

Fundamental Physics from the Planck Satellite

Planck 2013 results. XXII. Constraints on inflation

Planck 2013 Results. XXIV. Constraints on primordial non-Gaussianity

Planck 2013 results. XXIII. Isotropy and Statistics of the CMB

Planck 2013 results. XXV. Searches for cosmic strings and other topological defects

Planck 2013 results. XXVI. Background geometry and topology of the Universe

CMB in Canada: @CITA Boomerang, Acbar, CBI1,2, WMAP, Planck, ACT, Spider, Blast, & ACTpol, ABS, QUIET2; GBT-Mustang2, CARMA/SZA, SCUBA2, ALMA, CCAT. CMB@CIFAR:+ APEX, SPT, SPTpol, EBEX

the Planck Collaboration, including individuals from more than 100 scientific institutes in Europe, the USA and Canada



Planck is a project of the European Space Agency, with instruments provided by two scientific Consortia funded by ESA member states (in particular the lead countries: France and Italy) with contributions from NASA (USA), and telescope reflectors provided in a collaboration between ESA and a scientific Consortium led and funded by Denmark.

Planck 2013 results. XII. Component separation

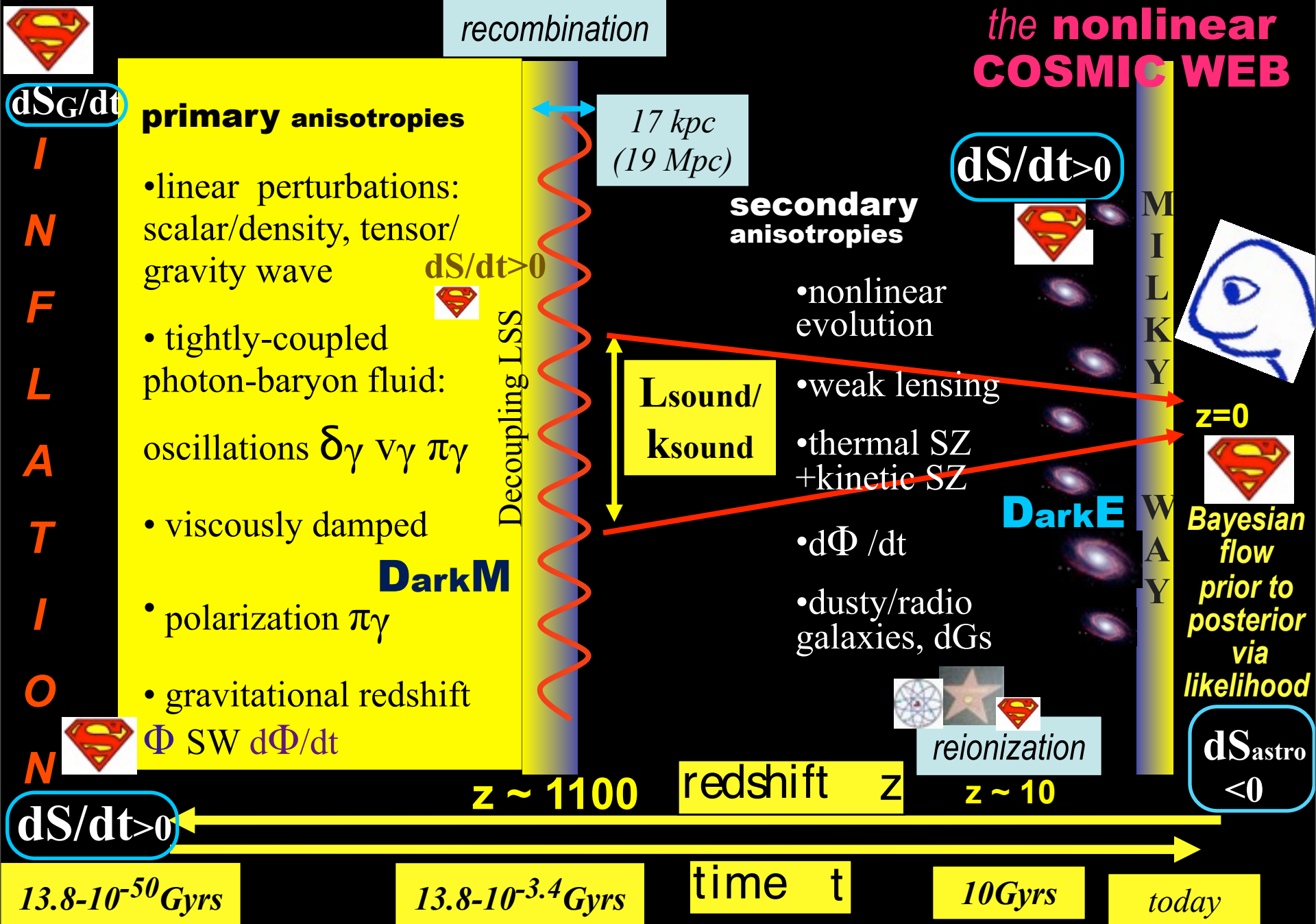
Planck 2013 results. XV. CMB power spectra and likelihood

Planck 2013 results. XVI. Cosmological parameters

Planck 2013 results. XVII. Gravitational lensing by large-scale structure

Planck 2013 results. XXVII. Doppler boosting of the CMB: Eppur si muove

Planck 2013 results. XIX. The integrated Sachs-Wolfe effect



∃ acceleration then & now $\{a, H \sim \rho^{1/2}/M_p, \epsilon = -d \ln H / d \ln a = 1 + q = 3/2(1 + w)\}$

∃ inflation then ($a \sim e^{-67}$ to $e^{-67-55++} < 10^{-35}$ s) & now ($a \sim 1$ to $e^{-1+} 10^{17}$ s)

∃ dark potential energy then $V_{de} \lesssim (10^{25.3} \text{ ev})^4$ & now $V_{de} \sim (10^{-2.9} \text{ ev})^4$

∃ dark kinetic energy then $K_{de} \lesssim (.003) V_{de}$ & now? $K_{de} \sim (-0.1 !! \text{ to } 0) V_{de}$

modified gravity = de: conformally equivalent to Einstein gravity + late-time
inflaton + fifth forces matter-de interaction ($\sim \rho_m - 3p_m = \text{Trace } T_m$)

∃ (zero-point) **quantum fluctuations** => the origin of observed cosmic structure

∃ curvature fluctuations. **scalar: adiabatic** + isocons, **tensor: gravity wave**

∃ **phonons** in early U $\ln(\rho a^{3(1+w)})/3(1+w) = \text{scalar adiabatic} + \text{inflaton is a collective field}$

the driven "vacuum" accelerates. but differentially? yes, both then & now
we compute it, but we don't really understand it: vacuum tightly coupled to gravity

we know more about early-inflaton dynamics than late-inflaton dynamics!!

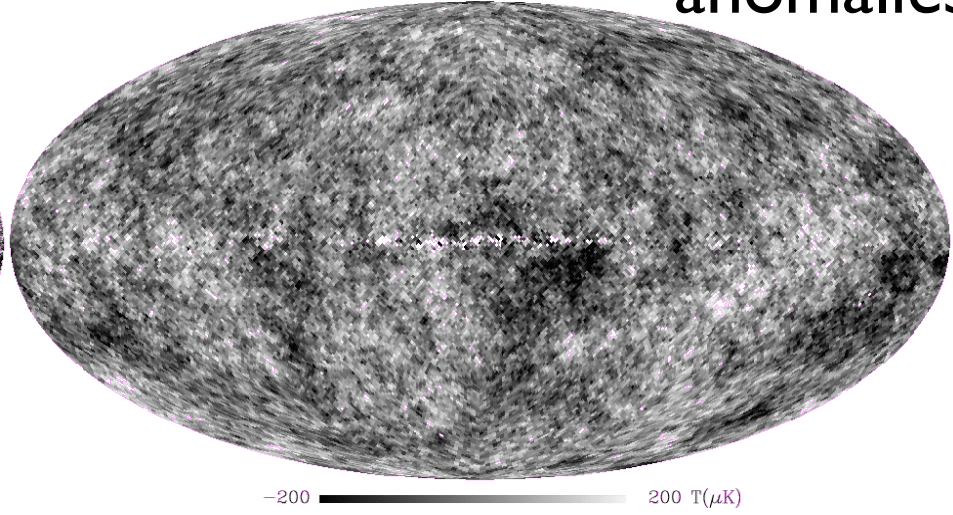
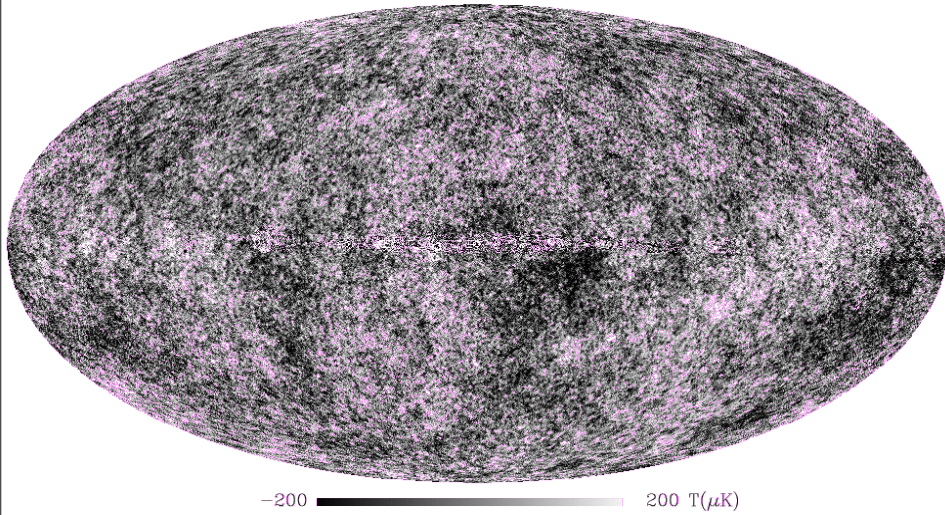
10 e-folds then cf. 1 e-fold now: because resolution (comoving wavenumber k) is
related to dynamics ($H a$) then, but not now

the **quantum fluctuations here & now** are not important for cosmic structure

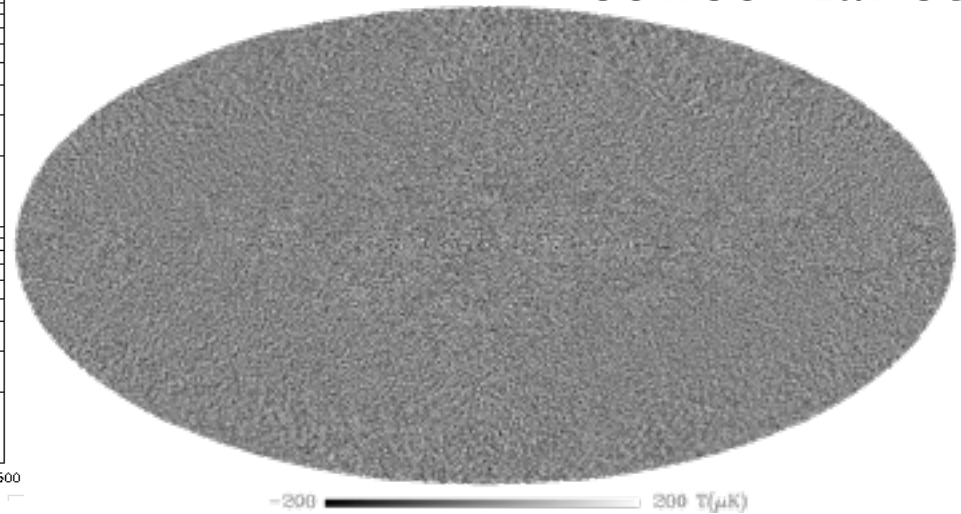
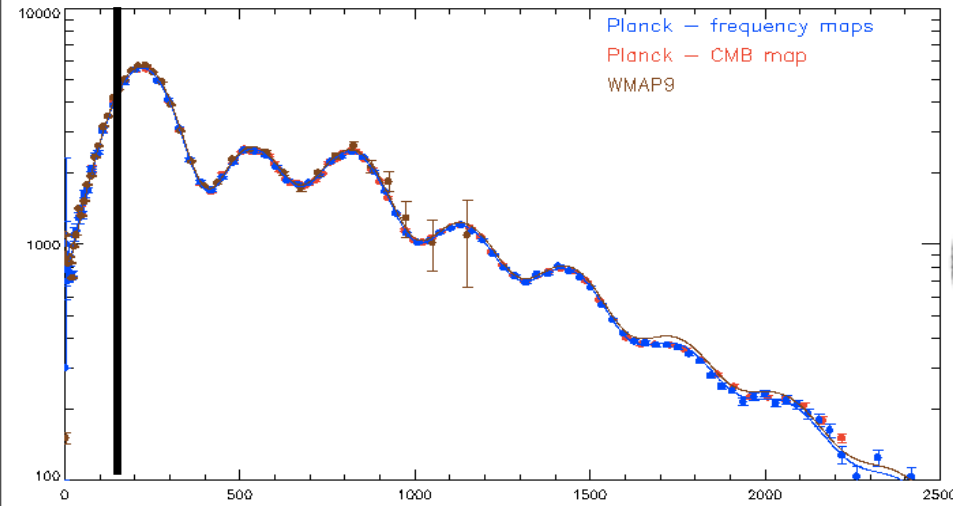
full Planck resolution

Planck smoothed to 1deg fwhm

**L < 134
anomalies**



**L > 134
concordance**



small scale leftover = where most of Planck's information resides > 120X

SIMPLICITY

at $a \sim e^{-7} \sim 1/1100 \Rightarrow$
 at $a \sim e^{-67-60} \sim 1/10^{30+25}$

reveals primordial sound waves in matter

\Rightarrow learn **contents & structure** at 380000 yr, $a \sim e^{-7}$
 \Rightarrow infer the structure far far earlier $a \sim e^{-67-60}$

7⁺ numbers

Early Universe **STRUCTURE**

“red” noise: 2 numbers at $a \sim e^{-67-55}$

$$\ln \text{Power}_s \sim \ln 22.0 \times 10^{-10} \pm 0.025$$

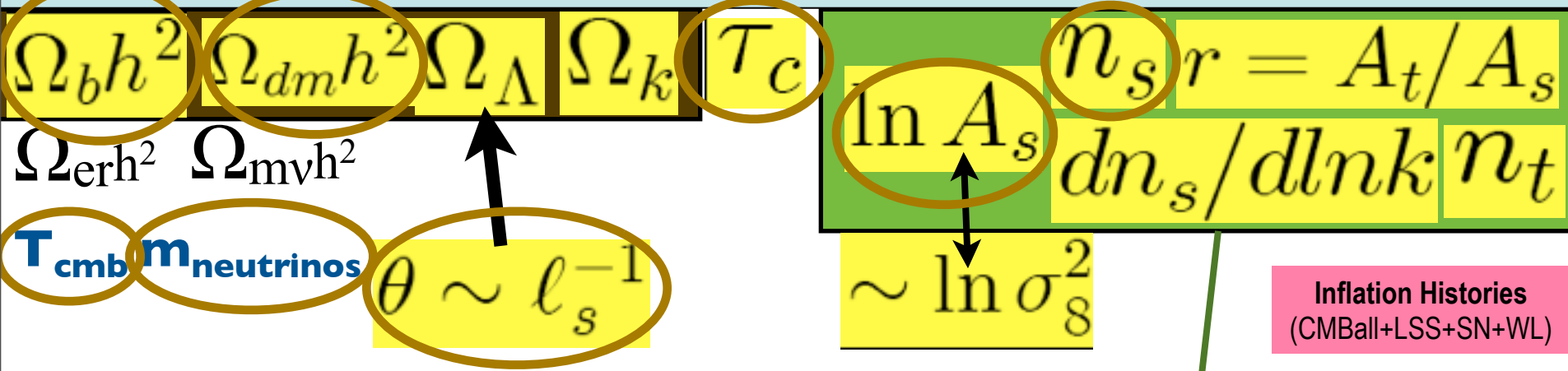
$$n_s = 0.9608 \pm 0.0054 \quad 5\sigma \text{ from } 1$$

$$-0.014 \pm 0.009$$

$$r < 0.12$$

95% CL on *running* $dn_s/d\ln k$, *running of running*, r = Tensor-to-Scalar ratio (GW), *isocurvature modes* for axions (<3.9%), baryons, neutrinos, curvatons (<0.25%)

Standard Parameters of Cosmic Structure Formation



P1.3 like, ACT12 final spectra & params, 1500 sq deg, ~600 for params, *SPT12 2540 sq deg*
Calabrese+13 ACT12+SPT12+WMAP9

standard inflation space: n_s $dn_s/d\ln k$ $r=T/S$ @k-pivots f_{nl}

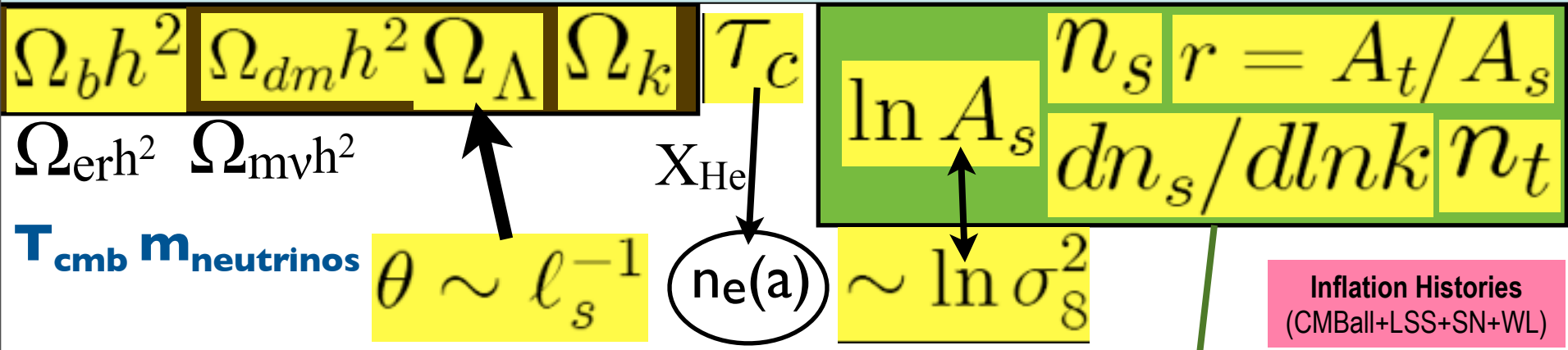
$\ln Power_s \sim \ln 22.0 \times 10^{-10} \pm 0.025$ **P1.3+ $\ln 22 \times 10^{-10} \pm 0.028$ **A12+S12+w9****

