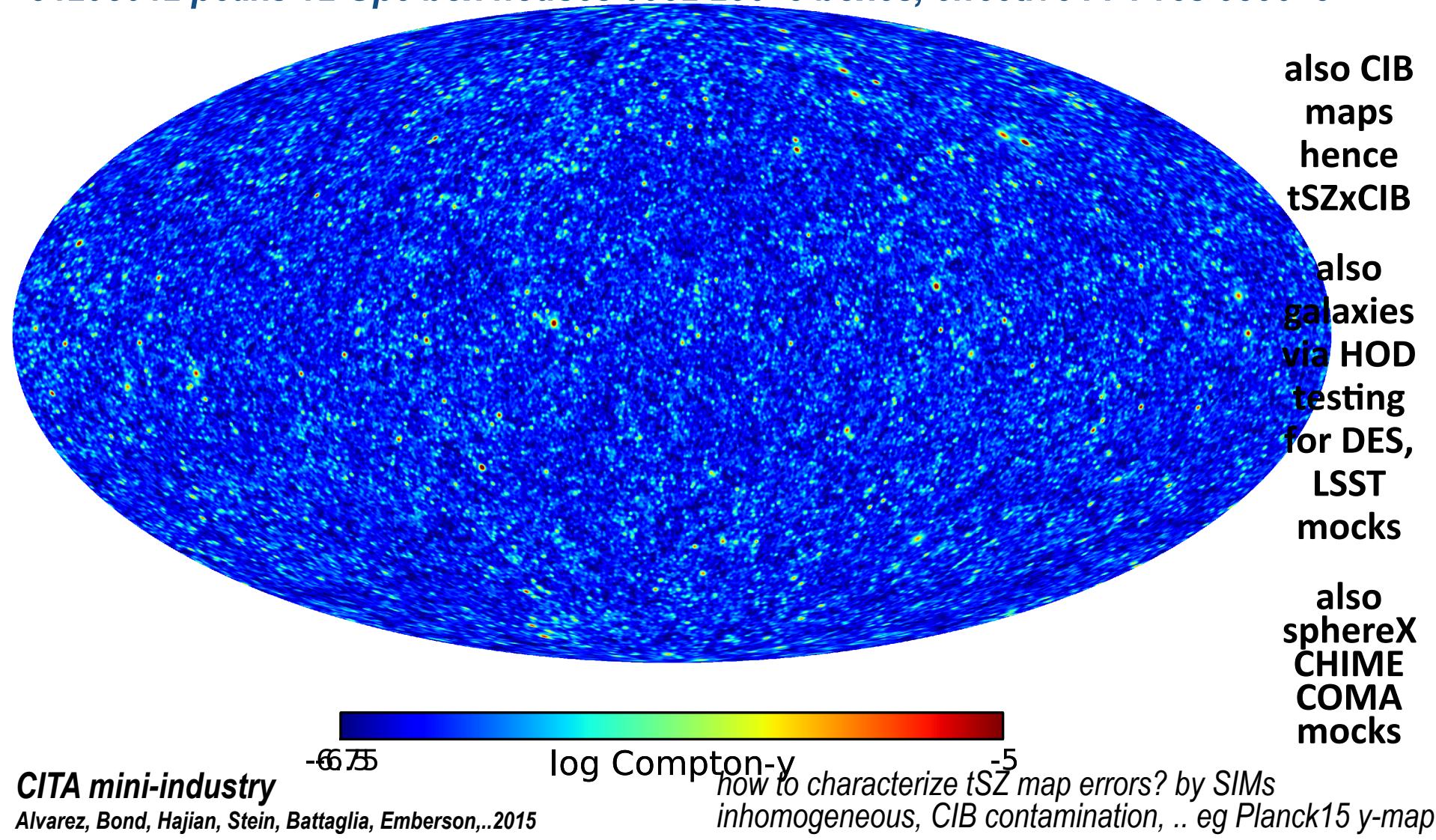
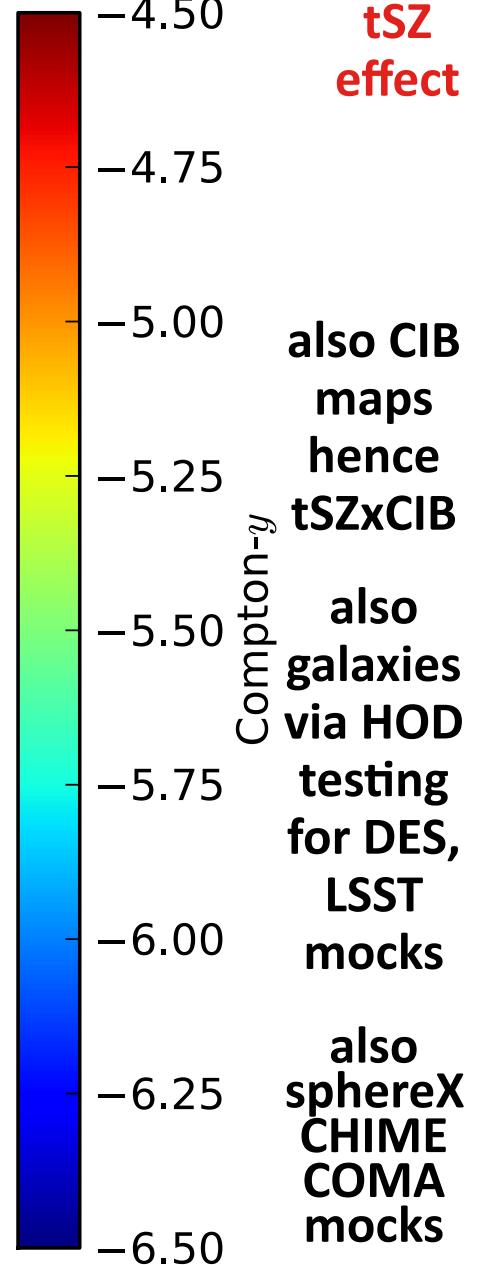
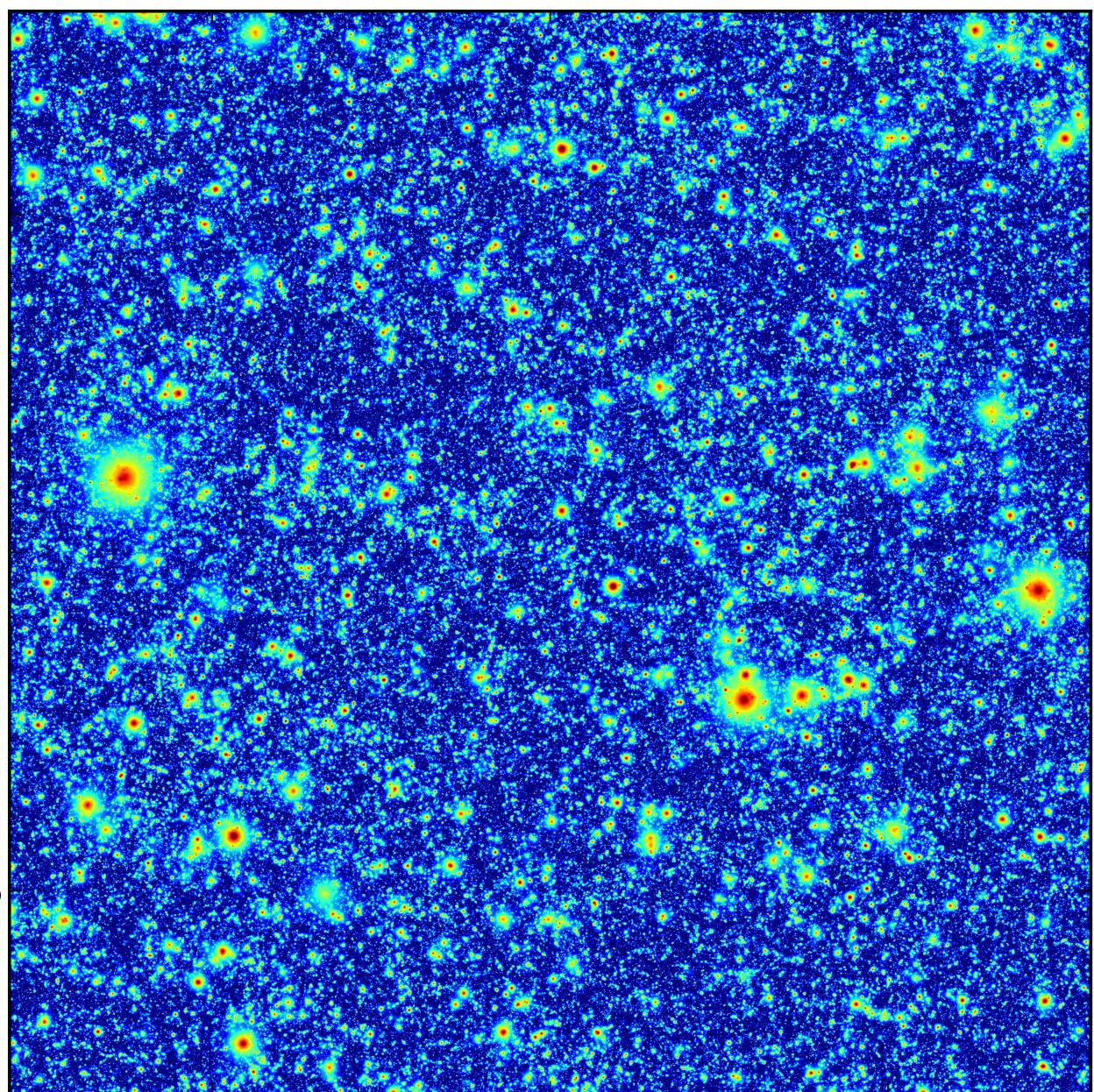


