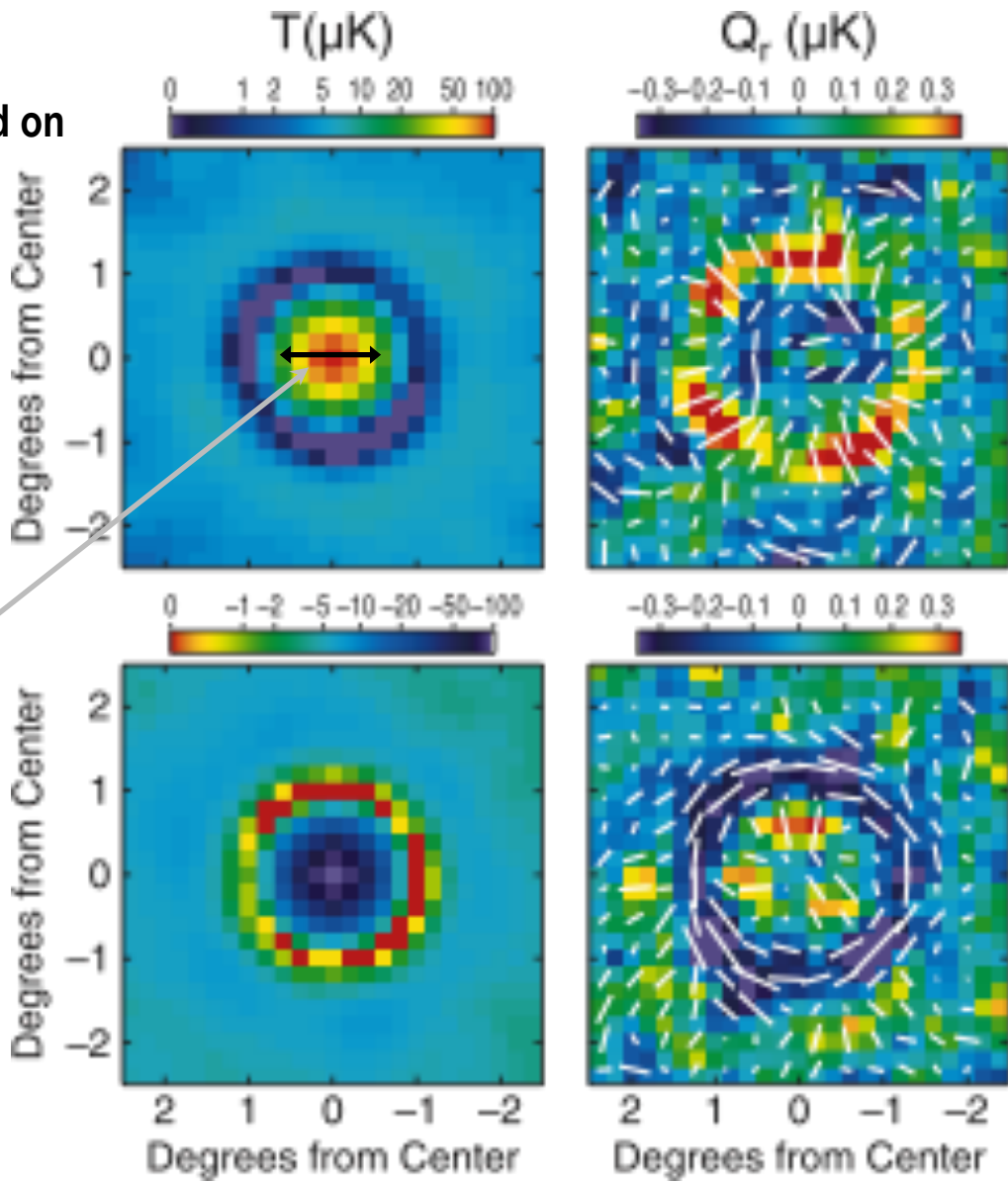
Planck2014, 2015 ACTpol, ABS, Spider, AdvACT, GLP, ..

CMB Peak Statistics

temperature stacked on temperature Peaks

polarization rotated & stacked on temperature Peaks

CMB Polarization BAO in the CMB – WMAP9



BAO scale:
 145.8 ± 1.2 Mpc

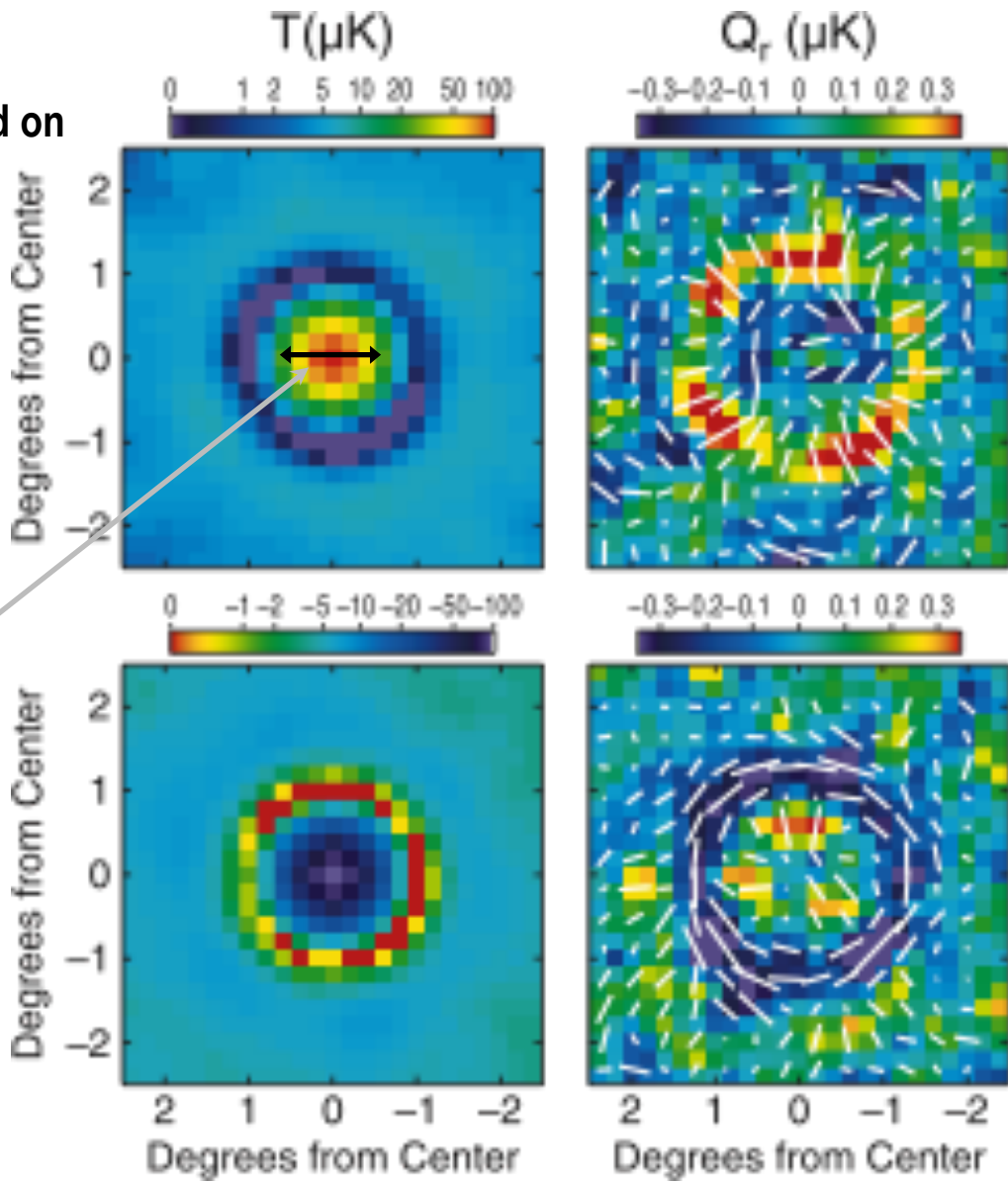


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CMB Polarization BAO in the CMB – WMAP9



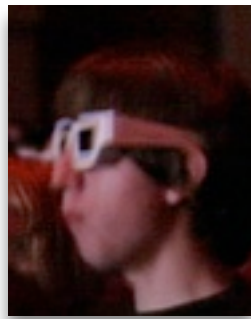
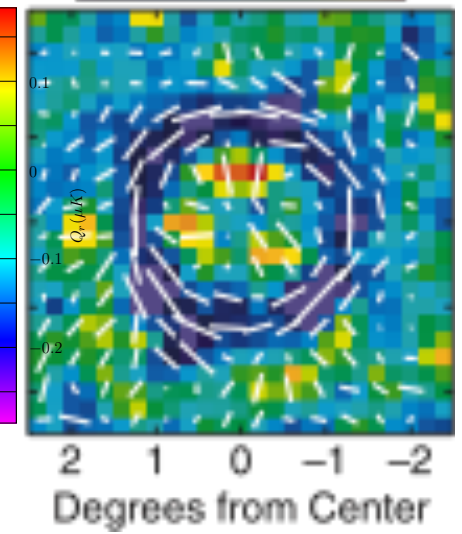
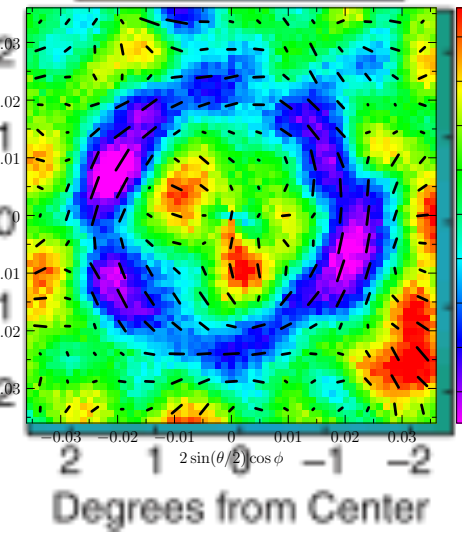
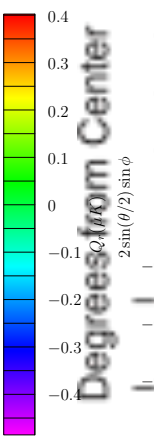
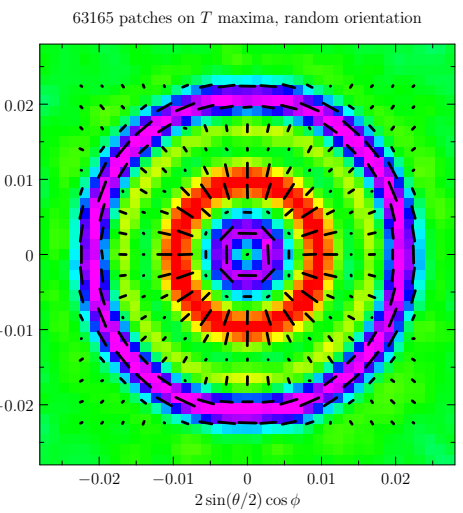
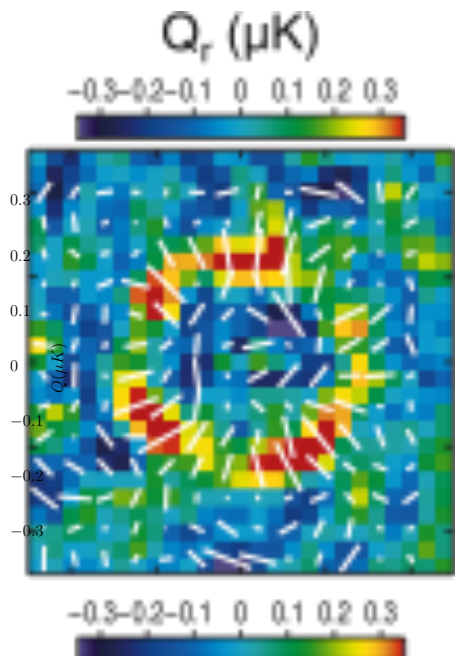
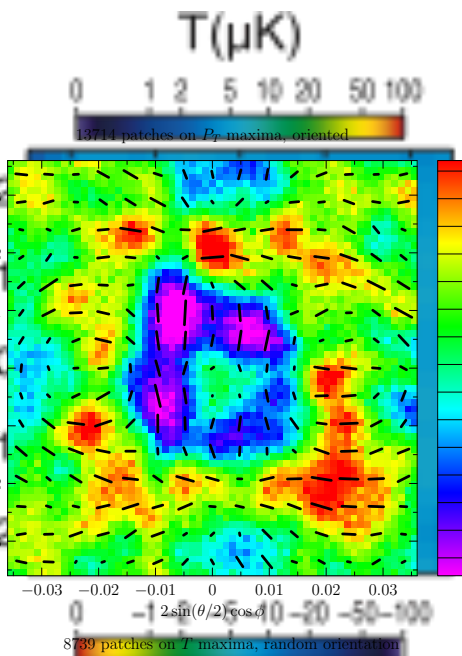
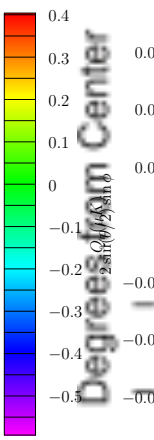
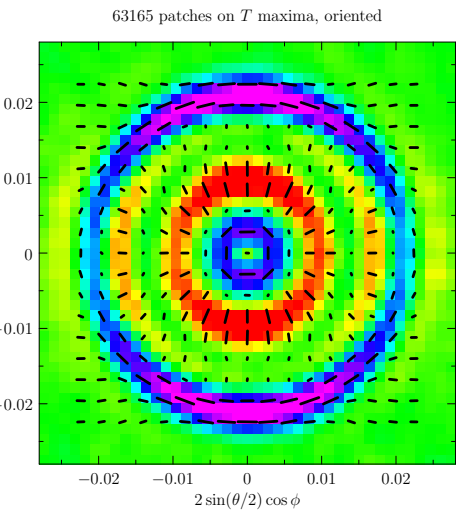
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CMB Peak Statistics

polarization rotated & stacked on temperature Peaks

CMB Polarization BAO in the CMB – WMAP9



B mode of polarization cf. E mode

linear scalar fluctuations create only E patterns

strain from CMB lensing tides distorts E pattern into a bit of B ^{SPT}

anisotropic strain from gravity waves => E & B

BICEP KECK



photons under strain

BICEP2 collaboration 2014

380 sq deg
 $f_{\text{sky}}=0.009$

512 antenna coupled TES bolometers
150 GHz for 3 seasons
cross-correlate with BICEP1, 100 GHz,
preliminary cross-correlate with KECK

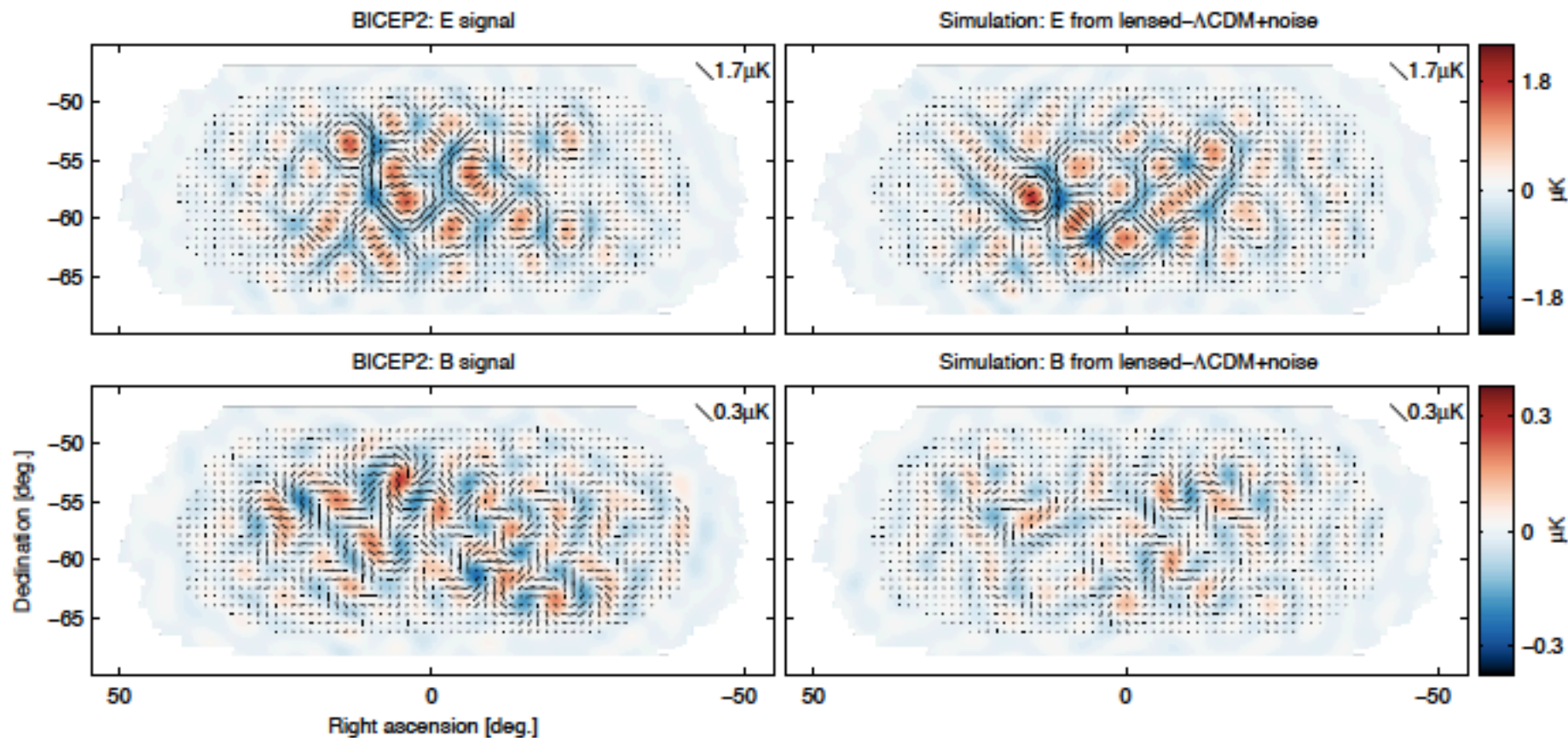
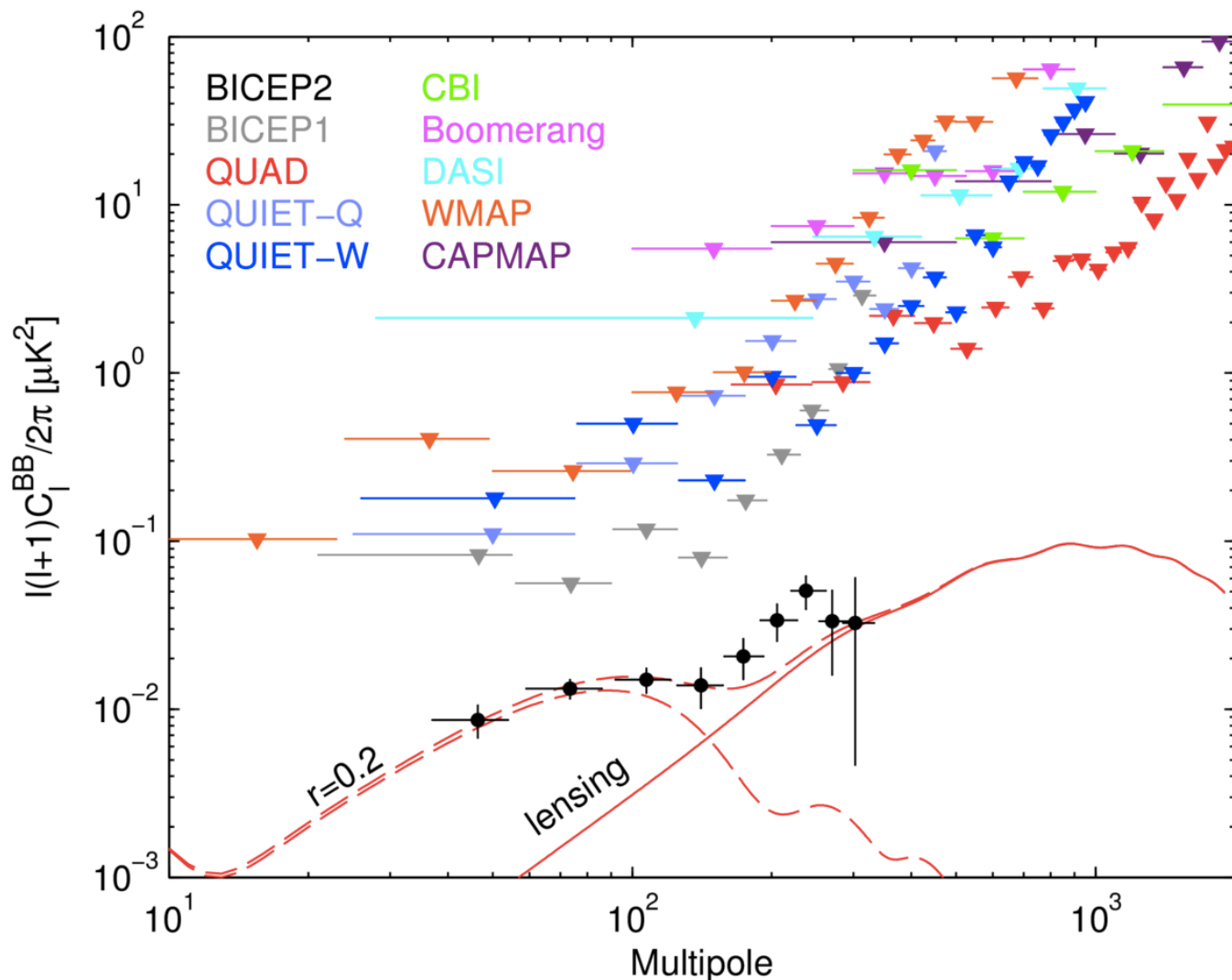


FIG. 3.— *Left:* BICEP2 apodized E -mode and B -mode maps filtered to $50 < \ell < 120$. *Right:* The equivalent maps for the first of the lensed- Λ CDM+noise simulations. The color scale displays the E -mode scalar and B -mode pseudoscalar patterns while the lines display the equivalent magnitude and orientation of linear polarization. Note that excess B -mode is detected over lensing+noise with high signal-to-noise ratio in the map ($s/n > 2$ per map mode at $\ell \approx 70$). (Also note that the E -mode and B -mode maps use different color/length scales.)

BICEP2 collaboration 2014 non-lensing B mode => $r=0.20 \pm 0.07 \pm 0.05$

cf. P13: r from TT

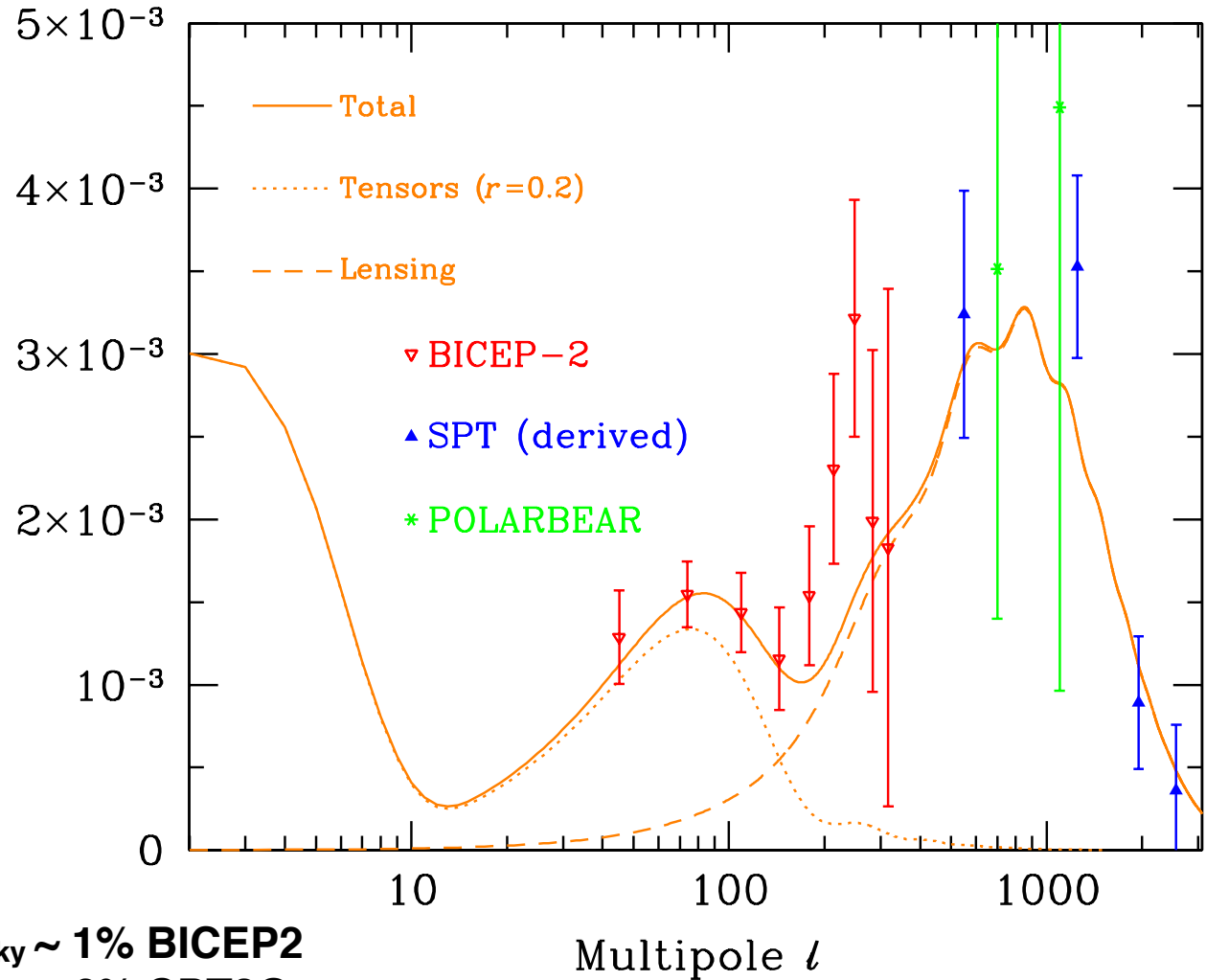
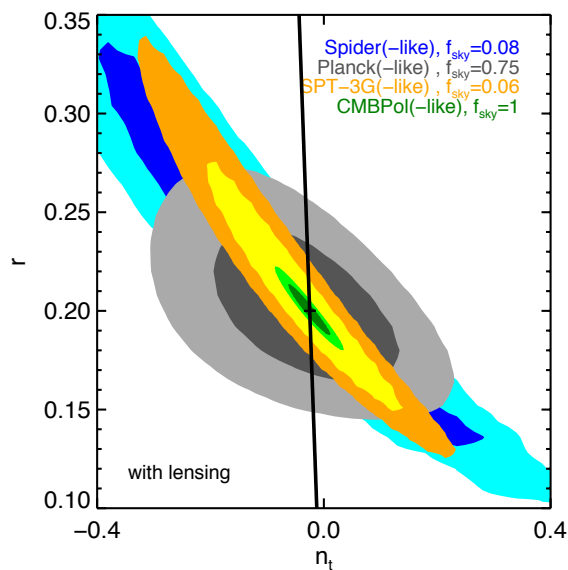
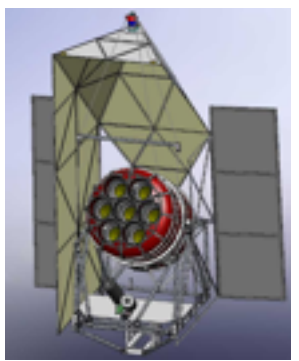
< 0.12 95% CL



$r = \text{GW power} / \text{scalar-curvature power} \approx 0.13 V / (2 \times 10^{16} \text{Gev})^4$
Potential Energy scale is the GUT level!