$n_s = 0.9608 \pm 0.0054$ (P1.3+WP+hiL+BAO**) 0.9678 ± 0.0088 **A12+S12+w9****
 ± 0.002 (P2.5ext)

$dn_s/d\ln k = -0.014 \pm 0.009$ (P1.3+WP, P1.3+WP+hiL+BAO**) -0.028 ± 0.010 **SPT12+****
 -0.003 ± 0.013 (**ACT12+ WMAP7+BAO+H0**)

$r < 0.12, 0.11, 0.16, 0.11, 0.13$ (95% CL: **P1.3+WP, P1.3+WP+hiL+BAO, A12, S12, W9)**
 $< 0.007-0.013$ (P2.5ext) $f_{nl}: 2.7 \pm 5.8$ local $\Rightarrow \pm 5$ (Pext) $f_{nl}: -42.3 \pm 75.2$ equil -25.3 ± 39.2 ortho

Standard Parameters of Cosmic Structure Formation



new parameters: trajectory probabilities for early-inflatons & late-inflatons (partially) blind cf. informed "theory" priors

P1.3 like, ACT12 final spectra & params, 1500 sq deg, ~600 for params, SPT12 2540 sq deg Calabrese+13 ACT12+SPT12+WMAP9

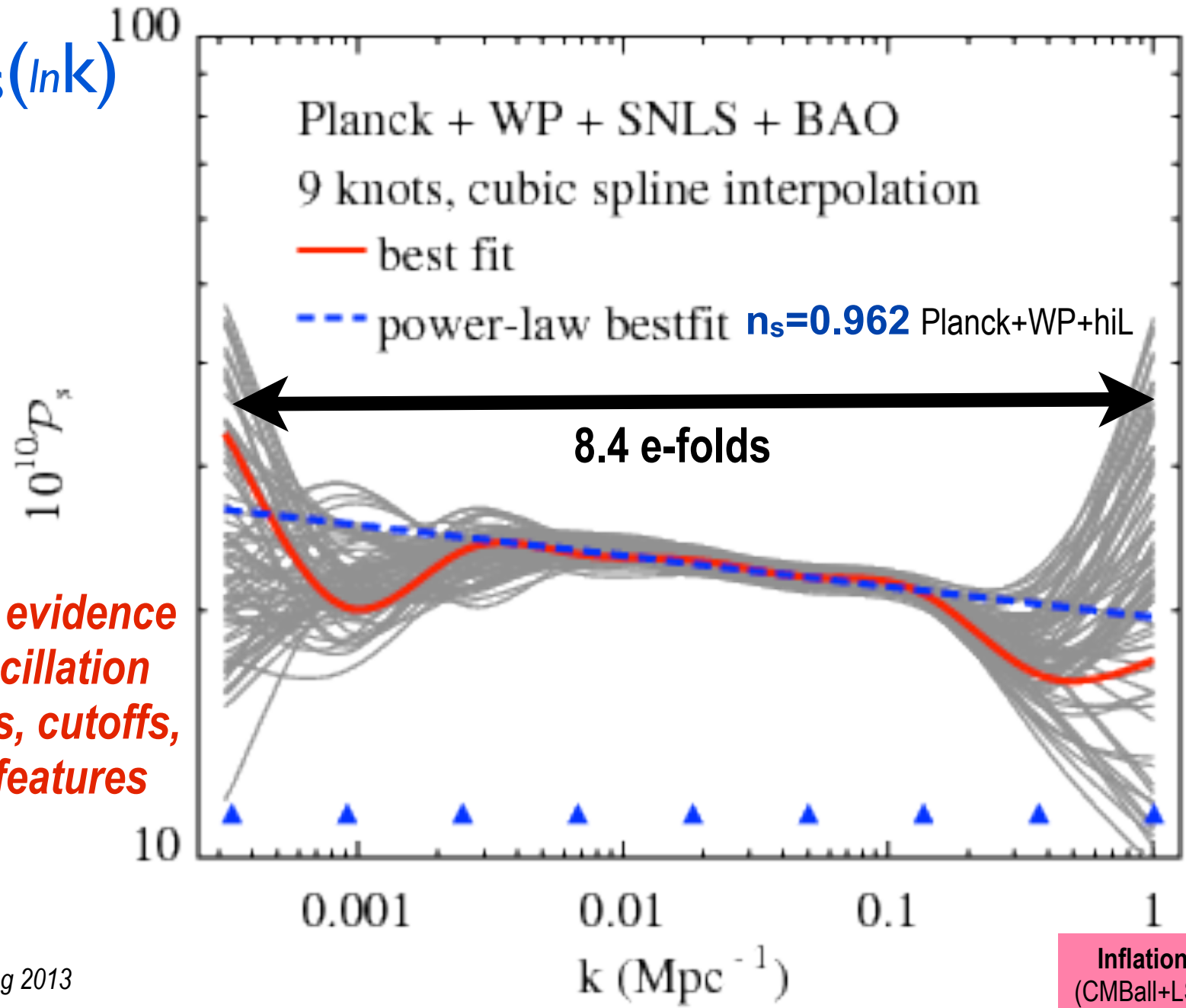
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trajectories: In Primordial power spectra ($\ln k$), $n_s(k)$, $\epsilon(Ha) = 3/2(1+w_t)$, $V(\psi)$, $\psi(a)$

scan $\ln P_s(\ln k)/A_s$, $\ln A_s = \ln P_s(k_{pivot,s})$, $r(k_{pivot,t})$; consistency \Rightarrow reconstruct $\epsilon(\ln H a)$, $V(\psi)$

$\ln P_s(\ln k)$



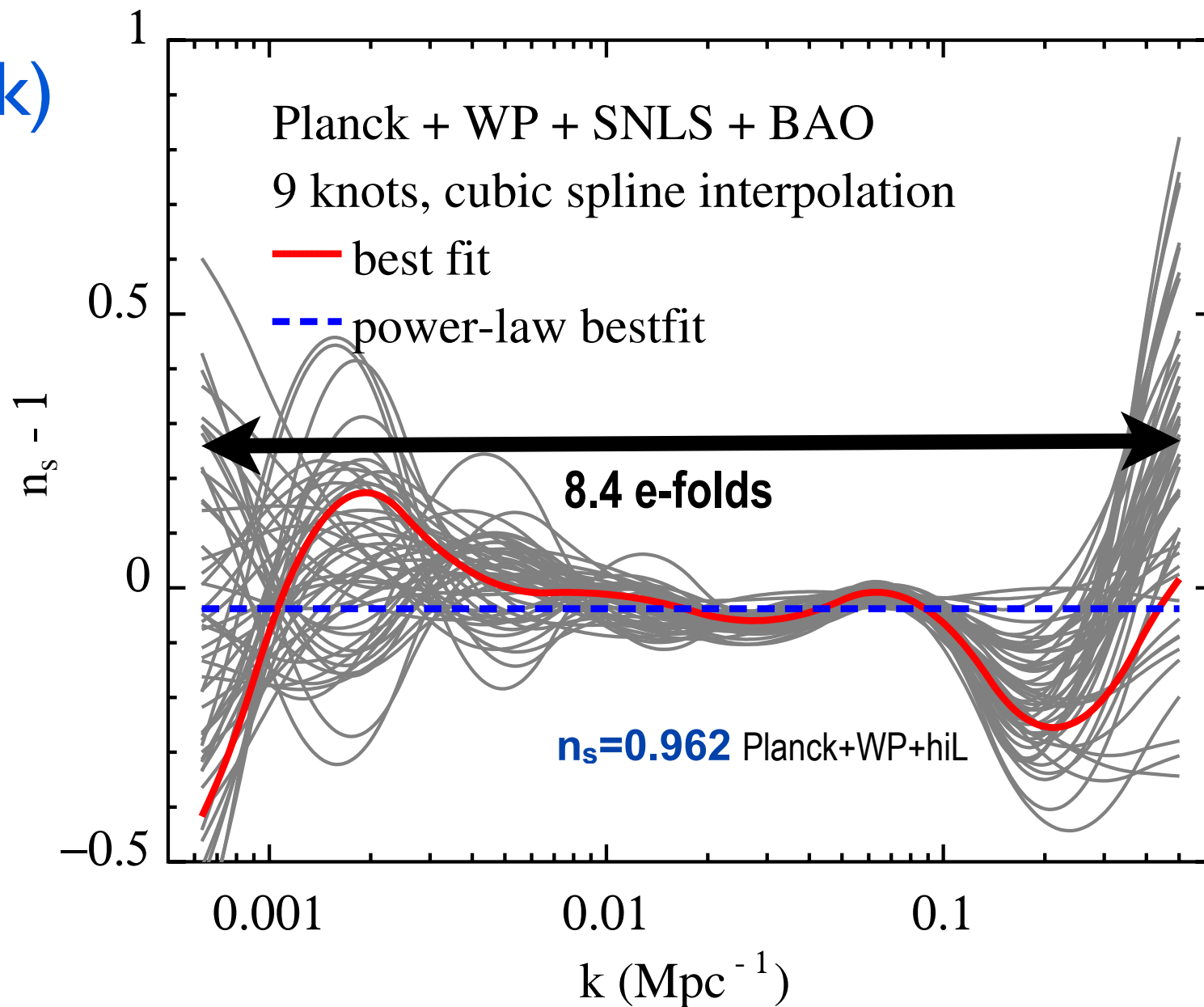
no strong evidence for oscillation patterns, cutoffs, local features

Bond, Huang 2013

Inflation Histories (CMBall+LSS+SN+WL)

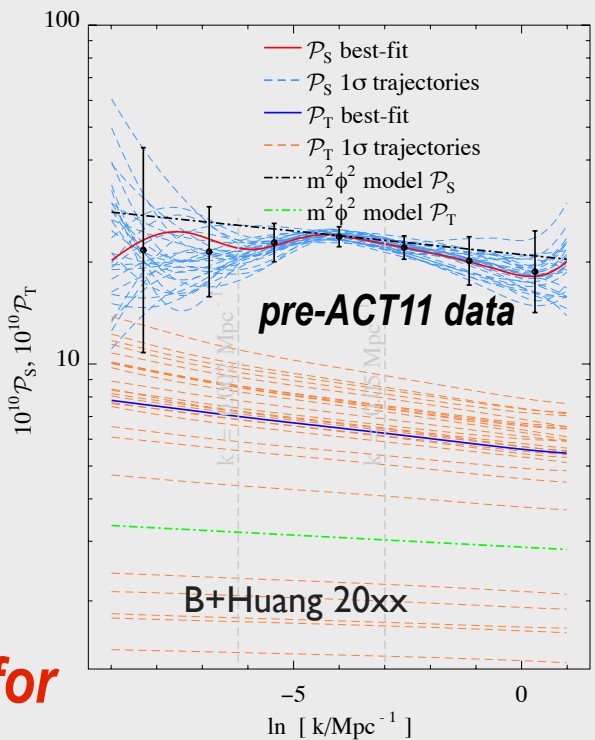
scan $\ln P_s(\ln k)/A_s$, $\ln A_s = \ln P_s(k_{pivot,s})$, $r(k_{pivot,t})$; consistency \Rightarrow reconstruct $\epsilon(\ln H a)$, $V(\psi)$

$n_s(\ln k)$

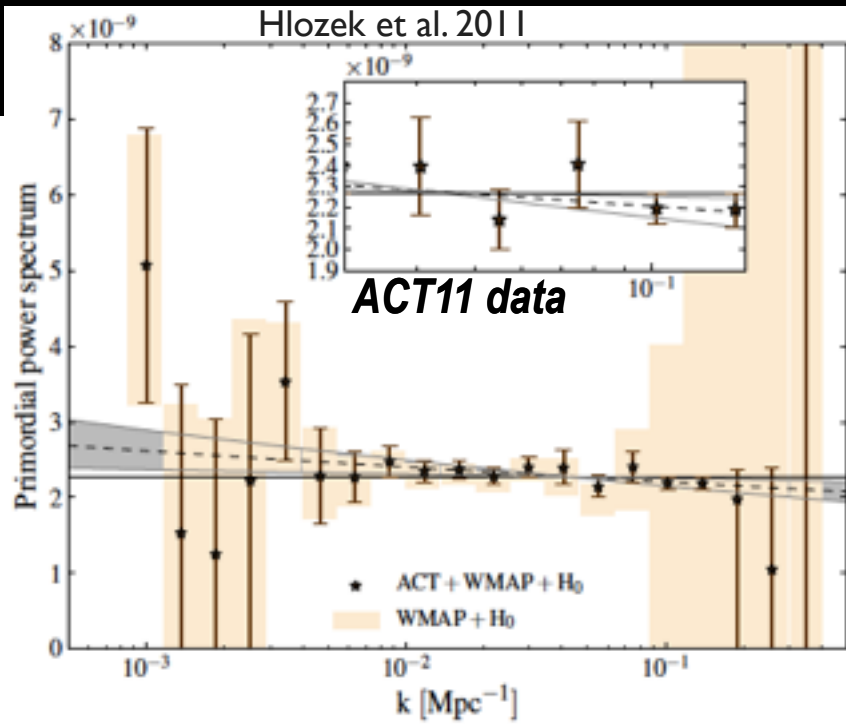


early-U, NOW

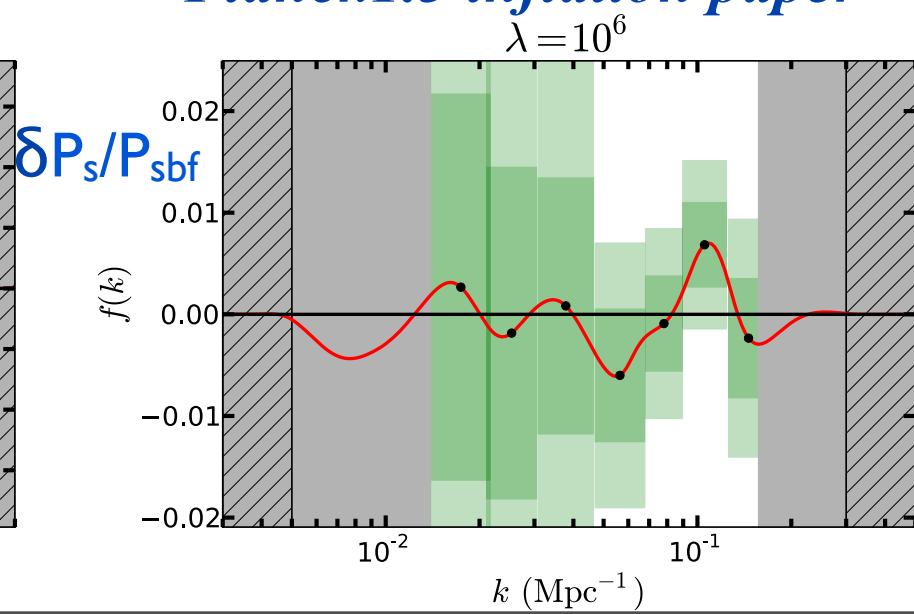
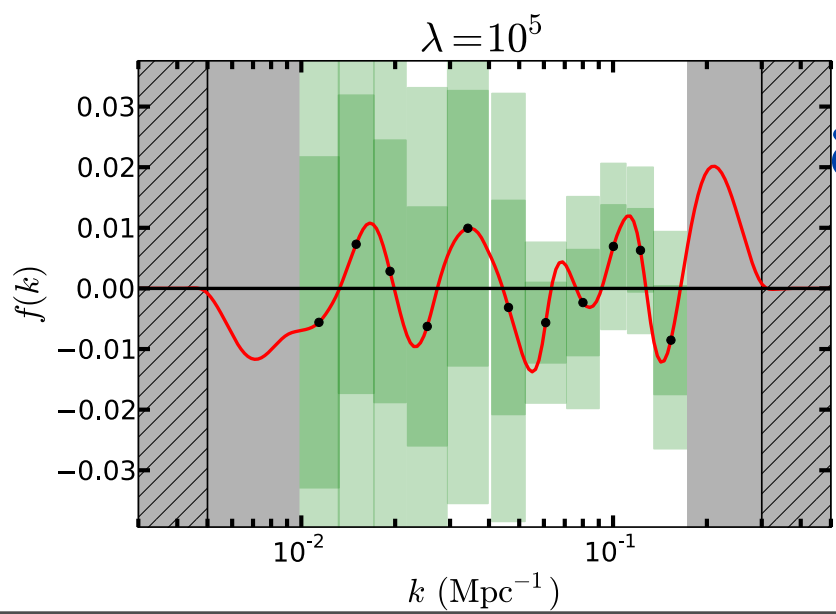
semi-blind & informed reconstruction of Scalar / Tensor power spectra, acceleration histories