the Cosmic Web of Clusters, seen thru Compton
cooling of high pressure electrons by the CMB via peak patch sims tSZ effect
Lightcone Simulation of Clusters > $1.0 \times 10^{13} M_{\text{sun}}$ to $z=2.5$ in projected BBPS pressure
~40 minutes on 256 cores on SciNet, 30000 core IBM GPC !!
84298042 peaks 12 Gpc box houses 9952 256^3 boxes, effective FFT res 5856^3



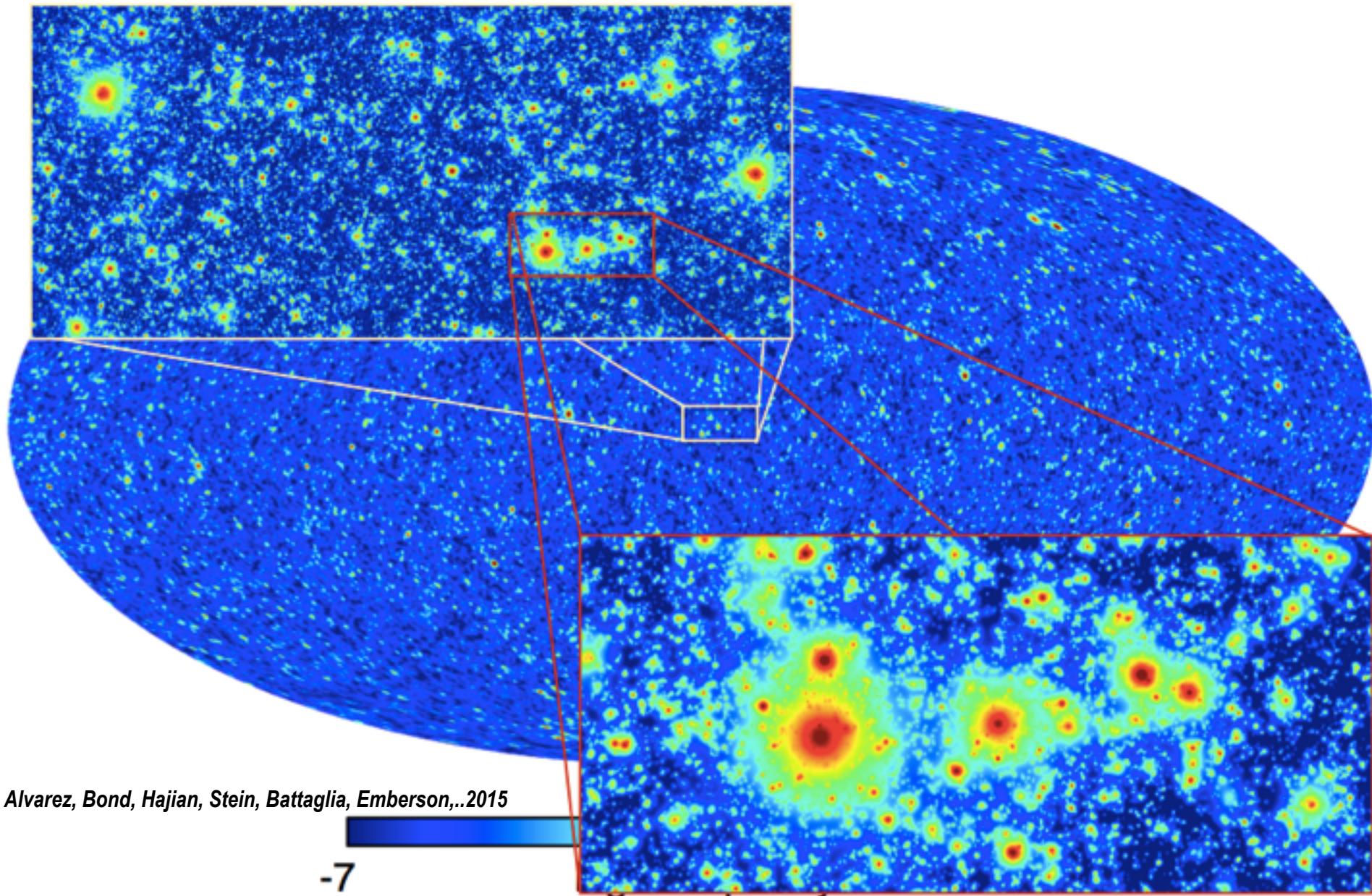
lots of structure. open: what is the relation of gastrophysics at higher z (~ groups) cf. lower z (~ cls)



the Cosmic Web of Clusters, seen thru Compton cooling of high pressure electrons by the CMB via *peak patch sims*