We are working heavily on Planck polarization, E Nov 2014, B TBD

Spider collaboration, LDB flight Fall 2014 +/-0.2 supposed to fly Fall 13, but US sequester stopped it



**f_{sky} ~ 1% BICEP2
 = 6% SPT3G
 = 8% SPIDER
 ~ 70% + PLANCK
 ~ 50% AdvACT**

**Spider24days+Planck2.5yr:
 r-n_t matrix-forecast
 for r=0.2 input
 (2σ_r ~0.04 including fgnds)**

similar r-forecasts for ABS+, Keck, AdvACT,..

ultra-early Universe sound spectrum $\ln P_s(\ln k)$

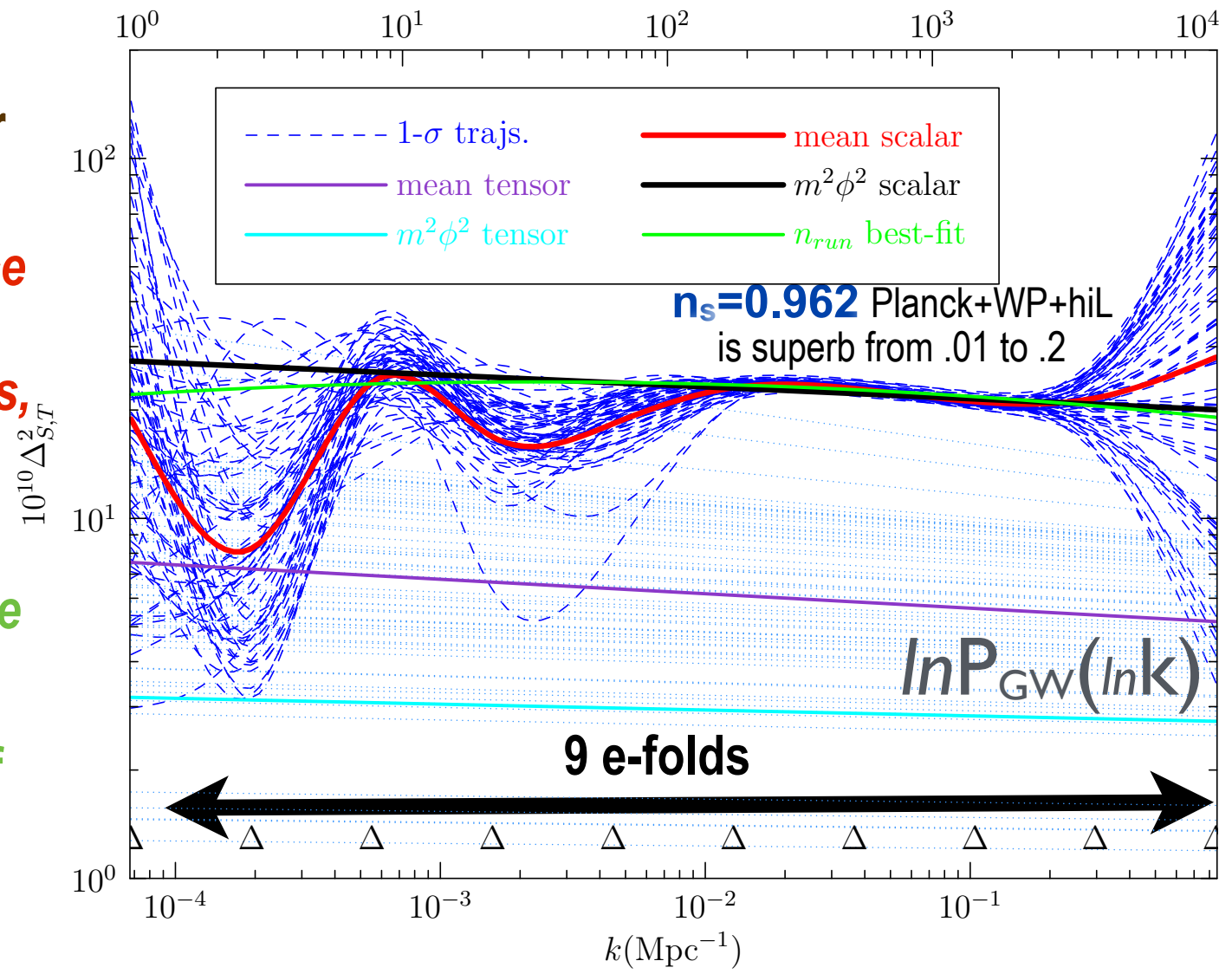
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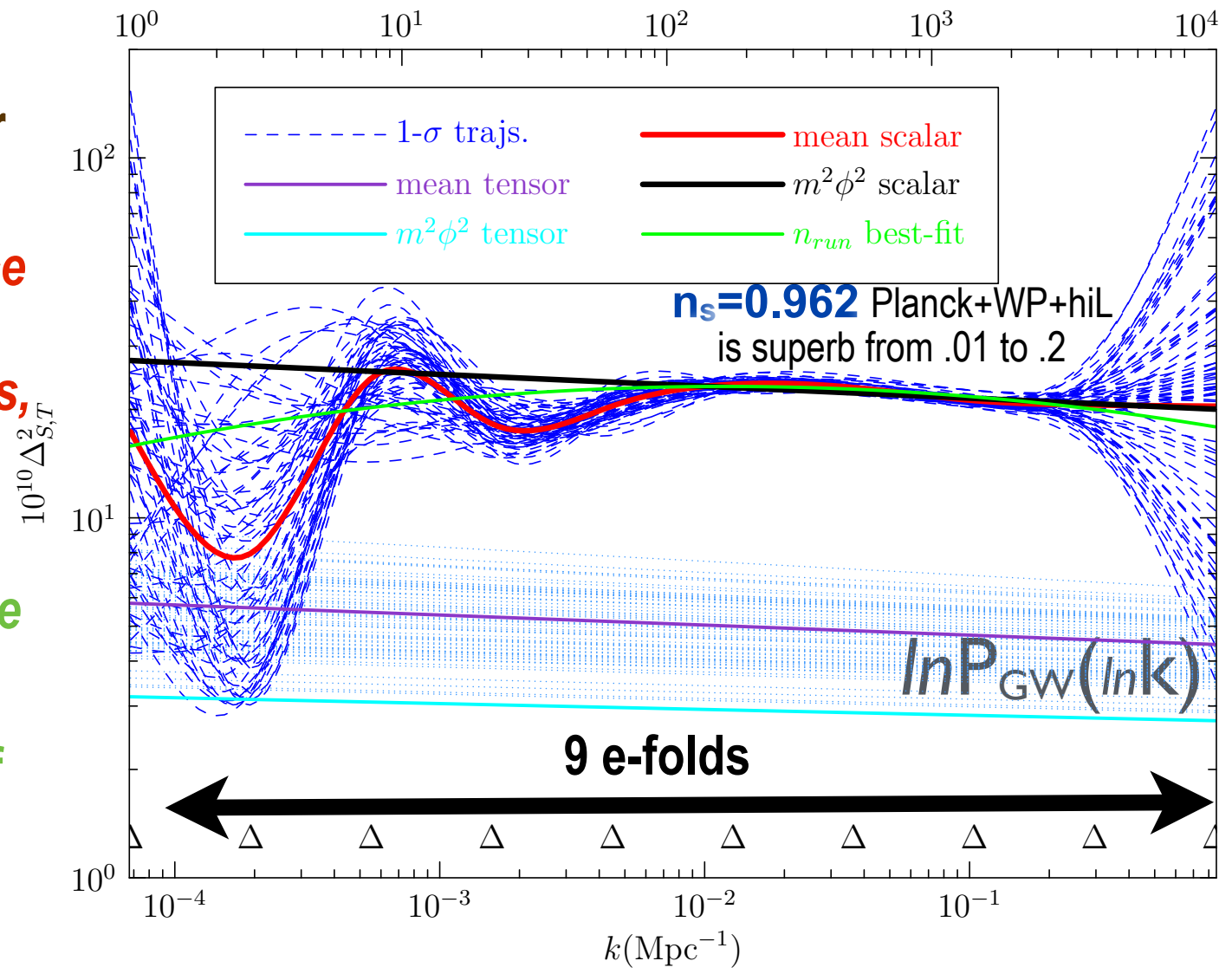
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**but hints of
change on large
 $L < 100$ scales**

**PS: running of
 P_s is a bad fit**



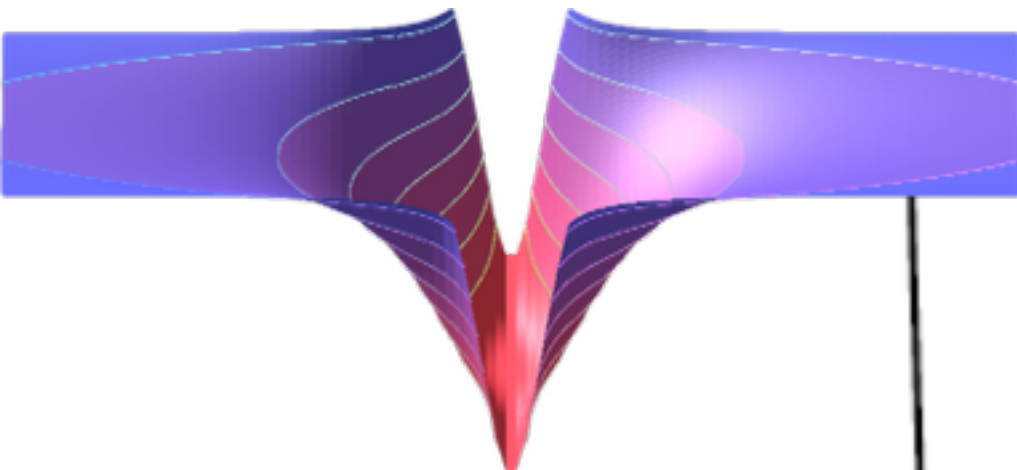
Bond, Braden, Huang, Frolov, Vaudrevange 2014

scan $\ln P_s(\ln k)/A_s$, $\ln A_s = \ln P_s(k_{pivot,s})$, $r(k_{pivot,t})$; consistency \Rightarrow reconstruct $\epsilon(\ln H a)$, $V(\psi)$

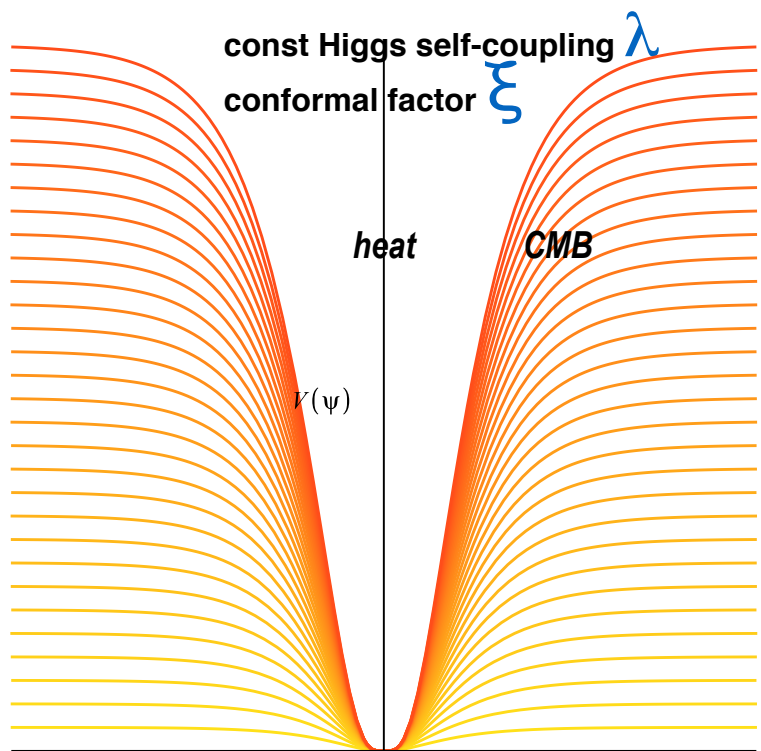
what is the inflaton's potential?

how was *matter & entropy* generated at the end of acceleration = inflation?

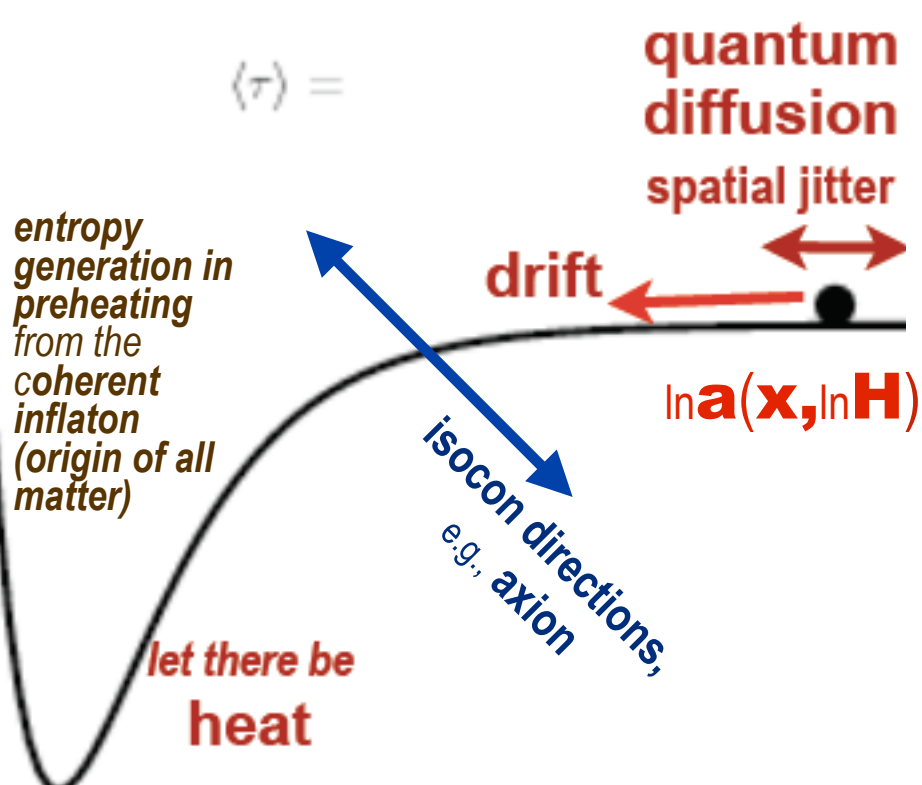
Relate it to the Higgs & standard model?



detecting $r \sim 0.2 \Rightarrow$
shape cannot be too flat



conformal potential-flattening SBB89 ψ

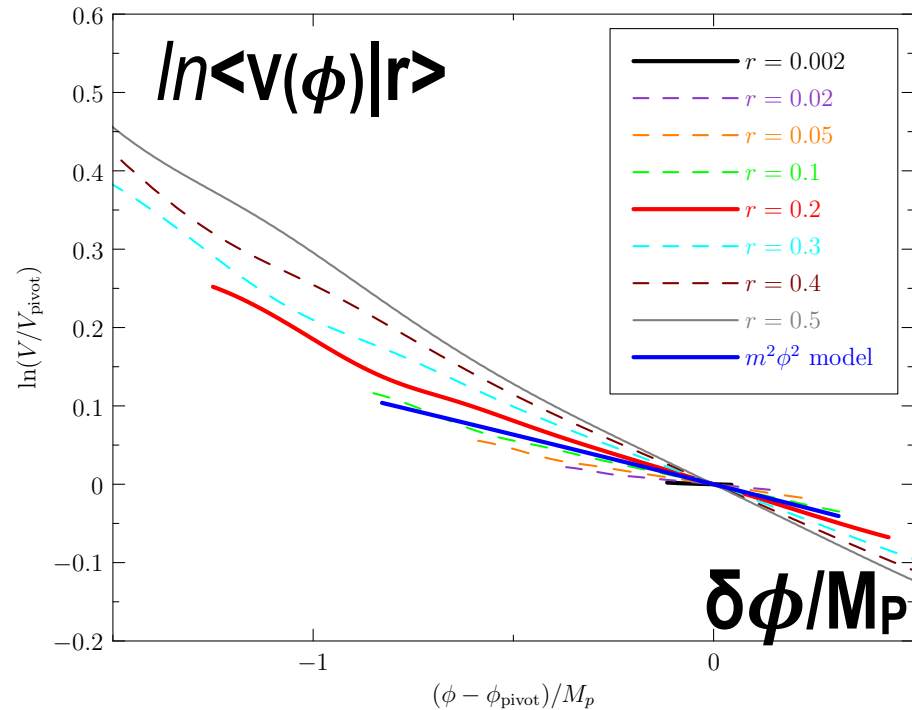
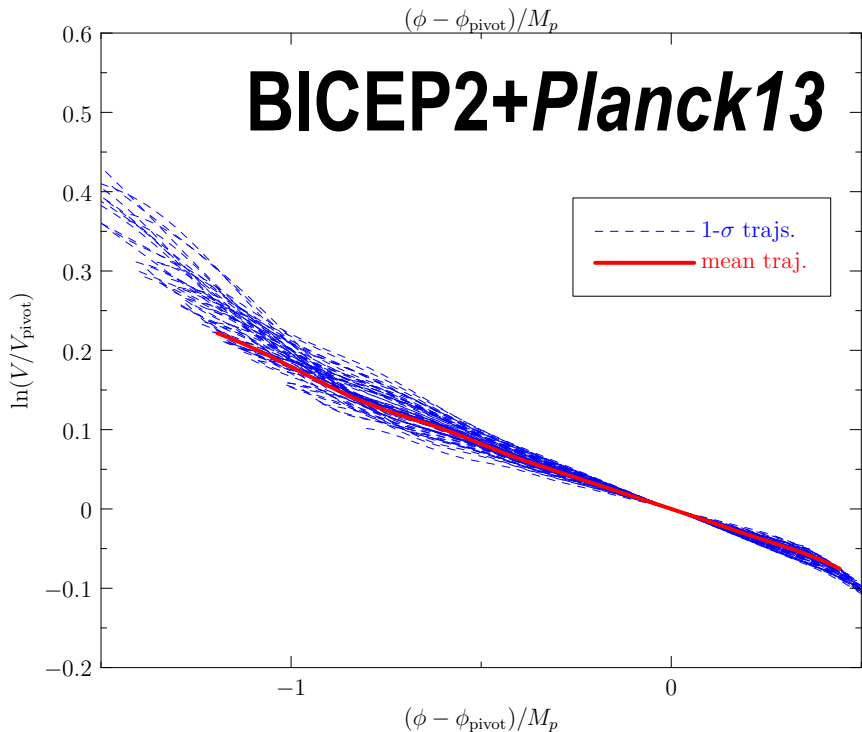
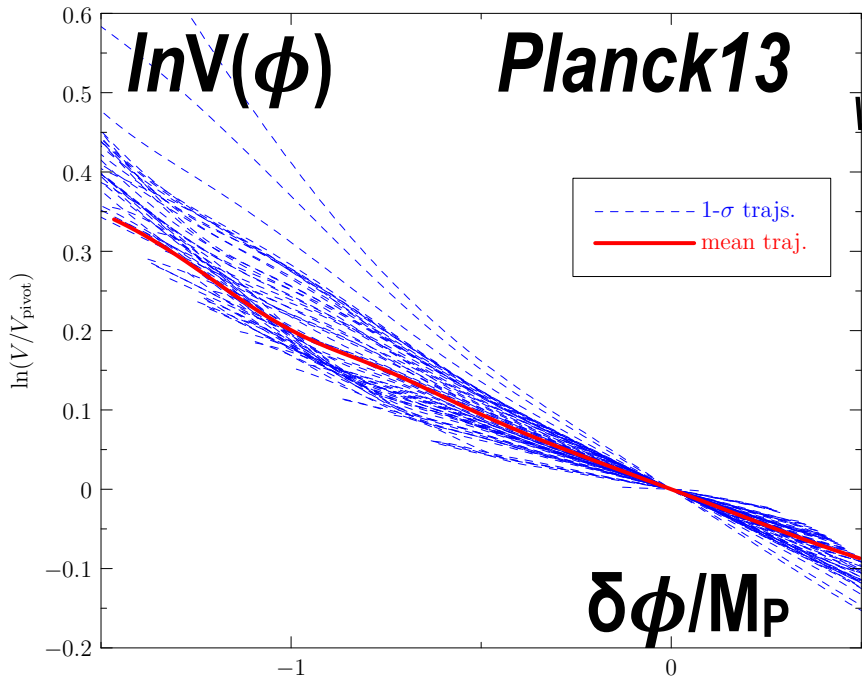


what is the inflaton's $V(\phi)$?

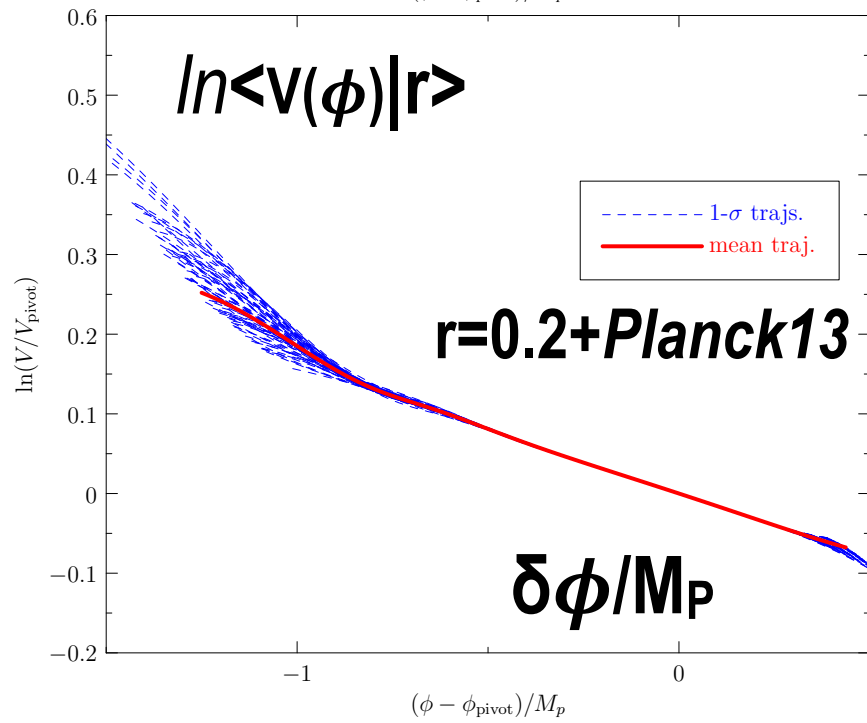
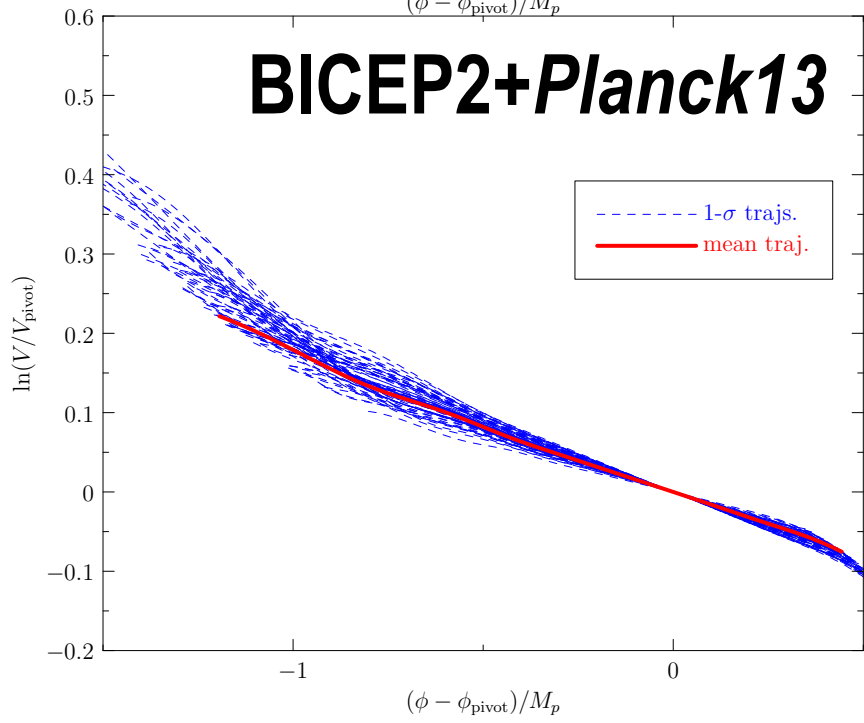
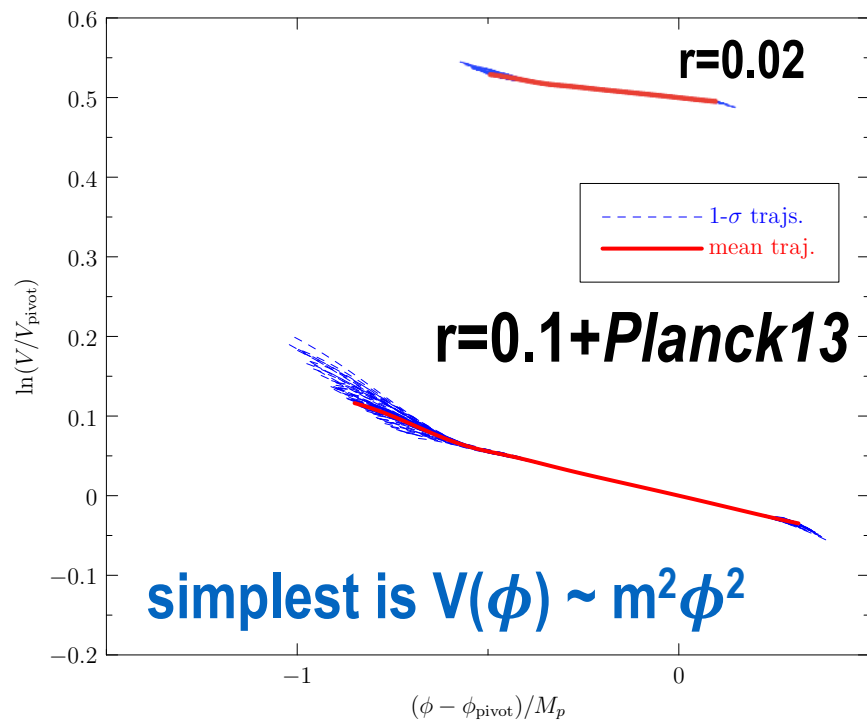
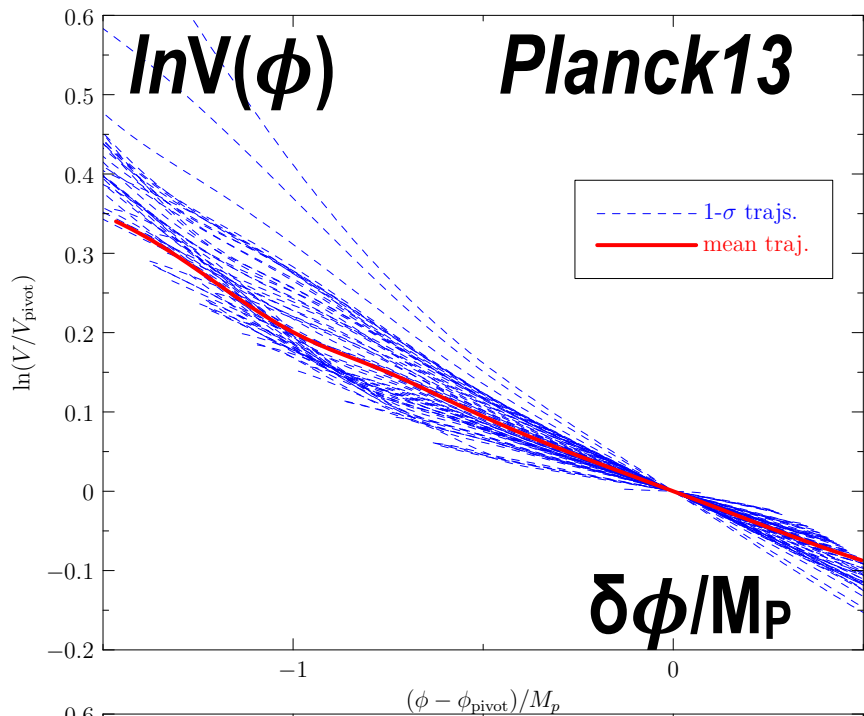
we reconstruct the scalar curvature power (isotropic strain) & the early universe acceleration histories as well

detecting $r \sim 0.2 \Rightarrow$
 $V(\phi)$ shape cannot be too flat over the observable range