no evidence for oscillation patterns



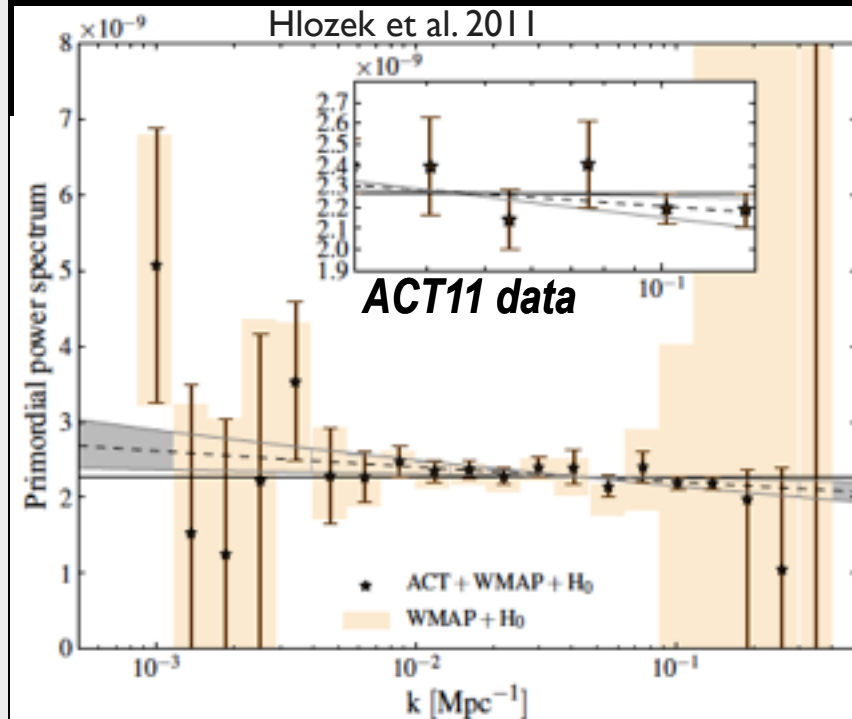
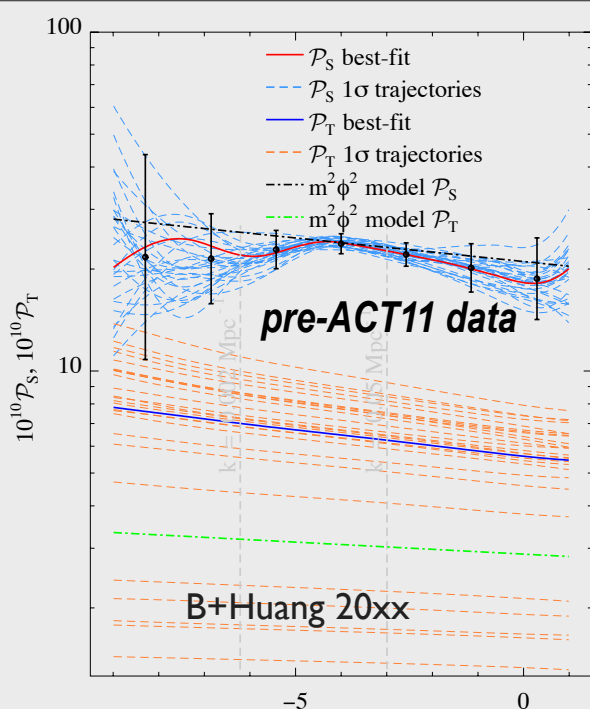
Planck1.3 inflation paper



$\delta P_S / P_{\text{sb}f}$

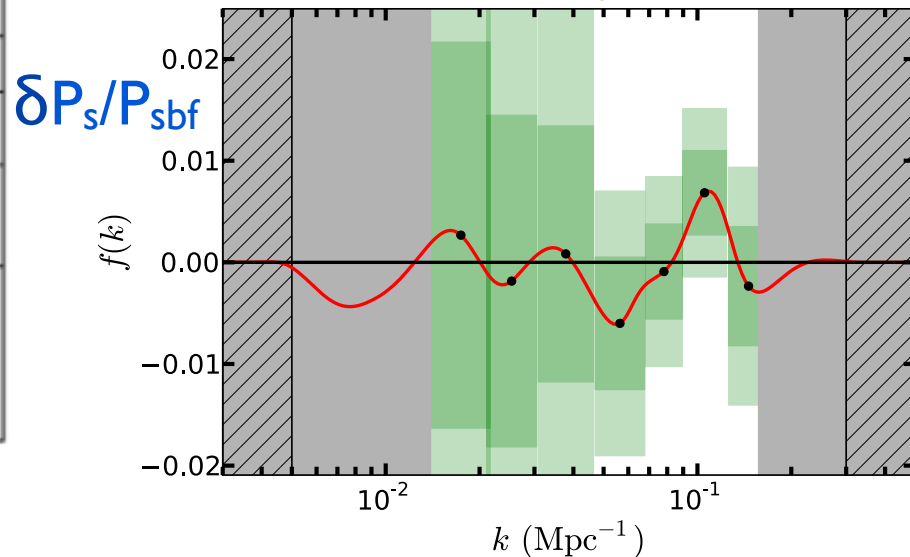
early-U, NOW

semi-blind & informed reconstruction of acceleration histories & S/T power spectra

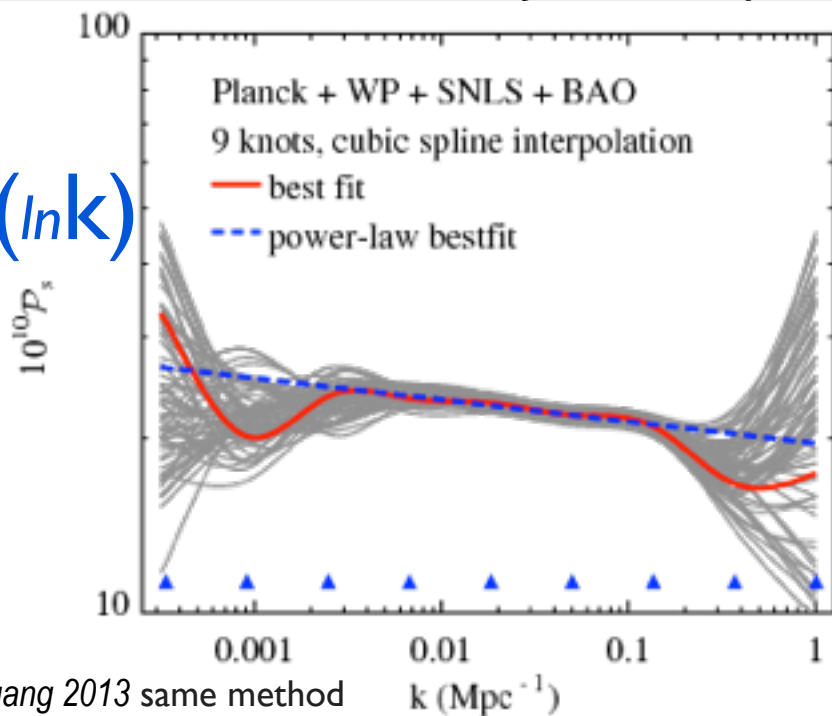


Planck1.3 inflation paper

$$\lambda = 10^6$$

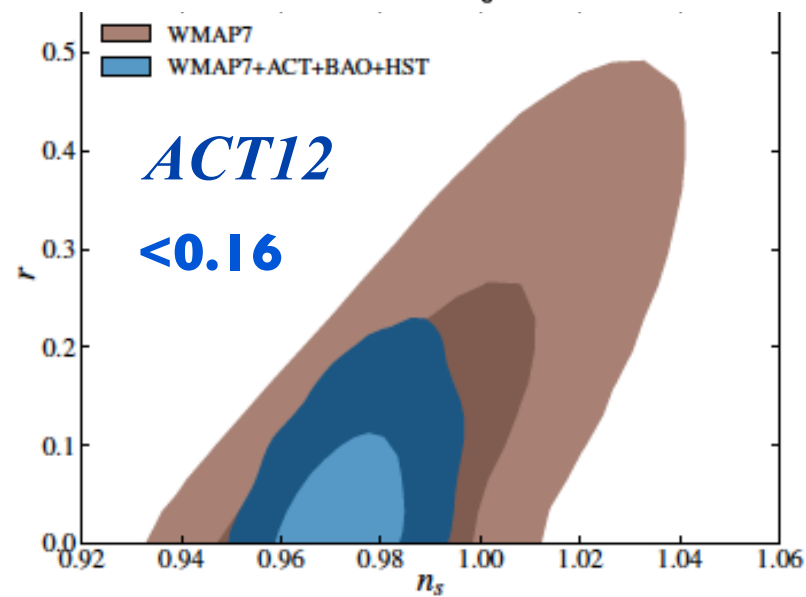
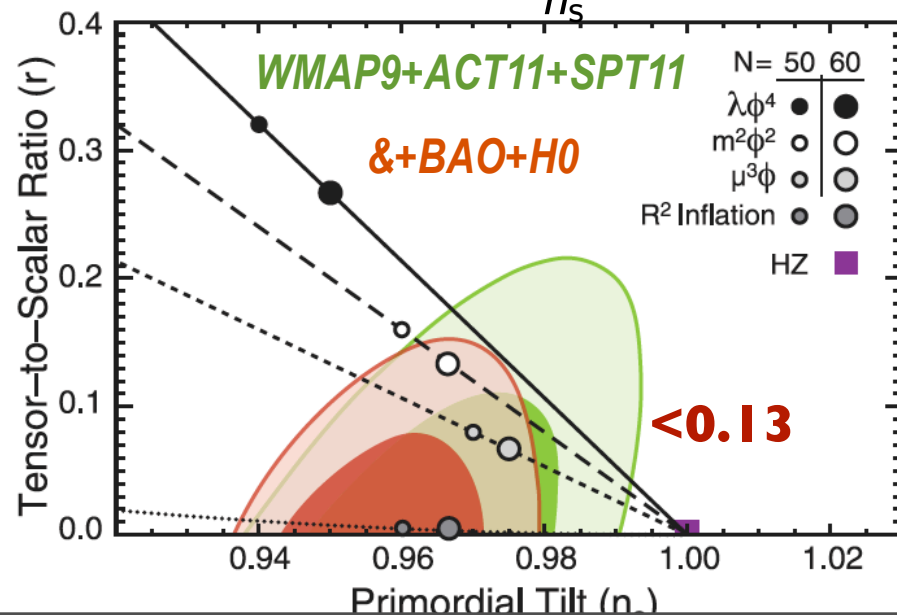
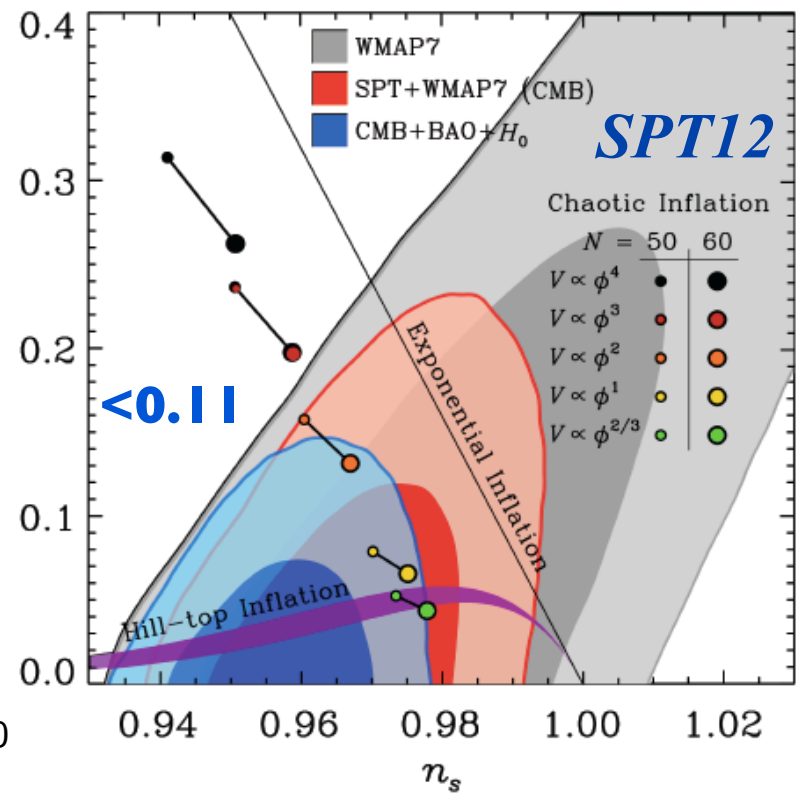
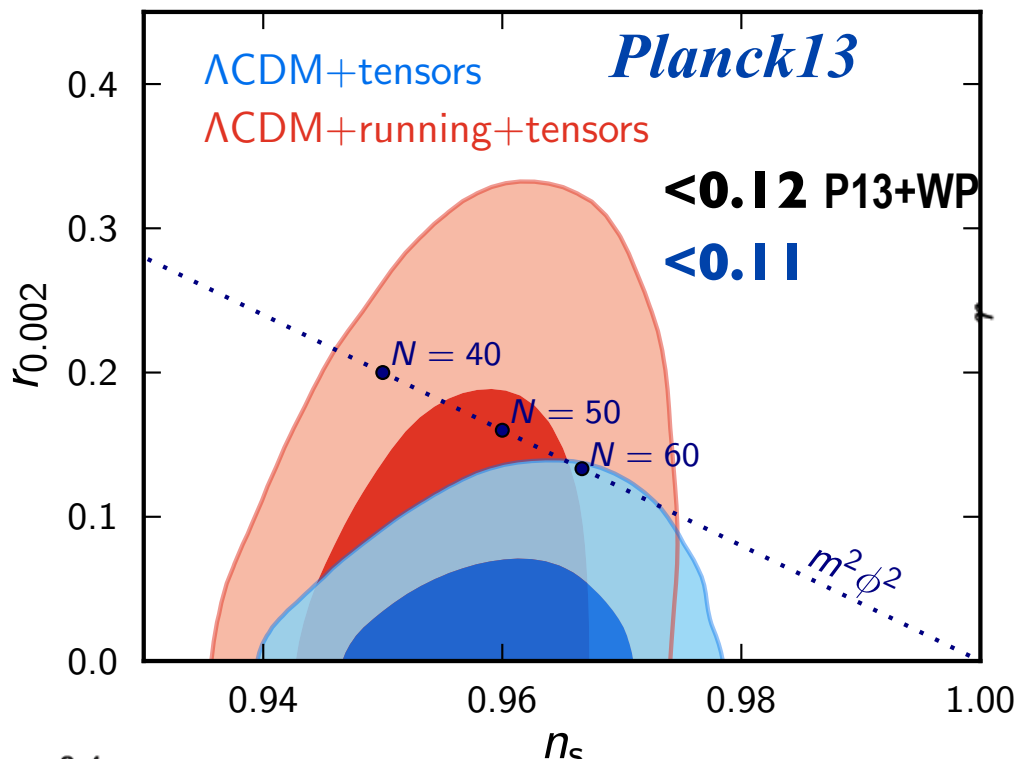


$\ln \mathcal{P}_s(\ln k)$



Bond, Huang 2013 same method

early-U Gravity Wave Constraints

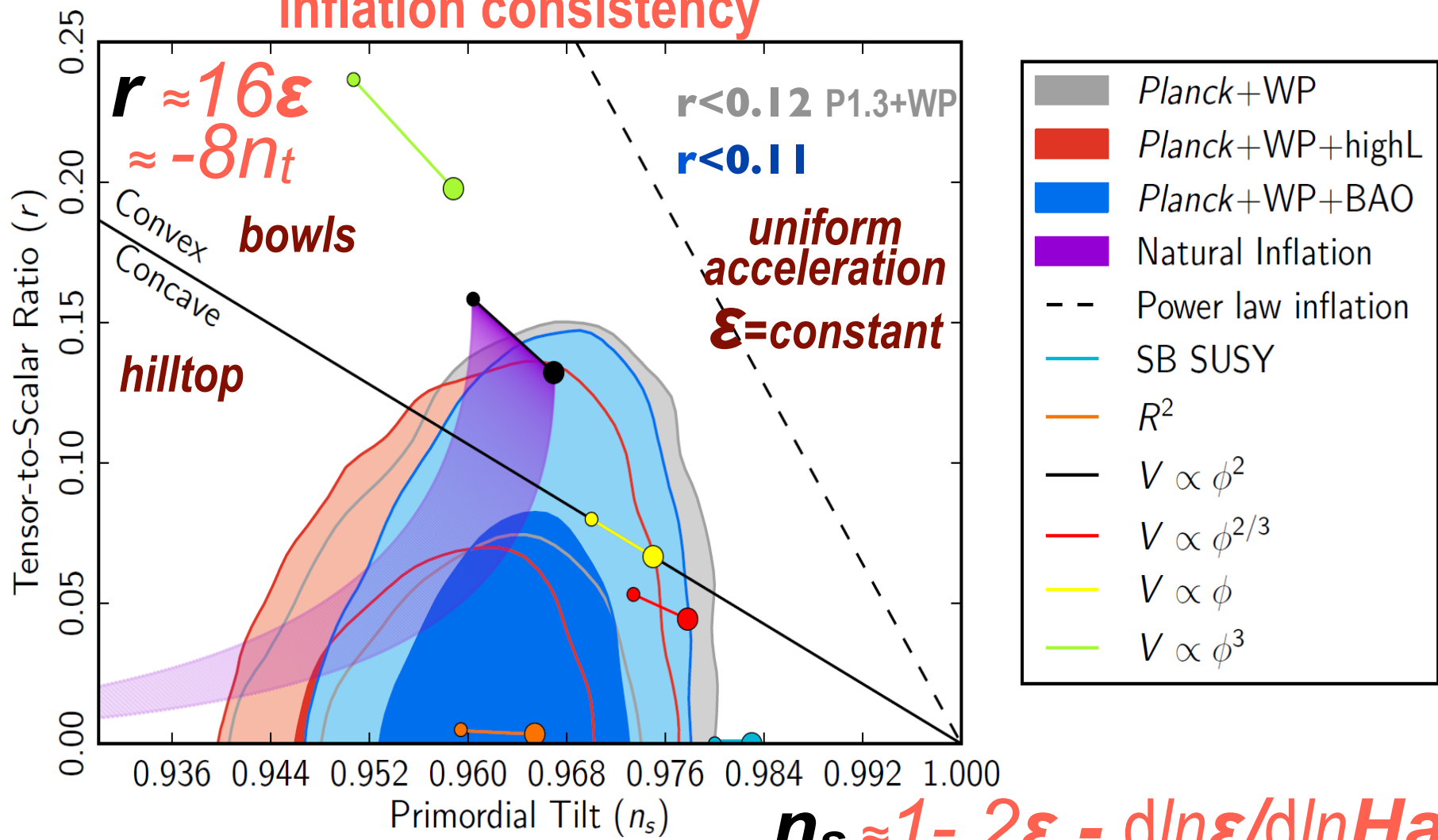


Consistent with single field slow roll, standard kinetic term & vacuum (with f_{NL} upper limits)

uniform acceleration line $\epsilon \equiv 3KE / (KE+PE) = \text{constant}$ is strongly ruled out

\Rightarrow early universe acceleration must change over observable scales (as well as to end inflation)

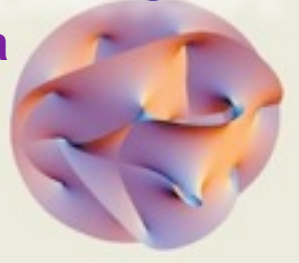
inflation consistency



r without B-mode p is delicate rule out: exponential potential models (power-law inf), the simplest hybrid inflationary models (Spontaneously Broken SUSY) & Φ^n , $n > 2$ monomial potentials of chaotic inflation some popular inflation survivors: Natural = pNGB, monodromy = driven pNGB, Roulette (shrinking holes in extra-dim), brane (separation), Higgs, flattened potentials = non-monomial, ...