tsz

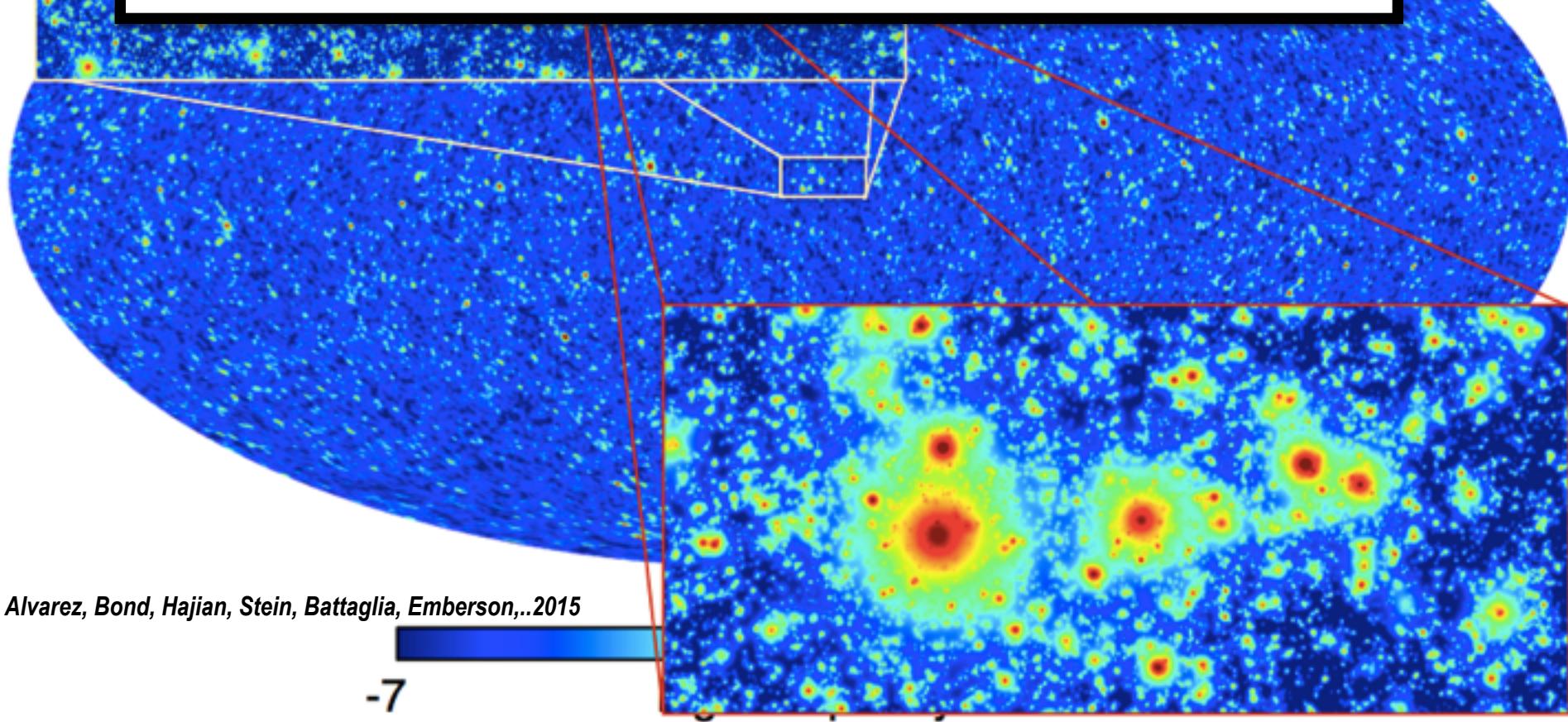
effect

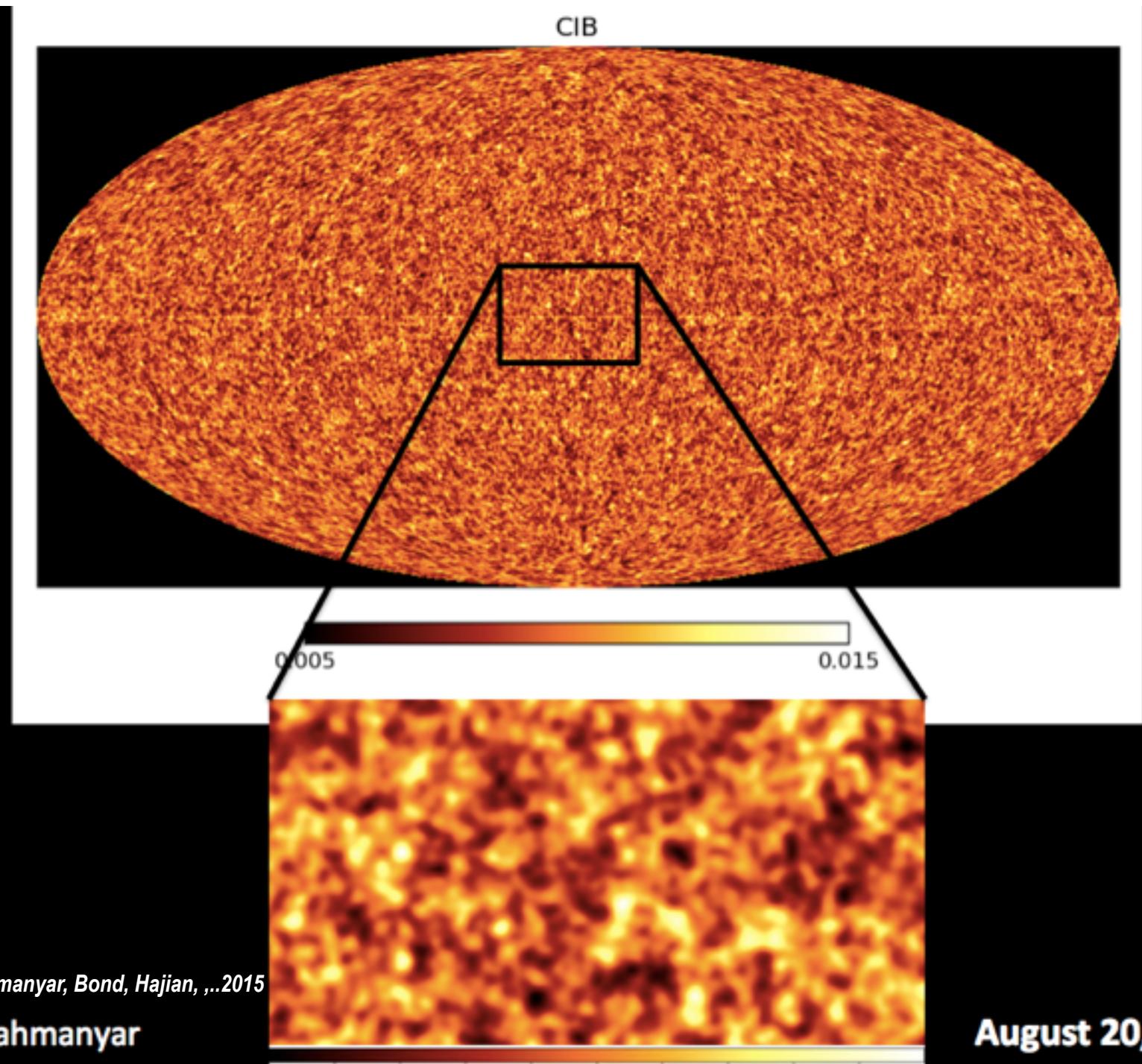


Alvarez, Bond, Hajian, Stein, Battaglia, Emberson,..2015

the Cosmic Web of Clusters, seen thru Compton
cooling of high pressure electrons by the CMB via *peak patch sims* tsz effect

modelling the fluctuations about mean
pressure fields is TBD but measurable /
measured in BBPS gasdynamical sims