Reconstructed mean potential (without BICEP constraint)



simplest is $V(\phi) \sim m^2\phi^2$

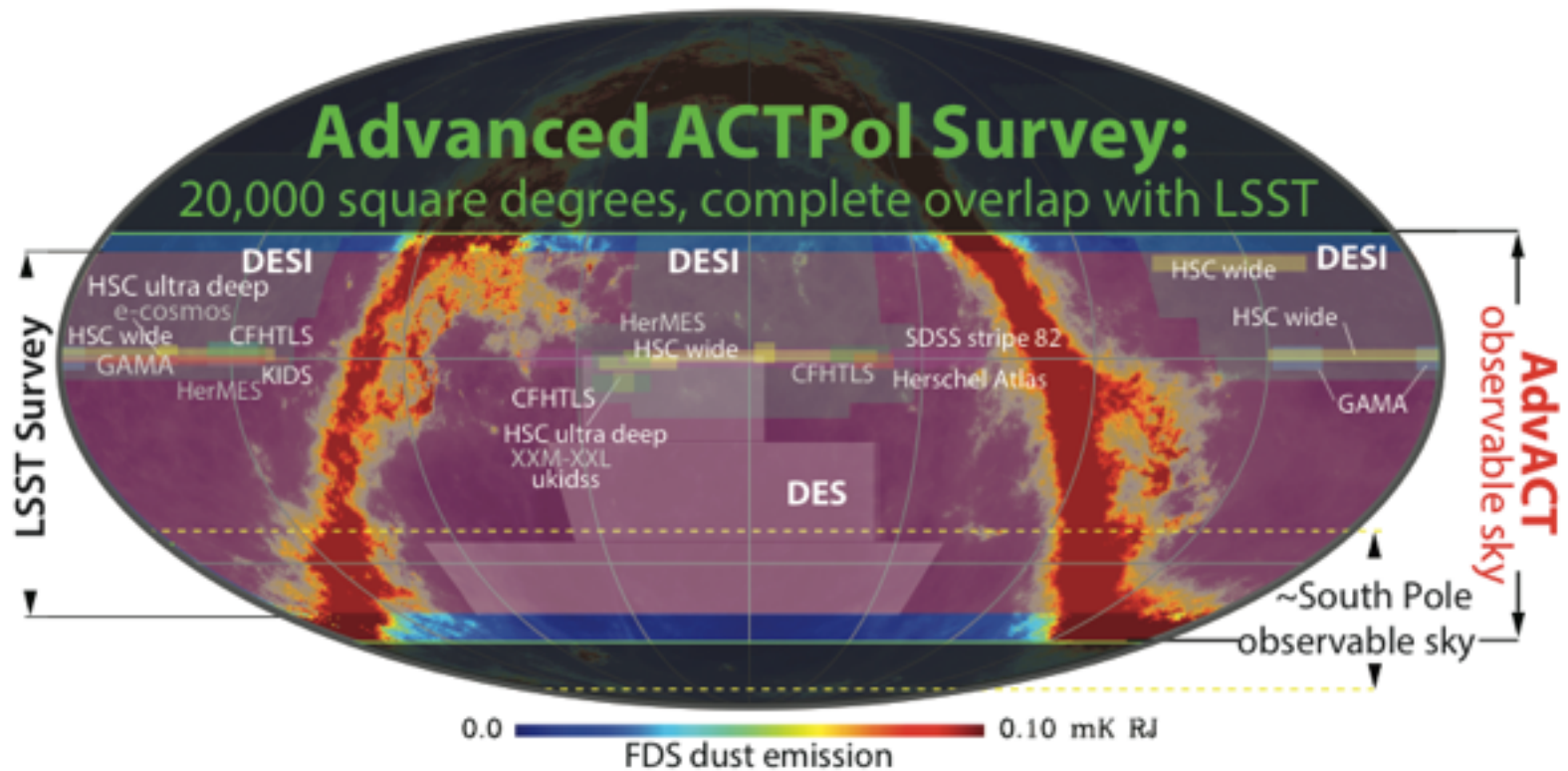


The ACT Collaboration

ACT, now ACTpol, => Advanced ACTpol



Advanced ACTPol (AdvACT) Observations

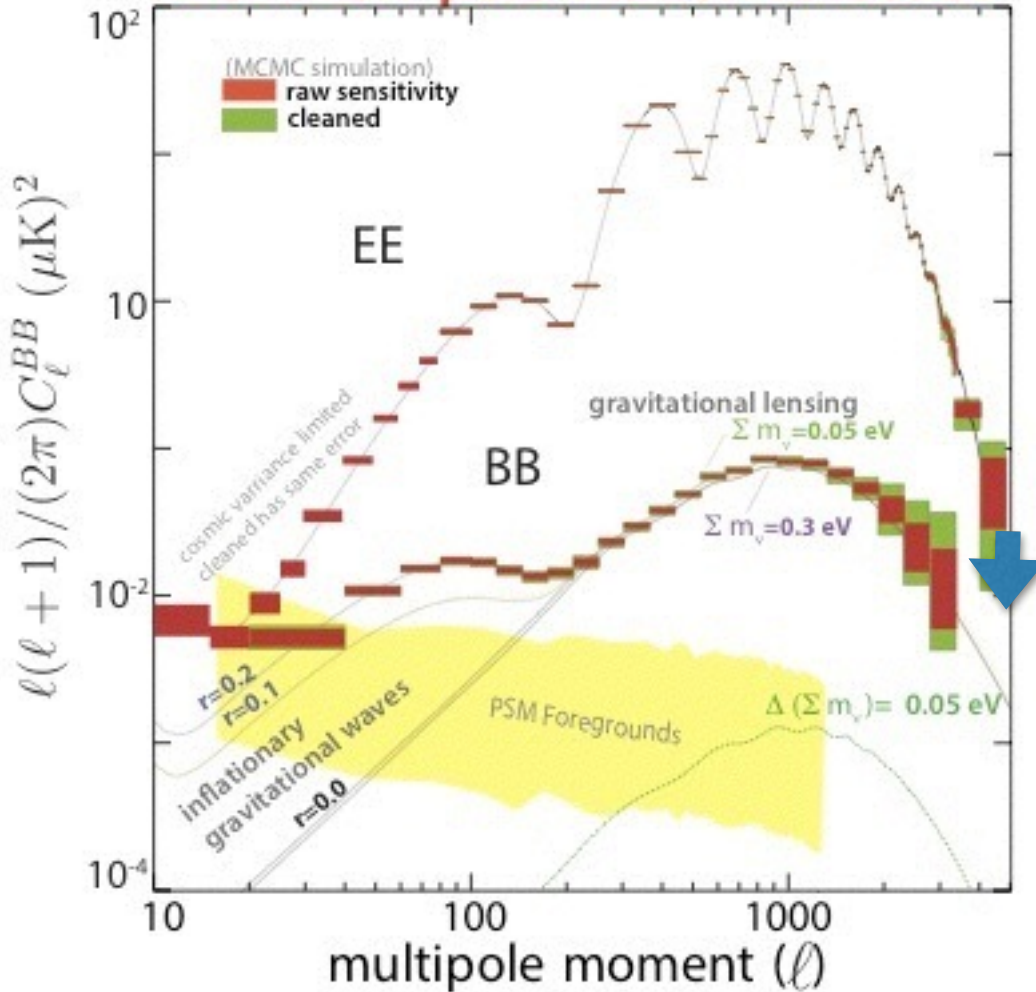


- $\sim 20,000 \text{ deg}^2$ survey ($f_{\text{sky}} \sim 0.5$) with complete LSST overlap as well as DES, ALMA, and other observatories located in Chile
- Substantial overlap with spectroscopic surveys (SDSS, PFS, DESI)



AdvACT: Power Spectra

AdvACT polarization forecast



High S/N B-mode detections for $r > 0.01$ are measured in independent frequency bands (90 & 150 GHz) and on many patches across the sky. This provides important cross-checks on any detected signal

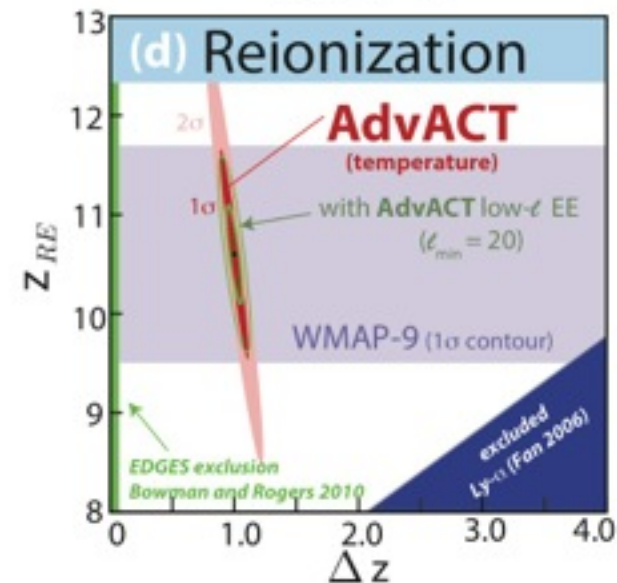
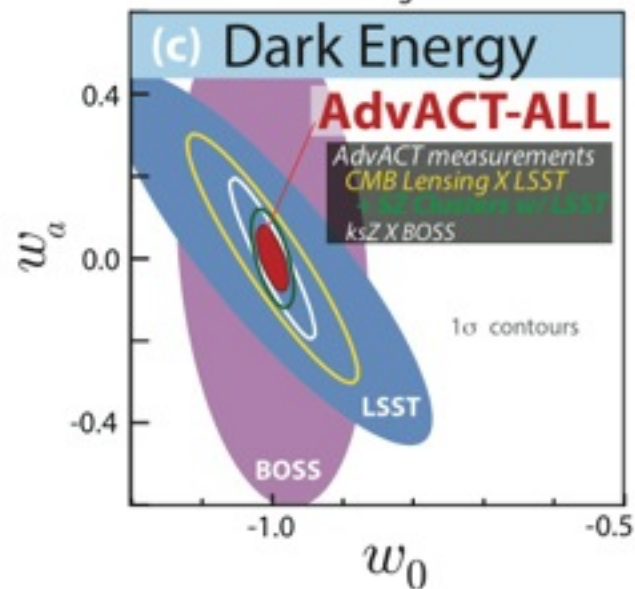
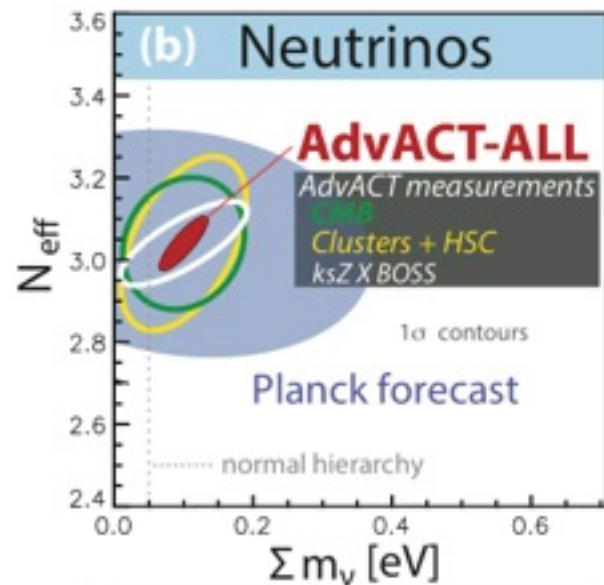
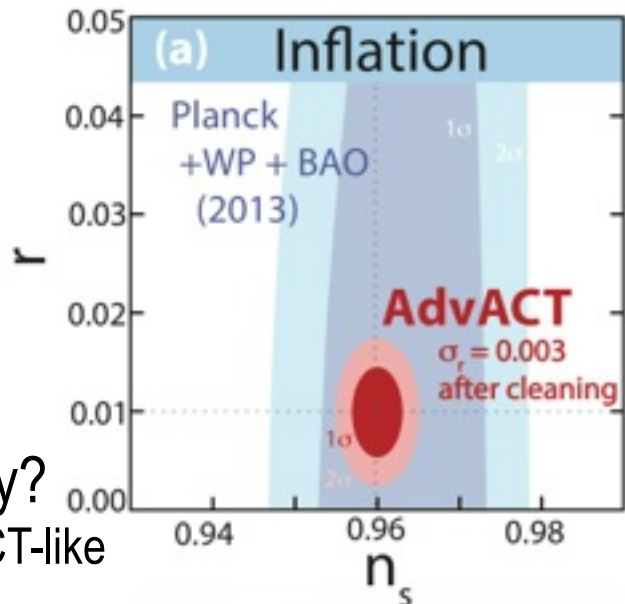
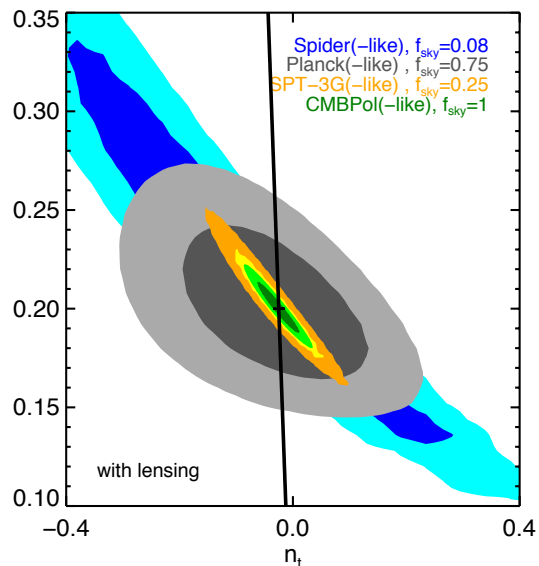
Also shown:

- Error bars before and after foreground cleaning
- Varying amplitudes of the gravitational lensing signal for different values of the sum of the neutrino masses
- Planck forecasts

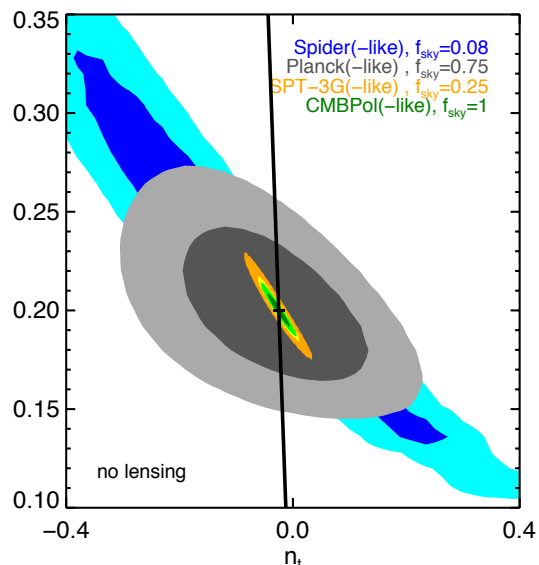
Error bars above shown for $r = 0.2$



AdvACT: Cosmological Forecasts & Planck2.5, Spider, future SPT3g, CMBpol



testing tensor consistency?
 better $f_{\text{sky}}=25\%$ for spt3g/AdvACT-like
 than current 6% goal for spt3g



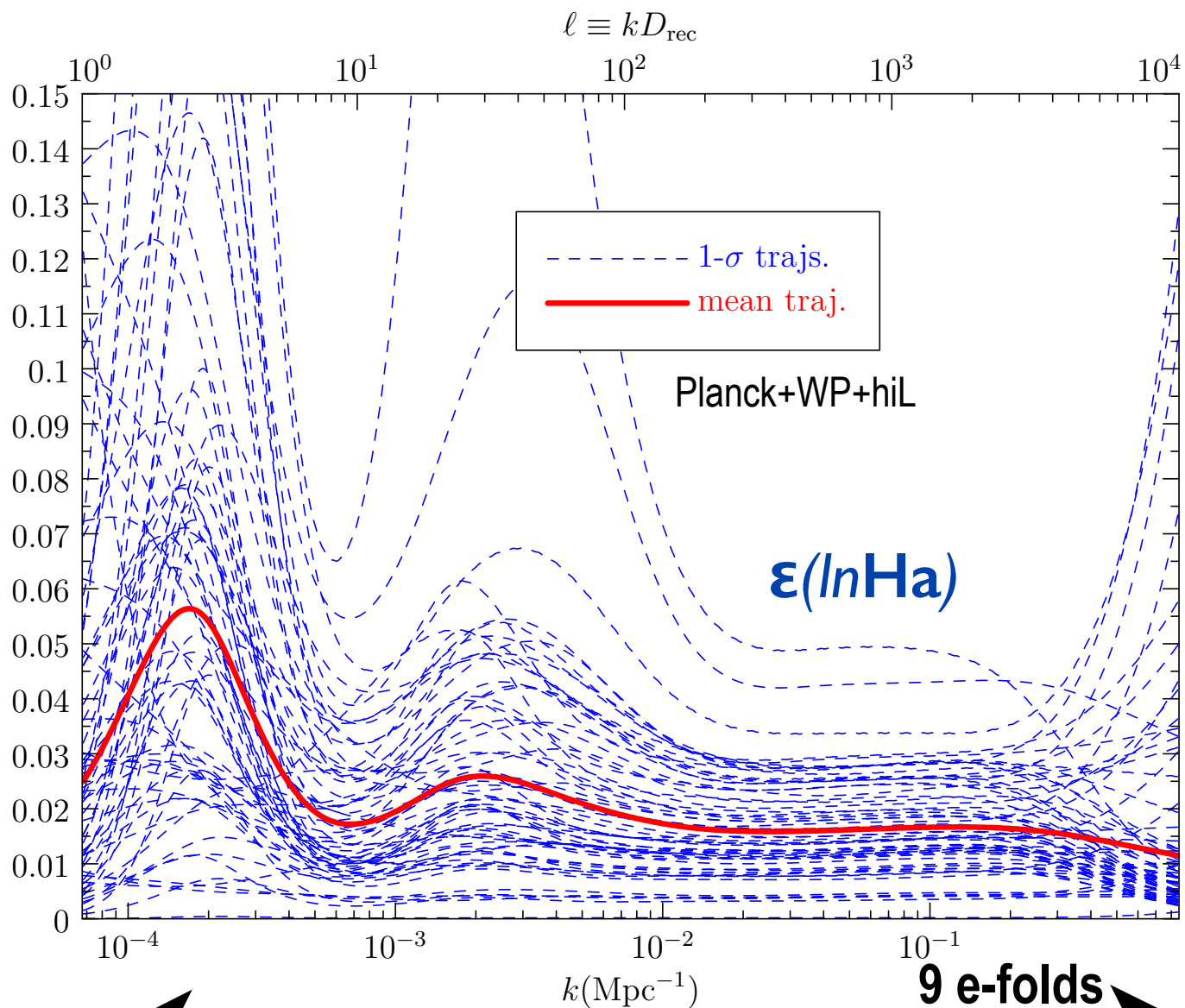
ultra-early Universe sound spectrum $\epsilon = -d \ln H / d \ln a$

new parameters:
trajectory probabilities for early-inflatons

no strong evidence for oscillation patterns, cutoffs, local features

but hints of change on large $L < 100$ scales

PS: running of P_s is a bad fit



9 e-folds

Bond, Braden, Huang, Frolov, Vaudrevange 2014

scan $\ln P_s(\ln k) / A_s$, $\ln A_s = \ln P_s(k_{pivot,s})$, $r(k_{pivot,t})$; consistency \Rightarrow reconstruct $\epsilon(\ln H a)$, $V(\psi)$

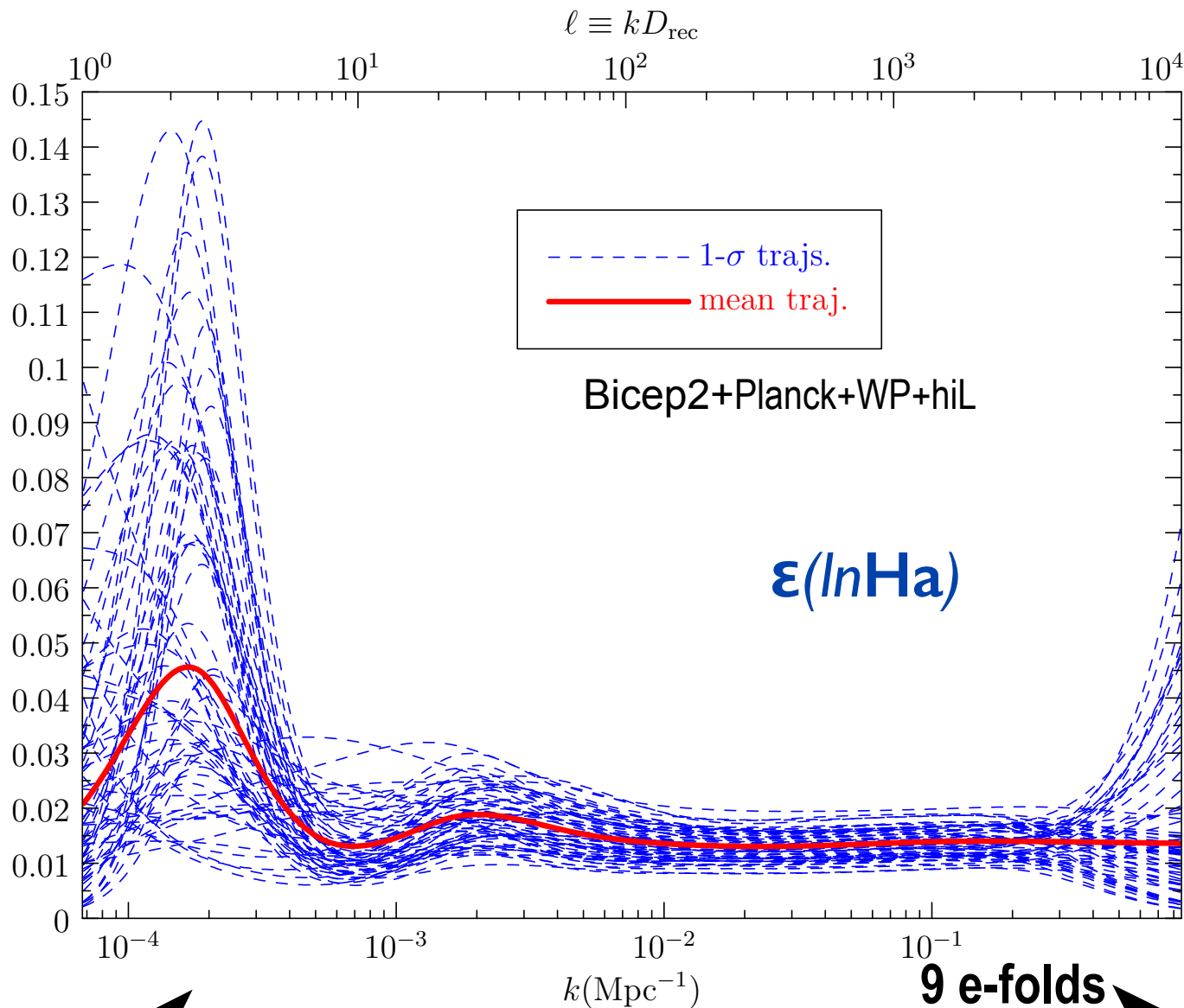
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Bond, Braden, Huang, Frolov, Vaudrevange 2014