Old view: Theory prior = delta function of THE correct one and only theory
New: Theory prior = probability distribution of late-flows on an *action* LANDSCAPE

6/7 tiny extra dimensions



1980

R^2 -inflation

Old Inflation

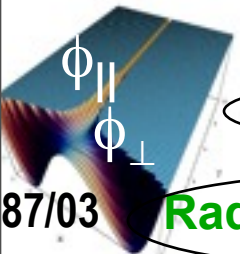
Chaotic inflation

New Inflation

Double Inflation

Power-law inflation

SUGRA inflation



87/03

Radical BSI inflation

running (nee variable M_P) inflation

Extended inflation

1990

Natural pMGB inflation

Hybrid inflation

Hybrid inflation

Hybrid inflation

Hybrid inflation

KLS94 preheating

SUSY F-term inflation

SUSY D-term inflation

Assisted inflation

Brane inflation



SUSY P-term inflation

Super-natural Inflation

K-flaton

2000

2003 KKL

N-flaton

ekpyrotic/cyclic

D3,D7 brane inflation

DBI inflation

moving brane separations

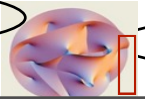
Racetrack inflation

Tachyon inflation

Warped Brane inflation

moduli fields

monodromy
Higgs inflation



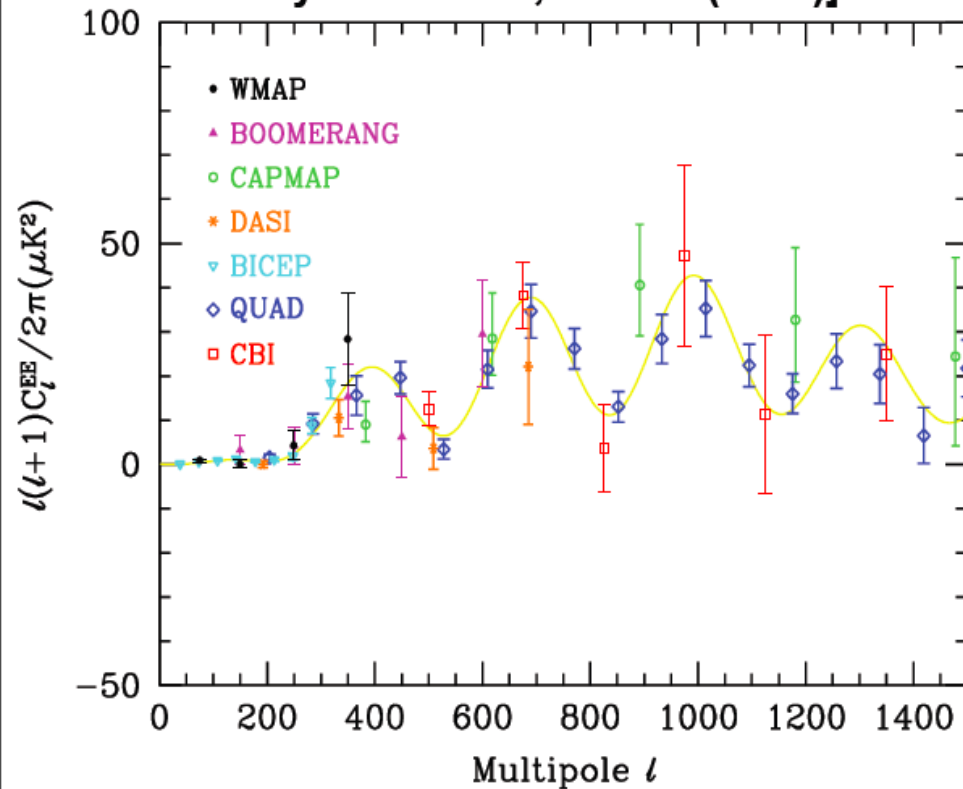
Roulette inflation Kahler moduli/axion

fibre inflation

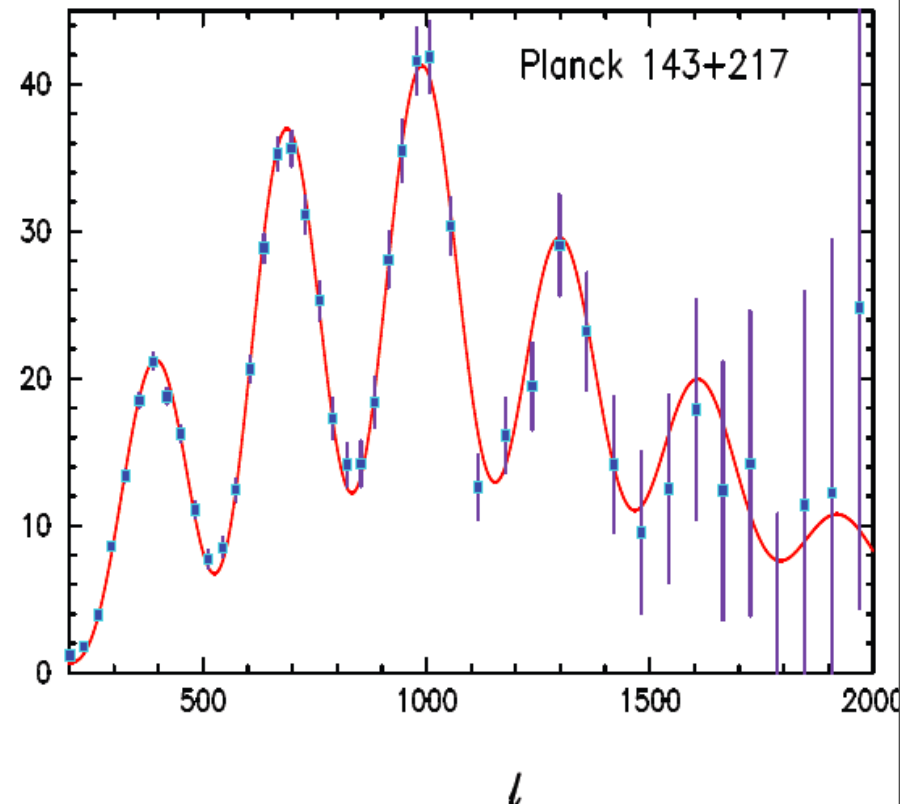
**best-fit P1.3yr TT model predicts the polarization. works perfectly at all frequency cross correlations
strengthens the case for the Galactic/extragalactic nuisance parameter model being accurate
teaser for 2014**

EE polarization

[J. Beringer et al. (Particle Data Group),
Phys. Rev. D86, 010001 (2012)]



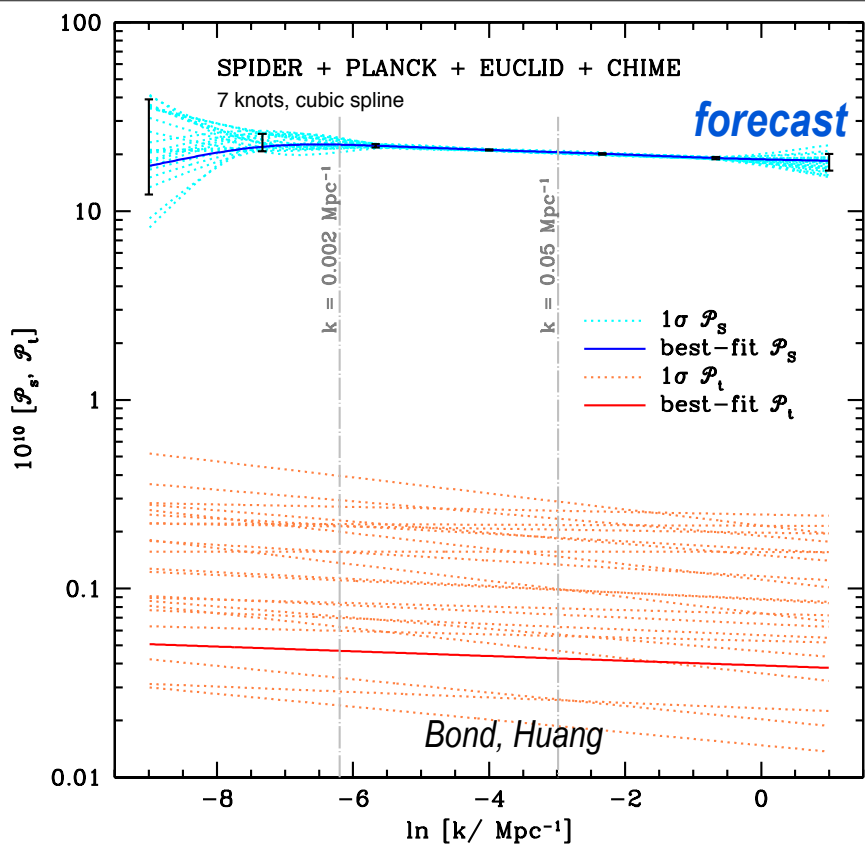
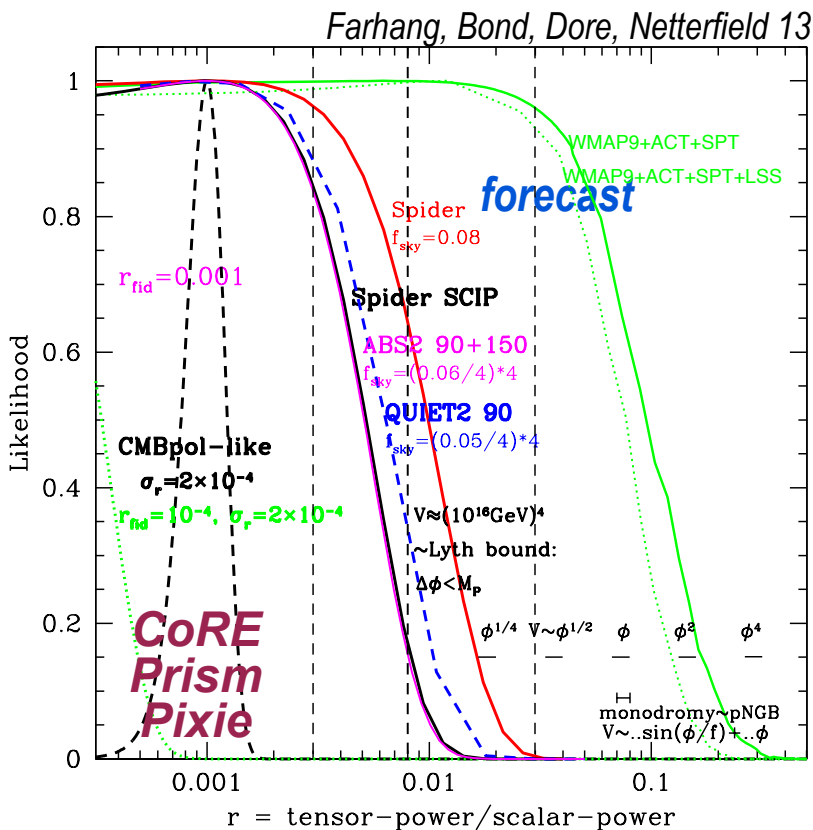
[Planck 2013 results. XVI.
Cosmological parameters]



a long path to constrain the B-mode of polarization at the $r = .02$ to $.05$ level of P2.5 forecasts

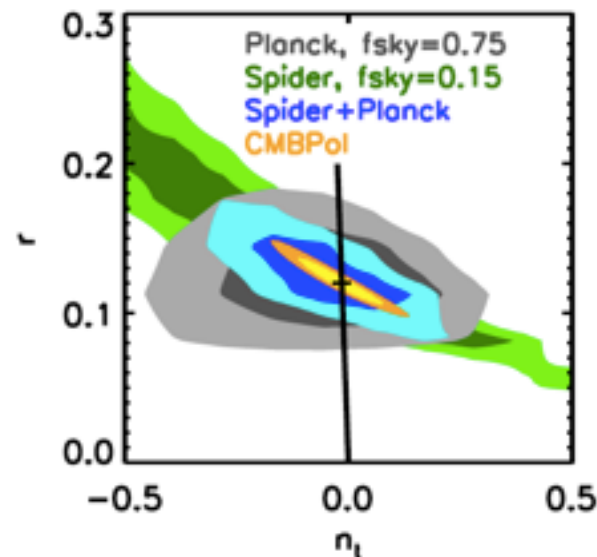
Spider24days+Planck2.5yr: r-n_t matrix-forecast for r=0.12 input for m²φ² (2σ_r ~0.02 including fgnds)

similar r-forecasts for ABS+, Quiet2, Keck, ..



can get B-mode shapes but without the precision needed to check

-n_t ≈ r/8 consistency



early-inflaton DE acceleration trajectories then

Bond, Huang 2013

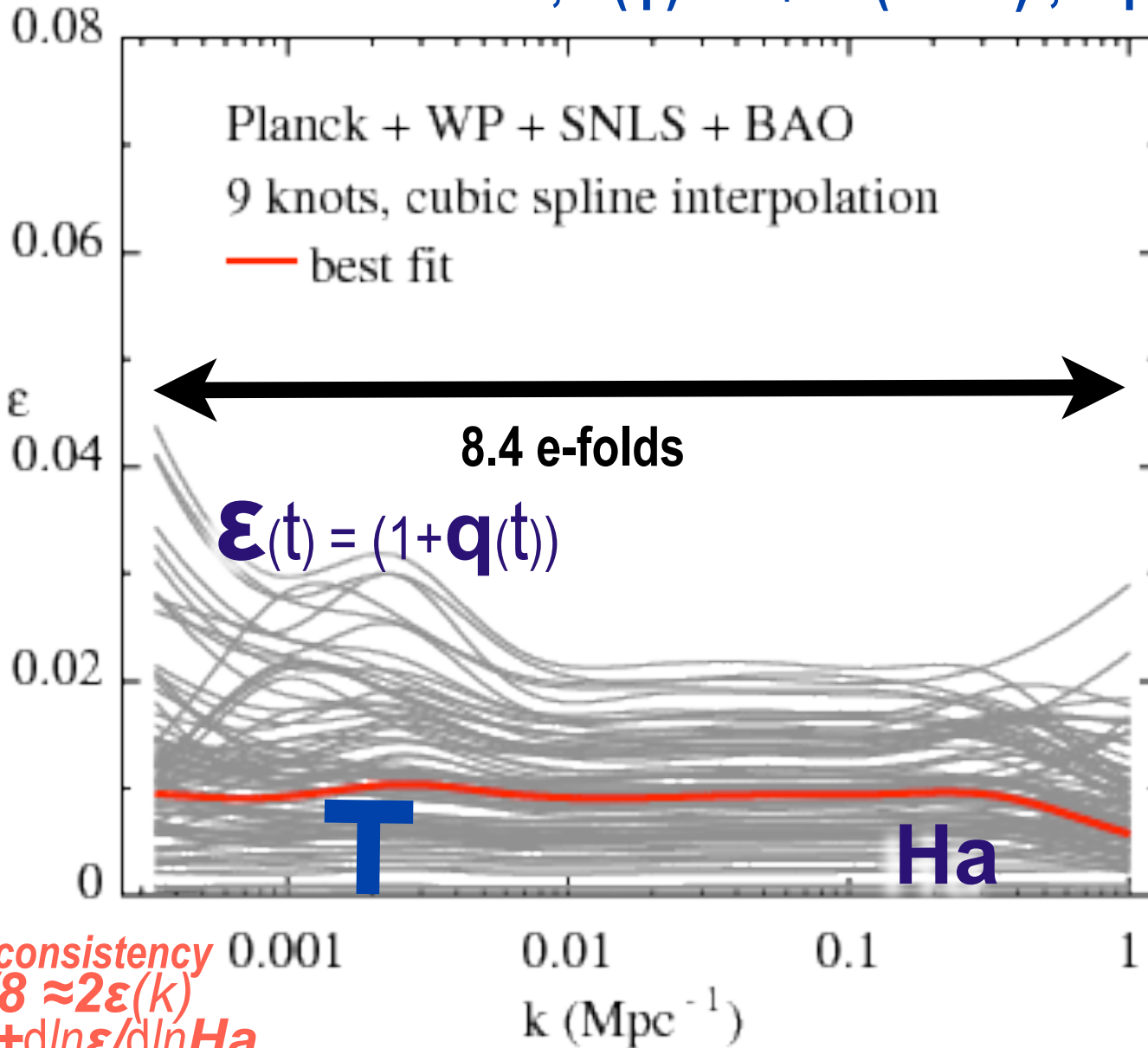
$$\epsilon = -d \ln H / d \ln a ; V(\psi) \approx 3M_P^2 H^2 (1 - \epsilon/3) ; d\psi / d \ln a = \pm \sqrt{\epsilon}$$