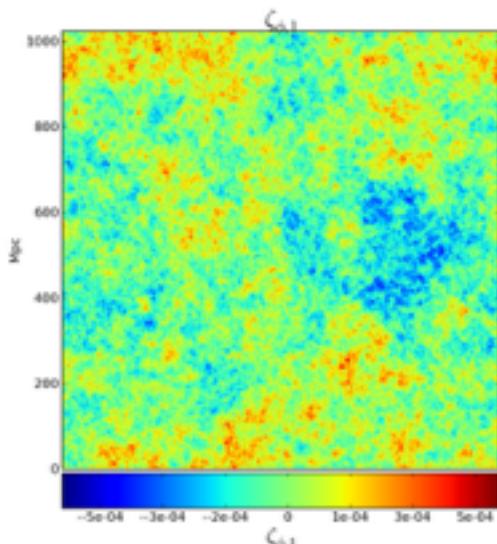
CIB
maps
hence
tSzxCIB

galaxies
via HOD

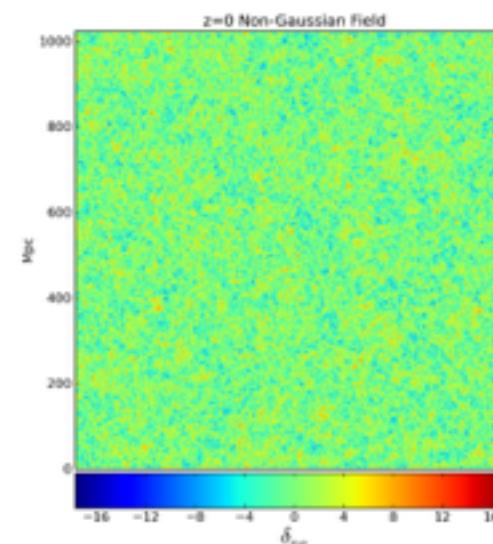
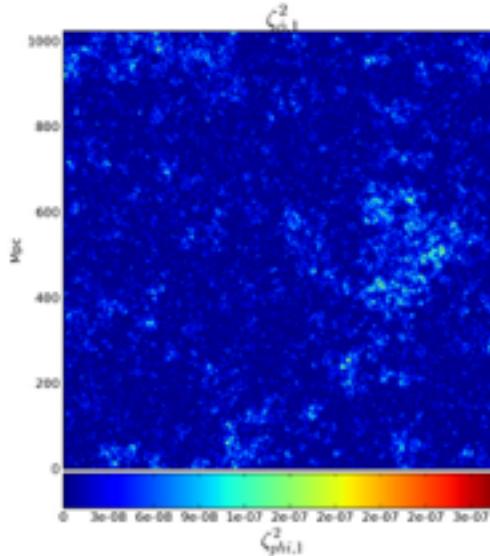
Planck
CIB halo
model
..so far

**Simulating primordial
non-Gaussianity for
clusters, galaxy
surveys, CIB, HI, ...**

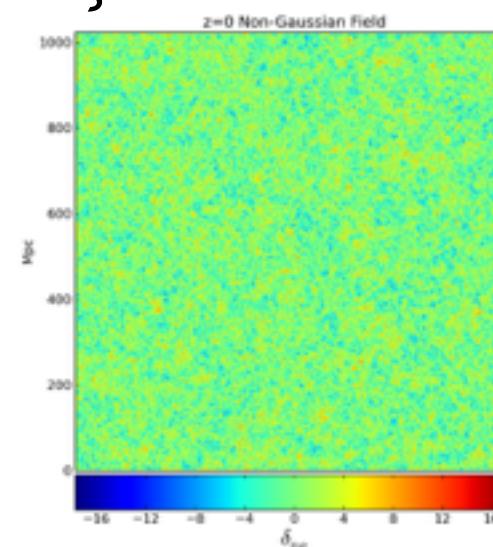
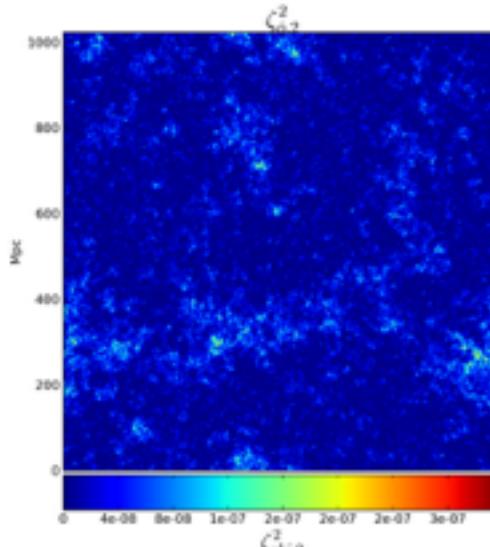
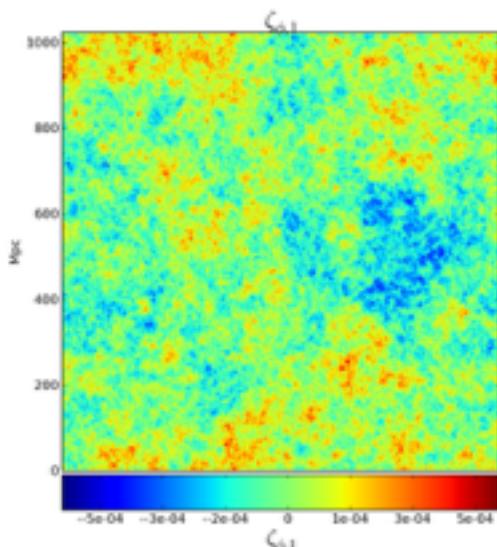
**Simulating non- Λ
dark energy models**
*modifying hot gas physics, dust CIB
physics, HOD, neutral hydrogen, CO, ..*