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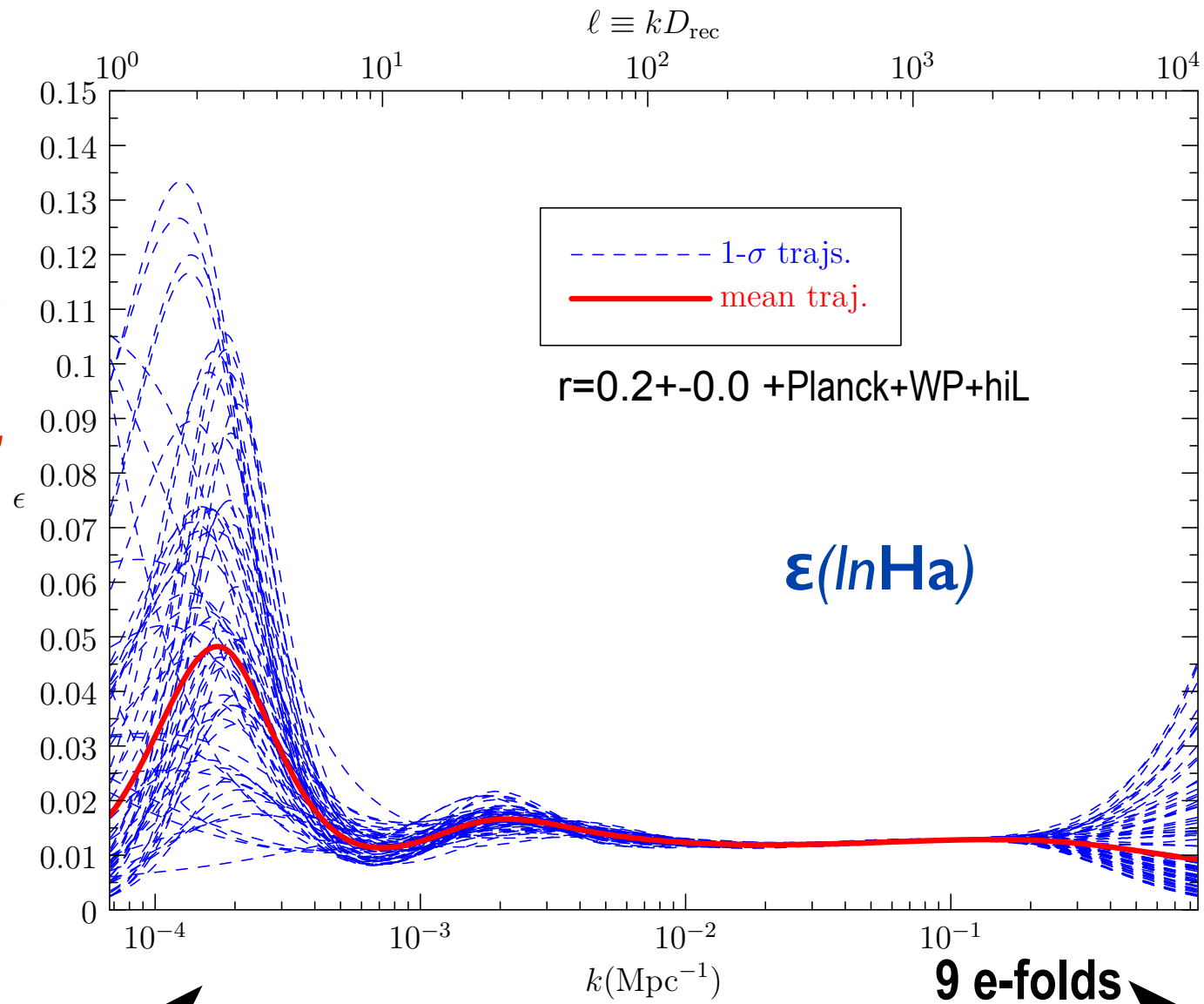
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Bond, Braden, Huang, Frolov, Vaudrevange 2014

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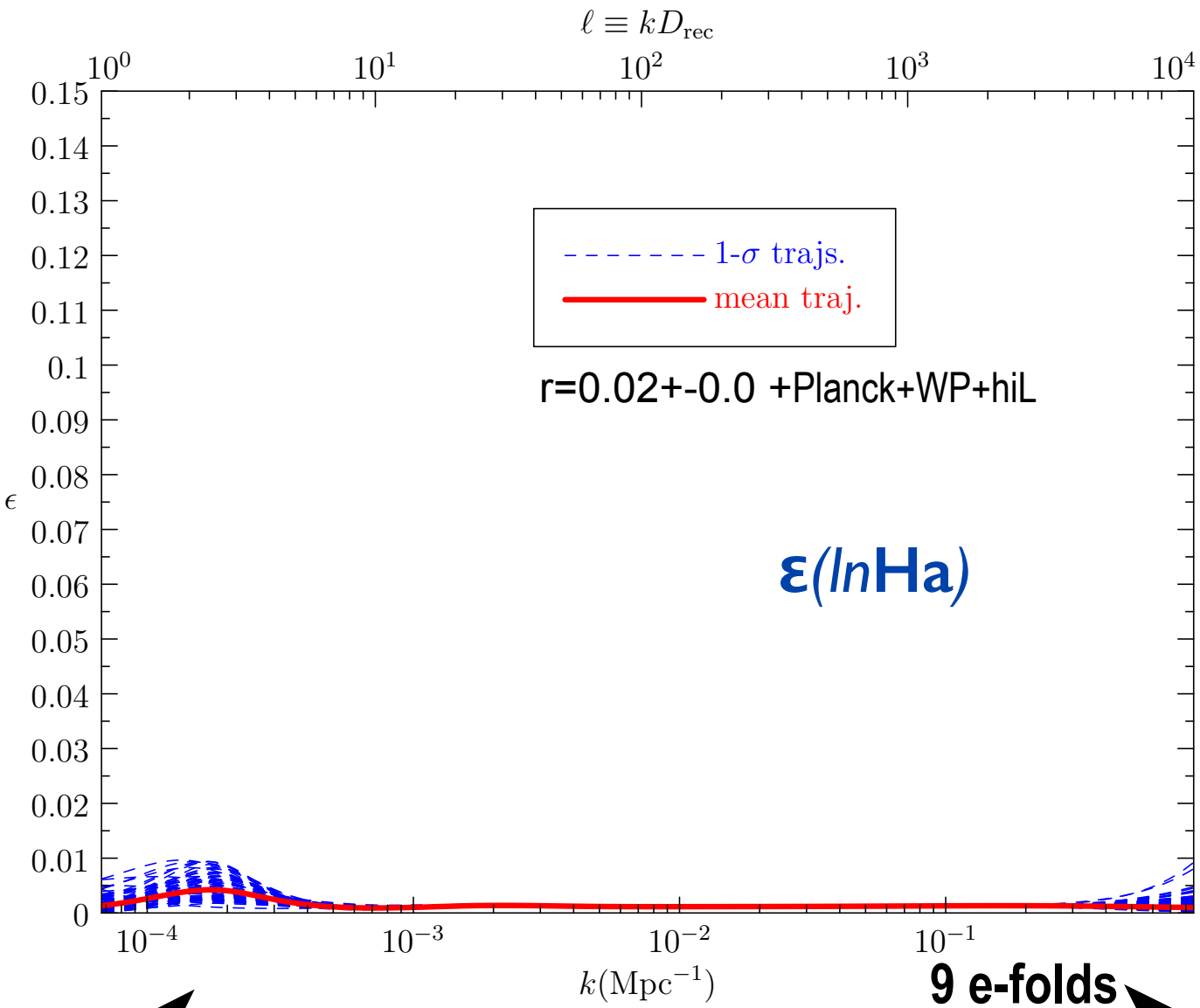
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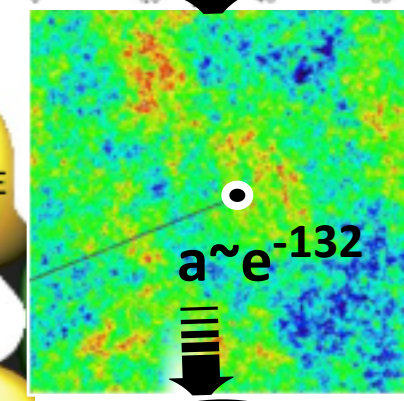
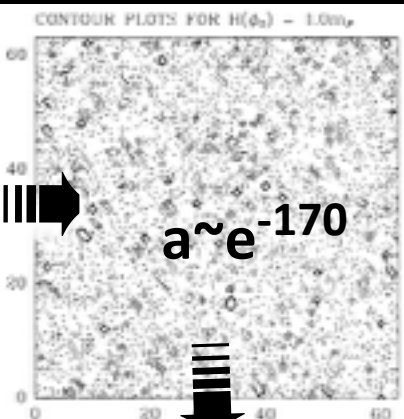
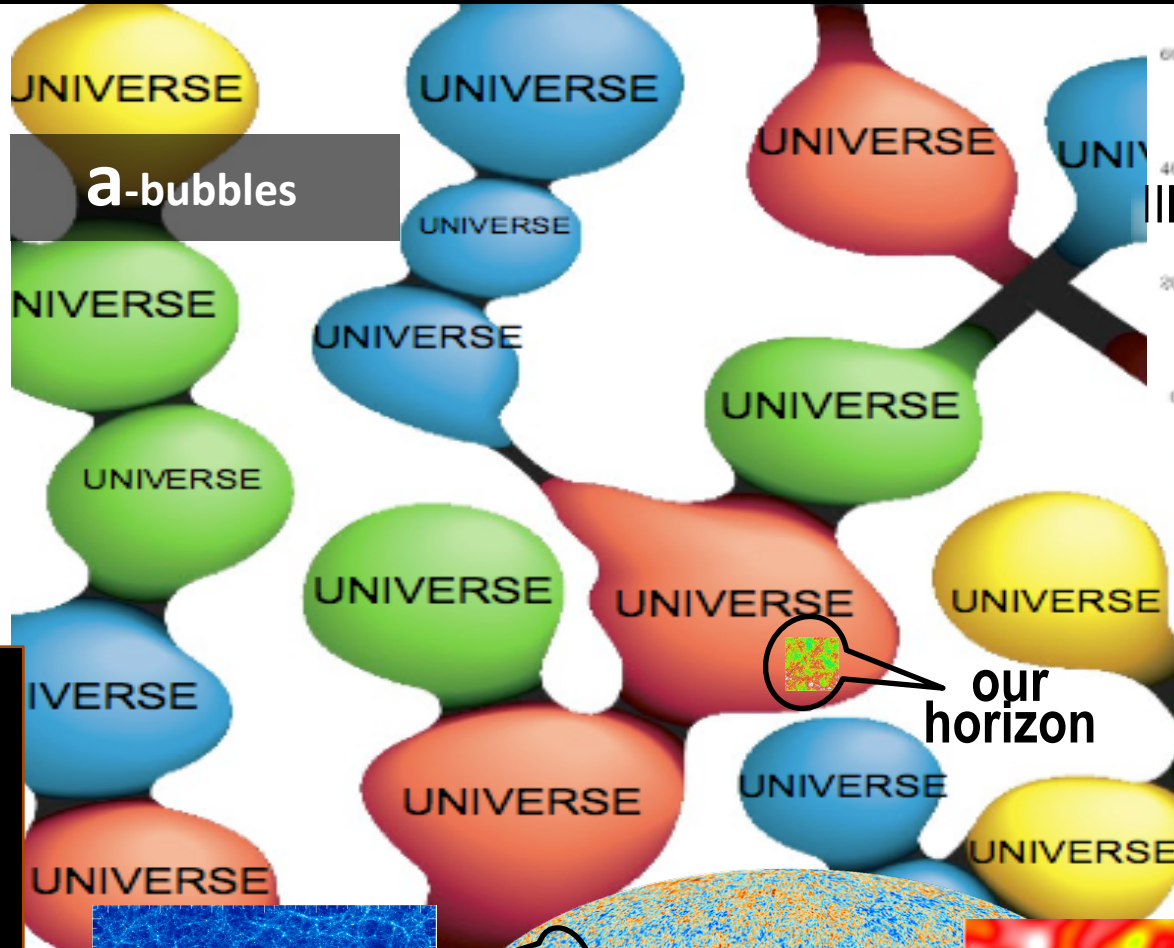
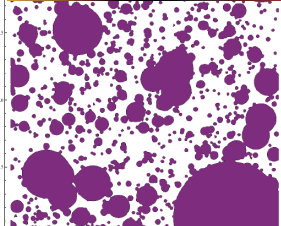
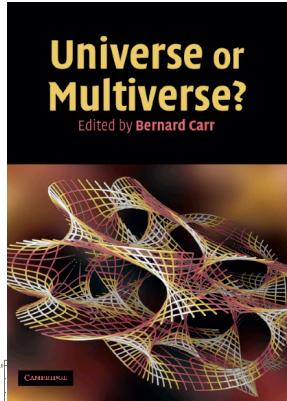
A dark field of galaxies with a central bright star. The image shows a vast field of galaxies, including spiral, elliptical, and irregular shapes, scattered across a black background. A prominent bright star with a four-pointed diffraction pattern is located in the center-left. The text is overlaid on the top half of the image.

from **COMPLEXITY**
to **SIMPLICITY**
to **COMPLEXITY**
to **SIMPLICITY**:
the Universe at Large

beyond our Horizon

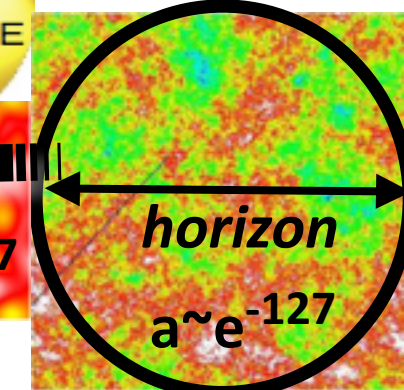
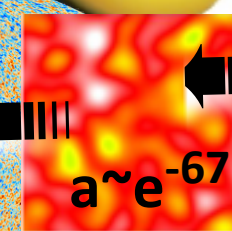
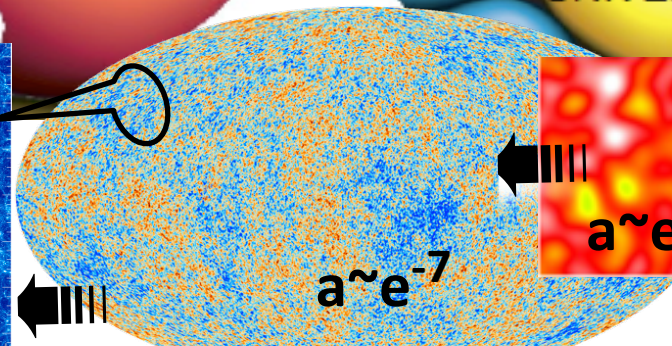
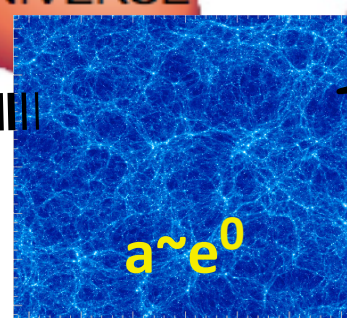
ultra-Ultra Large Scale Structure of the Universe

Horizons: the ultimate-speed constraint on light & information



END
a future DE-Void

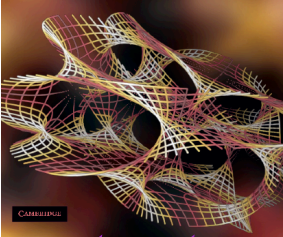
$a \sim e^{+++}$



ultra-Ultra Large Scale Structure of the Universe

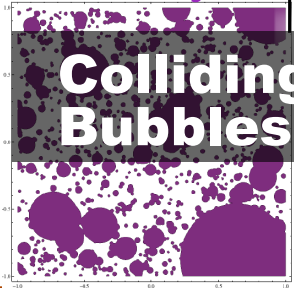
Horizons: the ultimate-speed constraint on light & information

Universe or Multiverse?
Edited by Bernard Carr



quantum tunnels = bubbly-U

Colliding Bubbles



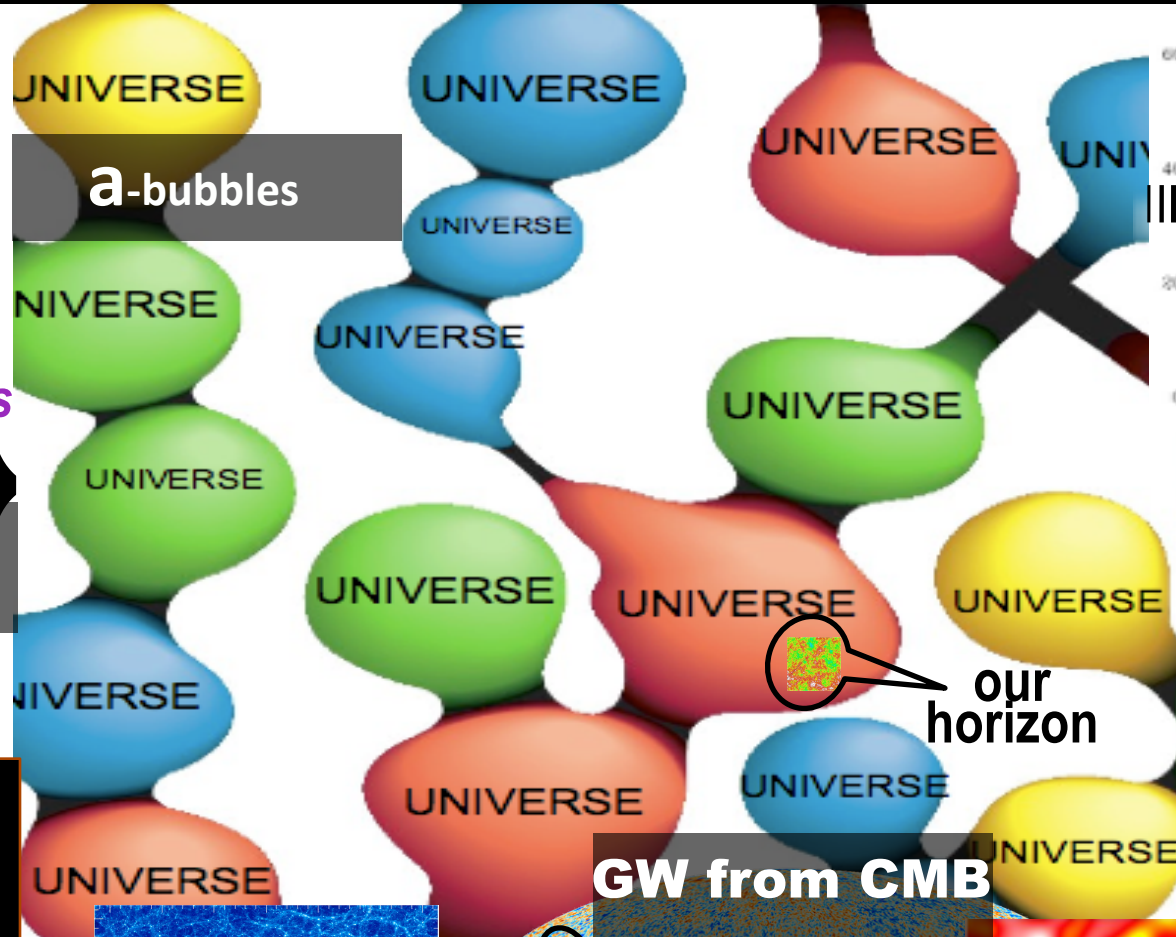
END

a future DE-Void



Dark Energy Trajectories

$a \sim e^{+++}$



Gastrophysical Simulations of the Cosmic Web

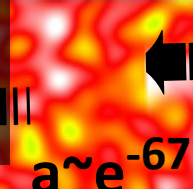
$a \sim e^0$

GW from CMB

recombination history

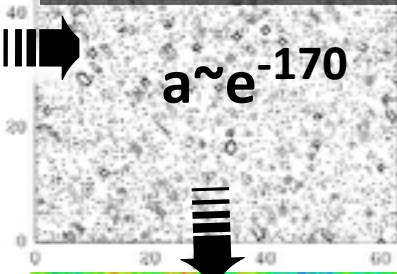
$a \sim e^{-7}$

Post-inflation Matter & Entropy



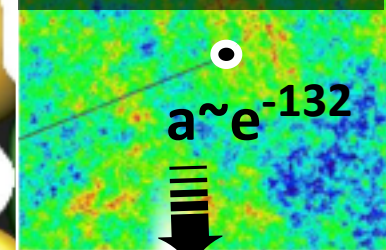
$a \sim e^{-67}$

Stochastic Inflation

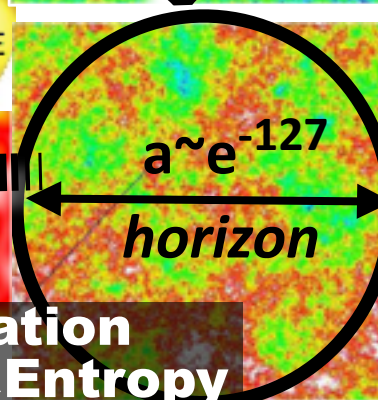


$a \sim e^{-170}$

Isocurvature Simulations



$a \sim e^{-132}$



$a \sim e^{-127}$

horizon

END

Let there be.....

Early **Dark Energy** from $e^{-170?}$ to e^{-67}

2+1 numbers quantum **noise** e^{-127} to e^{-67} in phonons (inflaton) & GW

Heat: matter & **radiation** $a \sim e^{-67}$

Dark Matter, light nuclei $a \sim e^{-21}$ to e^{-35}

Cosmic **Light**: 1st light released, 1st atoms $a \sim e^{-7}$

1st stars $a \sim e^{-3}$, 1st heavy nuclei (O, C, Fe,..)

Galaxies $> e^{-2.2}$

Earth $a \sim e^{-0.34}$

1st human writing $a \sim e^{-0.0000004}$

Late **Dark Energy** to e^{+++}

Let there be.....

Early **Dark Energy** from $e^{-170?}$ to e^{-67}

semi **ETERNAL** Universe
most of it never Banged

2+1 numbers quantum **noise** e^{-127} to e^{-67} in phonons (inflaton) & GW

Heat: matter & **radiation** $a \sim e^{-67}$

Our little **Big Bang**

Dark Matter, light nuclei $a \sim e^{-21}$ to e^{-35}

Cosmic **Light**: 1st light released, 1st atoms $a \sim e^{-7}$

1st stars $a \sim e^{-3}$, 1st heavy nuclei (O, C, Fe,..)

Galaxies $> e^{-2.2}$

Earth $a \sim e^{-0.34}$

1st human writing $a \sim e^{-0.0000004}$

Late **Dark Energy** to e^{+++}

Will our bit of the Universe re-Bang?
NO... maybe

SIMPLICITY

at $a \sim e^{-7} \sim 1/1100 \Rightarrow$

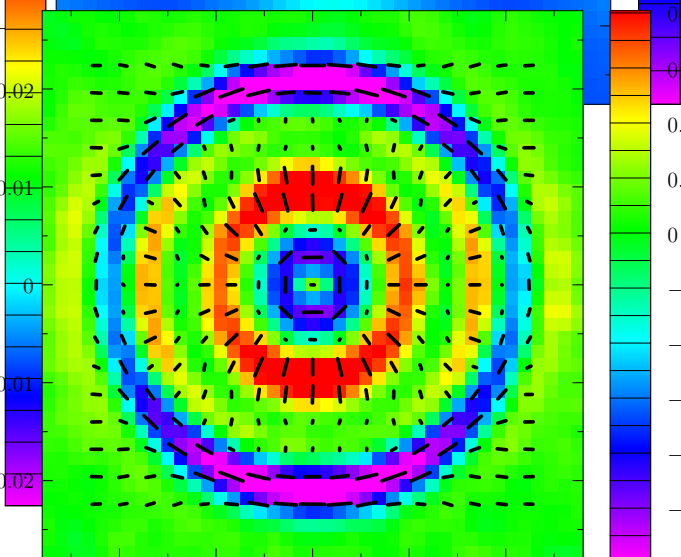
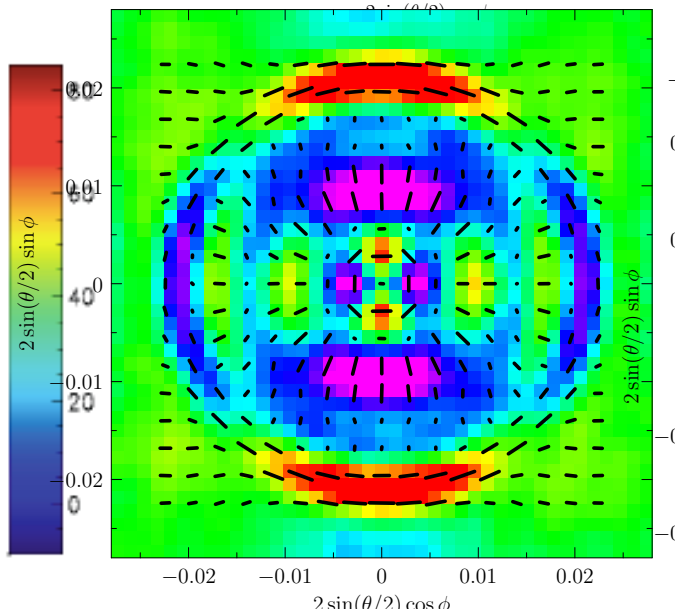
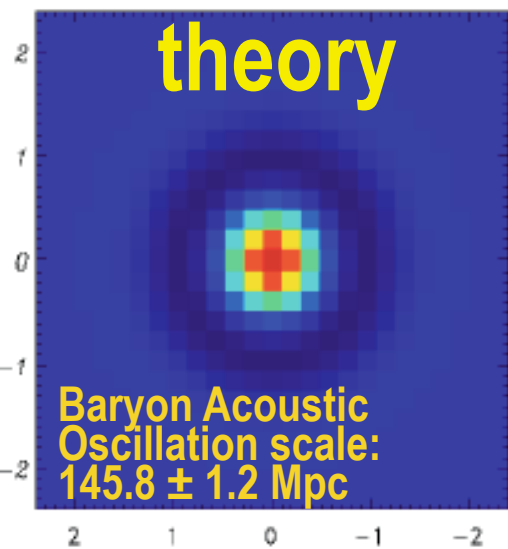
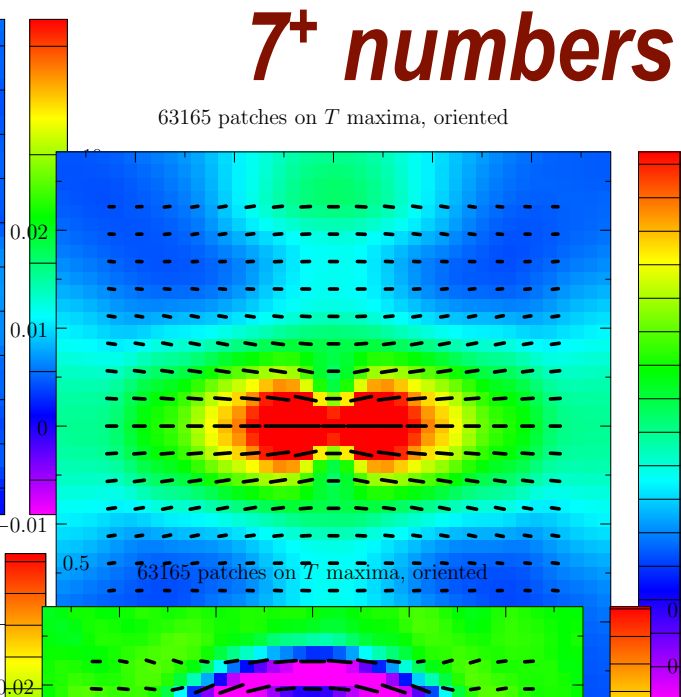
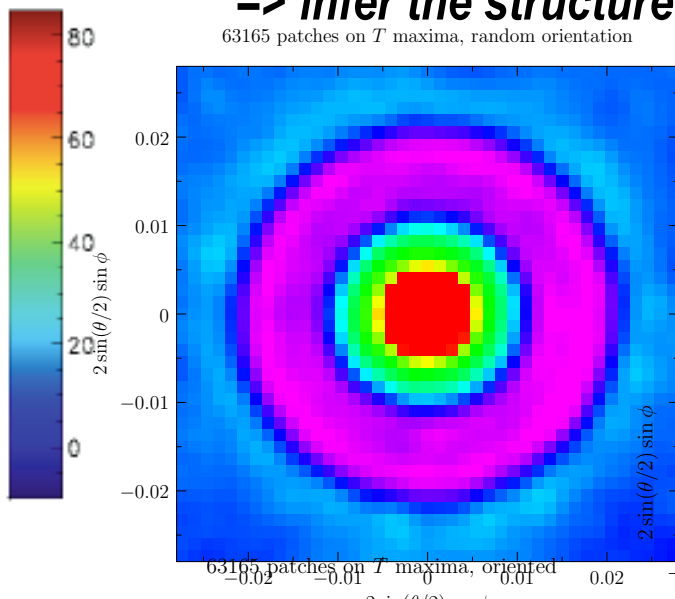
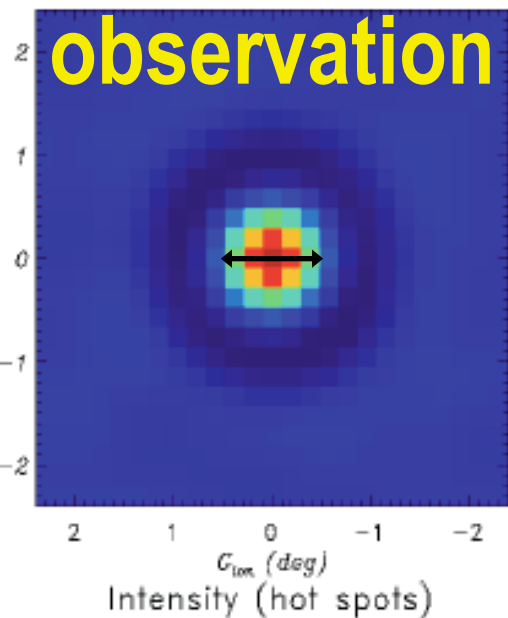
at $a \sim e^{-67+60} \sim 1/10^{30+25}$