aka

$$(1+W_{de})^{3/2}$$

then

(hydro)



resolution
 $\ln k \sim \ln H a$
dynamics

$$\epsilon \approx r / 16$$

$$\epsilon \approx V$$

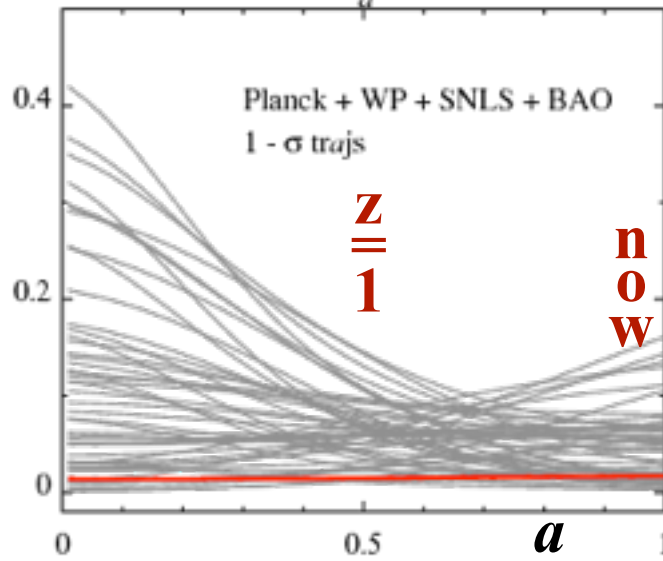
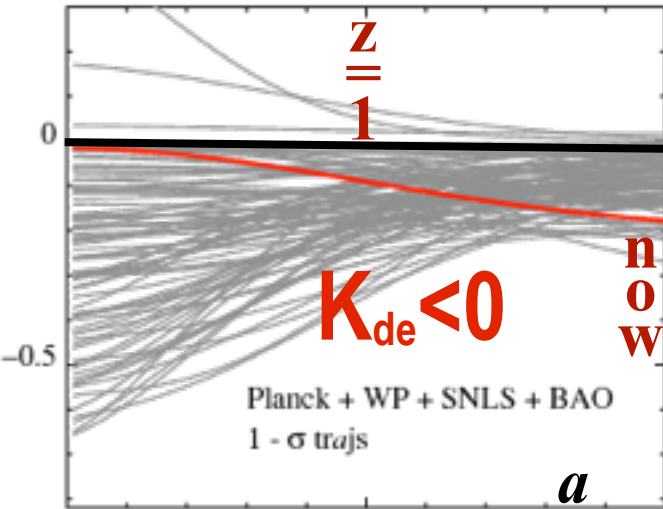
$$0.0005 (10^{16} \text{Gev})^4$$

inflation consistency
 $-n_t \approx r/8 \approx 2\epsilon(k)$
 $1-n_s \approx 2\epsilon + d \ln \epsilon / d \ln H a$

late-inflaton DE trajectories

informed = 3-parameter $W_{de}(a|\epsilon_s \epsilon_{de\infty} \zeta_s)$

Bond, Huang 2013



**$1+W_{de,0}$
= -0.13 ± 0.12**

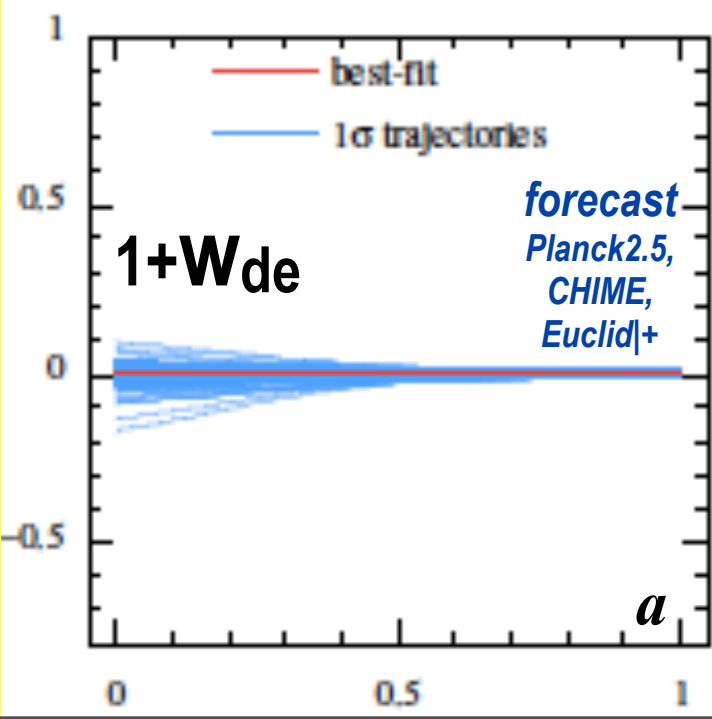
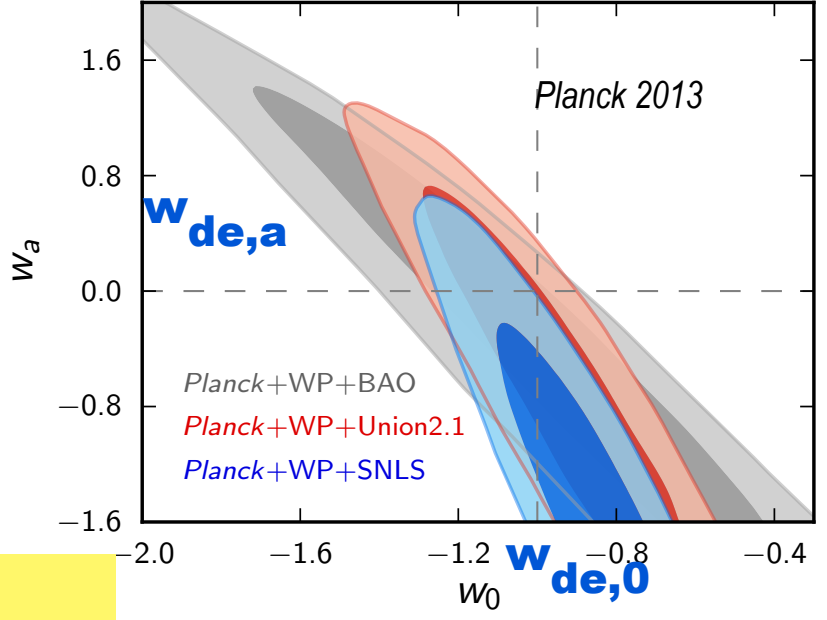
if $w_{de,a}$
 $V_{de}, \epsilon_{de\infty}$
 $\epsilon_s = (d \ln V / d \psi)^2 / 4$
@pivot a_{eq}

= $-0.25 + .20 - .26$

$P1.3+SNLS3 = 0.0 + .21$

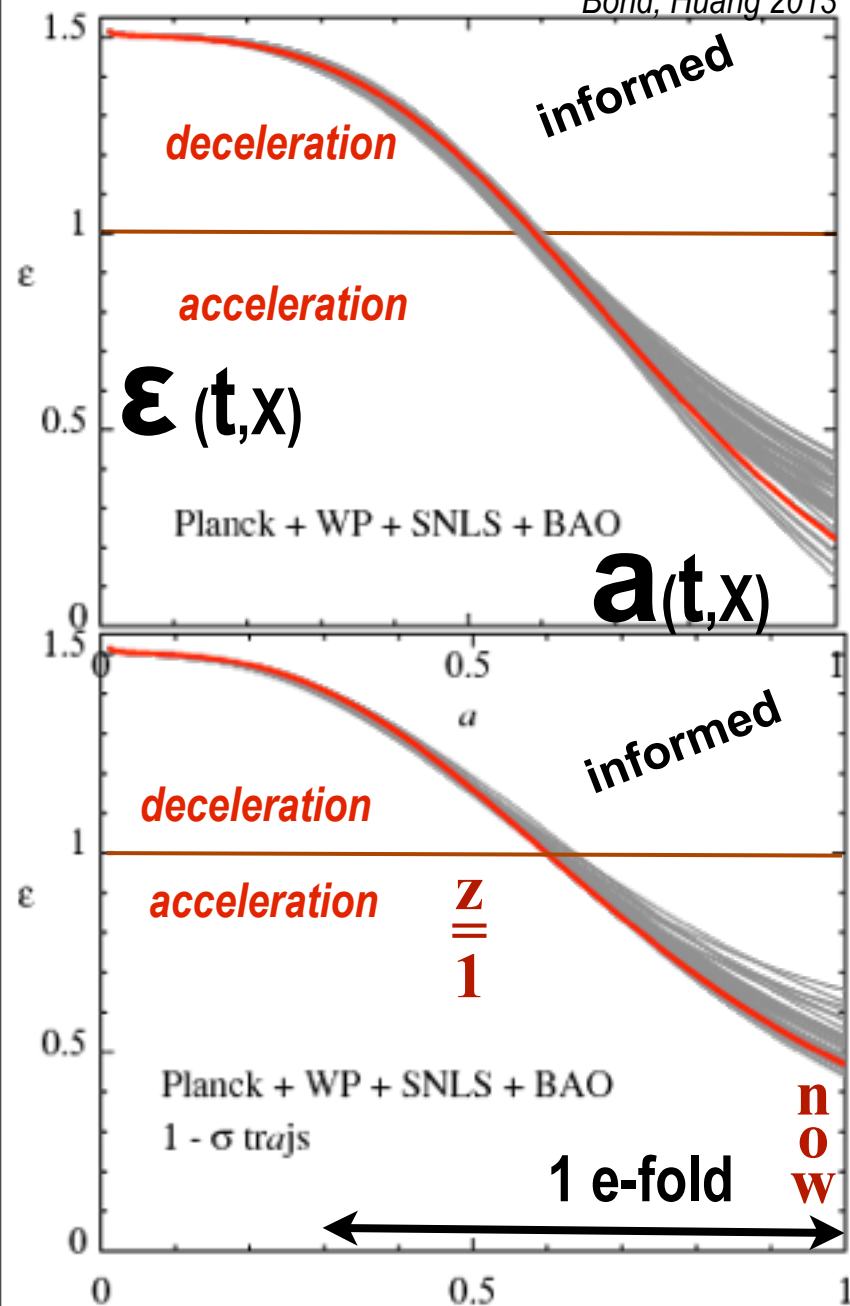
future $.005 + .031 - .025$

$$1+W_{de} = -d \ln p_{de} / d \ln a^3$$

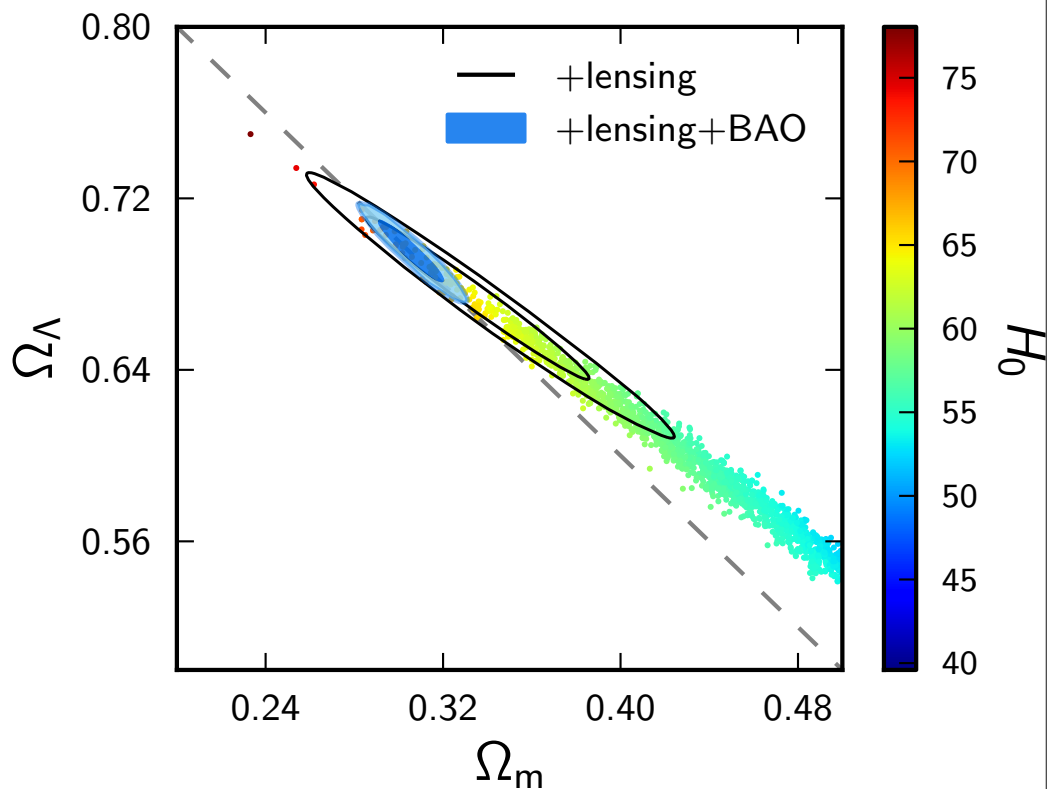


late-inflaton DE trajectories

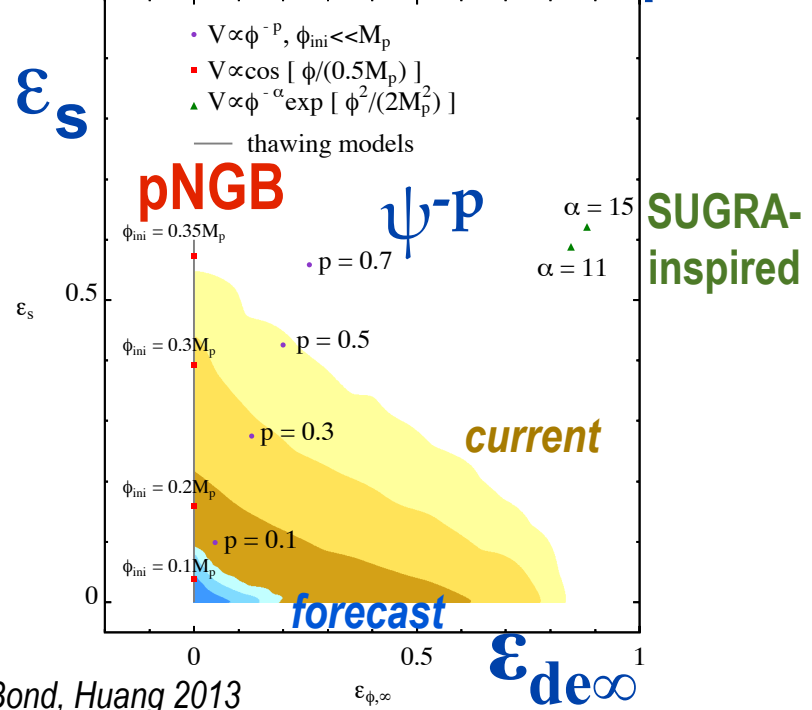
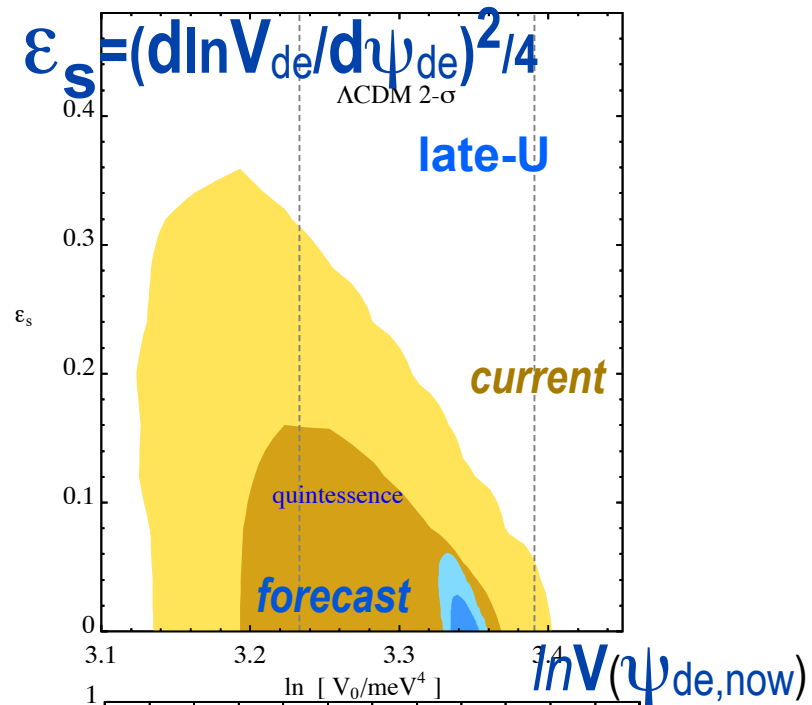
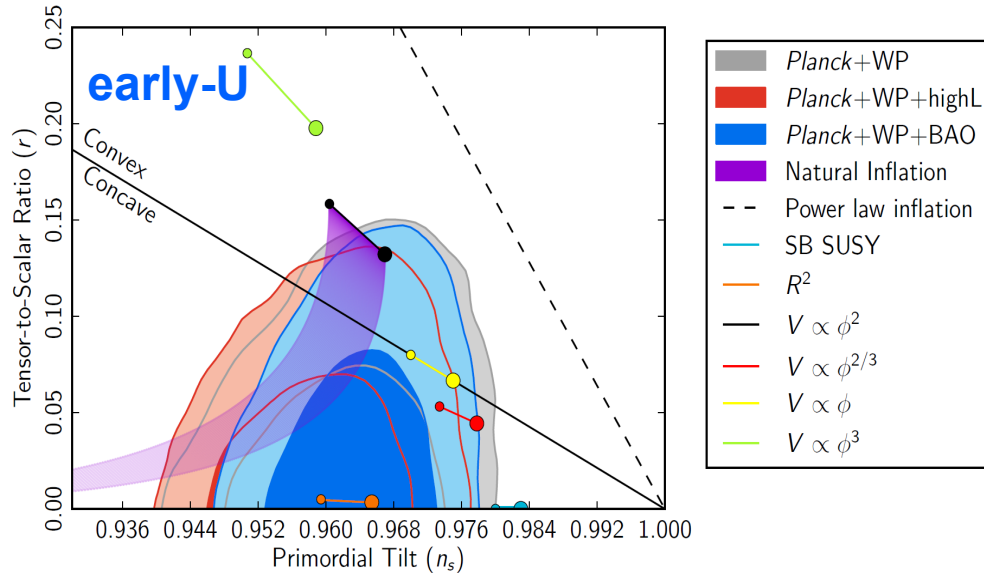
Bond, Huang 2013