conventional inflaton-induced correlated ζ^2
conventional but uncorrelated ζ^2



the non-Gaussian
initial density field



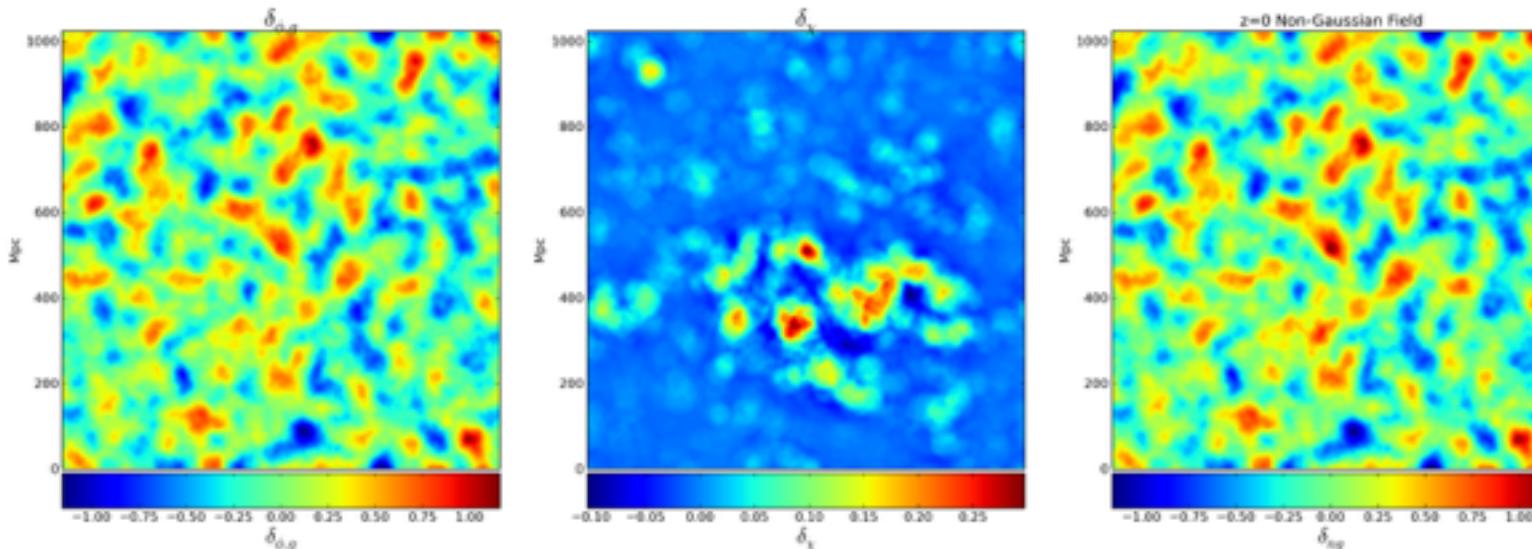
search with bispectrum & scale-dependent bias in power spectrum

$a_{\text{latt}}=1\text{Mpc}$, $N=1024$

LSS & nonGaussian mocks

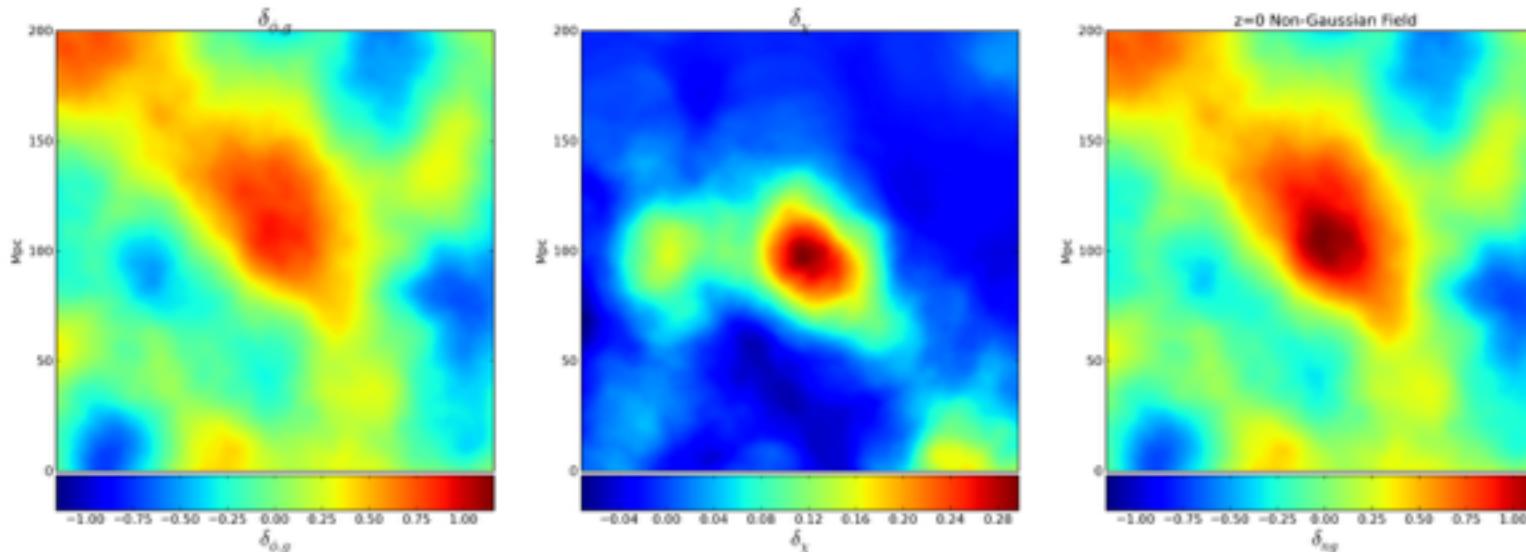
Gaussian Spike Model Smoothed on $R=32\text{Mpc}$

Alvarez, Bond, Huang, Stein, Braden, Frolov 14



modulated intermittent preheating nonG

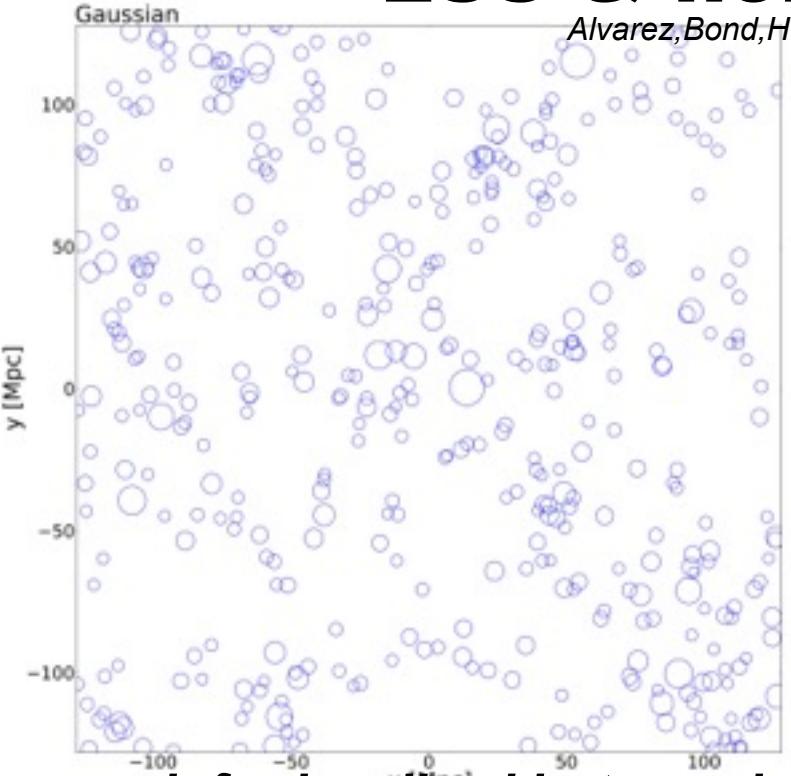
Gaussian Spike Model Smoothed on $R=32\text{Mpc}$