reveals *primordial sound waves in matter*

\Rightarrow learn **contents & structure** at 380000 yr, $a \sim e^{-7}$

\Rightarrow infer the structure far far earlier $a \sim e^{-67+60}$

7+ numbers



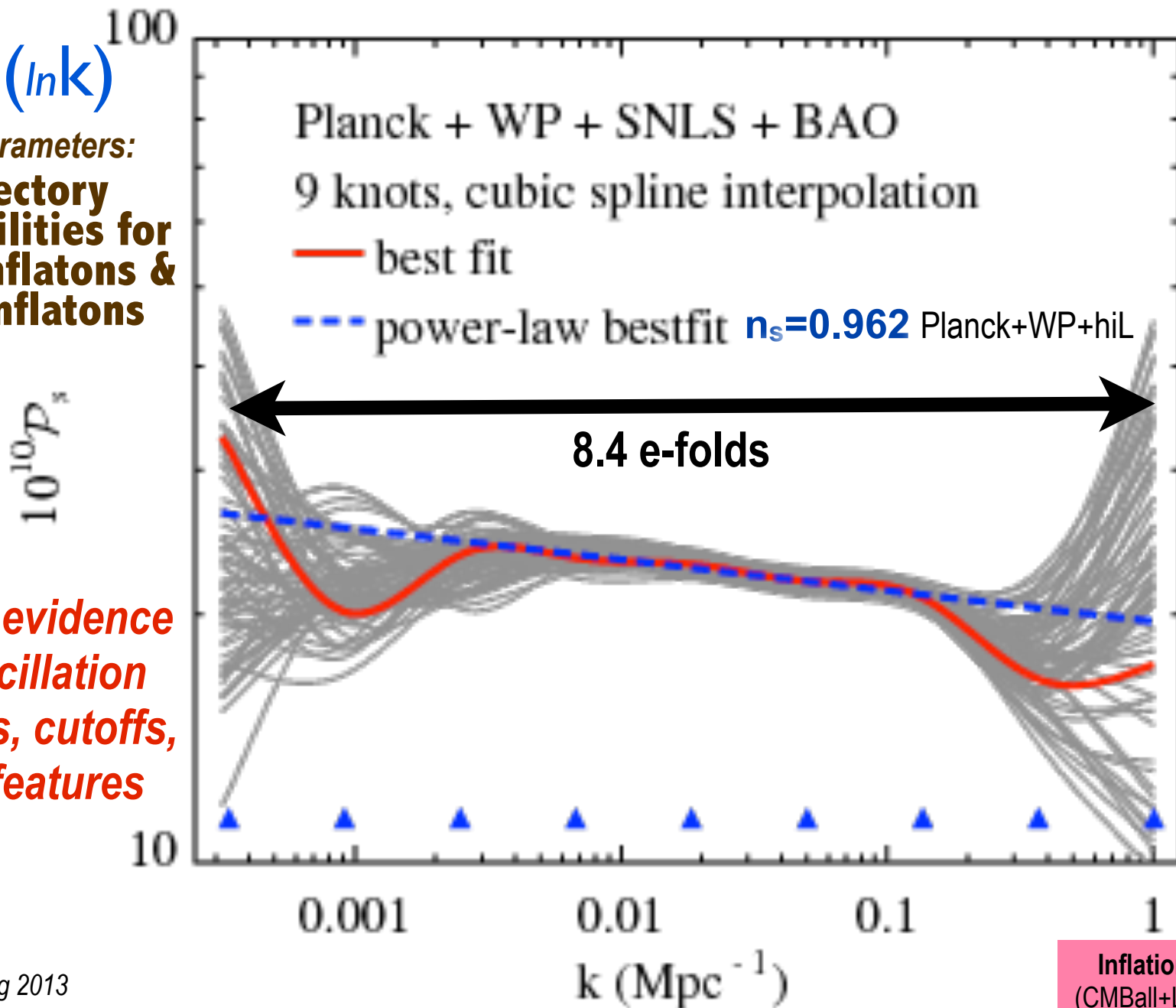
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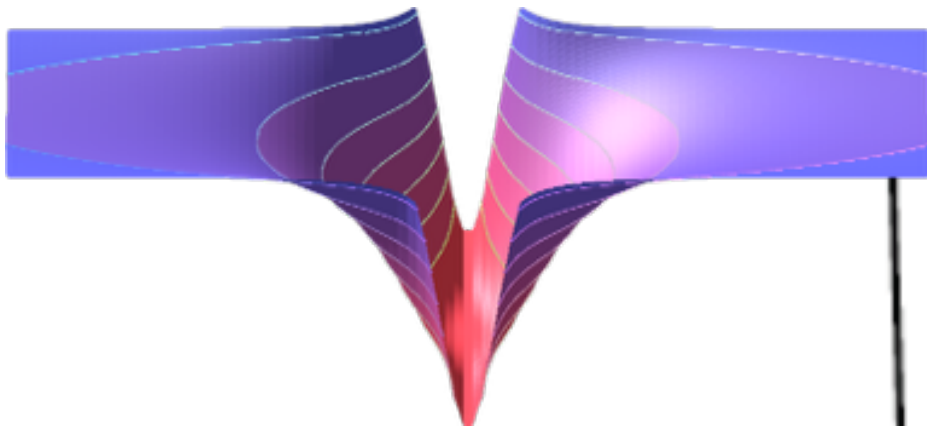


Inflation Histories
(CMBall+LSS+SN+WL)

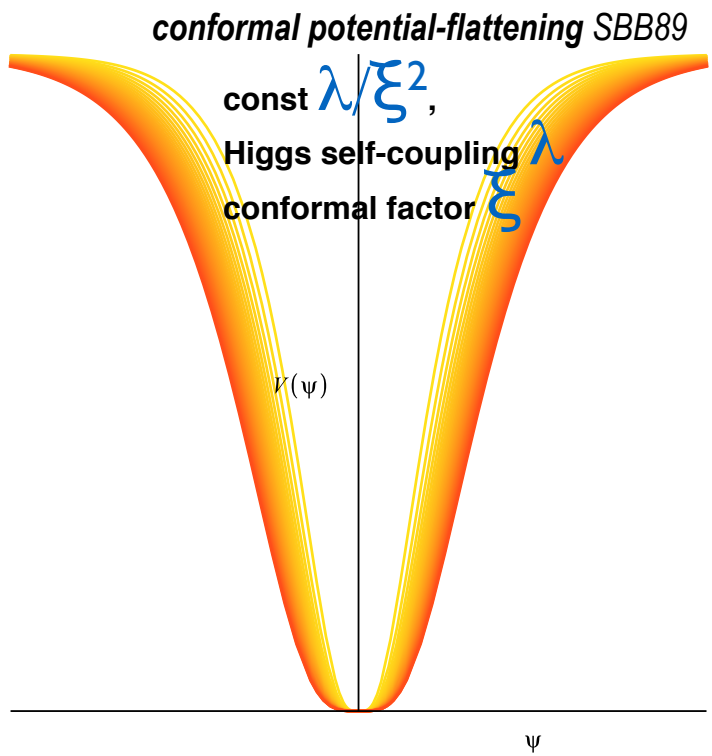
what is the inflaton's potential?

how was *matter & entropy* generated at the end of acceleration = inflation?

Relate it to the Higgs & standard model?



detecting $r \sim 0.2 \Rightarrow$
shape cannot be too flat



entropy generation in preheating from the coherent inflaton (origin of all matter)

let there be heat

$\langle \tau \rangle =$

quantum diffusion spatial jitter

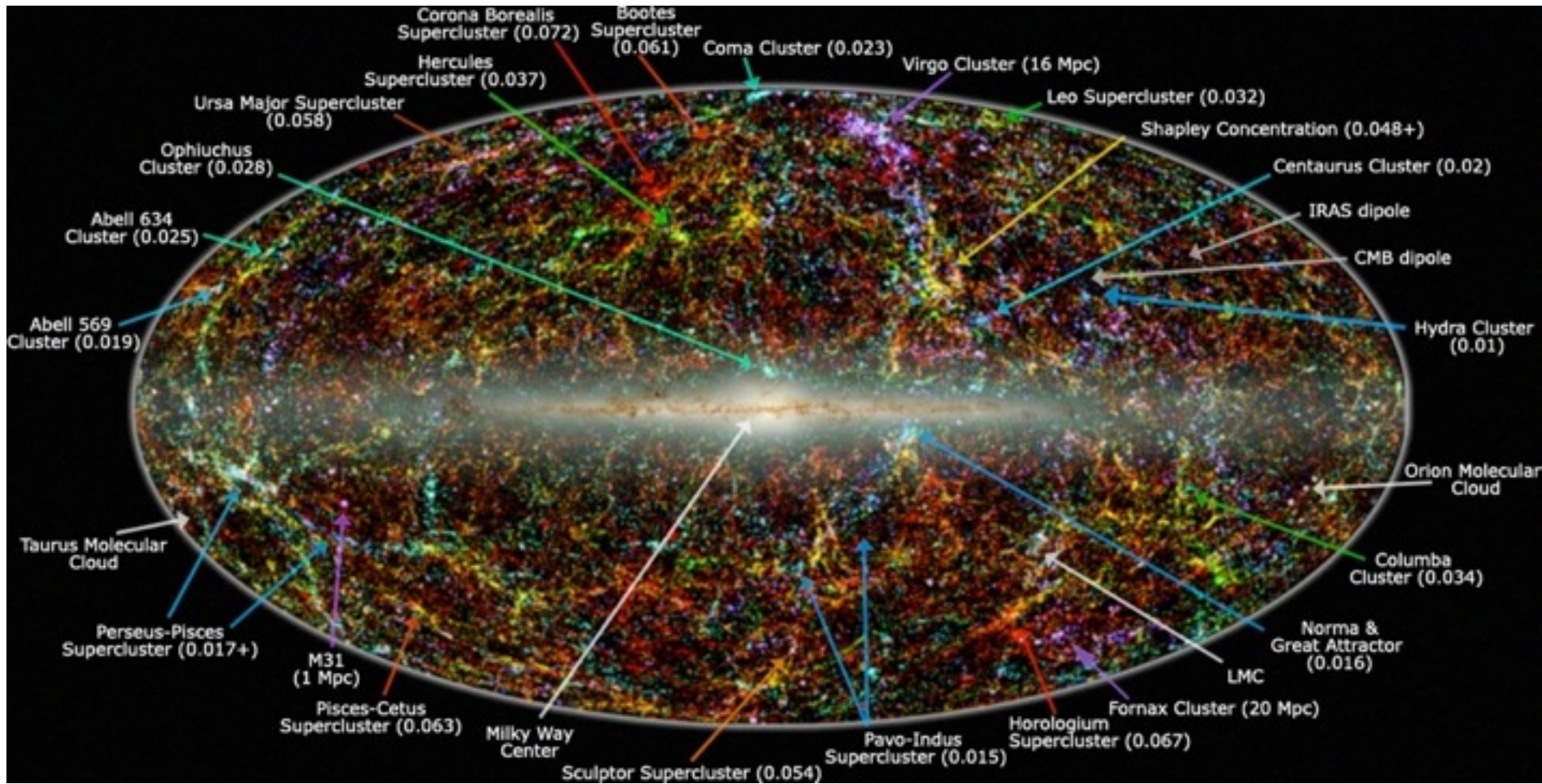
drift

$\ln a(\mathbf{x}, \ln H)$

isocon directions, e.g., axion

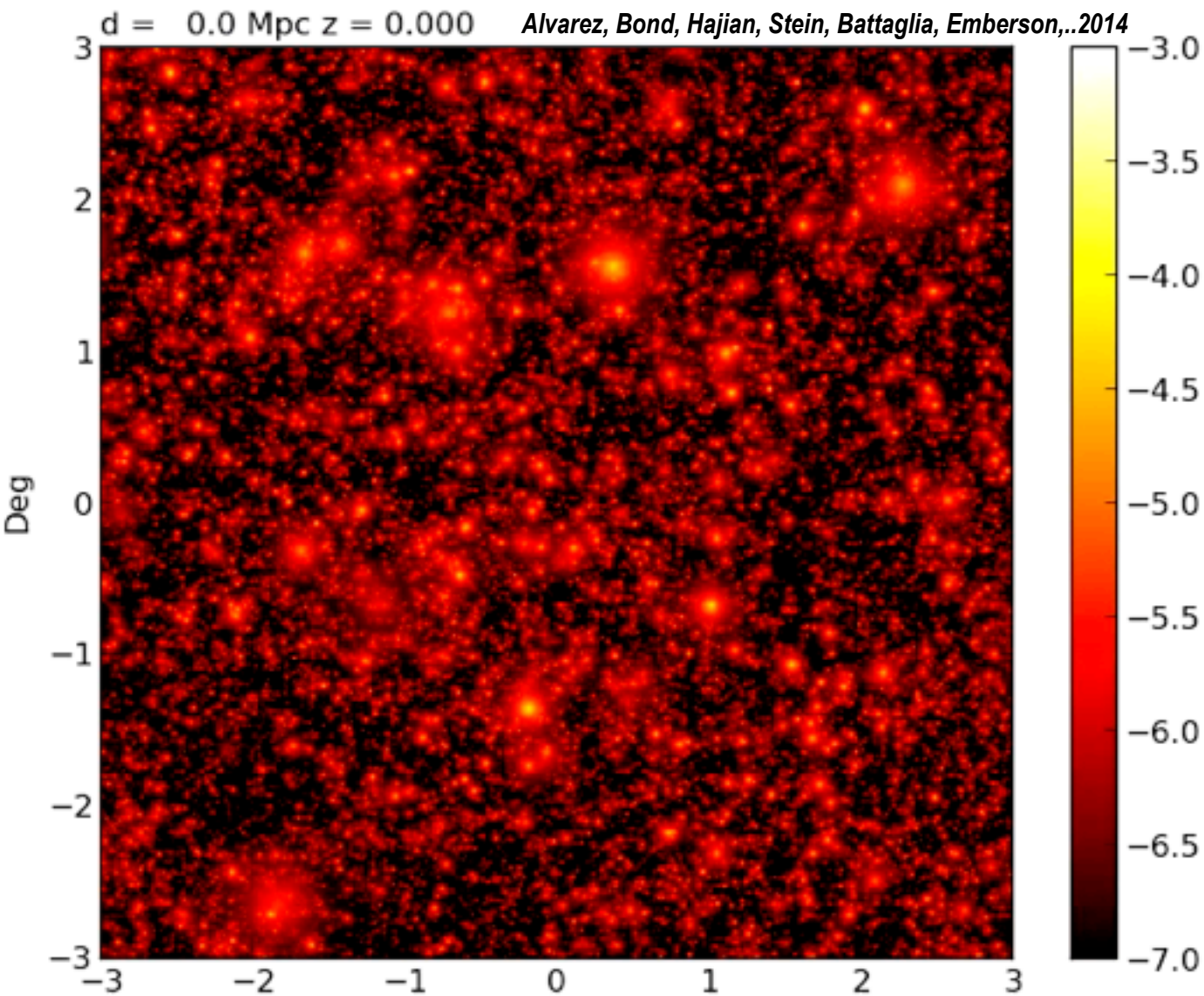
SERIAL INFLATION

cosmic web of nearby superclusters from 2mass+



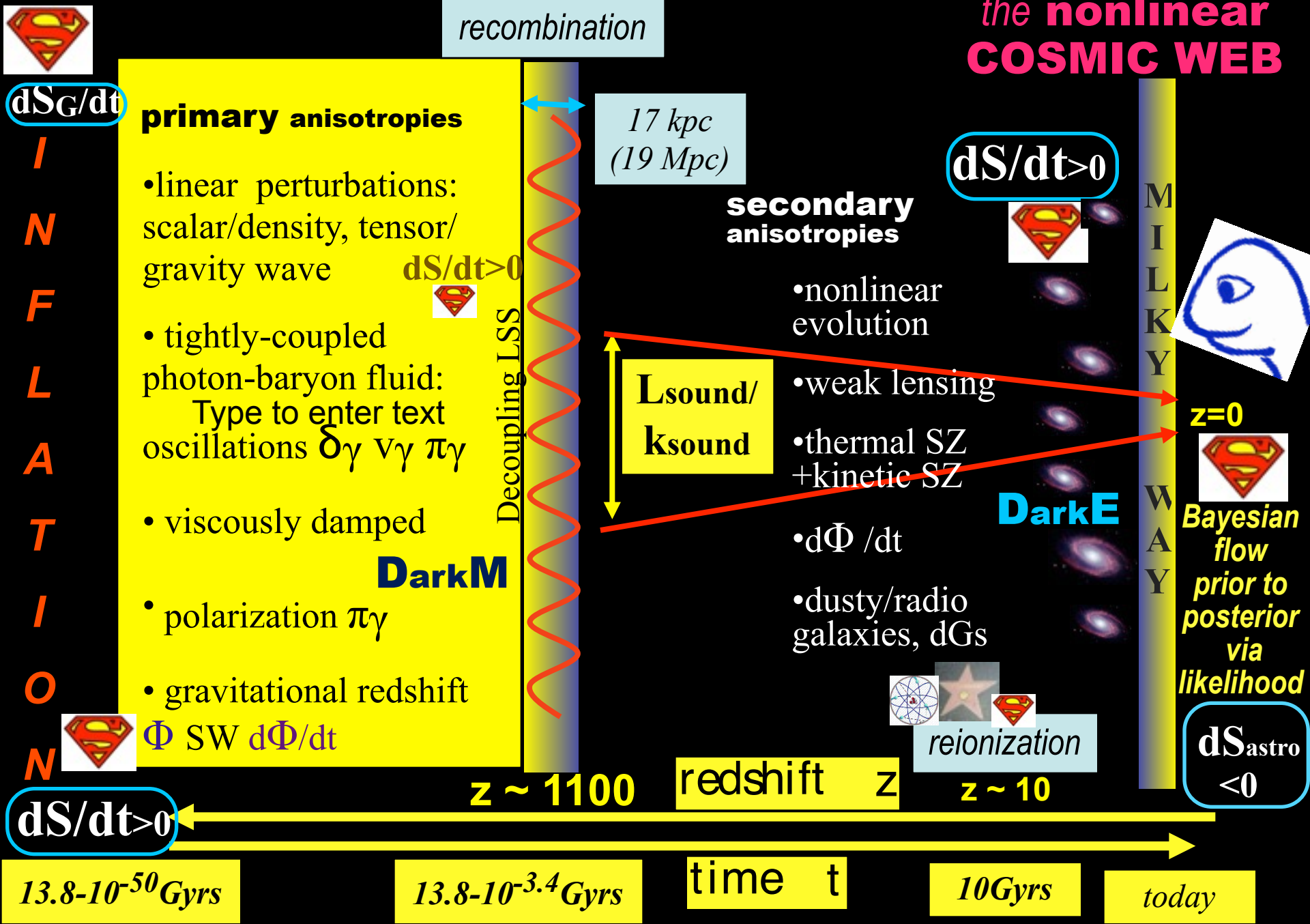
Mocking Heaven: lightcone sim for Λ CDM. 36 sq deg to $z=2$

Planck all-sky tSZ mock 1.5 hours on 256 cores on SciNet, 30000 core IBM GPC

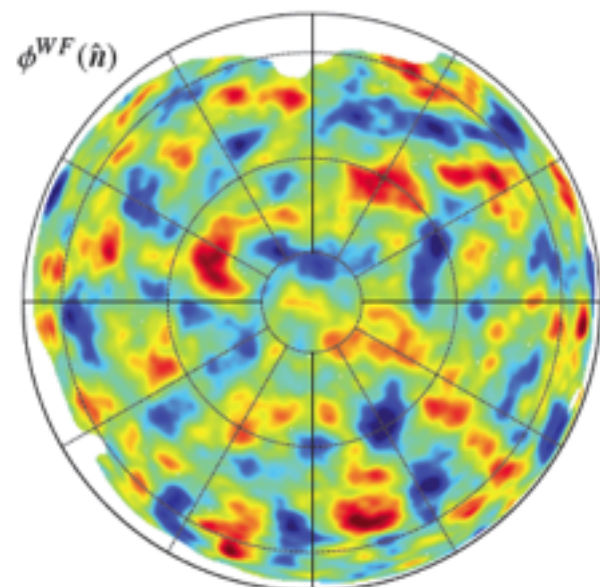


Planck, ACTpol, AdvACT, ALMA, CARMA, Mustang2 on GBT, eRosita.. COMA, CCAT.. CHIME

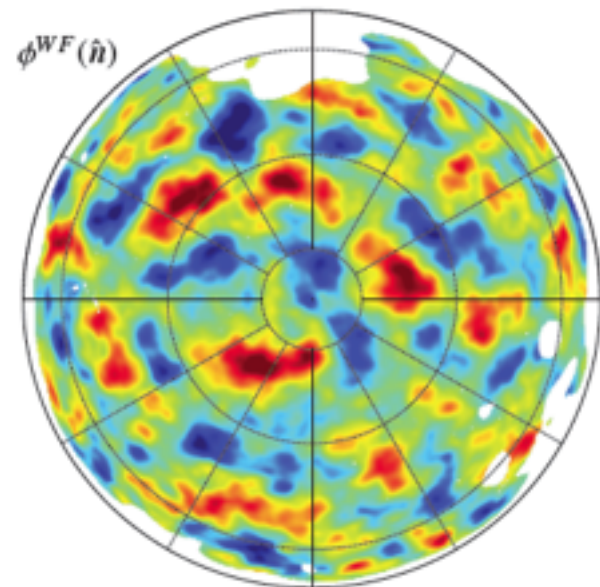
the **nonlinear** **COSMIC WEB**



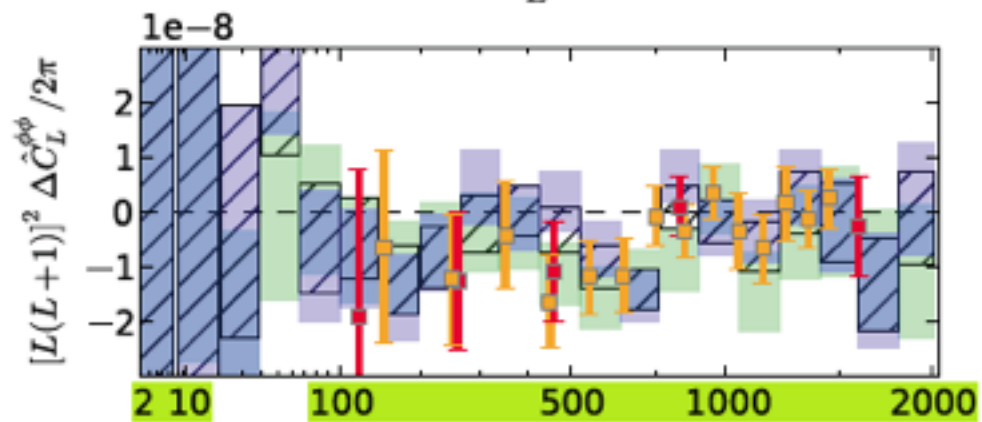
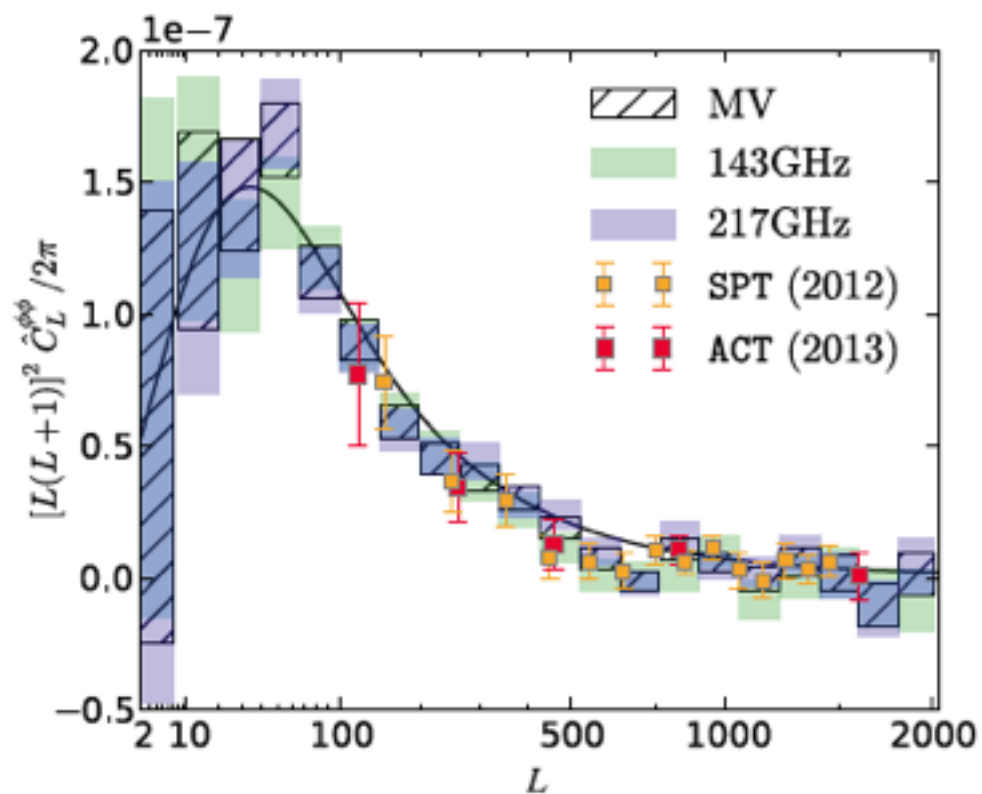
CMB Lensing: Planck13 cf. ACT12 and SPT12, good agreement