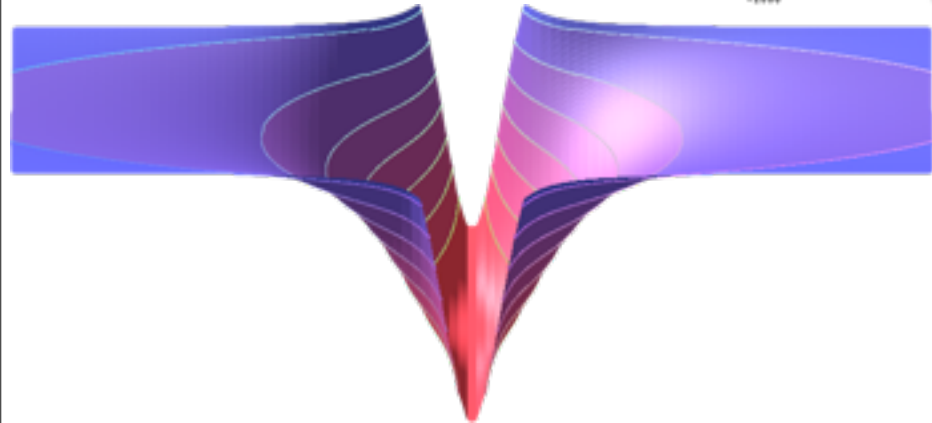
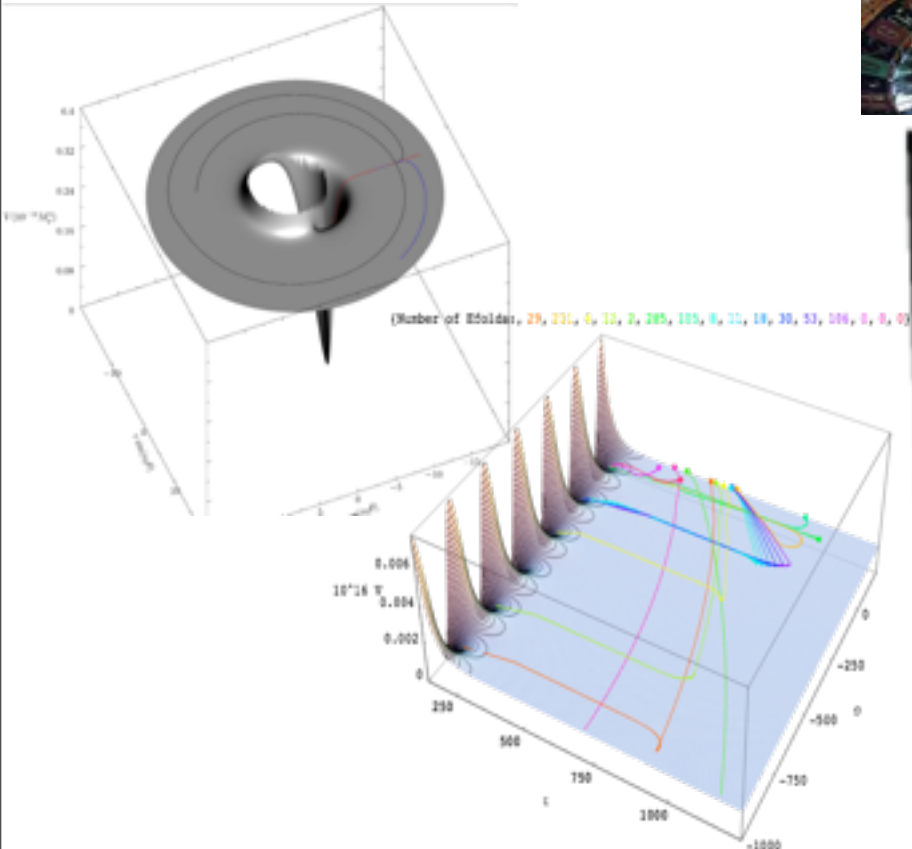
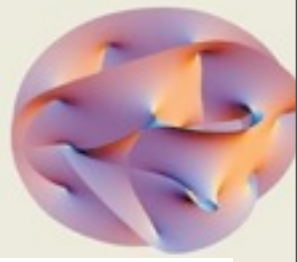
lensing breaks geometrical degeneracy:
 Planck alone gives dark energy cf. Planck+BAO



introduce a late-U DE plot littered with theory models similar to the early-U r - n_s plot. with HBK10/BH11 parameterization of the DE trajectories this can be done.



non-Gaussianity



Preheating After
Roulette Inflation

$$\langle \tau \rangle =$$

quantum
diffusion
spatial jitter

drift

$$\ln a(\mathbf{x}, \ln H)$$

entropy
generation in
preheating
from the
coherent
inflaton
(origin of all
matter)

let there be
heat

isocon directions,
e.g., axion

SEMILINEAR INFLATION

primordial nonGaussianity

nonG 3-point-correlation-pattern measure

f_{nl} : 2.7 ± 5.8 local for Newton potential *cf.* ± 5 (Pext)

$\Rightarrow f_{NL*} = 0.44 \pm 3.5$ for phonons/3-curvature

$-f_{nl}$: 42.3 ± 75.2 equil

-25.3 ± 39.2 ortho

$$\zeta_{NL}(x) = \zeta_G(x) + f_{NL*} (\zeta_G^2(x) - \langle \zeta_G^2 \rangle)$$

local smooth.
use optimal pattern estimators

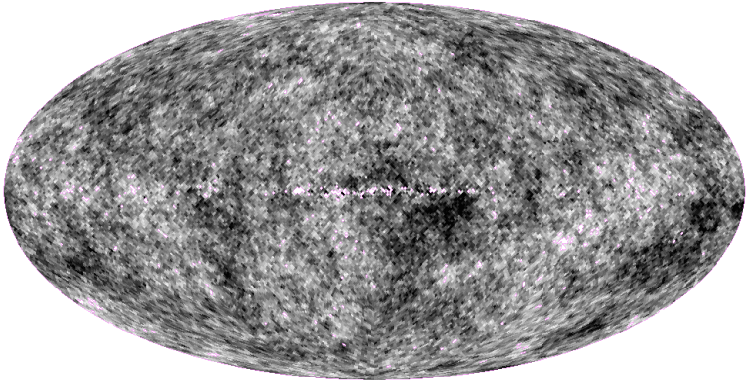
cf. DBI inflation: non-quadratic kinetic energy

$\zeta_{NL}(x)$ = equilateral pattern & orthogonal pattern

scale (k) dependent patterns: connecting to power spectrum broken scale invariance. hint?

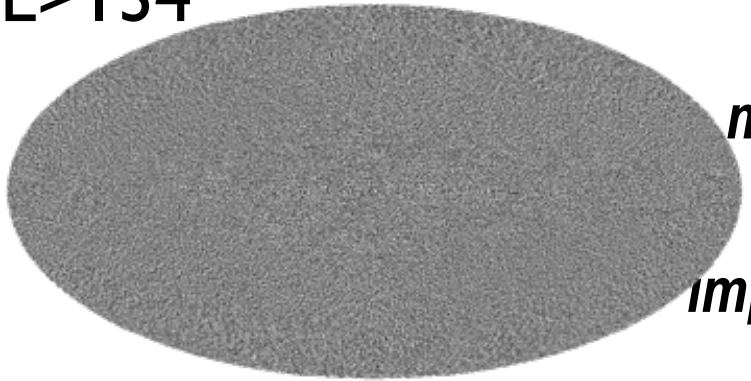
cosmic/fundamental strings/defects

from end-of-inflation & preheating chaos



$L > 134$

-200 200 $T(\mu K)$



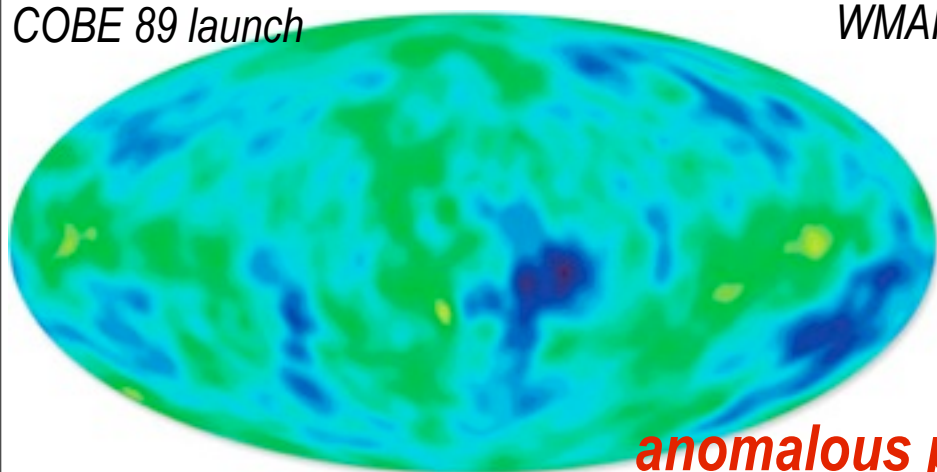
-200 200 $T(\mu K)$

most nonG info from high L: why Planck improved so much over WMAP9

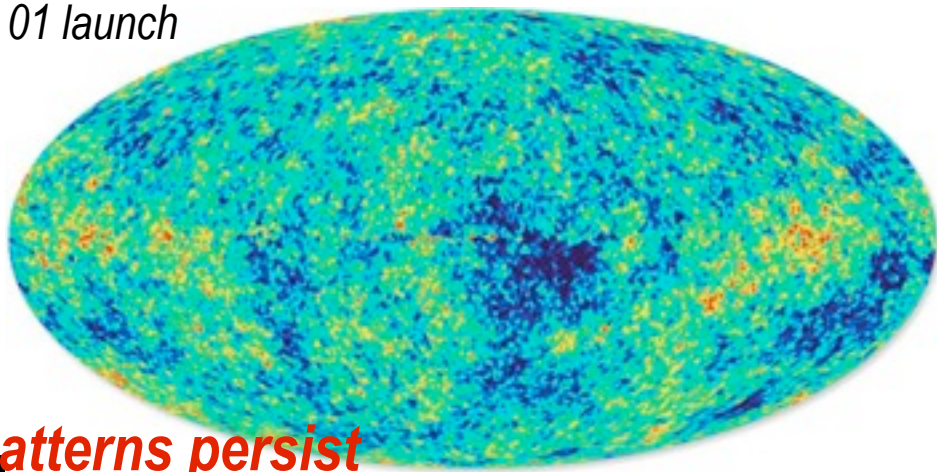
\downarrow $F_{NL}(\chi_b(x))$
intermittent CMB power bursts from super-bias of a GRF modulating field landscape scan

phonon $\sim \zeta_{NL} = \ln(\rho a^{3(1+w)}) / 3(1+w) \Rightarrow f_{NL*} = 3/5 f_{NL} - 1$

COBE 89 launch

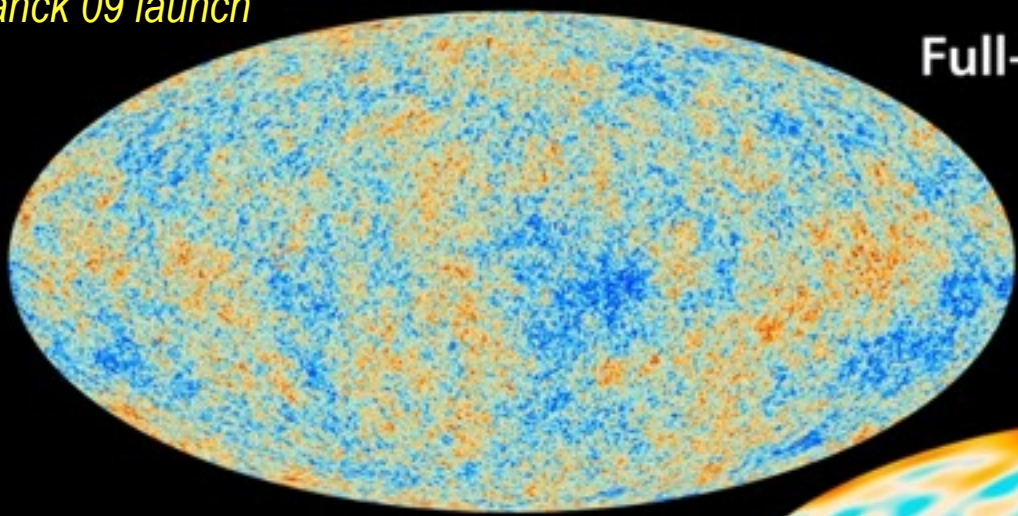


WMAP 01 launch



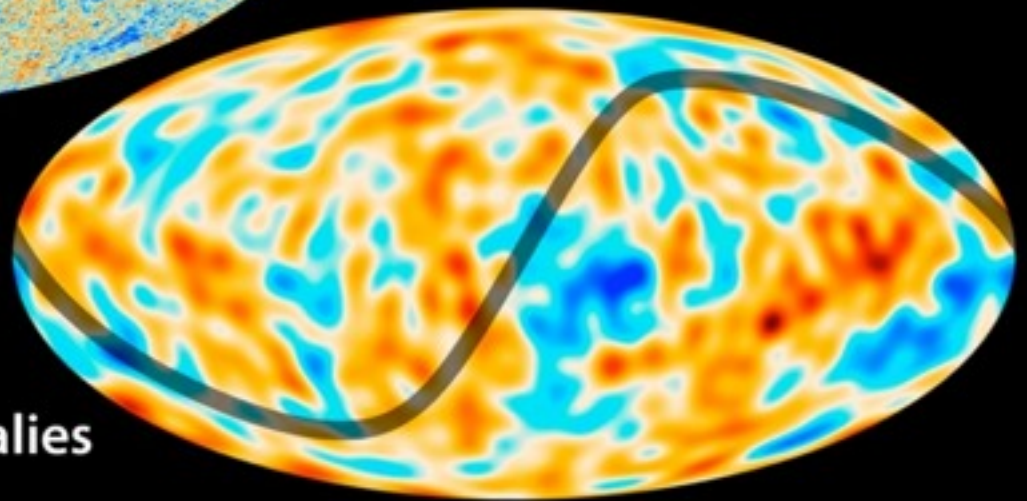
anomalous patterns persist

Planck 09 launch



Full-Sky Map

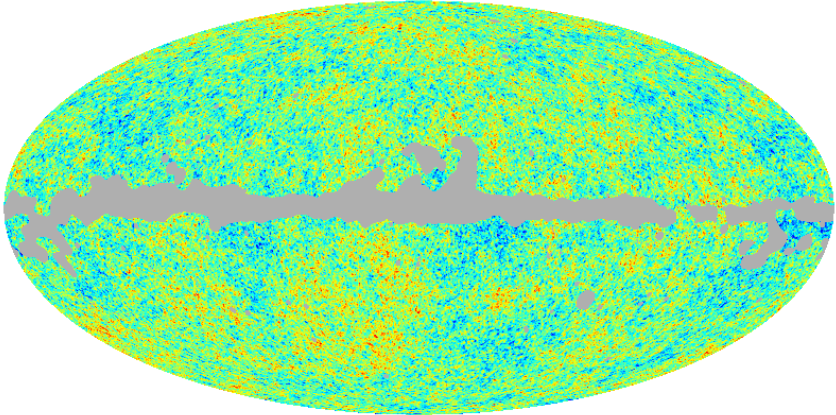
NonGaussian 3-point-pattern measure
 $f_{NL}: 2.7 \pm 5.8$ local $\Rightarrow \pm 5$ (Pext)
 $-f_{NL}: 42.3 \pm 75.2$ equil
 -25.3 ± 39.2 ortho & f_{NL}^{eff}



a homogeneous, anisotropic Bianchi VII_h model

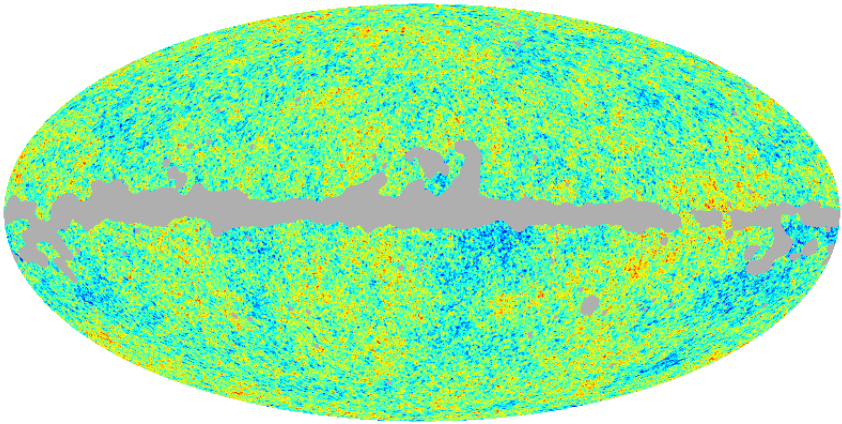
Anomalies

Fluctuation CMB Sky



-500. +500.