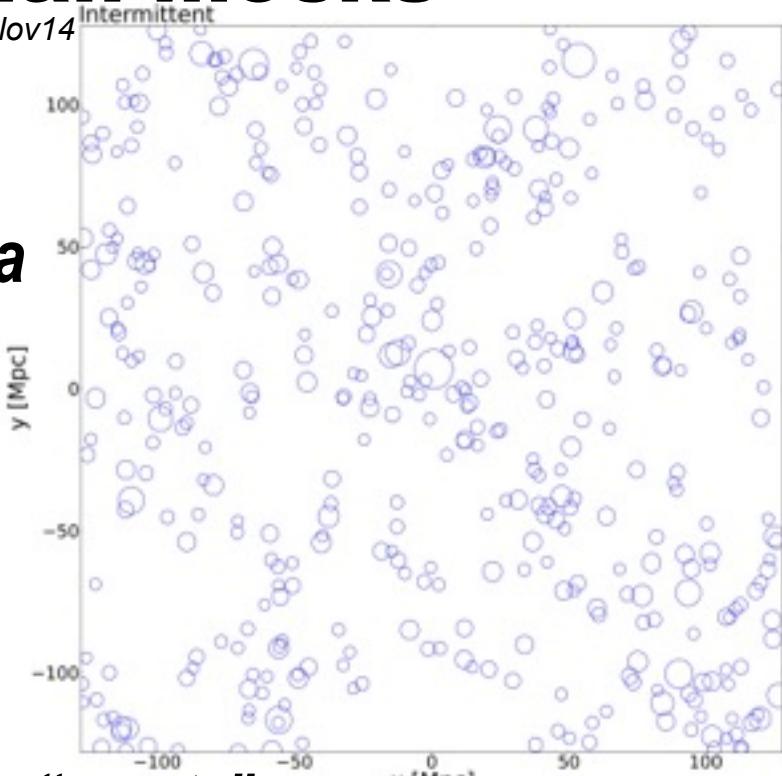
search for localized but very large scale rare “events” e.g., hierarchical peaks

LSS & nonGaussian mocks

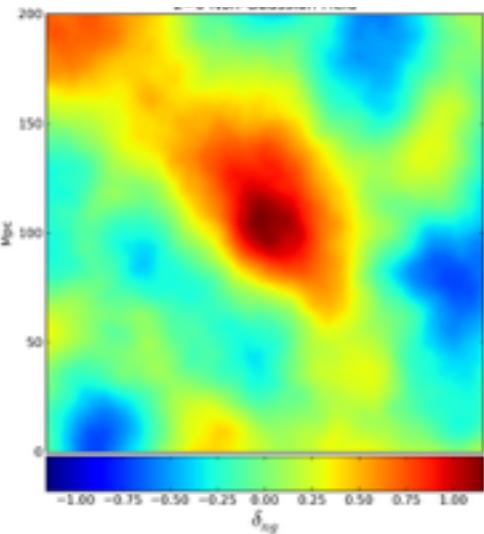
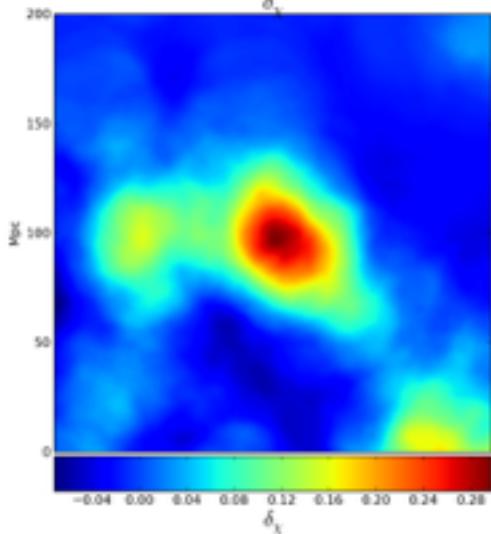
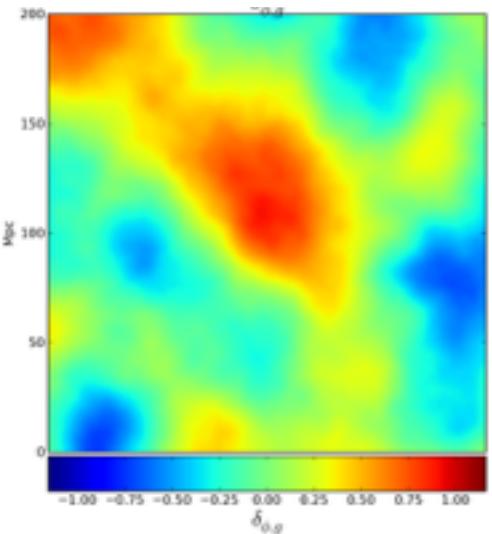
Alvarez, Bond, Huang, Stein, Braden, Frolov14



*halo nonG
patterns
galaxies via
HoD*



search for localized but very large scale rare “events” e.g., hierarchical peaks



usefulness of the pk patch simulation method for mocking **AdvACT?**

Planck y -map applications

non- G C_L^{SZ} error statistics

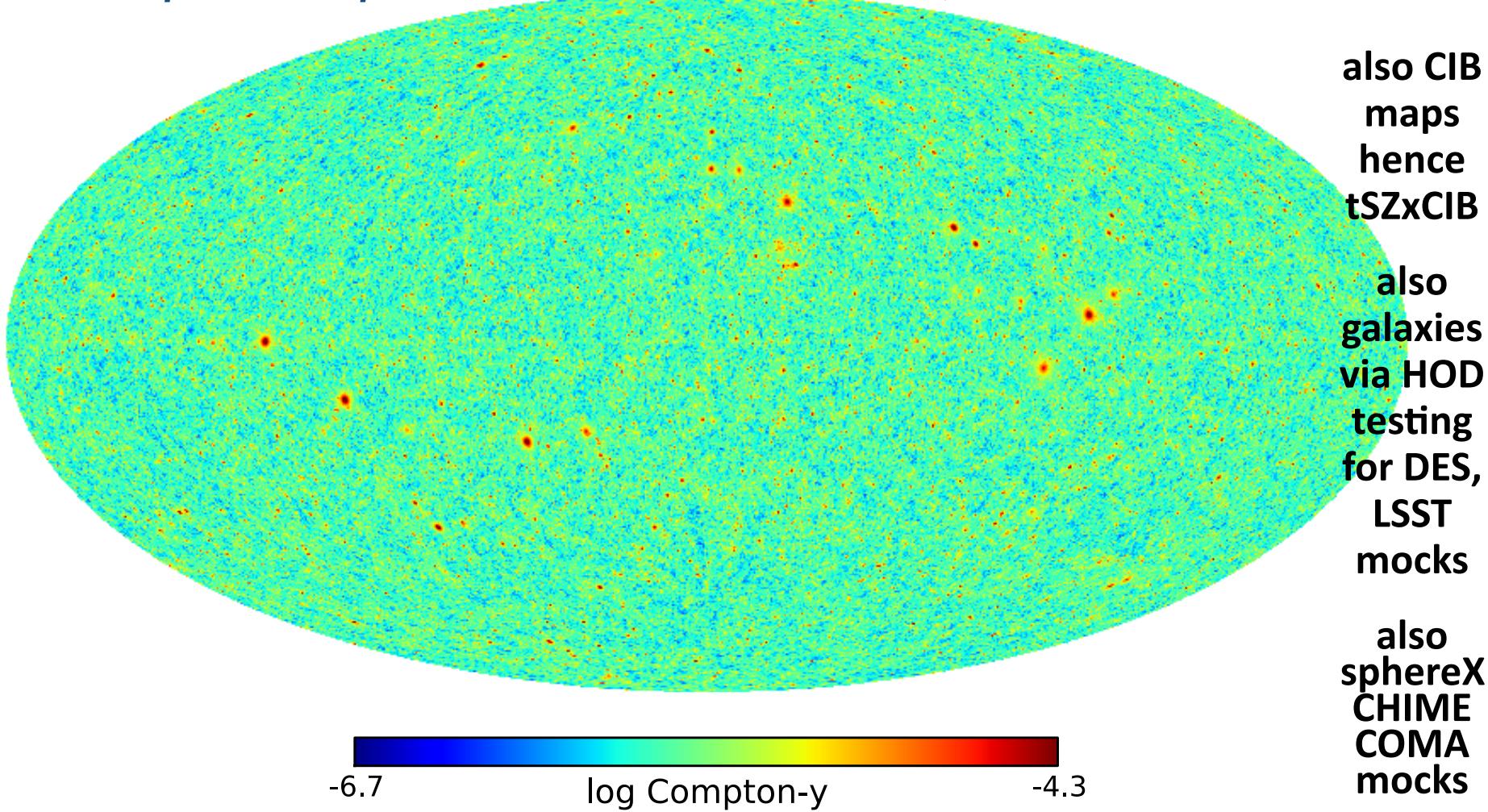
1,2,...N-point distributions in maps BM93/96