Galactic North



Galactic South



SIMPLICITY

at $a \sim e^{-7} \sim 1/1100 \Rightarrow$
at $a \sim e^{-67-60} \sim 1/10^{30+25}$

Planck2013 CMB map

reveals *primordial sound waves in matter*

\Rightarrow learn **contents & structure** at 380000 yr, $a \sim e^{-7}$

\Rightarrow infer the structure far far earlier $a \sim e^{-67-60}$

7+ numbers

Early Universe **STRUCTURE**

“**red**” **noise** in *phonons/strain*: 2 numbers at $a \sim e^{-67-55}$

$$\ln \text{Power}_s \sim \ln 22.0 \times 10^{-10} \pm 0.025$$

$$n_s = 0.9608 \pm 0.0054 \quad 5\sigma \text{ from } 1$$

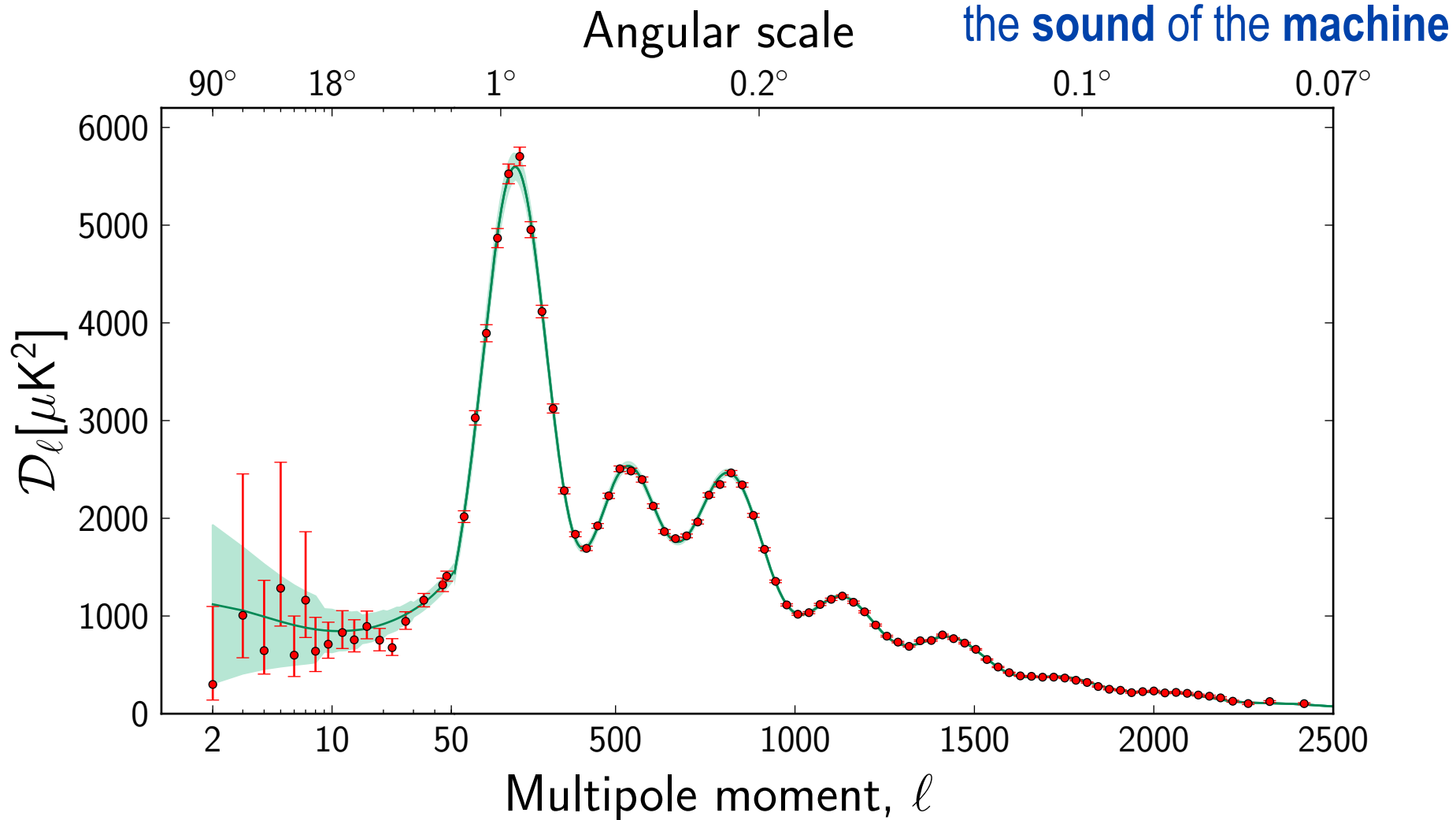
TBD: Full Mission + Polarization, Planck2014-15 + ACTpol, Spider, ..

$$-0.014 \pm 0.009$$

BICEP2

$$r < 0.12 \quad r = 0.20 + 0.07 - 0.05$$

95% CL on *running* $dn_s/d\ln k$, *running of running*, r = Tensor-to-Scalar ratio (GW), *isocurvature modes* for axions (<3.9%), baryons, neutrinos, curvatons (<0.25%)



Excellent agreement between the Planck temperature spectrum at high L and the predictions of the tilted Λ CDM model. Checks with polarization data provide full support to this conclusion.

extensive grid of cosmic models strongly constrain the x in tilted Λ CDM + x , x = subdominant deviations

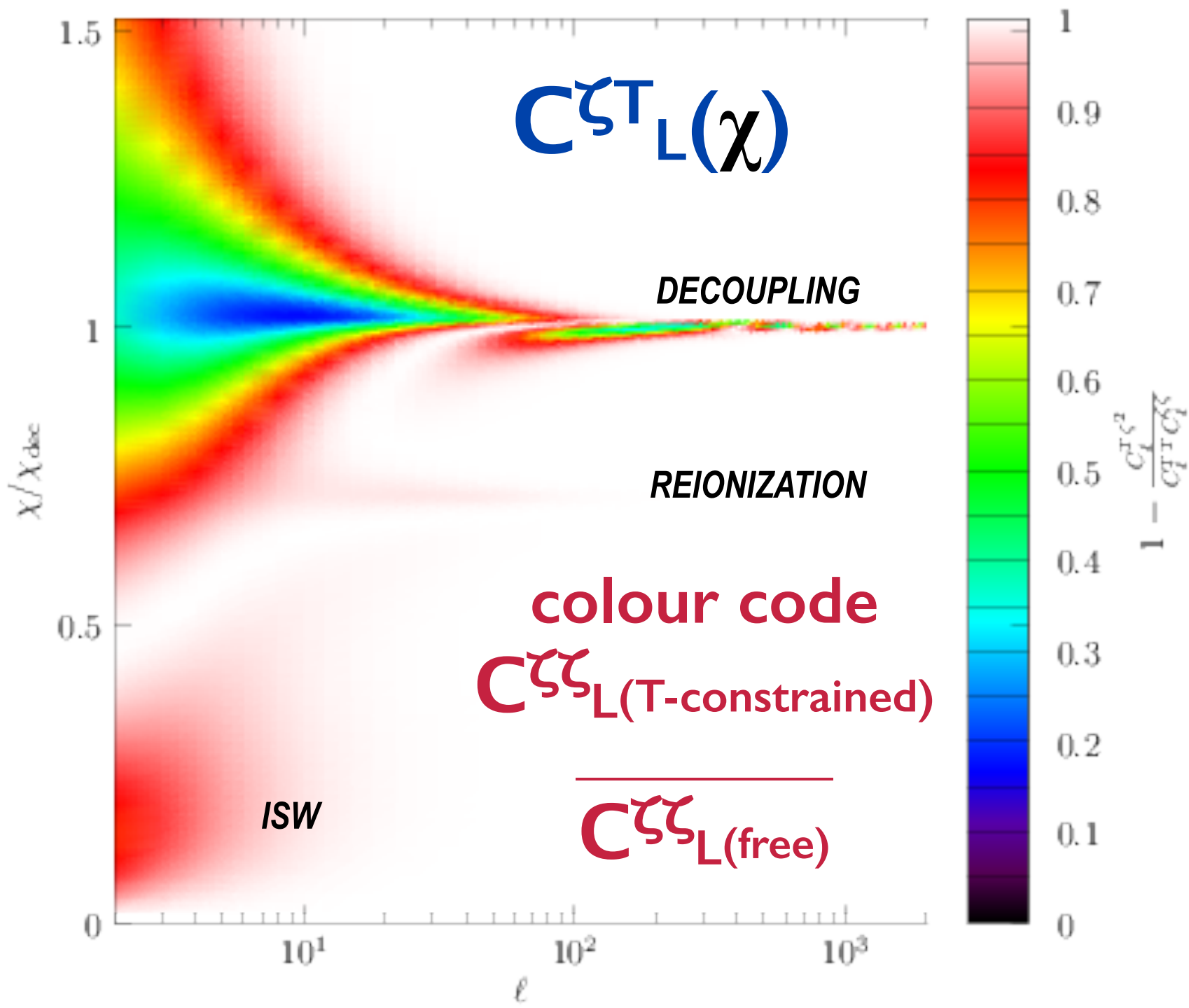
Planck basic parameters (Ω_b , H_0 ...), agree with BBN, BAO measure of acoustic scale. but H_0 lower than HST, small age change

No evidence for additional neutrino-like relativistic particles beyond the three families of neutrinos in the standard model.

The first 30 multipoles are low for the standard Λ CDM, with no obvious explanation. primordial fluctuation modification?

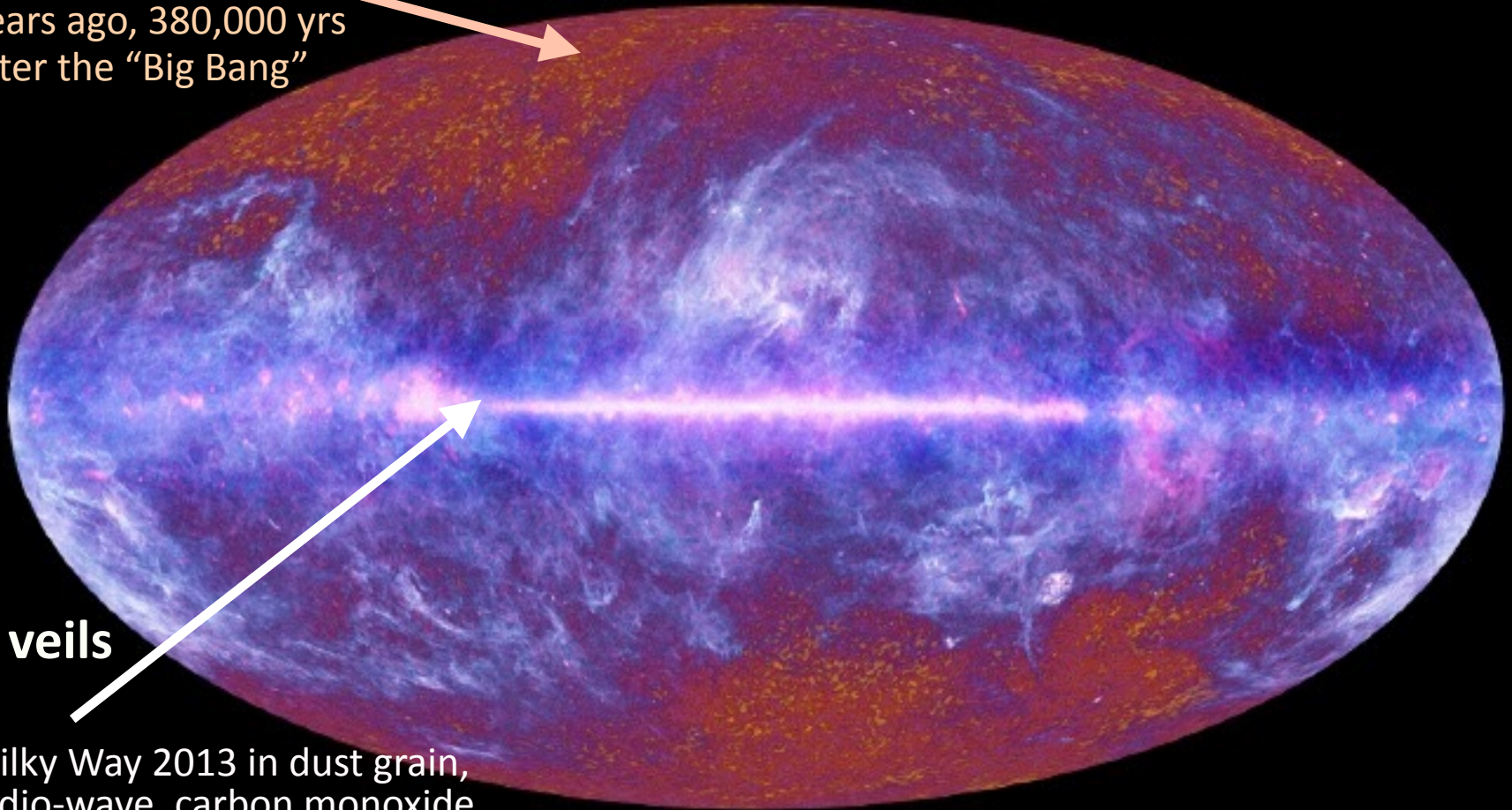
Exact scale invariance ruled out, $n_s < 1$, at $>4\sigma$ Planck alone, $>5.4\sigma$ Planck + WMAP polarization

No substantial evidence for beyond basic single field slow roll, Bunch-Davies vacuum, standard kinetic term inflation. no f_{NL}



COMPLEXITY of here & now

the primordial light,
released 13.8 billion
years ago, 380,000 yrs
after the "Big Bang"



7 veils

Milky Way 2013 in dust grain,
radio-wave, carbon monoxide
emissions; plus stellar, X-ray,
gamma ray, cosmic ray
emissions ...

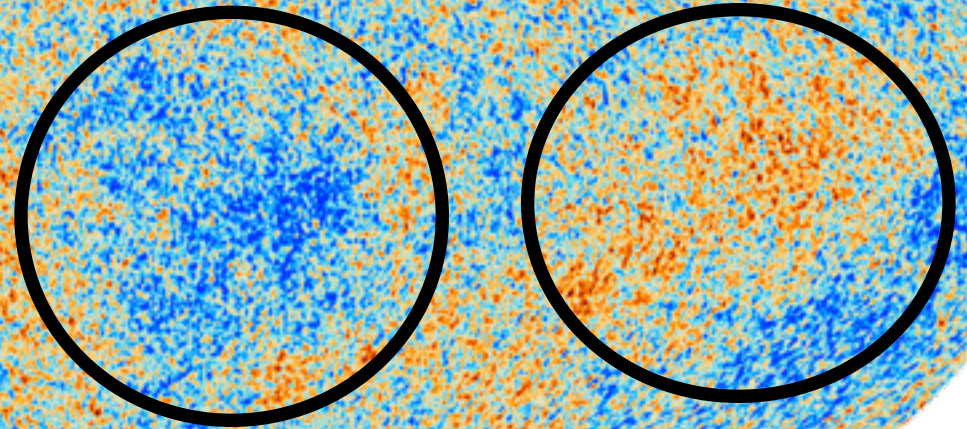
Planck's primordial light unveiled, March 21, 2013

reveals the **SIMPLICITY** of primordial cosmic structure

7⁺ numbers, 2+1 are inflation numbers

Gaussian to high precision for high multipole,
anomalies at low multipoles, non-Gaussian, anisotropic

=> inflation COMPLEXITY at $t \sim 10^{-36}$ seconds?



+ anomalies

**the rare
cold spot**

**hemisphere
difference in
power ~7% at
low resolution**

Grand Unified Theory of Anomalies? TBD
intermittent strain-power bursts (in curvature)?

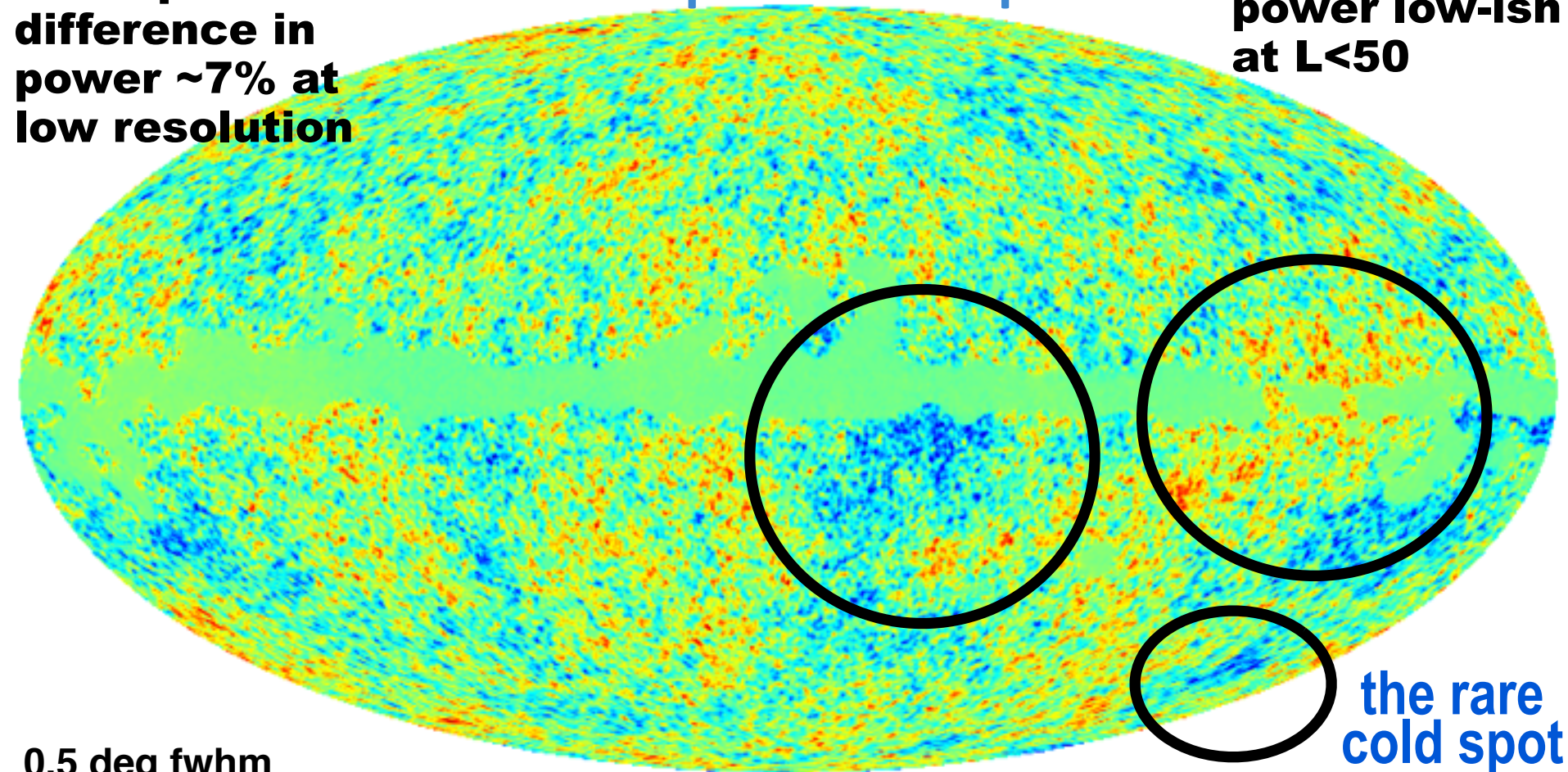
Gaussian to high precision for high multipole,
anomalies at low multipoles, non-Gaussian, anisotropic
anomalies => inflation COMPLEXITY at $t \sim 10^{-36}$ seconds?

mean temperature, 1000 realizations, smooth scale fwhm = 30 arcmin,

temperature map

power low-ish
at $L < 50$

hemisphere
difference in
power $\sim 7\%$ at
low resolution



0.5 deg fwhm

-355.

+340.

Grand Unified Theory of Anomalies? TBD

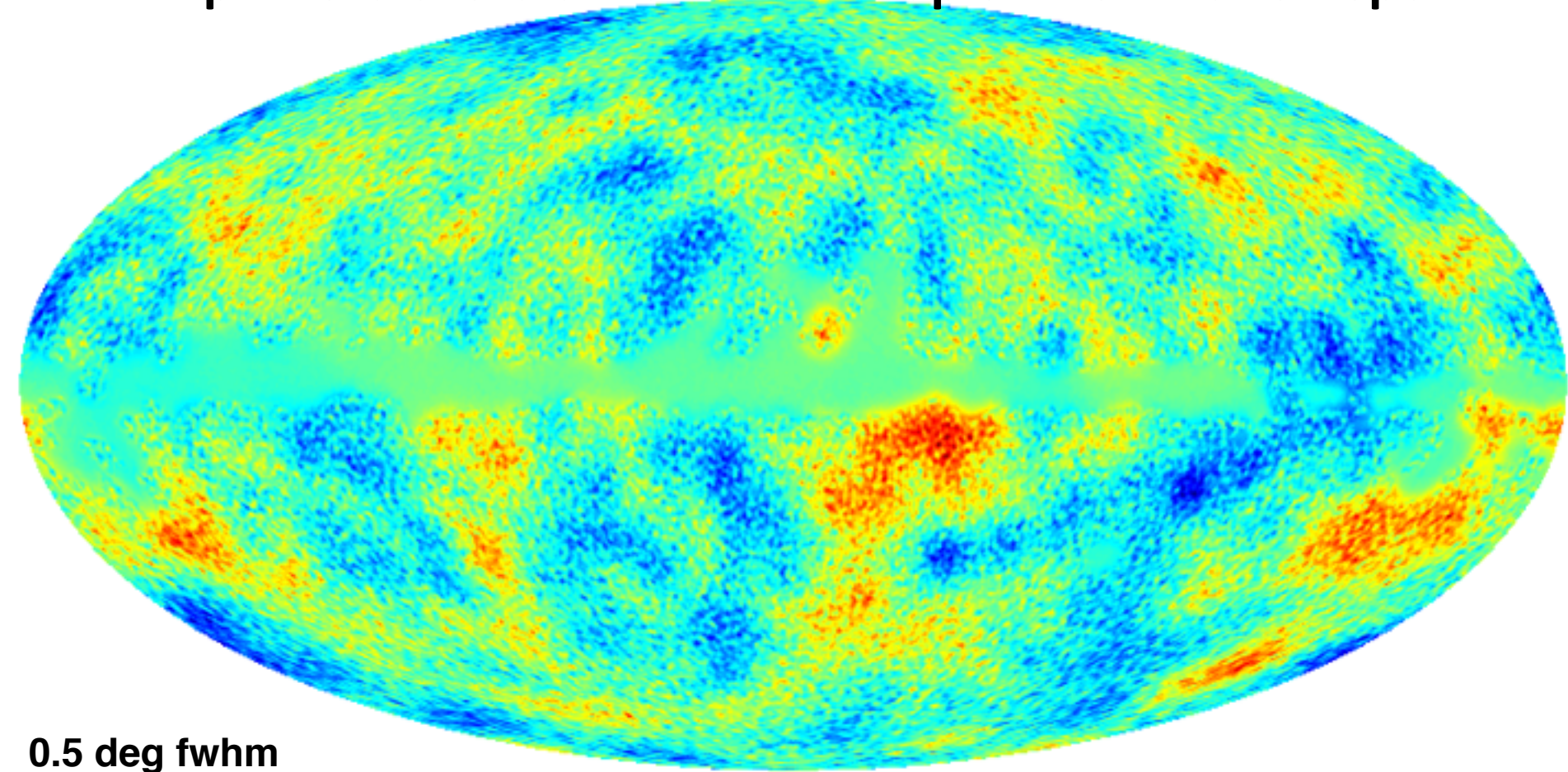
intermittent strain-power bursts (in curvature)?