real CMB Sky

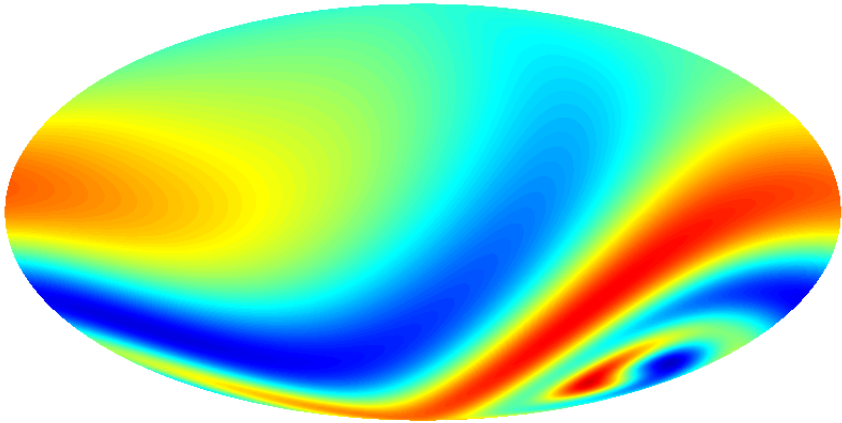


-500. +500.

=

mean field CMB Sky

+



-60.0 +60.0

homogeneous, anisotropic Bianchi VII_h model: ultralarge scale rotation/vorticity and shear, fit parameters violently disagree with Planck13 parameters. but maybe there is a grand unified theory of anomalies, as this tries to do.

Grand Unified Theory of Anomalies TBD

SIMPLICITY

at $a \sim e^{-7} \sim 1/1100 \Rightarrow$

at $a \sim e^{-67+60} \sim 1/10^{30+25}$

reveals primordial sound waves in matter

\Rightarrow learn **contents & structure** at 380000 yr, $a \sim e^{-7}$

\Rightarrow infer the structure far far earlier $a \sim e^{-67+60}$

hot & cold spots agree with BE87 Gaussian stats $n_{pk}(<v)$

PLANCK2013: 826', 105 peaks, coldest -4.97σ

WMAP7: 800', 105 peaks, coldest -4.87σ significance 1:300

7+ numbers

Early Universe **STRUCTURE**

“red” noise: 2 numbers at $a \sim e^{-67+55}$

WHITEN \Rightarrow MASK

\Rightarrow FILTER BANK

(SSG42 filter) \Rightarrow

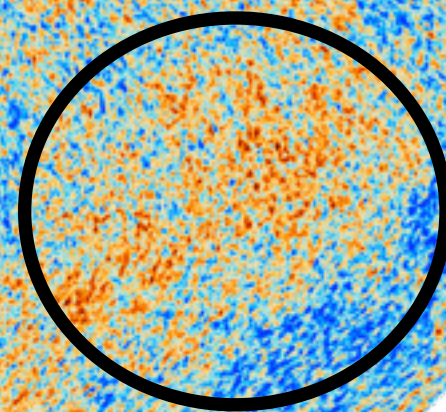
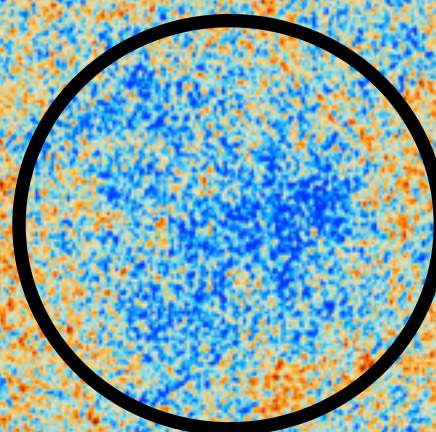
EXTRACT PEAKS

(hierarchical peak patches)

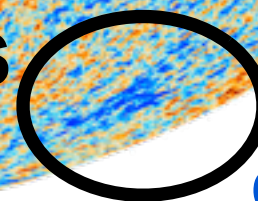
filter = extra dimension

scale space analysis

the ADS of our CFT



+ anomalies



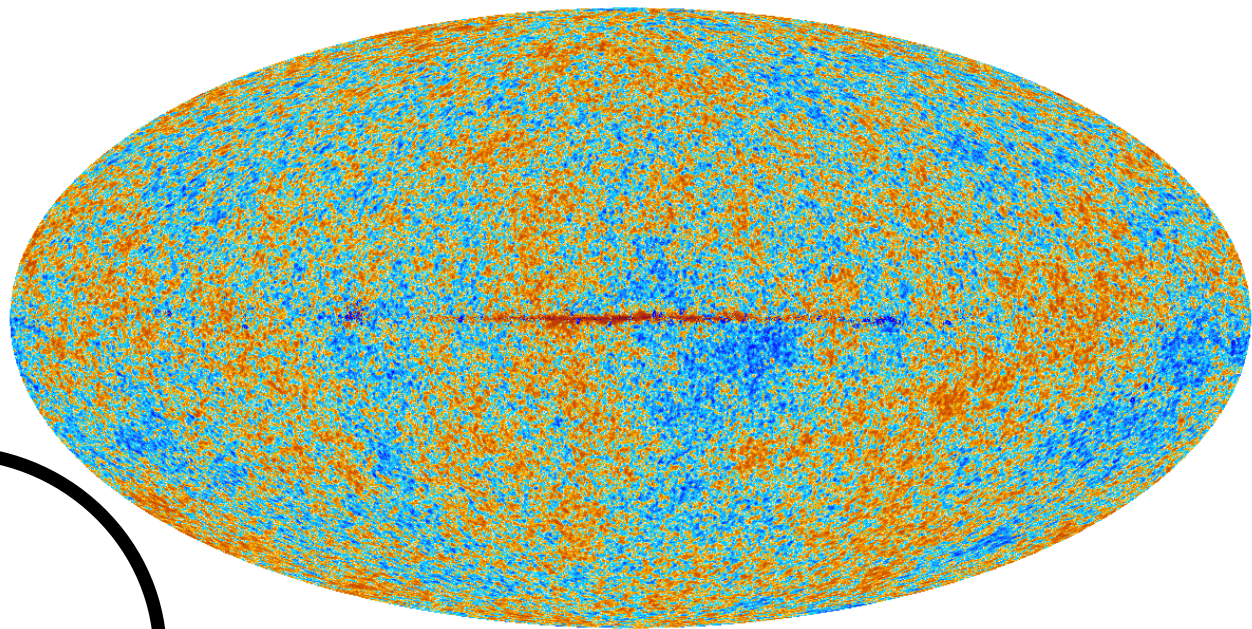
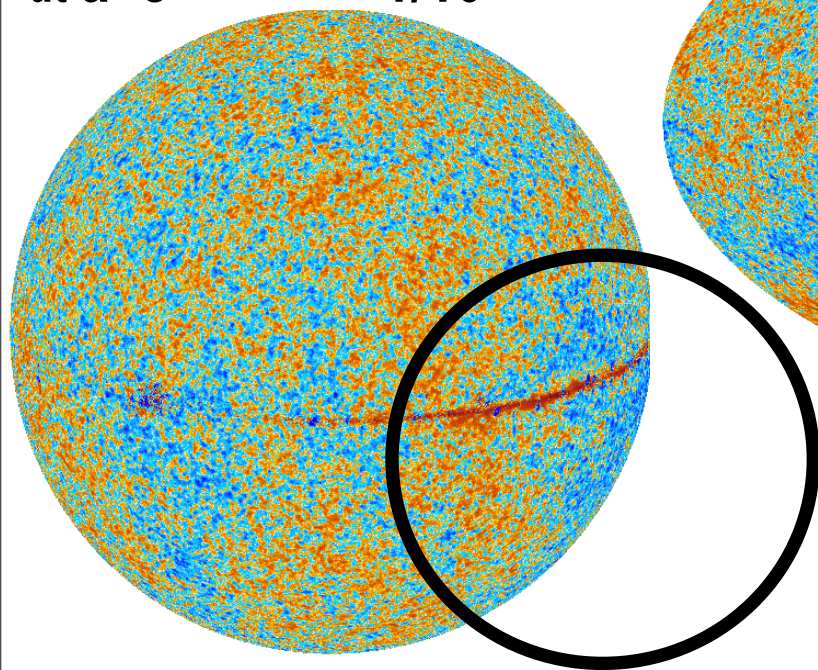
the rare cold spot

COMPLEXITY at $a \sim e^{-67}$?

SIMPLICITY

at $a \sim e^{-7} \sim 1/1100 \Rightarrow$

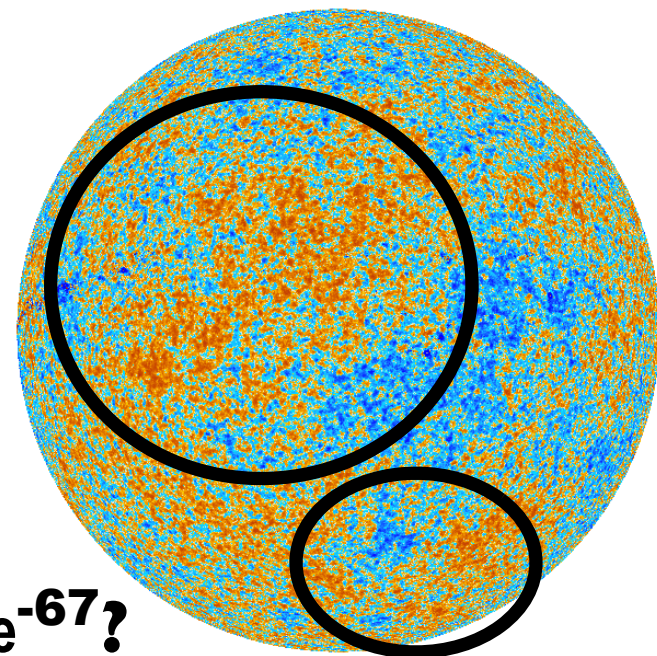
at $a \sim e^{-67+60} \sim 1/10^{30+25}$



Anomalies in Polarization? TBD

+ anomalies

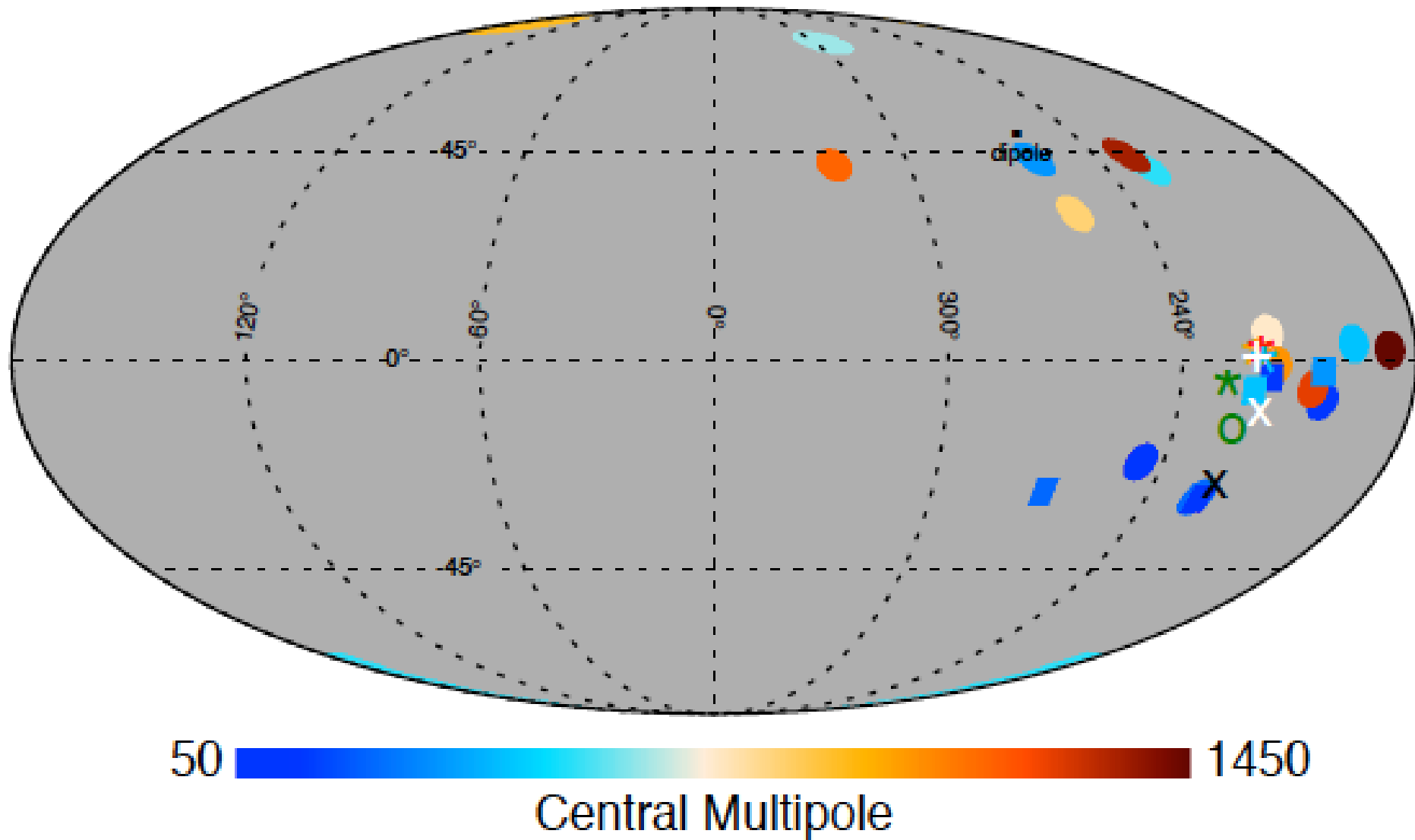
the rare
cold spot



COMPLEXITY at $a \sim e^{-67}$?

power spectrum asymmetry: dipole near Galactic Equator points towards LSS anomaly

$L < 400$ ~7% anomaly, $L > 400$ sub-percent (maybe, see Duncan Hanson's talk)



power spectrum asymmetry:
dipole near Galactic Equator points
towards LSS anomaly.