kSZ , κ_{lens} , ...

cross-correlations Xray-tSZ, Lens-tSZ, BCG-tSZ, CIB-tSZ, ..

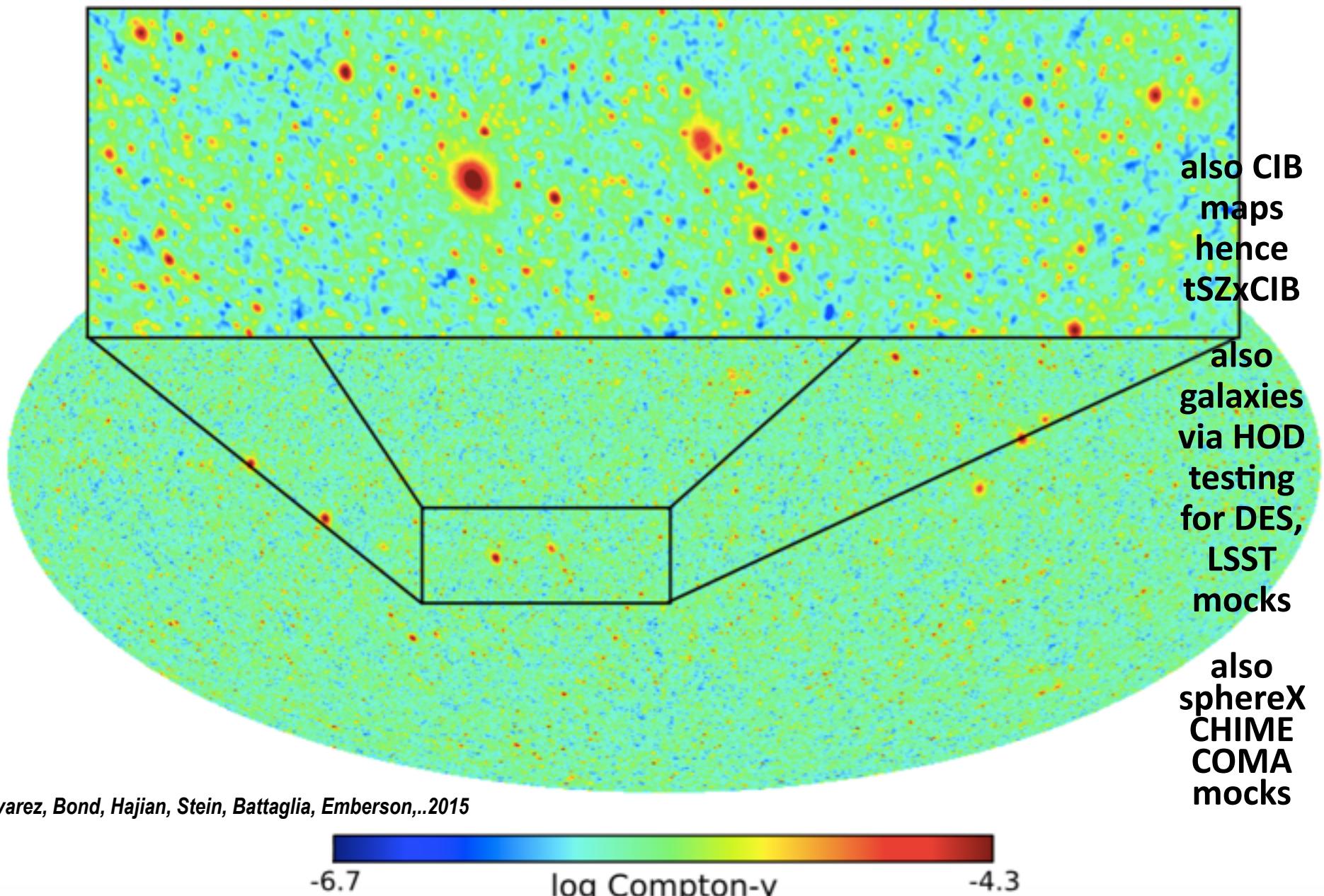
END

the Cosmic Web of Clusters, seen thru Compton
cooling of high pressure electrons by the CMB via peak patch sims **effect**
tsz
Lightcone Simulation of Clusters $> 1.0 \times 10^{13} M_{\text{sun}}$ to $z=2.5$ in projected BBPS pressure
~40 minutes on 256 cores on SciNet, 30000 core IBM GPC !!
84298042 peaks 12 Gpc box houses 9952 256^3 boxes, effective FFT res 5856^3



the Cosmic Web of Clusters, seen thru Compton cooling of high pressure electrons by the CMB via *peak patch sims*