reveals map of **primordial isotropic strain / phonons**

$\int d\text{visibility}(\text{distance}) < \text{Trace}(\alpha) | \text{Temp} > \text{ (angles, distance)}$

mean zeta, 1000 realizations, smooth scale fwhm = 30 arcmin,

=> **primordial scalar curvature map of the inflation epoch**



0.5 deg fwhm



Reconstructing the Early Universe

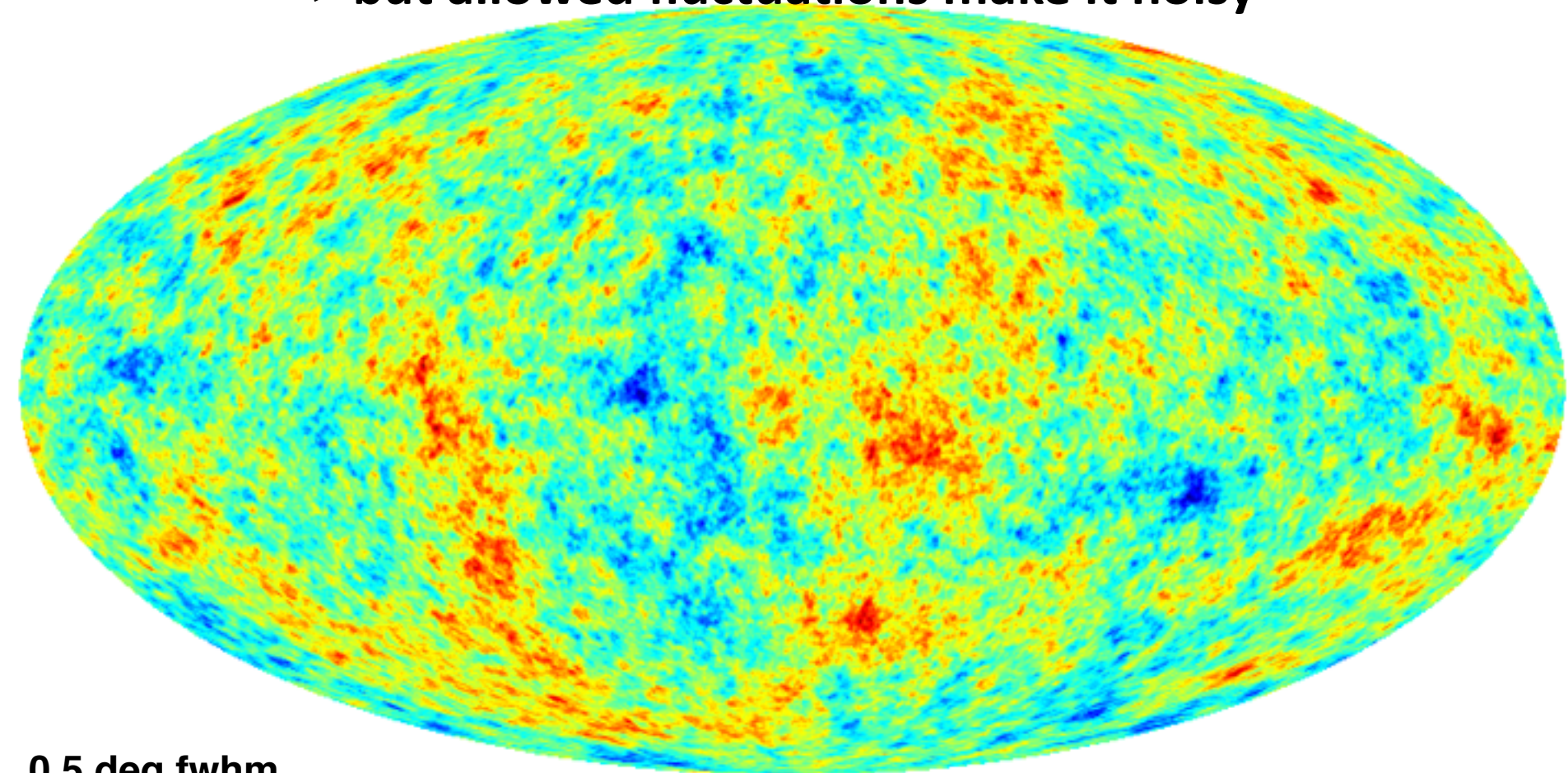
visibility mask

reveals map of **primordial isotropic strain / phonons**

$$\int d\text{visibility}(\text{distance}) \langle \text{Trace}(\boldsymbol{\alpha}) | \text{Temp} \rangle + \delta \text{Trace}(\boldsymbol{\alpha})$$

one realization of fullsky zeta, fwhm = 30 arcmin

=> **but allowed fluctuations make it noisy**



0.5 deg fwhm

-8.61

+7.54

Reconstructing the Early Universe

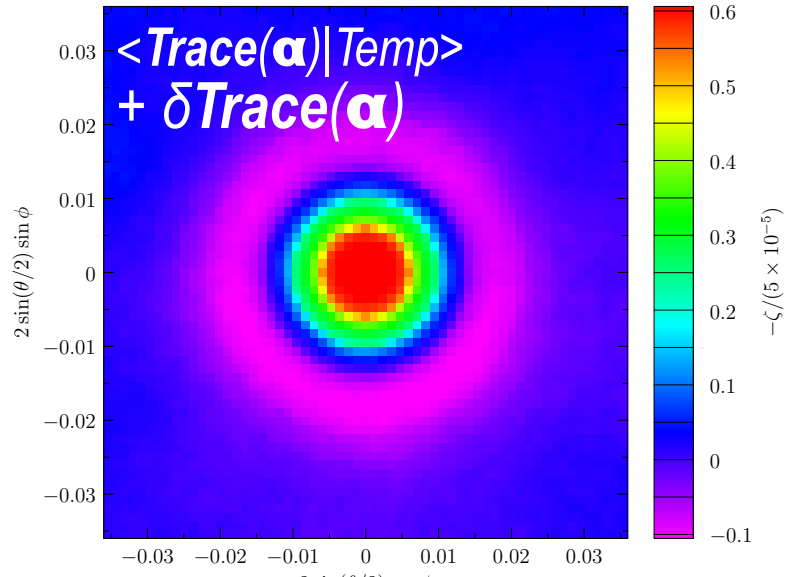
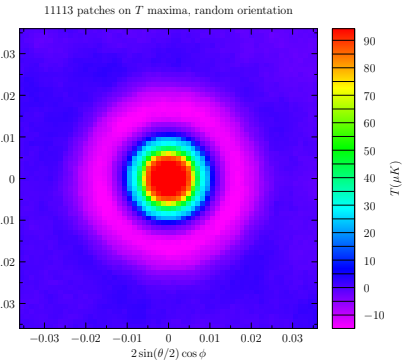
visibility mask

CMB Peak Statistics

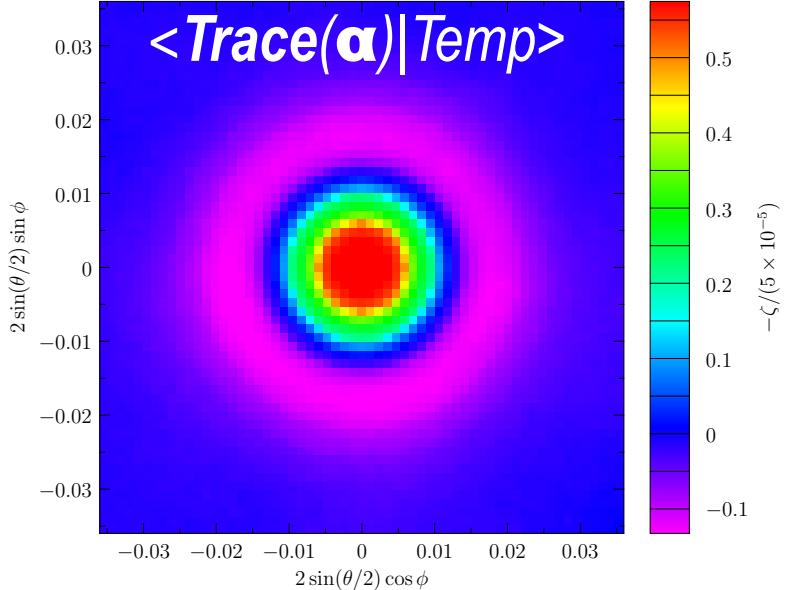
curvature stacked on temperature Peaks

Primordial curvature stacked Planck 2013

stacking a realization of ζ map, 11113 patches on T maxima, random orientation



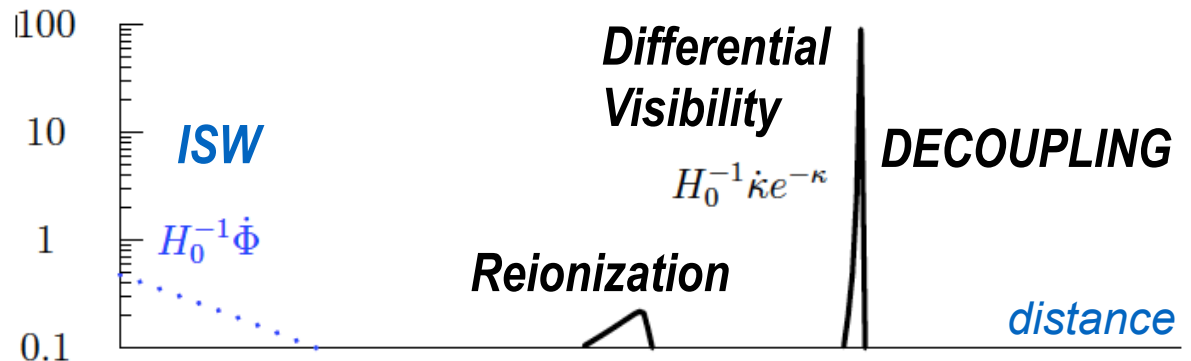
stacking mean ζ map, 11113 patches on T maxima, random orientation



reveals map of **primordial isotropic strain / phonons**

$$\int d\text{visibility}(\text{distance}) \langle \text{Trace}(\boldsymbol{\alpha}) | \text{Temp} \rangle + \delta \text{Trace}(\boldsymbol{\alpha})$$

=> but allowed fluctuations make it noisy



CMB-probe no tomography (radial distance (redshift)):

CMB-probe ~ differential visibility
at decoupling/recombination (all L)
reionization/reheating (low L)

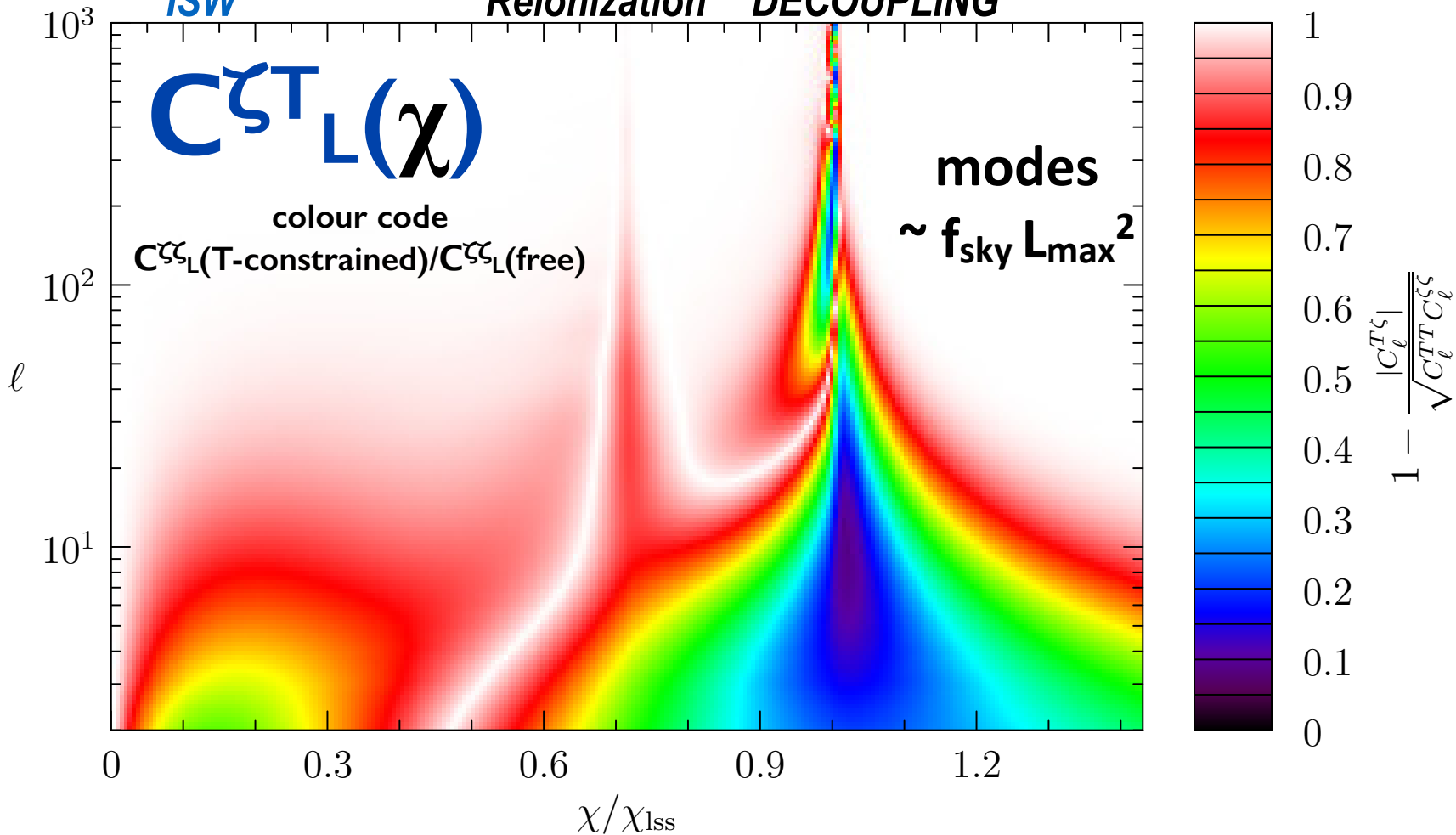
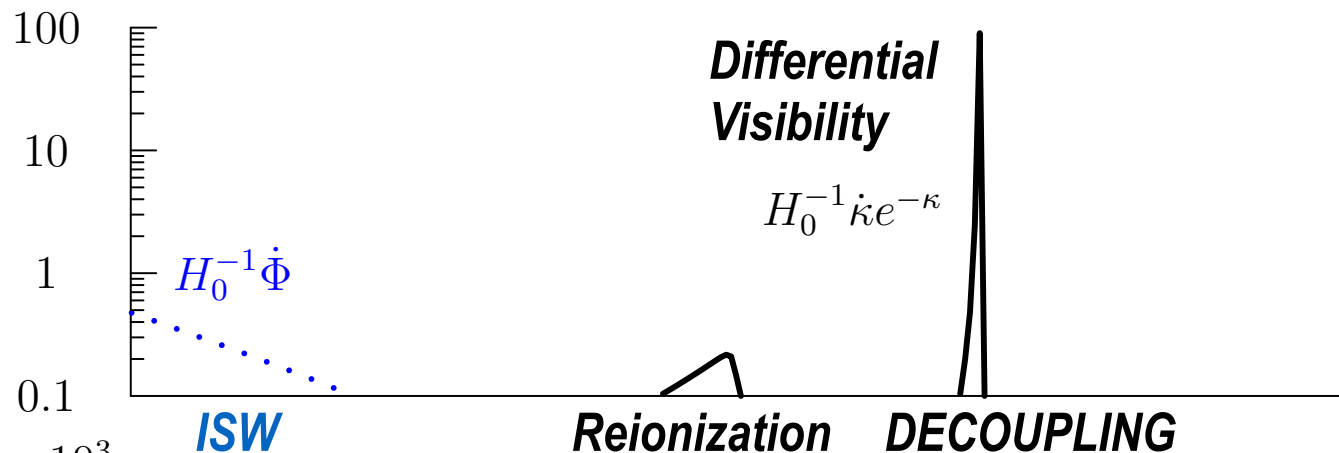
CMB-probe ~ changing gravitational potential

Integrated Sachs Wolfe effect (low L), Rees-Sciama effect (hi L)

available modes: $f_{\text{sky}} L_{\text{max}}^2 - f_{\text{sky}} L_{\text{min}}^2$ $L_{\text{max}} \sim L_{\text{damp}}$

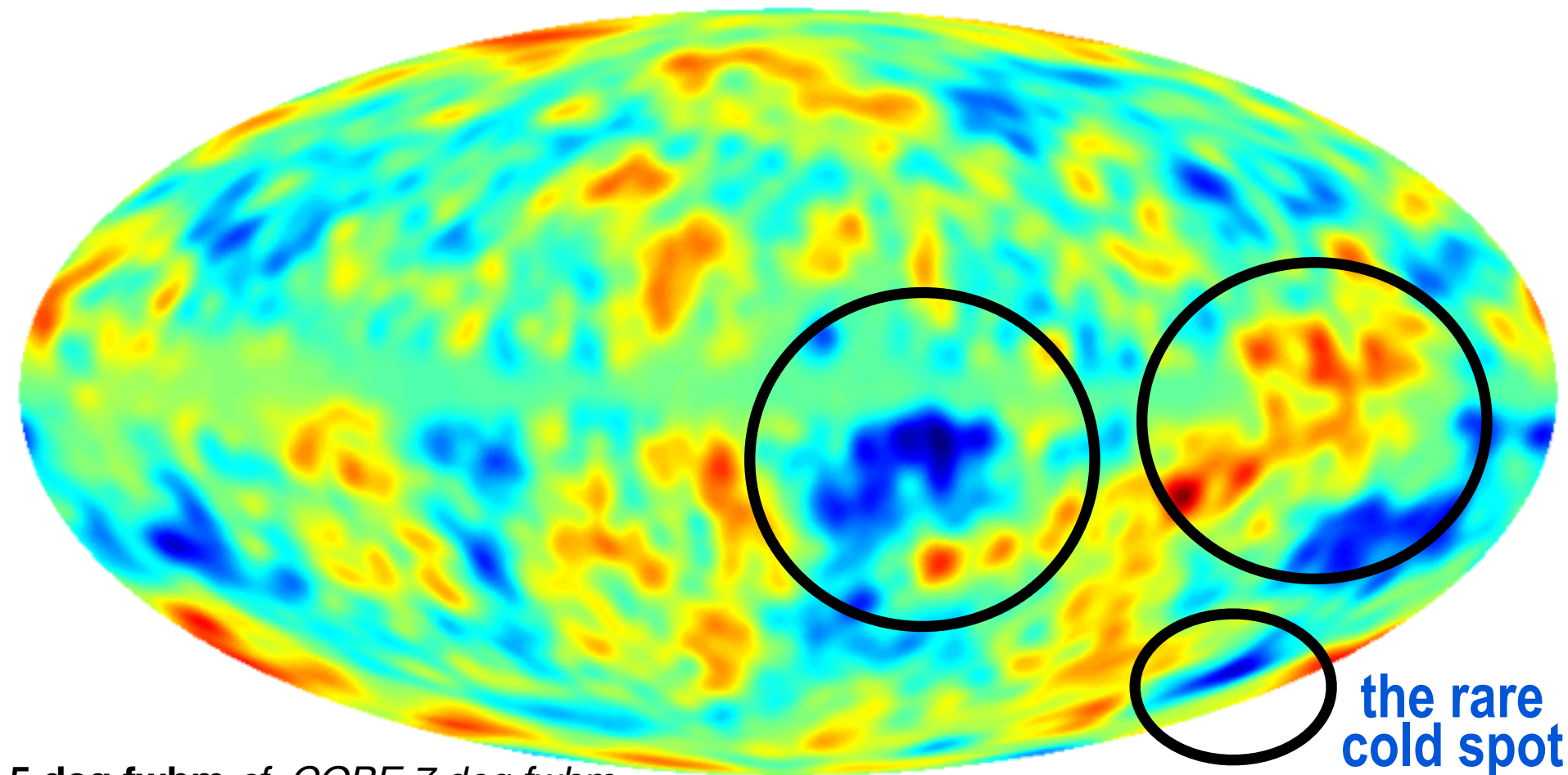
Large Scale Structure Galaxy Surveys

available modes ~ $f_{\text{sky}} L_{\text{max}}^2 k_{\text{max}} d_{\text{max}}$
 $\sim f_{\text{sky}} (k_{\text{max}}^3 d_{\text{max}}^3)$, $k_{\text{min}} \sim 2\pi/d_{\text{max}}$ $V_{\text{com}} \sim d_{\text{max}}^3$



temperature map

mean temperature, 1000 realizations, smooth scale fwhm = 300 arcmin,



-151.

+145.

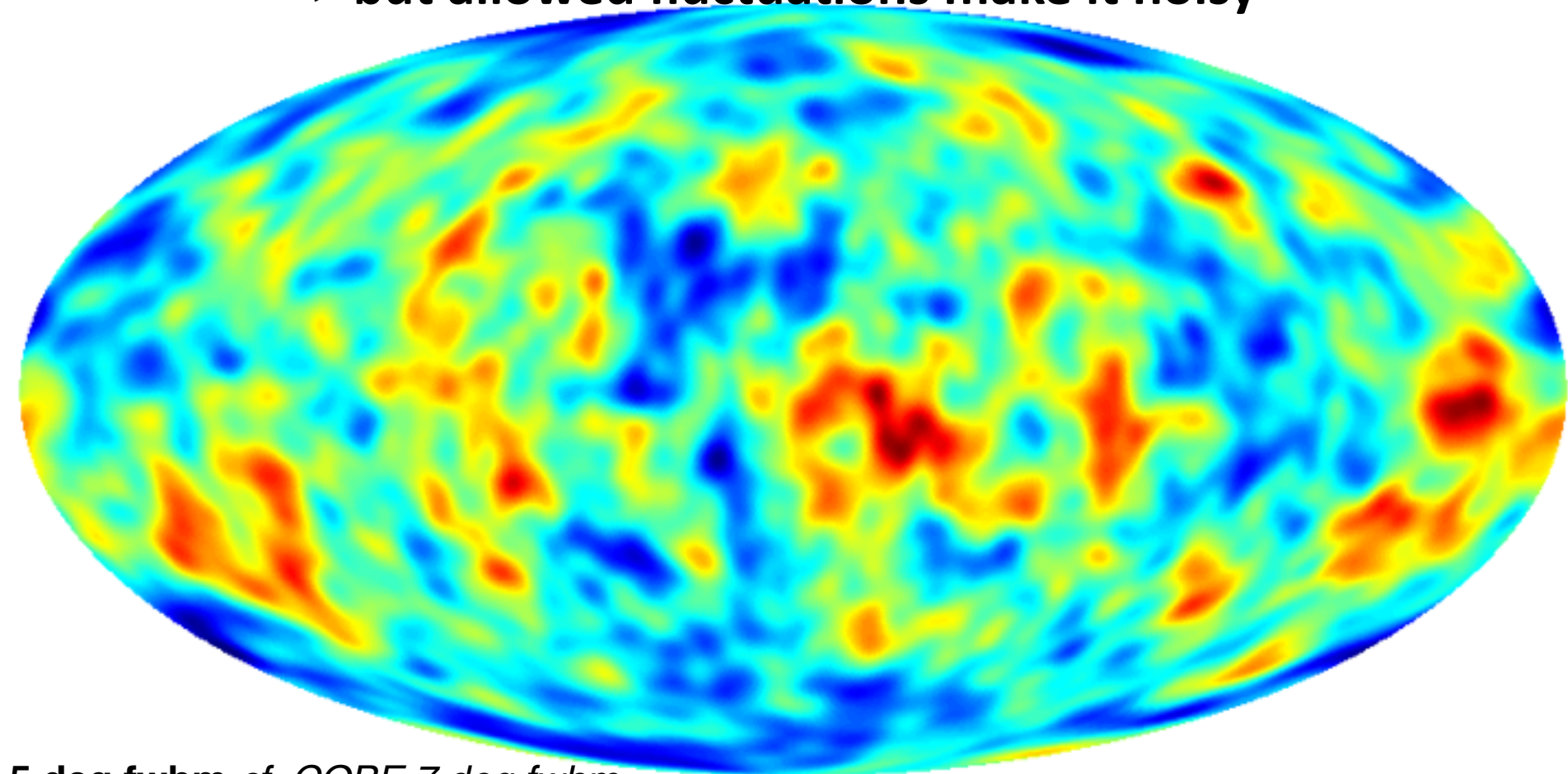
Temperature changes
in micro-degrees

reveals map of **primordial isotropic strain / phonons**

$$\int d\text{visibility}(\text{distance}) \langle \text{Trace}(\boldsymbol{\alpha}) | \text{Temp} \rangle + \delta \text{Trace}(\boldsymbol{\alpha})$$

one realization of fullsky zeta, fwhm = 300 arcmin

=> **but allowed fluctuations make it noisy**



5 deg fwhm *cf. COBE 7 deg fwhm*



Reconstructing the Early Universe

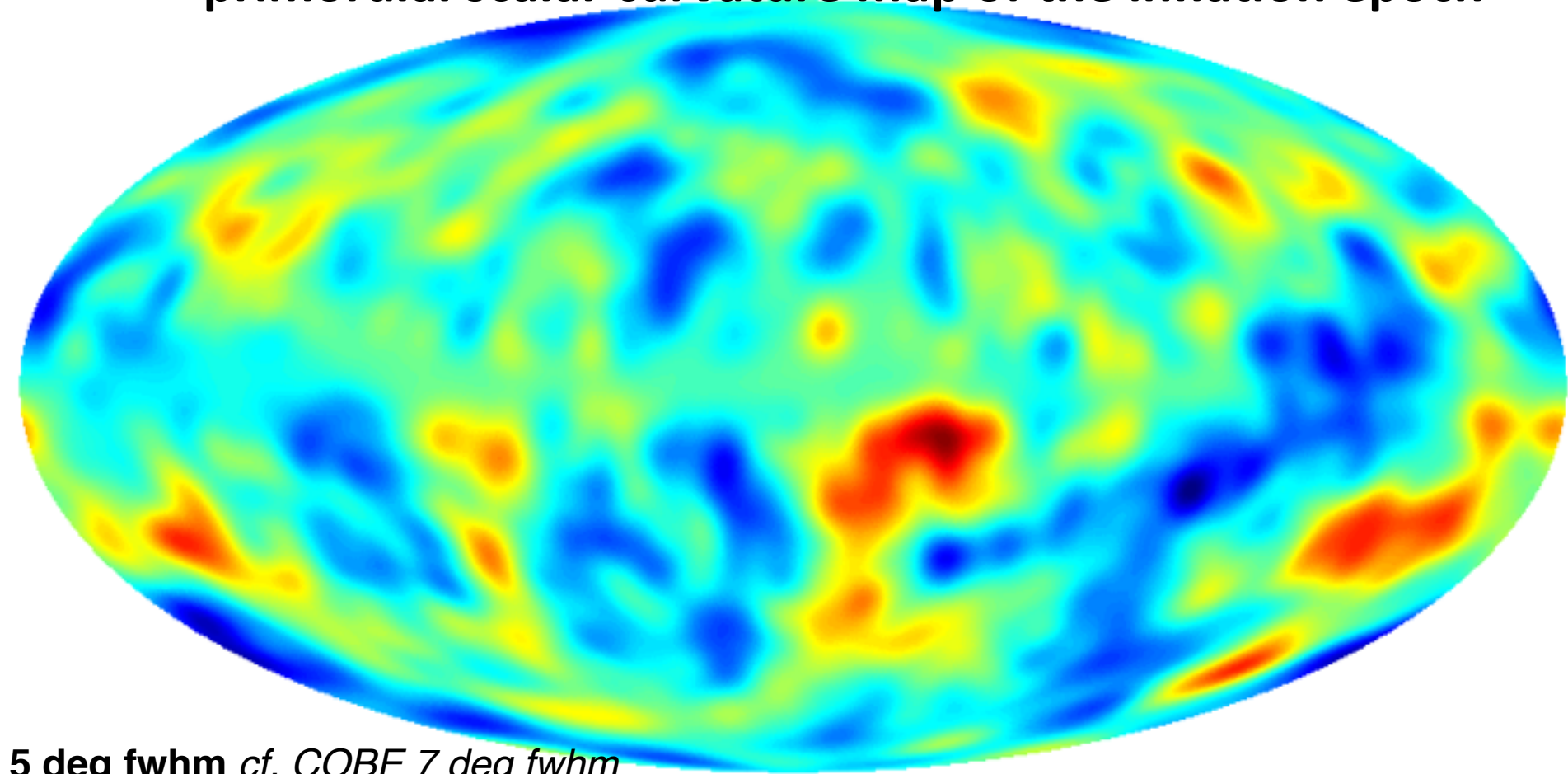
visibility mask

reveals map of **primordial isotropic strain / phonons**

$\int d\text{visibility}(\text{distance}) \langle \text{Trace}(\boldsymbol{\alpha}) | \text{Temp} \rangle$ (angles, distance)