Low L asymmetry is firm P13 & WMAP,
high L subject to Doppler boost correction

Challinor & Lewis 02, Hanson+ 09, **Planck2103 XXVII,**
Doppler Boosting of the CMB

dipole modulation $\Delta T(\mathbf{q}) \Rightarrow (1 - (x \coth(x/2) - 1) \mathbf{q} \cdot \mathbf{v}) \Delta T(\mathbf{q})$,
 $x = hv/T$

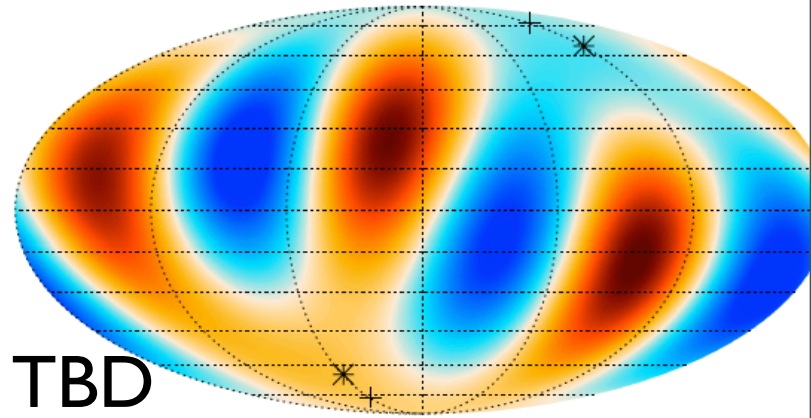
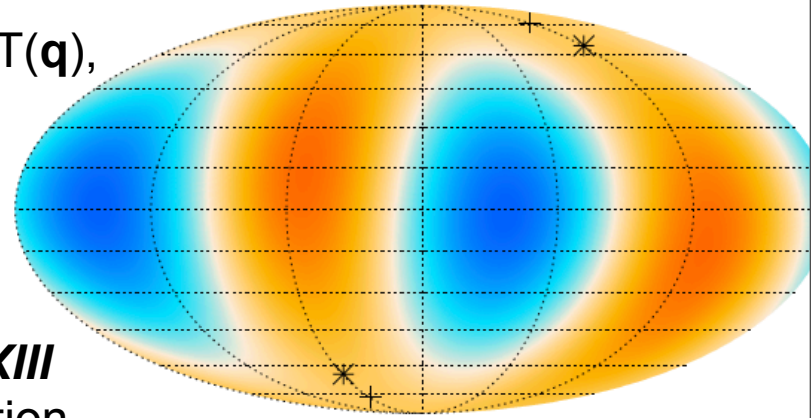
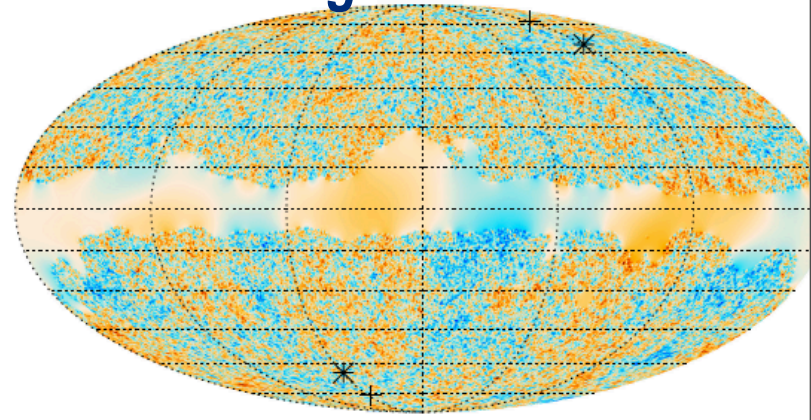
aberration $\mathbf{q} \Rightarrow \mathbf{q} + \nabla(\mathbf{q} \cdot \mathbf{v})$

5σ detection of kinematic dipole effects

influence on high L power asymmetry (cf. **P13 XXIII**
Isotropy & Statistics TBD) dipole power modulation
<0.2% with $L_{\max} = 2000$?

low L (<400) power asymmetry is robust

octupole quadrupole alignment
within ~ 10 deg



Anomalies in Polarization? TBD

Grand Unified Theory of Anomalies TBD

primordial non-Gaussianity

$$\zeta_{\text{NL}}(\mathbf{x}) = \zeta_{\text{G}}(\mathbf{x}) + \mathbf{f}_{\text{NL}} * (\zeta_{\text{G}}^2(\mathbf{x}) - \langle \zeta_{\text{G}}^2 \rangle)$$

local smooth. use optimal pattern estimator
cf. DBI inflation: non-quadratic kinetic energy

cosmic/fundamental strings/defects
from end-of-inflation & preheating

$$\zeta_{\text{NL}}(\mathbf{x}) = \zeta_{\text{G}}(\mathbf{x}) + \mathbf{F}_{\text{NL}}(\chi_{\text{b}}(\mathbf{x}))$$

modulating preheating

$\mathbf{f}_{\text{NL}}^{\text{eff}}$ + cold spots

$$\zeta_{\text{NL}}(\mathbf{x}) = \zeta_{\text{G}}(\mathbf{x}) + \mathbf{F}_{\text{NL}}(g_{\text{b}}(\mathbf{x}))$$

phonon $\sim \zeta_{\text{NL}} = \ln(\rho a^{3(1+w)})/3(1+w) \Rightarrow \mathbf{f}_{\text{NL}}^* = 3/5 \mathbf{f}_{\text{NL}} - 1 = 0.44 \pm 3.5$

Full-Sky Map

NonGaussian 3-point-pattern measure

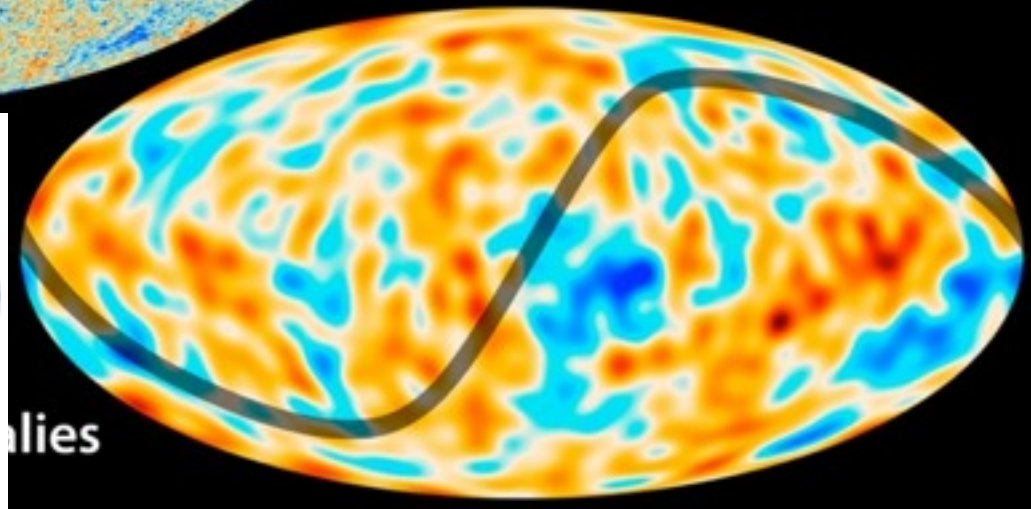
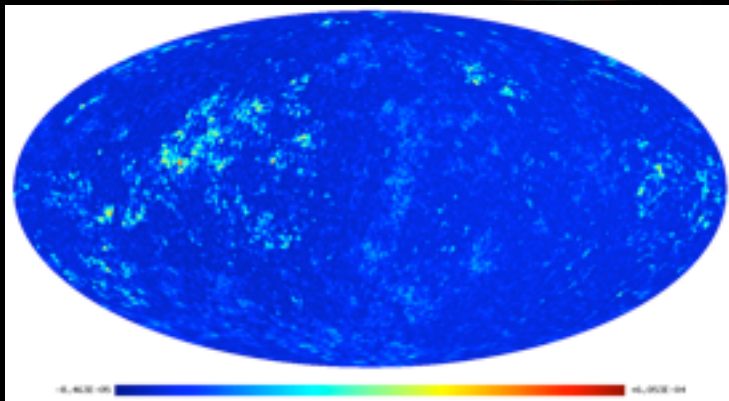
f_{NL} : 2.7 ± 5.8 local cf. ± 5 (Pext)

$\Rightarrow f_{\text{NL}}^* = 0.44 \pm 3.5$

$-f_{\text{NL}}$: 42.3 ± 75.2 equil

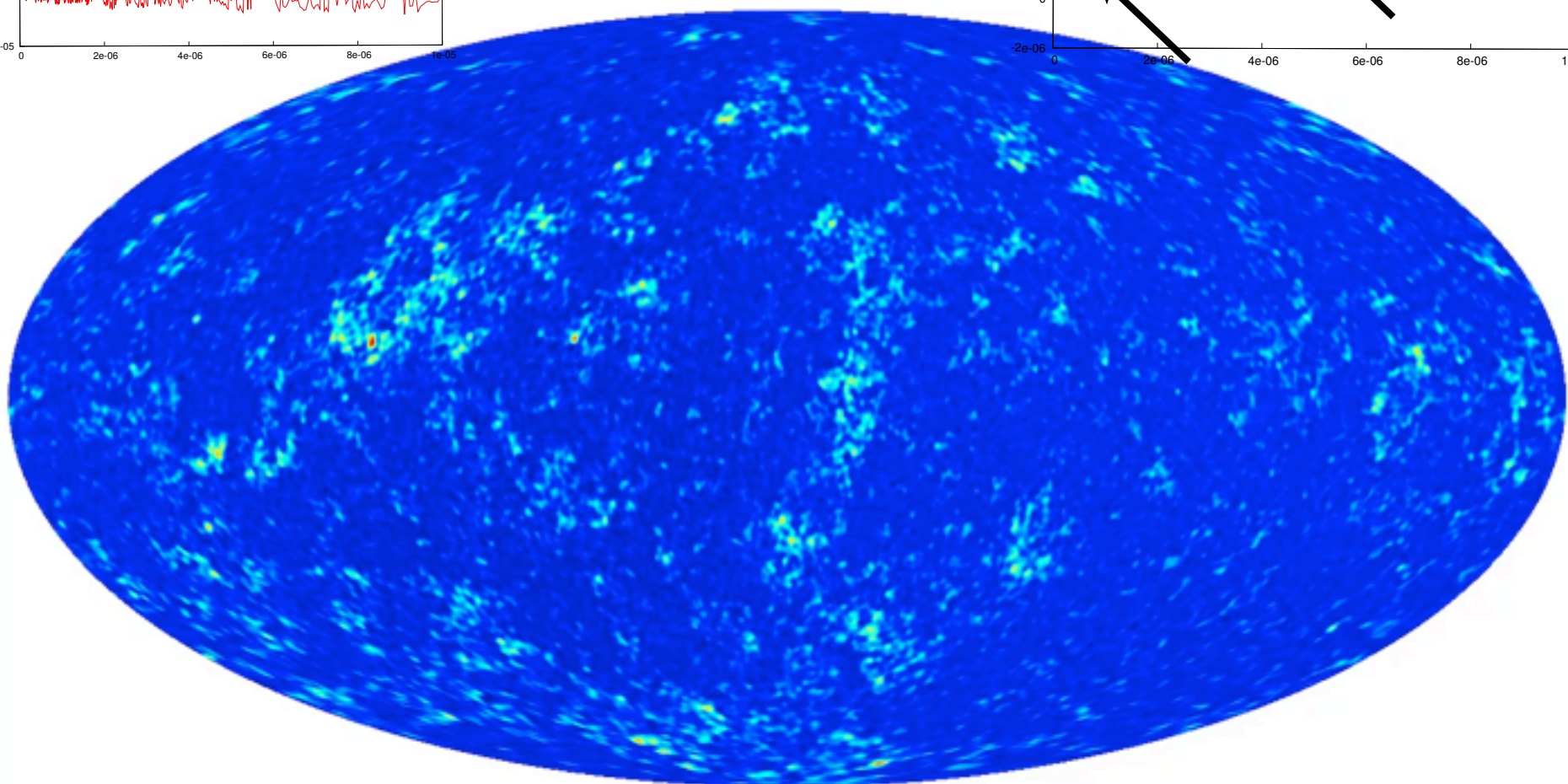
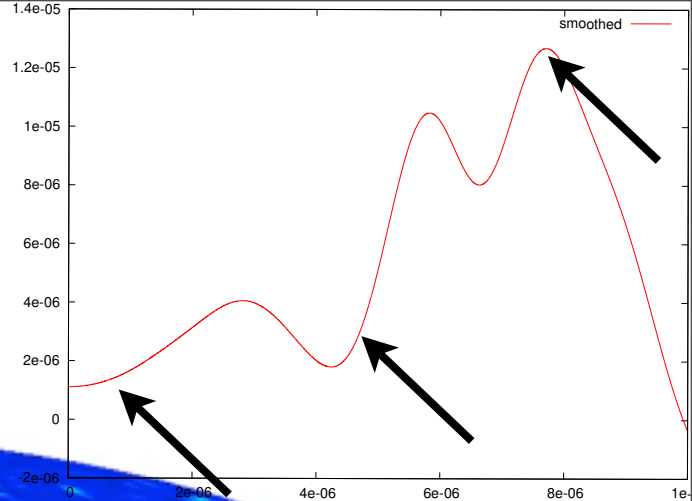
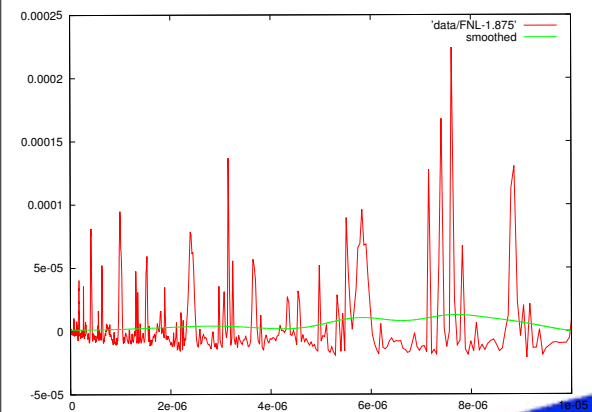
-25.3 ± 39.2 ortho

super-bias of ULSS & LSS fields
modulating preheating:
intermittency from rare event
nonG tails



subdominant structure change as we scan $\chi > h$

Bond, Braden, Frolov, Huang13

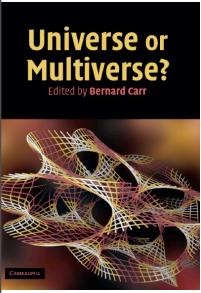


-1.203E-05

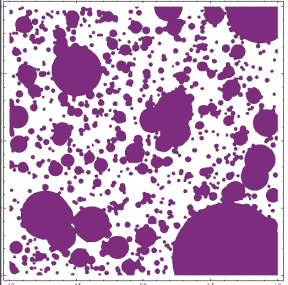
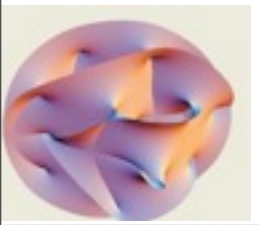
+7.520E-05

conclusions:
nothing definitive
yet **for anomalies,**
may just lead to
potential & >horizon
constraints
but amusing patterns do arise

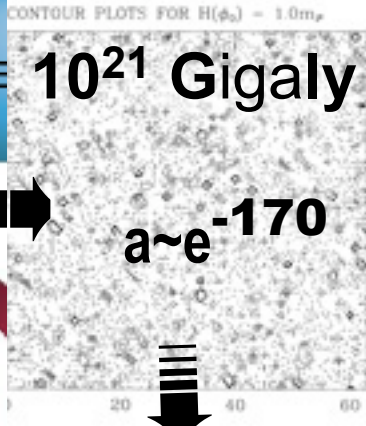
Horizons: the ultimate-speed constraint on light & information



higher dimensions 6?

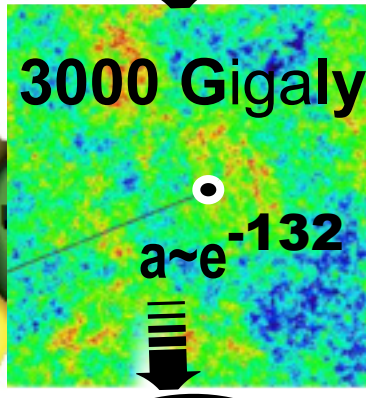


quantum tunnels = bubbly-U



10^{21} Gigaly

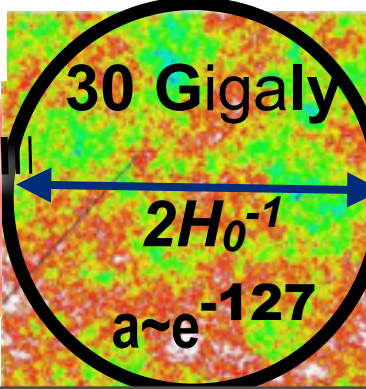
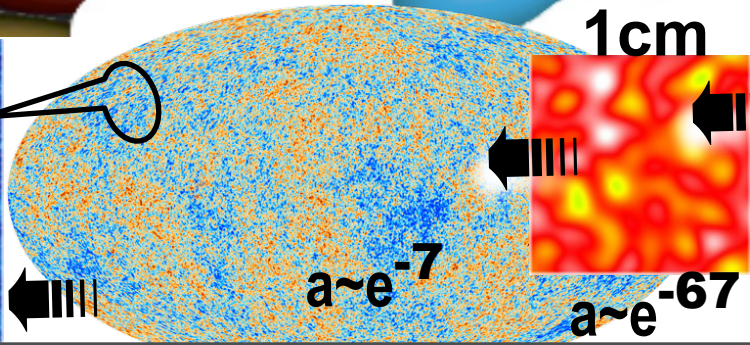
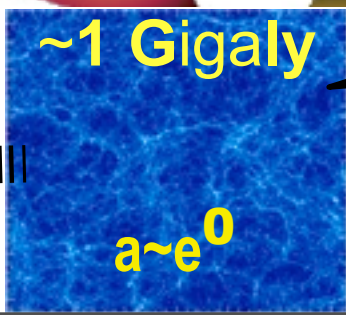
$a \sim e^{-170}$



3000 Gigaly

$a \sim e^{-132}$

a future DeVoid
 O dark dark dark. They all go into the dark, The vacant interstellar spaces, the vacant into the vacant ... So the darkness shall be the light



Fundamental Physics from the Planck Satellite

Planck 2013 results. XXII. Constraints on inflation

Planck 2013 Results. XXIV. Constraints on primordial non-Gaussianity

Planck 2013 results. XXIII. Isotropy and Statistics of the CMB

Planck 2013 results. XXV. Searches for cosmic strings and other topological defects

Planck 2013 results. XXVI. Background geometry and topology of the Universe

END

the Planck Collaboration, including individuals from more than 100 scientific institutes in Europe, the USA and Canada



Planck is a project of the European Space Agency, with instruments provided by two scientific Consortia funded by ESA member states (in particular the lead countries: France and Italy) with contributions from NASA (USA), and telescope reflectors provided in a collaboration between ESA and a scientific Consortium led and funded by Denmark.

Planck 2013 results. XII. Component separation

Planck 2013 results. XV. CMB power spectra and likelihood

Planck 2013 results. XVI. Cosmological parameters

Planck 2013 results. XVII. Gravitational lensing by large-scale structure

Planck 2013 results. XXVII. Doppler boosting of the CMB: Eppur si muove

Planck 2013 results. XIX. The integrated Sachs-Wolfe effect