tsz
effect

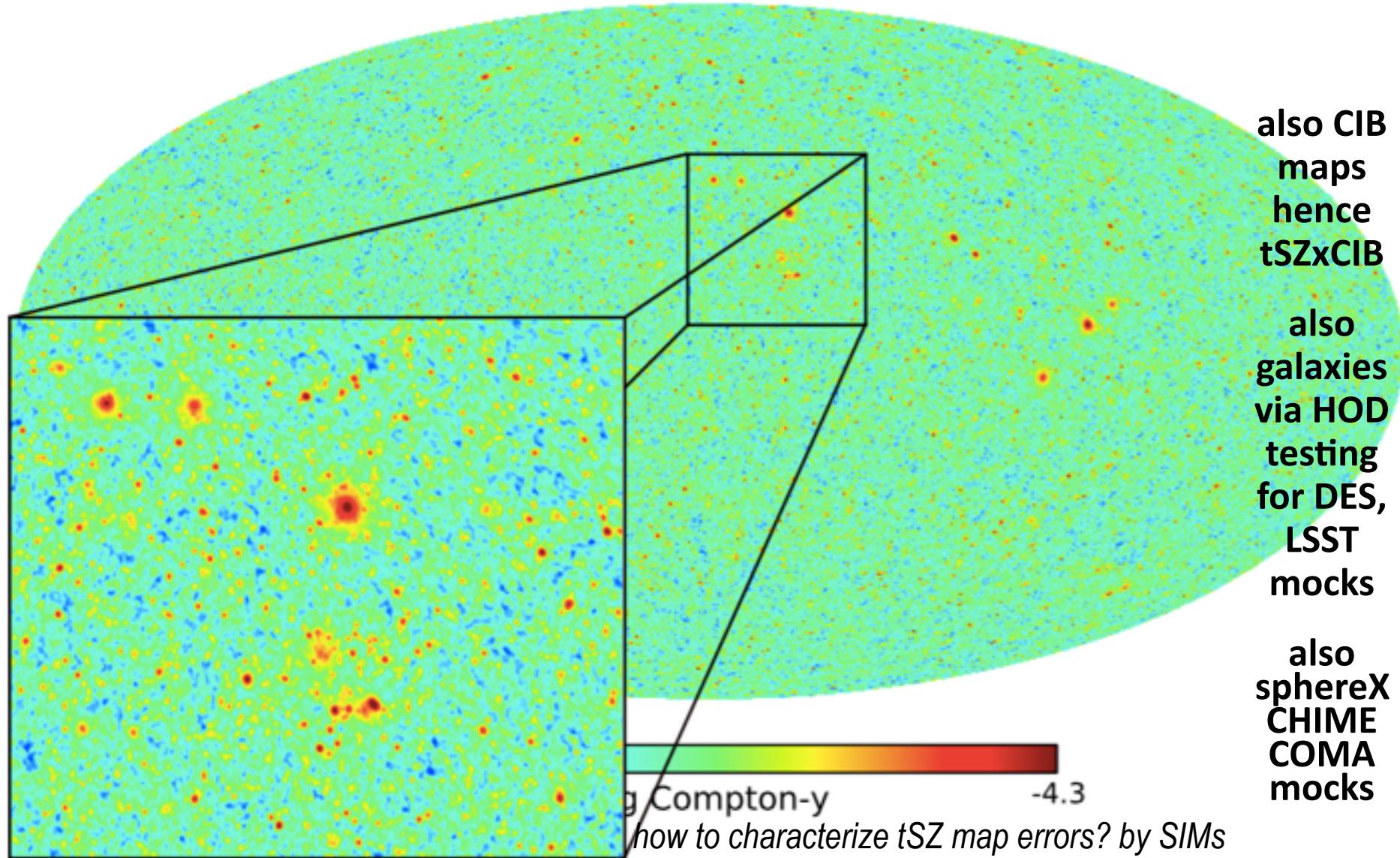


the Cosmic Web of Clusters, seen thru Compton cooling of high pressure electrons by the CMB via *peak patch sims*

tsz

effect

Lightcone Simulation of Clusters > $1.0 \times 10^{13} M_{\text{sun}}$ to $z=2.5$ in projected BBPS pressure

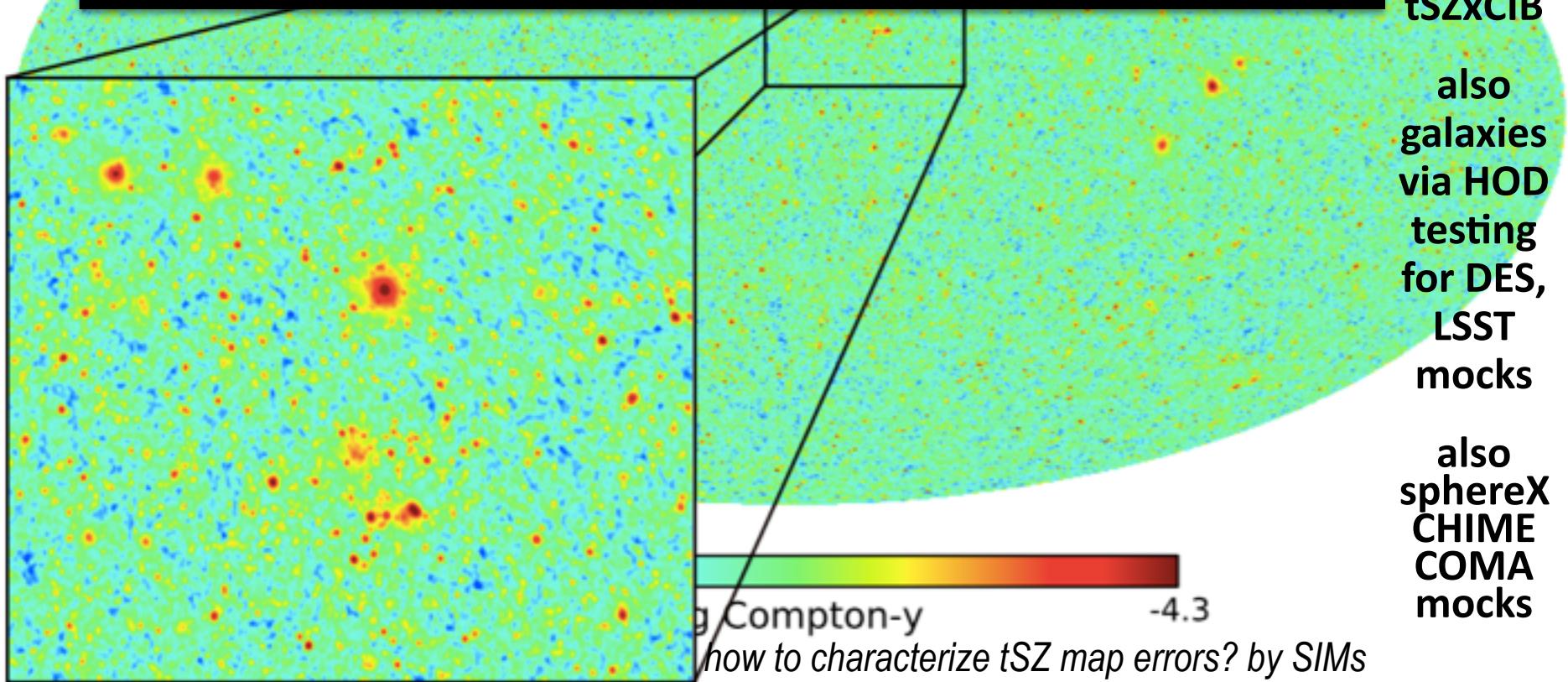


the **Cosmic Web of Clusters**, seen thru Compton
cooling of high pressure electrons by the CMB via *peak patch sims*

tsz
effect

Lightcone Simulation of Clusters > $1.0 \times 10^{13} M_{\text{sun}}$ to $z=2.5$ in projected BBPS pressure

modelling the fluctuations about mean
pressure fields is TBD but measurable /
measured in BBPS gasdynamical sims



how to characterize tSZ map errors? by SIMs

inhomogeneous, CIB contamination, .. eg Planck15 y-map