mean zeta, 1000 realizations, smooth scale fwhm = 300 arcmin,

=> **primordial scalar curvature map of the inflation epoch**



5 deg fwhm *cf. COBE 7 deg fwhm*



Reconstructing the Early Universe

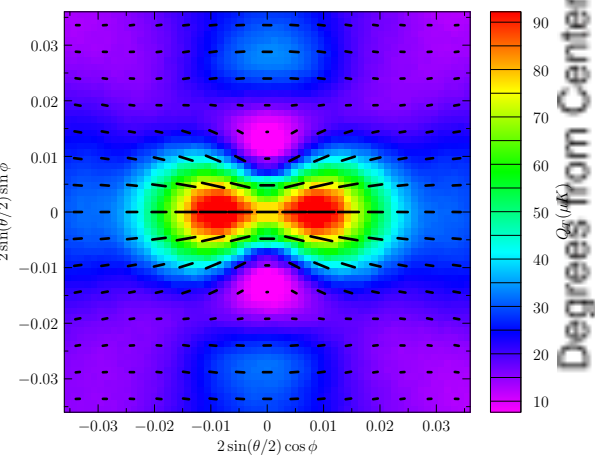
visibility mask

CMB Peak Statistics

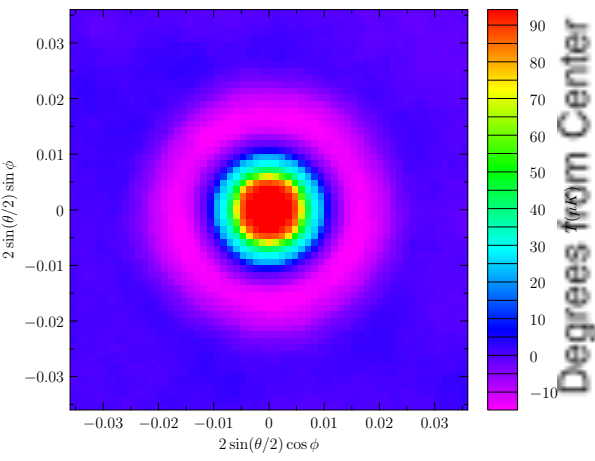
temperature stacked on temperature Peaks

polarization rotated & stacked on temperature Peaks

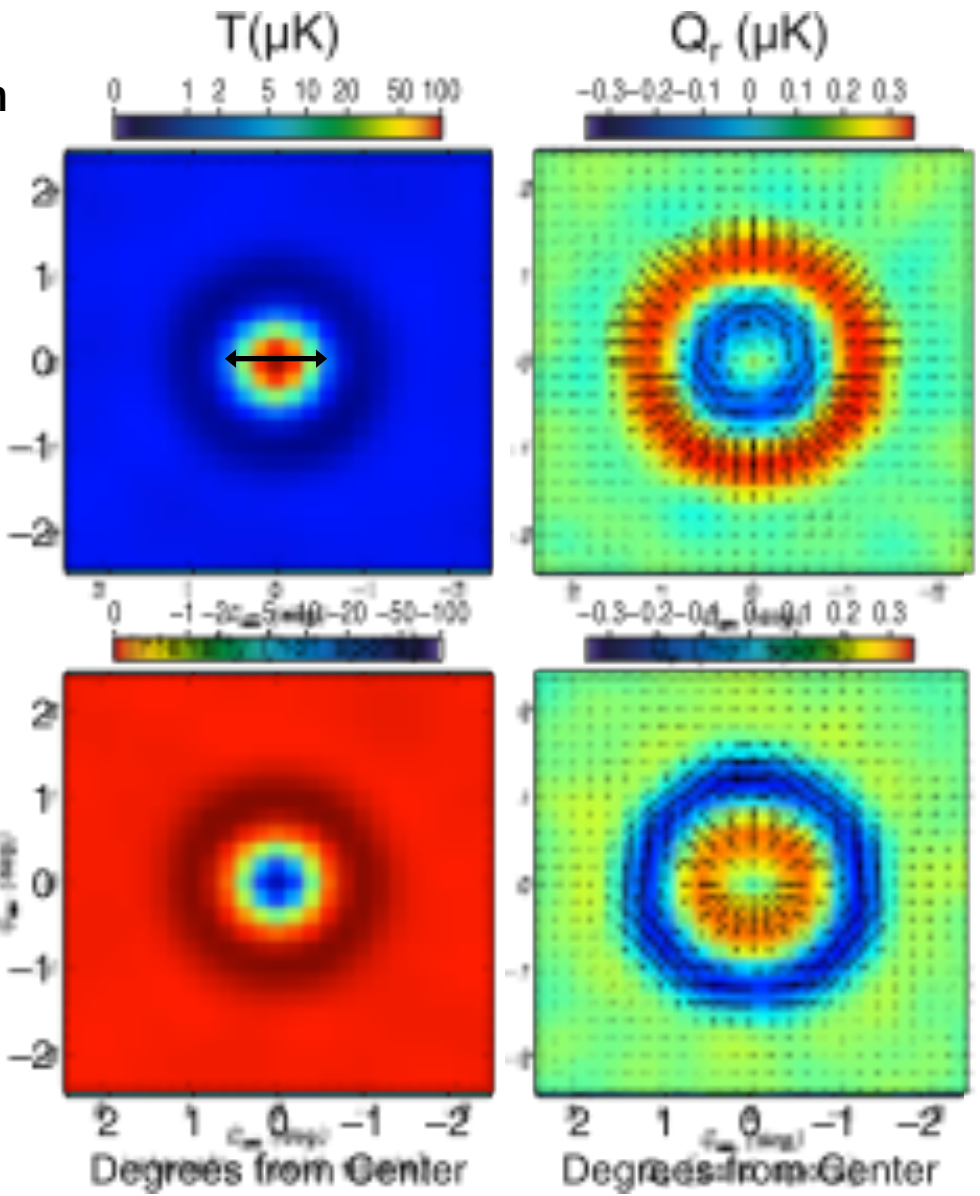
10825 Q_T patches on T maxima, oriented



11113 patches on T maxima, random orientation



CMB Polarization BAO in the CMB - Planck2013



Planck2013 teaser for Planck2014 polarization release

E mode patterns



no B here

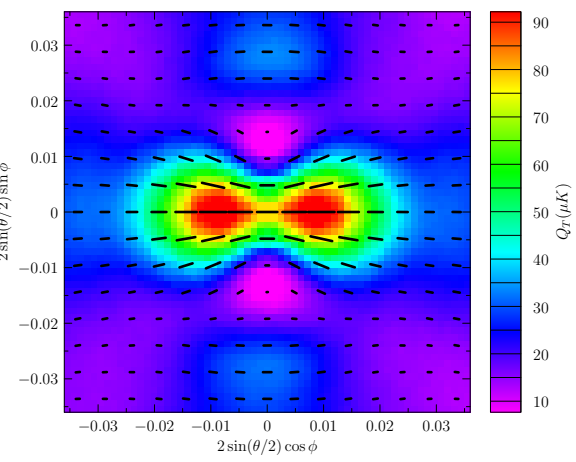
Planck2014, 2015 ACTpol, ABS, Spider, AdvACT, GLP, ..

CMB Peak Statistics

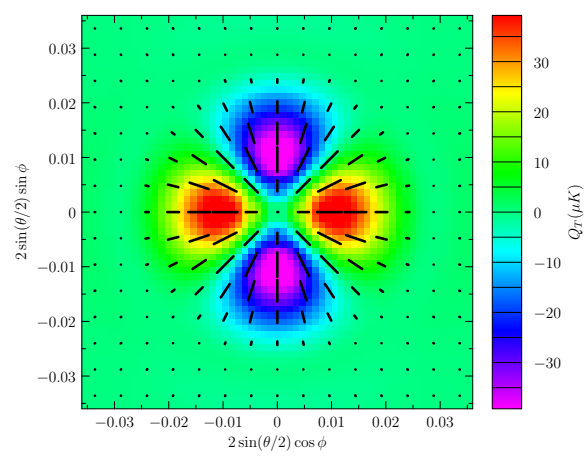
temperature stacked on temperature Peaks

polarization rotated & stacked on temperature Peaks

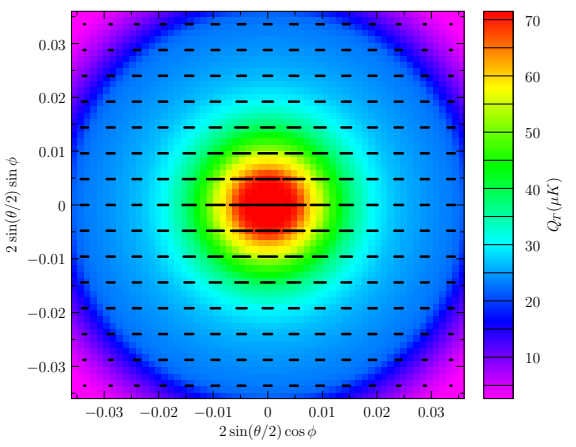
10825 Q_T patches on T maxima, oriented



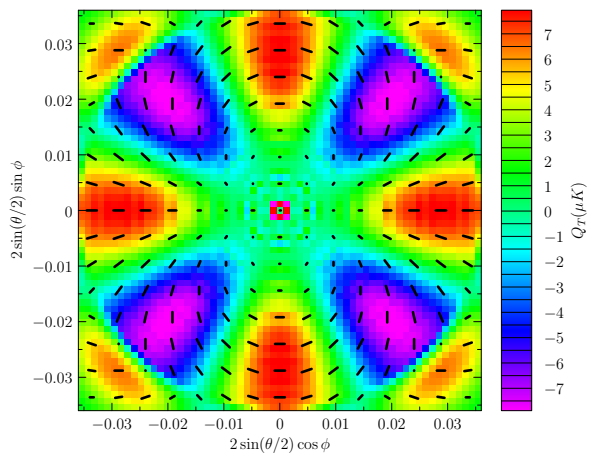
10825 Q_T patches on T maxima, oriented, $m = 2$ component



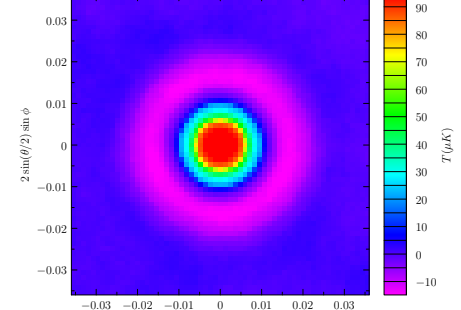
10825 Q_T patches on T maxima, oriented, $m = 0$ component



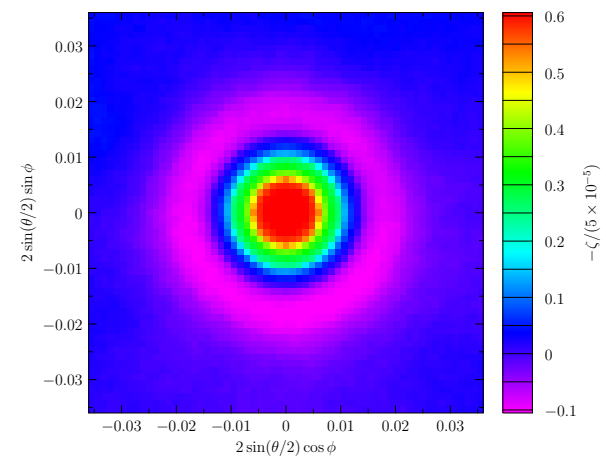
10825 Q_T patches on T maxima, oriented, $m = 4$ component



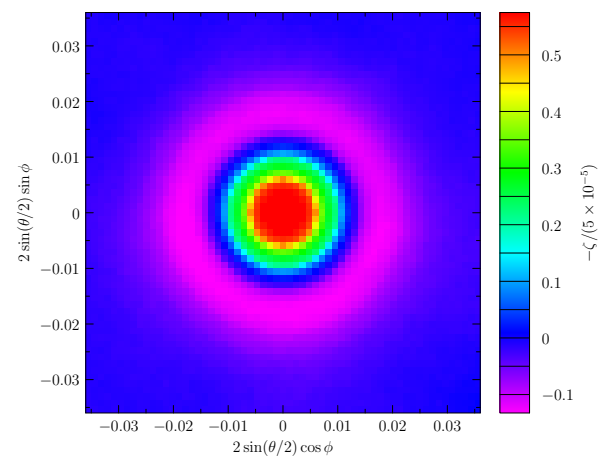
CMB Polarization Planck 2013



stacking a realization of ζ map, 11113 patches on T maxima, random orientation



stacking mean ζ map, 11113 patches on T maxima, random orientation

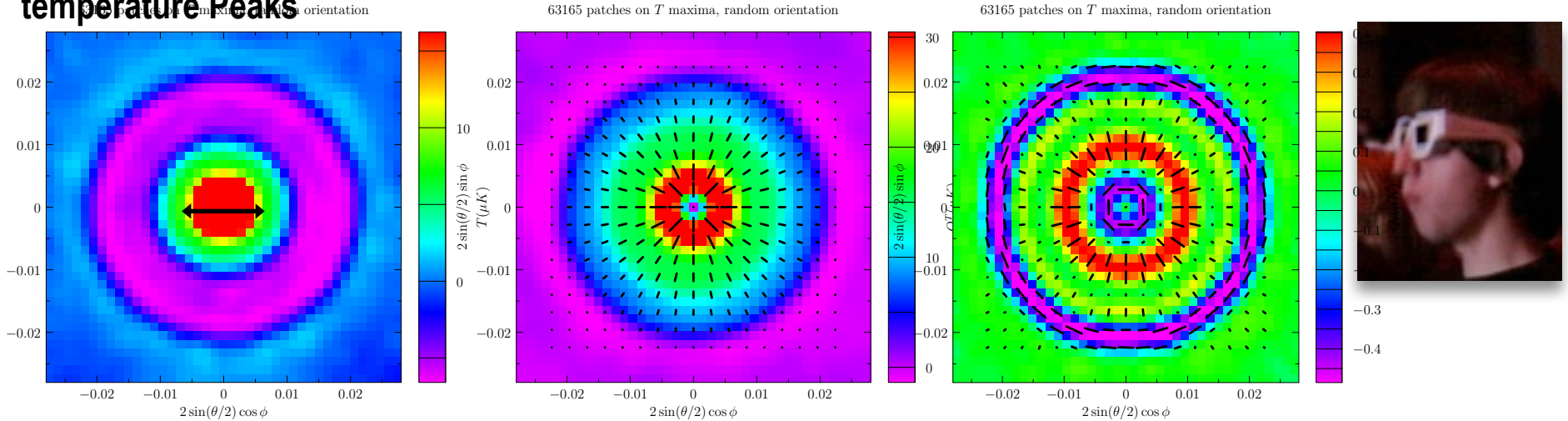


CMB Peak Statistics

temperature stacked on temperature Peaks
 polarization rotated & stacked on temperature Peaks

CMB Polarization

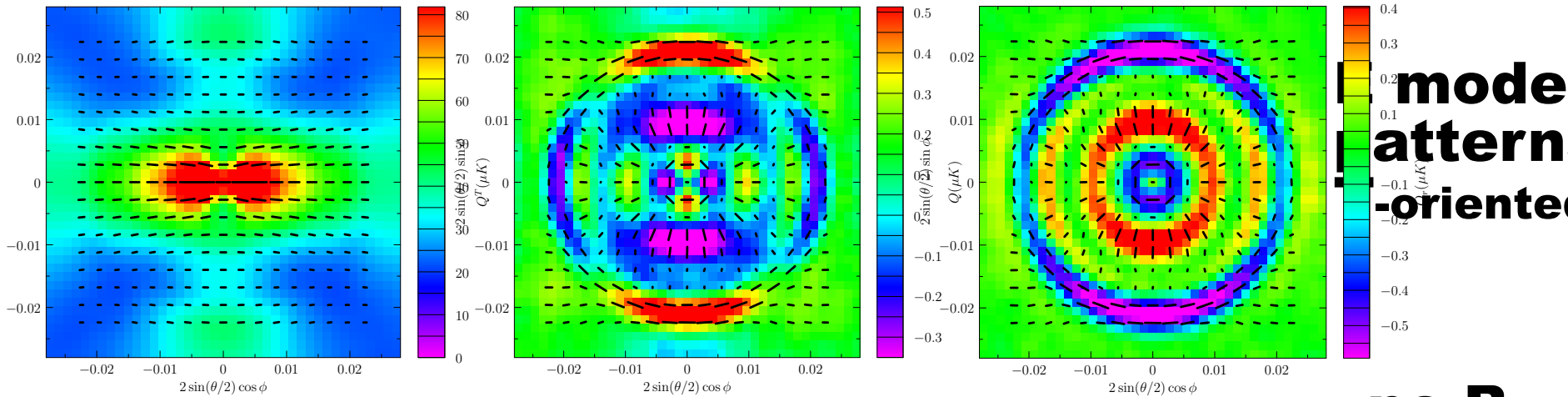
sample temperature and polarization patterns for **Planck2014: oriented peaks**



63165 patches on T maxima, oriented

63165 patches on T maxima, oriented

63165 patches on T maxima, oriented



mode patterns oriented

no B here

Planck2014, 2015 ACTpol, ABS, Spider, AdvACT, GLP, ..

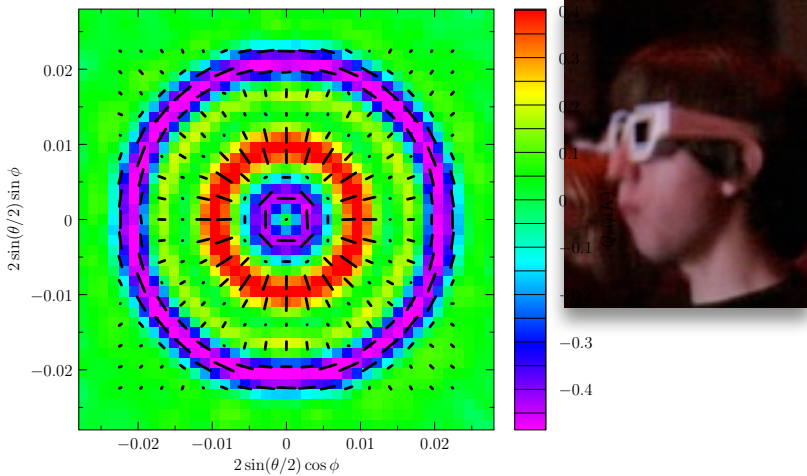
CMB Peak Statistics

temperature stacked on temperature Peaks
 polarization rotated & stacked on temperature Peaks

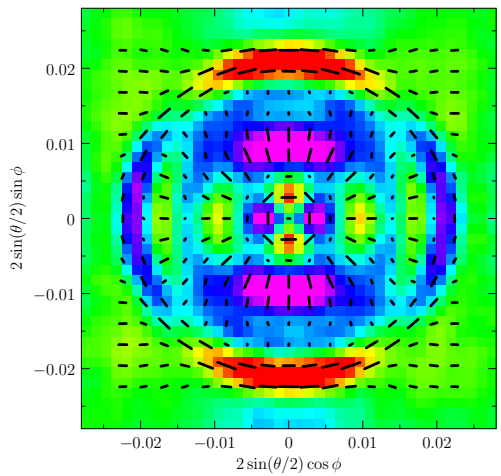
CMB Polarization

sample temperature and polarization patterns for **Planck2014: oriented peaks**

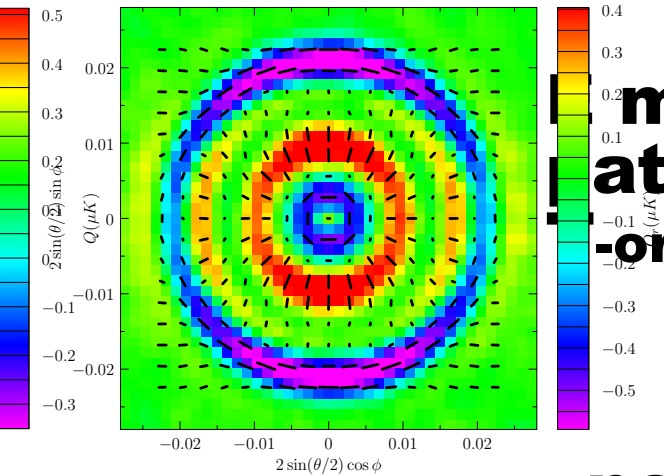
63165 patches on T maxima, random orientation



63165 patches on T maxima, oriented



63165 patches on T maxima, oriented



mode patterns oriented

no B here

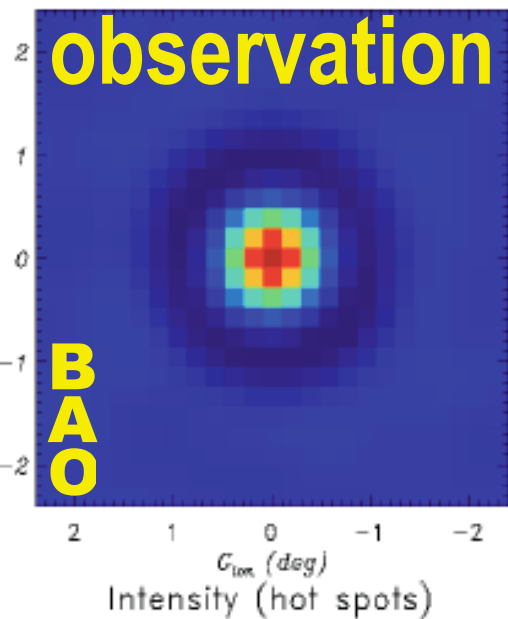
Planck2014, 2015 ACTpol, ABS, Spider, AdvACT, GLP, ..

SIMPLICITY

reveals *primordial sound waves in matter*

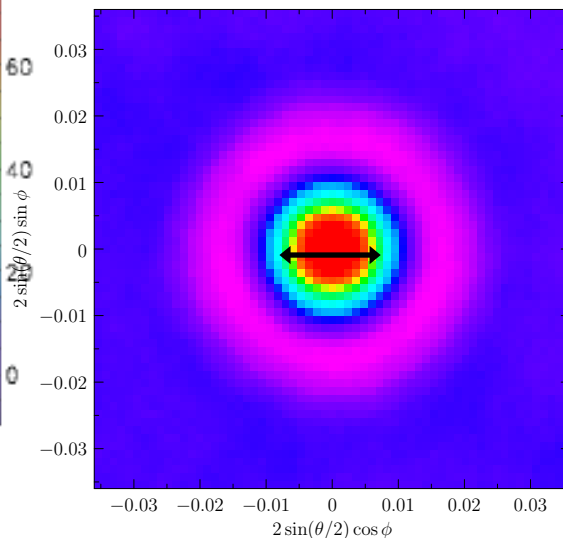
at $a \sim e^{-7} \sim 1/1100 \Rightarrow$

at $a \sim e^{-67+60} \sim 1/10^{30+25}$



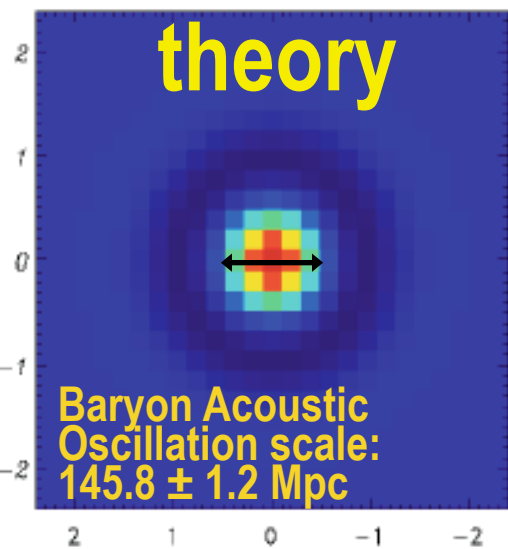
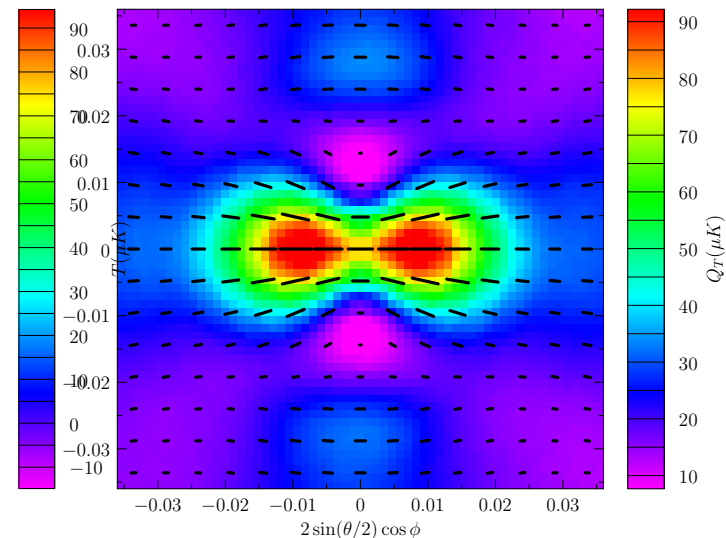
B2FHAP14, Planck2013

11113 patches on T maxima, random orientation

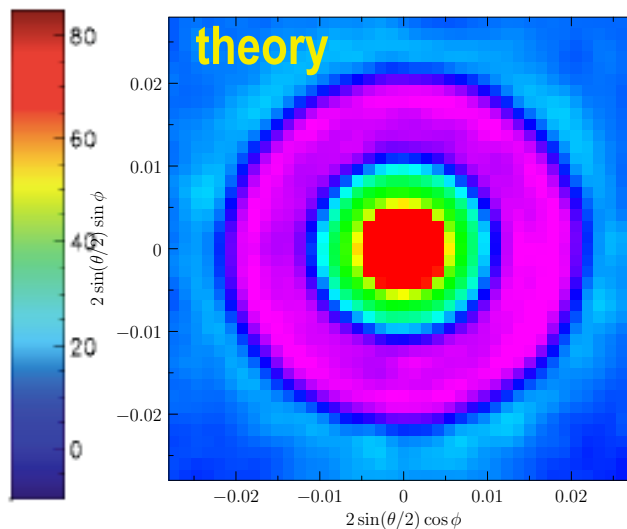


oriented peaks, anisotropic CMB strain

10825 Q_T patches on T maxima, oriented



63165 patches on T maxima, random orientation



63165 patches on T maxima, oriented

